

# Large uncertainties in global hydroxyl projections tied to fate of reactive nitrogen and carbon

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**The hydroxyl radical (OH) sets the oxidative capacity of the atmosphere, and thus the removal rate of pollutants and reactive greenhouse gases. While observationally derived constraints exist for global annual mean present-day OH abundances and interannual variability, OH estimates for past and future periods rely primarily on global atmospheric chemistry models. These models disagree  $\pm 30\%$  in mean OH and in its changes from the pre-industrial to late 21<sup>st</sup> century, even when forced with identical anthropogenic emissions. A simple steady-state relationship that accounts for ozone photolysis frequencies, water vapor, and the ratio of reactive nitrogen to carbon emissions explains temporal variability within most models, but not inter-model differences. Here, we show that departure from the expected relationship reflects the treatment of reactive oxidized nitrogen species (NO<sub>y</sub>) and the fraction of emitted carbon that reacts within each chemical mechanism. Our findings imply a need for observational constraints on NO<sub>y</sub> partitioning and lifetime as well as the fate of carbon-containing reaction intermediates in order to better constrain models and reduce uncertainties in OH projections.**

hydroxyl | atmospheric oxidative capacity

The hydroxyl radical (OH) is a keystone chemical species in the atmosphere, determining the removal rate of many trace gases of importance to climate, composition and human and ecosystem health (1). For example, reaction with OH in the troposphere\* is the dominant sink for methane, a powerful greenhouse gas and precursor for tropospheric ozone, a major surface pollutant and greenhouse gas itself (2). Understanding what drives variability in OH is therefore critical for forecasting future changes in the self-cleansing capability of the atmosphere. The fundamental chemistry of background OH has been well known for decades (3–5). Nevertheless, global atmospheric chemistry models show large disagreement in mean OH and its transient response to specified changes in emissions (6, 7).

Fig. 1 shows global mean OH and its temporal evolution within the ensemble of simulations that participated in the Atmospheric Chemistry-Climate Model Intercomparison Project (ACCMIP) (10). The ACCMIP ensemble is a comprehensive suite of global three-dimensional atmospheric chemistry models driven by identical anthropogenic emission scenarios for the period 1850 to 2100 (8, 9). Despite applying identical anthropogenic emissions in all ACCMIP models, tropospheric mean abundances of OH range  $\pm 30\%$  relative to the multi-model mean during the last decade of the historical simulation (gray-shaded interval of Fig. 1). The models also disagree as to whether OH increases or decreases across any prescribed

emission scenario, except for a small window between 1980 and 2010 (when all increase).

Our best estimates of global mean abundance and interannual variability from OH rely on proxy measurements, particularly methyl chloroform (11), as the high reactivity and short lifetime of OH make direct measurement difficult and impractical for constraining spatial and temporal variability (12). On average, global atmospheric chemistry models cannot reproduce meridional gradients in CO and other long-lived reactants, implying possible errors in simulated OH spatial and seasonal distributions (6, 13). They also overestimate global mean OH with respect to observational constraints from the methane and methyl chloroform lifetimes (6, 7), and underestimate the magnitude of interannual variability in OH inferred from proxy observations ( $0.5 \pm 0.4\%$  of year-to-year changes in the ACCMIP ensemble versus  $2.3 \pm 1.5\%$  derived from methyl chloroform (11)). These findings highlight gaps in our understanding of the processes that determine the oxidative capacity of the atmosphere and its variability, thereby hindering our ability to accurately predict its future evolution.

In contrast to the inter-model discrepancies in OH abundances, trends and variability, the models tend to produce consistent simulations for key tropospheric species that are intimately coupled with OH, such as tropospheric ozone (14). This implies that some—or even all—models are capturing mean abundances and spatial and temporal trends of longer-lived species at least partly for the wrong reasons.

## Significance Statement

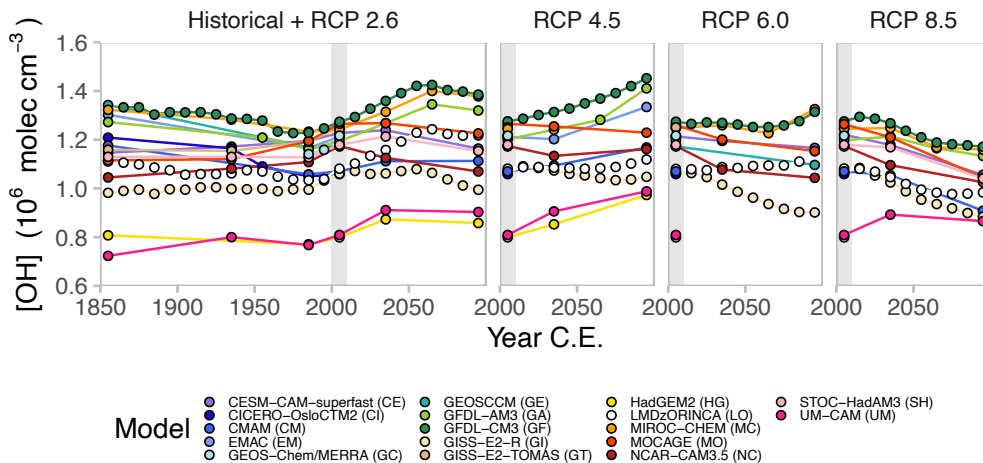
Reaction with OH is the dominant loss mechanism for many atmospheric gases of interest for air quality, climate change, and stratospheric ozone. Understanding how and why OH may change in the future is therefore paramount for predicting changes in the societal impacts associated with such changes. Future projections of models strongly disagree in how OH responds to changing emissions and climate – even in the sign of the change. Here, we demonstrate inter-model differences in OH are best explained by disparate implementations of chemical and physical processes that affect reactive oxides of nitrogen and organic chemical species. Targeted observations can reduce uncertainty in the chemical budgets of these key species to increase confidence in future projections of composition and its impacts.

L.T.M. designed the research; L.T.M. performed the research; A.M.F. and D.T.S. advised the research; L.T.M. wrote the paper; All authors contributed to the discussion.

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\* The lowermost several kilometers of the atmosphere in contact with the Earth's surface.



**Fig. 1.** Large disagreement in decadal mean tropospheric OH and its transient evolution in global atmospheric chemistry models. The models shown prescribe identical anthropogenic emissions from a historical reconstruction (8) and four possible future “Representative Concentration Pathway” (RCP) scenarios (9). Different colors represent different models. Gray rectangle highlights the period 2000-2010 in each scenario. Full model names shown in legend with a two-letter abbreviation shown in parentheses used in subsequent figures.

54 The original analysis of OH in the ACCMIP simulations  
 55 noted that the change in OH over time in a given model cor-  
 56 related with the ratio of change in its burden of reactive ni-  
 57 trogen oxides ( $\text{NO}_x \equiv \text{NO} + \text{NO}_2$ ) to change in its carbon  
 58 monoxide (CO) burden, but did not provide a mechanistic  
 59 explanation (6). Here, we re-examine the ACCMIP model  
 60 ensemble through the lens of fundamental OH chemistry to  
 61 explain the disparate behavior between the models. The key  
 62 areas of uncertainty we identify provide a target for future ob-  
 63 serving strategies to advance most rapidly our understanding  
 64 as formalized in the models used to project future atmospheric  
 65 abundances of pollutants and reactive greenhouse gases.

### 66 A Steady-State Relationship for OH

67 The key reactions controlling OH are highly and nonlinearly  
 68 coupled to one another (3-5, and see *SI* Fig. S1). In the  
 69 GEOS-Chem chemistry-transport model (CTM; see Materials  
 70 and Methods section), primary production of OH by photol-  
 71 ysis or photolysis-initiated chains contributes to approxi-  
 72 mately 60 % of total production in the present day, with most of that  
 73 source resulting from photolysis of ozone in the presence of wa-  
 74 ter vapor. The remaining 40 % results from chemical recycling  
 75 of the  $\text{HO}_x$  family ( $\text{HO}_x \equiv \text{OH} + \text{peroxy radicals}$ ) via reac-  
 76 tion of peroxy radicals with reactive oxides of nitrogen ( $\text{NO}_x$ )  
 77 or ozone. Loss of OH is primarily via reaction with reduced  
 78 carbon species such as CO (40 %) and methane (15 %), with  
 79 products of methane oxidation, other non-methane volatile  
 80 organic compounds (NMVOCs) and their oxidation products,  
 81 ozone,  $\text{H}_2$ ,  $\text{HO}_x$  species, and other minor species account-  
 82 ing for the remainder. Here, we refer to the sum of CO, methane  
 83 and all NMVOCs as total “reactive carbon.”

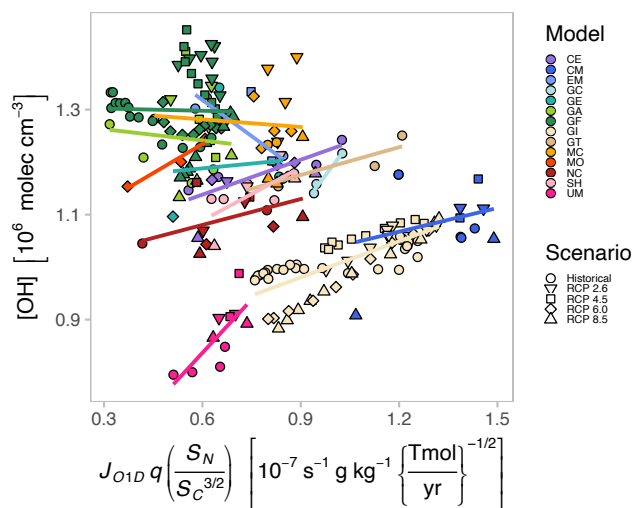
84 The equations that describe the dominant OH photochem-  
 85 istry are simple enough that they yield an analytic solution  
 86 for OH when one assumes “steady state,” in which the sum of  
 87 the source terms is balanced by all loss terms (see *SI* Sect. S1).  
 88 At steady state, variations in OH should vary proportional to

$$89 \quad J_{\text{O}(^1D)} q \frac{S_N}{S_C^{3/2}}, \quad [1]$$

90 where  $J_{\text{O}(^1D)}$  is the photolysis frequency of ozone to  $\text{O}(^1D)$   
 91 ( $\text{s}^{-1}$ ),  $q$  is the specific humidity ( $\text{kg H}_2\text{O kg air}^{-1}$ ), and  $S_N$   
 92 and  $S_C$  are the emission rates of  $\text{NO}_x$  ( $\text{Tmol N yr}^{-1}$ ) and

reactive carbon ( $\text{Tmol C yr}^{-1}$ ) species, respectively. The first  
 two terms limit the rate of total  $\text{HO}_x$  production whereas the  
 emission ratio determines the relative partitioning of  $\text{HO}_x$  be-  
 tween OH and peroxy radicals (with nitrogen oxides favoring  
 OH and reactive carbon favoring peroxy radicals). The expo-  
 nent in the denominator accounts for nonlinear feedbacks  
 in the coupled chemistry. Collectively, the relationship de-  
 scribes how the basic OH photochemistry changes in response  
 to changing emissions and climate parameters. Eq. [1] has  
 been previously used to explain global OH variability within  
 a single CTM across a wide range of climate and emissions  
 scenarios with high skill (15). However, in principal it should  
 hold across multiple models, as it derives from the analytic  
 solution to the steady-state OH photochemistry.

Large temporal and inter-model variability exists within  
 the decadal mean evolution of each component of Eq. [1] of  
 the ACCMIP ensemble (Fig. S2). The models employed a  
 variety of methods for determining photolysis rates, ranging  
 from look-up tables to online calculation with column radi-  
 ation codes (10). Most models agree in their mean photolysis  
 frequencies, although outliers exist in the ensemble. Changes  
 in photolysis frequencies with time are relatively small com-  
 pared to inter-model variability, although it has been pre-  
 viously shown that OH is especially sensitive to them (e.g.,  
 15-17). These temporal changes are driven primarily by  
 changes in overhead ozone abundances associated with anthro-  
 pogenic emissions of stratospheric ozone-depleting substan-  
 ces and greenhouse gases (18), with peak photolysis rates in  
 each model occurring during the height of stratospheric ozone  
 loss ca. 1980-2010. The models are in tight agreement in both  
 tropospheric mean water vapor abundance and its exponential  
 increase over time in response to temperature-driven increases  
 in saturation water vapor pressure (19).  $\text{NO}_x$  emissions are  
 dominated by the prescribed anthropogenic sources and there-  
 fore temporal variability dominates over inter-model variabil-  
 ity. However, slight differences do exist between models and  
 in time due to different parameterizations of climate-sensitive  
 natural sources such as lightning (20). In contrast, emissions  
 of reactive carbon species vary widely across the models and  
 in time. These variations are dominated by natural sources;  
 the models employed a wide range of biogenic NMVOC imple-  
 mentations with respect to species emitted (including none),



**Fig. 2.** Decadal mean OH as a function of the steady-state photochemical relationship. Variability within many individual models is accurately explained by Eq. [1]. However, large inter-model variability is not. Individual models are identified by color and emission scenario by shape. An ordinary linear regression line is shown for each model sub-population. See S/ Fig. S3 for individual model detail at annual time scales.

all methane is oxidized to CO (22). The large inter-model variability in NMVOC emissions are not correlated with inter-model variability in global mean OH, even though NMVOCs consume OH, because relatively short NMVOC lifetimes restrict their influence to be immediately downwind of sources, which are mostly in the continental boundary layer. Instead, differences in the portion of NMVOCs fully oxidized to CO (i.e.  $\epsilon_C$ ), a gas with a sufficiently long lifetime to mix into the free and remote troposphere, best explain inter-model variability in global mean OH.

From the perspective of reactive carbon, the atmosphere is a low-temperature furnace: reduced organic gases are emitted and subsequently oxidized through a chain of increasingly semi-oxidized species toward CO and CO<sub>2</sub>. However, increased oxidation tends toward lower-volatility species that may condense (23). Other intermediates are soluble and/or prone to surface uptake, and physical processes may siphon these from the atmosphere as well. All in all, physical loss is a greater atmospheric sink of NMVOC carbon than chemical conversion to CO and CO<sub>2</sub> in most models ( $\epsilon_C$  often <0.5; Figs. 3a and S4; also 24–26).

Atmospheric model developers must make choices that influence an individual model’s oxidative efficiency. We must balance a yet imperfect scientific understanding of atmospheric oxidation mechanisms (e.g., 27) with desired model capabilities and computational limitations. Near-explicit mechanisms using elementary reactions based on observational constraints for the oxidation of isoprene alone would require thousands of reactions and several hundred species (28), intractable for global atmospheric models. For these reasons, global models typically employ reduced chemical mechanisms with lumped surrogate species and empirically-derived stoichiometries in order to best reproduce observations within computational limits.

We highlight two end members of the ACCMIP ensemble by contrasting the hydrocarbon oxidation mechanisms of the Goddard Institute for Space Studies (GISS) and Geophysical Fluid Dynamics Laboratory (GFDL) models. In the GISS model (Fig. S5), gaseous NMVOCs are emitted as 2 explicit and 3 lumped species, transforming among 18 NMVOC species before conversion to CO. All oxidation pathways lead to CO, except for seven species (39 %) that may be physically removed, and a terminal pathway included to enable tuning of relative oxidation rates to observations (29). In contrast, the GFDL model (Fig. S6) emits NMVOCs as 10 explicit and 2 lumped species, transforms among 44 total NMVOC species, with 28 species (64 %) physically removable by wet or dry processes. Furthermore, the GFDL empirical oxidation stoichiometries include partial direct conversion of NMVOCs to CO<sub>2</sub>, bypassing CO. Given the greater opportunity for loss of intermediates, only 23 % of NMVOC emitted by the GFDL model is oxidized to CO, versus 53 % in the GISS model. Consequently, the GISS model has a greater fraction of its emitted reactive carbon reacting with OH, and thereby lower OH.

Despite large inter-model variability, the oxidative efficiency is relatively invariant in time for a given model (Fig. S7). This implies that the oxidative efficiency is primarily determined by the specific combination of species, reactions and sinks implemented in a given model. Nevertheless, minor temporal changes occur, reflecting shifts in the distribution and magnitudes of reactive carbon emissions and pre-

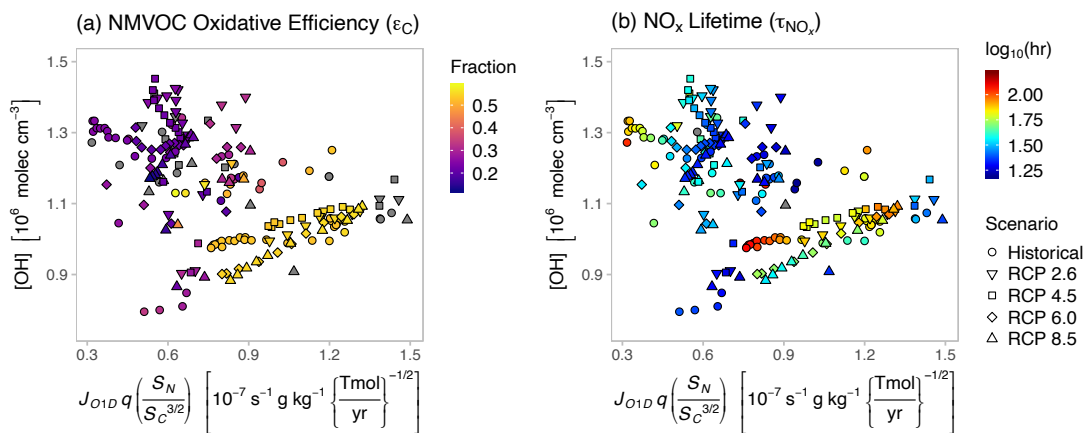
chemical mechanism, base emissions, and the response of emissions to changes in meteorology, e.g., the increase in isoprene emission by terrestrial plants at warmer temperatures (21).

Fig. 2 shows decadal mean OH in the ACCMIP simulations as a function of Eq. [1]. If differences in the four key parameters in Fig. S2 were sufficient to explain OH variability within and between the models then all points would fall on the same line with a positive slope. The OH steady-state predictor is usually sufficient to explain variability within individual models, although not in the models with the highest mean OH abundances. A large mode of residual inter-model variability dominates the ensemble variability. This inter-model variability lies largely orthogonal to that explained by the OH photochemistry, and is therefore likely independent.

Fig. 3 shows that the residual variability can be best explained by considering two new parameters derived from available ACCMIP model diagnostics. First, the large inter-model variability is primarily related to the NMVOC “oxidation efficiency,”  $\epsilon_C$ , which we define as the fraction of carbon atoms emitted in NMVOC compounds oxidized to CO before removal from the atmosphere (units of mol C per mol C). Models that have a higher fraction of their emitted carbon converted to CO tend to have lower OH (Fig. 3a). Second, departure from linearity within a single model is primarily related to shifts in the tropospheric lifetime of NO<sub>x</sub>,  $\tau_{\text{NO}_x}$  (units of time; Fig. 3b). Variations in these two parameters act to alter the magnitude of the HO<sub>x</sub> partitioning response to changes in emissions of NO<sub>x</sub> versus reduced carbon species in Eq. [1].

### Factors Driving Model Diversity

The most apparent correlation of inter-model variability in Fig. 3 is with the NMVOC oxidative efficiency ( $\epsilon_C$ ), with models having higher  $\epsilon_C$  tending to have lower OH. While the greatest sinks of OH are CO and methane, respectively, these do not vary much between the ACCMIP models by design; methane and anthropogenic CO are prescribed, and nearly



**Fig. 3.** Residual inter- and intra-model variability of Fig. 2 explained by two parameters. Panel (a) shows the points of Fig. 2 colored by their decadal mean “oxidative efficiency,” the fraction of emitted NMVOC carbon that is oxidized to CO before atmospheric removal. Panel (b) shows the points colored by their  $\tau_{\text{NO}_x}$ , i.e. the  $\text{NO}_x$  burden ( $\equiv \text{NO} + \text{NO}_2$ ) divided by its source. Gray-shaded points indicate insufficient model data archived for calculation.

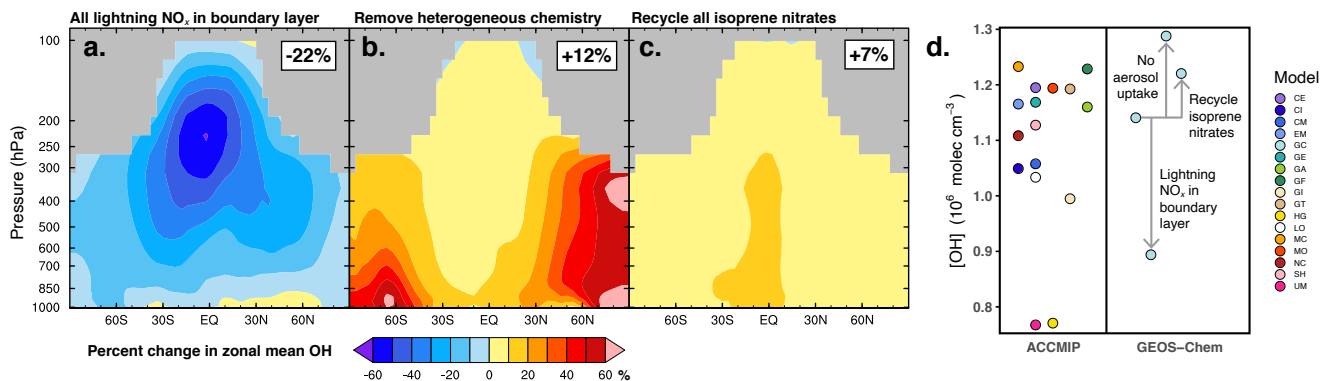
231 cipation. For example, the methane increase in the RCP 8.5  
 232 scenario consumes a larger fraction of OH, slowing the oxida-  
 233 tion rate of NMVOCs, allowing more time for intermediate  
 234 species to be lost via deposition, thereby reducing  $\epsilon_C$  in some  
 235 models.

236 In contrast to  $\epsilon_C$ , variability in  $\tau_{\text{NO}_x}$  contributes to both  
 237 inter- and intra-model variability in OH. The rapid cycling  
 238 of  $\text{NO}_x$  between  $\text{NO}$  and  $\text{NO}_2$  allows for the catalytic  
 239 production of relatively large quantities of OH (Fig. S8), other-  
 240 wise all OH would be quickly titrated from the atmosphere.  
 241 Longer  $\tau_{\text{NO}_x}$  enables relatively more OH production per  $\text{NO}_x$   
 242 molecule. Global mean  $\tau_{\text{NO}_x}$  varies by orders of magnitude  
 243 across models (Fig. 3b) and some models simulate temporal  
 244 changes by nearly as much, although others suggest little  
 245 change (Fig. S7). Therefore, variations in  $\tau_{\text{NO}_x}$  can influ-  
 246 ence both inter- and intra-model changes in global mean OH.  
 247 Fig. S9 demonstrates the sensitivity of the steady-state OH  
 248 abundance to mean  $\tau_{\text{NO}_x}$  in the GFDL model. The factors  
 249 controlling the lifetime of  $\text{NO}_x$  are many and can compete in  
 250 sign. We next demonstrate the sensitivity of global mean OH  
 251 to three key parameters with known inter-model disparities.

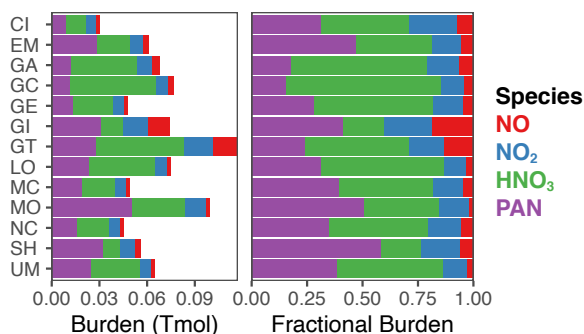
252 First, the ACCMIP models employ very different vertical  
 253 and meridional distributions of lightning  $\text{NO}_x$  (cf. Fig. 3 of  
 254 ref. 20). Because  $\tau_{\text{NO}_x}$  increases exponentially with altitude,  
 255 global mean  $\tau_{\text{NO}_x}$  is highly sensitive to the mean vertical injec-

tion height of lightning  $\text{NO}_x$ . Most models release lightning  
 256  $\text{NO}_x$  into the free troposphere. However, two ACCMIP  
 257 ensemble members (UM and HG) release their lightning  $\text{NO}_x$   
 258 primarily into the boundary layer. When we compare a sensi-  
 259 tivity simulation using the GEOS-Chem CTM with lightning  
 260  $\text{NO}_x$  released at the surface instead of at altitude, global mean  
 261 OH decreases by 22% (Fig. 4a), a sizable portion of the differ-  
 262 ence in OH between GEOS-Chem and the two ACCMIP  
 263 models releasing all lightning  $\text{NO}_x$  near the surface (Fig. 4d).

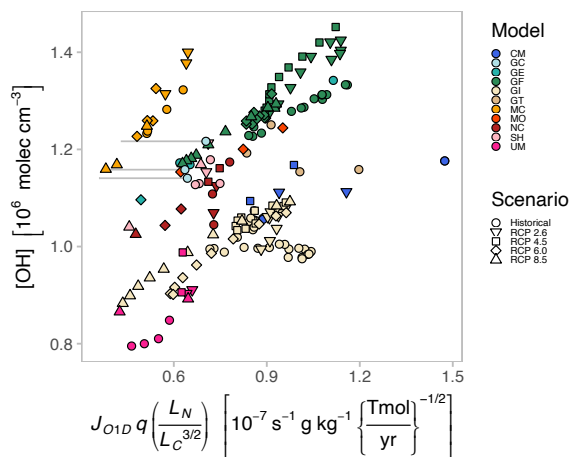
264 Second, heterogeneous reactions can vary substantially be-  
 265 tween global models. Uptake of gaseous species on the surface  
 266 of aerosol particles is a highly efficient means for sequestering  
 267  $\text{NO}_y$  and  $\text{HO}_x$  from the atmosphere, short-circuiting the  
 268 OH photochemistry included in Eq. 1. Most models include  
 269 hydrolysis of  $\text{N}_2\text{O}_5$  on aerosol particles, often the dominant  
 270 local sink for  $\text{NO}_y$ , with a known strong influence on global  
 271 OH (30, 31). However, models differ in implementation (e.g.,  
 272 some models vary the uptake coefficient  $\gamma$  with aerosol com-  
 273 position) as well as in the inclusion of other heterogeneous  
 274 losses. When we disable heterogeneous losses within GEOS-  
 275 Chem (normally  $\text{N}_2\text{O}_5$  hydrolysis with  $\gamma_{\text{N}_2\text{O}_5} \leq 0.02$  (32)  
 276 and  $\text{HO}_2$  uptake with  $\gamma_{\text{HO}_2} \leq 0.1$  (33)), global mean OH in-  
 277 creases by 12 % (Fig. 4b), the magnitude of many inter-model  
 278 differences. We note that the only difference in the gas-phase  
 279 chemistry of the two variants of the GISS model submitted  
 280



**Fig. 4.** Global mean OH is highly sensitive to the treatment of reactive nitrogen. Change in zonal mean OH in the GEOS-Chem CTM for the year 1980 for the following scenarios: (a) all lightning  $\text{NO}_x$  is released into the planetary boundary layer instead of at altitude, (b) all heterogeneous reactions are turned off, and (c) all nitrogen sequestered in isoprene nitrates is recycled back to  $\text{NO}_x$ . The box reports the percent change in the global OH burden. Panel (d) shows decadal mean OH in the ACCMIP models for the 1980s vs. that in GEOS-Chem and the three sensitivity simulations shown in panels (a)-(c). The  $x$ -axis is arbitrary.



**Fig. 5.** ACCMIP models show large disagreement in their reactive nitrogen budgets. Left panel shows the cumulative absolute burdens of the four dominant NO<sub>y</sub> species: peroxyacetylnitrate (PAN), HNO<sub>3</sub>, NO<sub>2</sub>, and NO. Decadal tropospheric mean shown for the 2000s from 13 global atmospheric chemistry models. The right panel shows the cumulative fractional burdens per model.



**Fig. 6.** Variability in decadal mean OH in the ACCMIP simulations is better captured when considering photochemical sinks of reactive nitrogen and carbon (as in this figure) than sources (as in Fig. 2). Instead of emission, we apply the total loss rate of CO and methane ( $L_C$ ), which is equivalent to the sum of the sources of CO, methane and the oxidative efficiency times the NMVOC source. Furthermore, we estimate the loss of NO<sub>x</sub> to nitric acid ( $L_N$ ). See Methods and Materials for details. Gray horizontal lines indicate the error in estimating  $L_N$  from archived monthly-mean values relative to online calculation with GEOS-Chem (not archived by the ACCMIP experiment).

281 to the ACCMIP study was that one included heterogeneous  
 282 uptake of HNO<sub>3</sub> on dust particles (GI; 34) and one did not  
 283 (GT), resulting in a longer  $\tau_{\text{NO}_x}$  and 20 % higher OH in the  
 284 latter (Fig. 4d).

285 Lastly, NO<sub>y</sub> is often stored in relatively long-lived reser-  
 286 voir species, which prolong and redistribute the influence of  
 287 NO<sub>x</sub> on oxidants. However, these reservoir species may also  
 288 be lost to deposition, which bypasses the assumption that all  
 289 NO<sub>x</sub> is lost via oxidation to HNO<sub>3</sub> intrinsic to Eq. 1. One par-  
 290 ticular area of model diversity in NO<sub>y</sub> reservoirs regards the  
 291 fate of isoprene nitrates, products of oxidation of isoprene by  
 292 NO<sub>x</sub>, especially at night. Some chemical mechanisms, includ-  
 293 ing that implemented in our version of GEOS-Chem, assume  
 294 that this nitrate is immediately lost to deposition, reflecting  
 295 observational evidence that this can be a major local NO<sub>y</sub> sink  
 296 (e.g., 35). Others allow for part of this nitrate to be recycled  
 297 back to NO<sub>x</sub>. When we perform a sensitivity test in GEOS-  
 298 Chem in which we allow all isoprene nitrates to be recycled  
 299 back to NO<sub>x</sub>, global mean OH increases by 7 % (Fig. 4c),  
 300 the magnitude of differences between GEOS-Chem and the  
 301 GFDL model (GA, GF), which included isoprene nitrate re-  
 302 cycling (albeit, <100 %; 36).

303 Disparate implementations of processes that affect  $\tau_{\text{NO}_x}$   
 304 therefore cause the NO<sub>y</sub> budget to be a key source of model  
 305 uncertainty with respect to OH. Fig. 5 shows that despite  
 306 nearly identical NO<sub>x</sub> emission magnitudes, the models show  
 307 large variability in their absolute burdens and relative parti-  
 308 tioning of the dominant NO<sub>y</sub> species. This is much greater  
 309 than seen in the longer-lived species such as ozone and CO.  
 310 Furthermore, Fig. 5 is missing other major NO<sub>y</sub> species not  
 311 archived for ACCMIP, e.g., HONO or HO<sub>2</sub>NO<sub>2</sub>, which likely  
 312 also vary between models.

313 Taken together, the sensitivity of global mean OH to  $\epsilon_C$   
 314 and  $\tau_{\text{NO}_x}$  implies that a key assumption of Eq. 1 does not  
 315 hold between models or within some models, i.e. the sources  
 316 of reactive nitrogen and carbon are not balanced by their  
 317 dominant sinks with respect to OH (oxidation to nitric acid  
 318 and carbon dioxide, respectively). Various processes siphon  
 319 emitted reactive nitrogen and carbon from the atmosphere  
 320 before they may influence OH photochemistry. It is variations  
 321 in the fraction of total nitrogen and carbon loss pathways  
 322 that involve OH that ultimately control OH abundance in the  
 323 models, and therefore the atmospheric chemistry and physics

324 that occurs between emission and loss. Fig. 6 shows that we  
 325 are better able to capture inter- and intra-model variability  
 326 in OH by shifting our focus from the sources of nitrogen and  
 327 carbon to the sinks ( $R = 0.5$  for  $n = 221$  decadal time slices).

## 328 Conclusions

329 Atmospheric chemistry model contributions to the ACCMIP  
 330 experiment strongly disagree in their global mean OH and its  
 331 transient evolution in response to prescribed changes in an-  
 332 thropogenic emissions (Fig. 1). A steady-state relationship  
 333 derived from the basic coupled photochemistry of the O<sub>x</sub>-  
 334 HO<sub>x</sub>-NO<sub>x</sub>-CO system explains OH variability within models  
 335 if the NO<sub>x</sub> lifetime ( $\tau_{\text{NO}_x}$ ) remains relatively constant (Fig. 2).  
 336 However, model OH is sensitive to inter-model differences in  
 337 the fraction of non-methane volatile organic carbon that is  
 338 completely oxidized through CO (and CO<sub>2</sub>),  $\epsilon_C$ , as well as  
 339  $\tau_{\text{NO}_x}$  (Fig. 3). The oxidative efficiency is generally fixed for a  
 340 given model, determined by the kinetics of the reactions and  
 341 species included in the model. In contrast, the NO<sub>x</sub> lifetime  
 342 is much more variable between models and even within some  
 343 models, and is especially sensitive to the altitude of lightning  
 344 emissions, heterogeneous chemistry, and treatment of reactive  
 345 nitrogen reservoir species (Fig. 4), although this list is not ex-  
 346 haustive. Variations in global mean OH between and within  
 347 global models are therefore better characterized by variations  
 348 in the relative loss pathways of reactive nitrogen and carbon  
 349 than in their emissions (Fig. 6).

350 We recommend future efforts diagnose comprehensive re-  
 351 active nitrogen budgets in models and evaluate with obser-  
 352 vations, especially in the tropical free troposphere. Likewise,  
 353 we recommend future observational campaigns target quan-  
 354 tification of the global source of CO from NMVOC oxidation.  
 355 Improved observational estimates of  $\tau_{\text{NO}_x}$  and  $\epsilon_C$  will help to  
 356 constrain models and improve projections of the future ox-  
 357 idizing power of the atmosphere.

**A. Data analyzed.** The Atmospheric Chemistry and Climate Model Intercomparison Project (ACCMIP) consisted of 16 atmospheric chemistry models driven by identical anthropogenic emissions and climate forcings over the 1850-2100 period (10). Most models were run as decadal time-slice experiments for a few core decades (1850s, 1980s, 2000s, 2030s, 2090s), although the GISS-E2-R model submitted its transient simulations for the Coupled Model Intercomparison Project Phase 5 (CMIP5). The ACCMIP and CMIP5 simulations were performed in support of the Fifth Assessment Report of the Intergovernmental Panel on Climate Change. The diagnostics saved for the ACCMIP study are archived at the British Atmospheric Data Centre (<http://badc.nerc.ac.uk>), as downloaded in Sept. 2019. Here we analyze the historical simulations and the four future RCP scenarios (9). All 3-d and 2-d variables were processed at monthly temporal resolution by integrating tropospheric totals or airmass-weighted mean values, as appropriate. The troposphere is defined with a monthly chemical tropopause of 150 ppbv ozone, consistent with earlier ACCMIP analyses (6, 7, 14). The monthly values are then averaged to obtain decadal means. We supplement the ACCMIP data with transient runs from the GFDL-CM3 contribution to the CMIP5 study that used the same emissions and forcings (37) and GEOS-Chem (next subsection).

Ozone photolysis is directly archived (*photoid*) or estimated from the archived OH production rate (*prodohj3*) using coefficients from the JPL Handbook (38). Specific humidity is archived as *hus*.  $S_N$  is archived as *eminox*.  $S_C$  is the sum of total emission of CO (*emico*), NMVOCs (*emivoc*) and methane, which we assume is equivalent to the total loss rate of methane (*lossch4*). Emission totals are adjusted to match values self-reported by the modeling groups in Lamarque et al. (10), as outlined in the SI Sect. S2. By assuming that direct emissions and chemical production of CO from methane and NMVOCs are balanced by its total chemical loss, then the oxidative efficiency can be estimated as  $\epsilon_C \equiv \frac{P(\text{CO})_{\text{NMVOC}}}{E_{\text{NMVOC}}} \approx \frac{L(\text{CO}) - E_{\text{CO}} - P(\text{CO})_{\text{CH}_4}}{E_{\text{NMVOC}}}$ . Total chemical loss of CO was archived as *lossco*, and again we use *emivoc* and *emico*. We estimate the source of CO from methane oxidation by assuming a 100 % conversion of oxidized methane from *lossch4* to CO (22).  $\tau_{\text{NO}_x}$  is determined as the total burden of NO (*vmrno*) and NO<sub>2</sub> (*vmrno2*) divided by the total source (*eminox*).

To adjust the steady-state relationship for Fig. 6, we estimate  $L_C$  as *lossco* + *lossch4*, which neglects carbon lost to intermediates. We estimate  $L_N$  as the production rate of HNO<sub>3</sub> from the pressure-dependent NO<sub>2</sub> + OH reaction using monthly archived NO<sub>2</sub>, OH, air density, and temperature, and rate constants from ref. (38).

**B. Original simulation description.** We use the default GEOS-Chem global CTM (v9-01-03; available at <http://www.geos-chem.org>) with its “tropchem” mechanism driven by MERRA reanalysis meteorological fields (39) degraded to 4° latitude by 5° longitude by 47 vertical levels (38 in the troposphere), and the same emissions used in the ACCMIP study. Methane is prescribed as a surface boundary condition and allowed to advect and react, consistent with the ACCMIP models. The base simulation is initialized over the year 1979, and then archived for 1980-2010. NO<sub>x</sub> sensitivity experiments are then performed by re-initializing over 1979 and archiving year 1980 for the 3 scenarios described in the main text.

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1

## 2 **Supplementary Information for**

### 3 **Uncertainties in global hydroxyl projections tied to fate of reactive nitrogen and carbon**

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#### 7 **This PDF file includes:**

8     Supplementary text

9     Figs. S1 to S9

10    References for SI reference citations

## 11 Supporting Information Text

### 12 S1. Derivation of the steady-state relationship

13 Here, we derive a solution of the steady-state system of fundamental OH chemistry to relate changes in simulated global mean  
14 OH to a simple function of four convolved key parameters: ozone photolysis rates, water vapor abundances, and emissions of  
15 reactive nitrogen and carbon species. This relationship was previously applied by Murray et al. (1), and is an expansion of an  
16 earlier derivation by Wang and Jacob (2). Figure S1 shows the coupling between the dominant reactions and species that  
17 influence OH abundances.

18 The primary source of HO<sub>x</sub> (≡ OH + peroxy) radicals to the troposphere is the photolysis of ozone to O(<sup>1</sup>D), followed by  
19 reaction with water vapor,



23 otherwise, O(<sup>1</sup>D) is quenched and reforms ozone,



25 Cycling of HO<sub>x</sub> species during oxidation of CO in the presence of NO<sub>x</sub> (≡ NO + NO<sub>2</sub>) leads to ozone production,



31 In addition to R5, the oxidation of NO to NO<sub>2</sub> also takes place by reaction with ozone,



33 Loss of HO<sub>x</sub> radicals is primarily by self-reaction of HO<sub>2</sub> and by oxidation of NO<sub>2</sub> by OH to nitric acid,



37 Ozone is also transported from the stratosphere and can be lost by deposition to the surface.

38 The steady state equations of the system described by reactions R1 through R9 for HO<sub>x</sub>, OH, O<sub>3</sub>, NO<sub>x</sub>, and CO concentrations  
39 are, respectively,

$$40 \quad 2 \frac{J_{\text{R1}} k_{\text{R2}} [\text{H}_2\text{O}]}{k_{\text{R3}} [\text{M}]} [\text{O}_3] = 2k_{\text{R8}} [\text{HO}_2]^2 + k_{\text{R9}} [\text{OH}] [\text{NO}_2] [\text{M}] \quad [1]$$

$$42 \quad k_{\text{R5}} [\text{HO}_2] [\text{NO}] + 2 \frac{J_{\text{R1}} k_{\text{R2}} [\text{H}_2\text{O}]}{k_{\text{R3}} [\text{M}]} [\text{O}_3] = k_{\text{R4}} [\text{OH}] [\text{CO}] + k_{\text{R9}} [\text{OH}] [\text{NO}_2] [\text{M}] \quad [2]$$

$$44 \quad k_{\text{R5}} [\text{HO}_2] [\text{NO}] + F_{\text{STE}} = \frac{J_{\text{R1}} k_{\text{R2}} [\text{H}_2\text{O}]}{k_{\text{R3}} [\text{M}]} [\text{O}_3] + k_d [\text{O}_3] \quad [3]$$

$$46 \quad S_N = L_N = k_{\text{R9}} [\text{OH}] [\text{NO}_2] [\text{M}] \quad [4]$$

$$48 \quad S_C = L_C = k_{\text{R4}} [\text{OH}] [\text{CO}] \quad [5]$$

49 where  $k_i$  or  $J_i$  is the reaction or photolysis rate constant for reaction  $i$ ,  $k_d$  is the ozone depositional loss frequency ( $\text{s}^{-1}$ ),  $F_{\text{STE}}$   
50 is the net transport rate of ozone from the stratosphere to the troposphere (molecules  $\text{cm}^{-3} \text{s}^{-1}$ ), and  $S_N$  and  $S_C$  are the  
51 source rates of NO<sub>x</sub> and CO, respectively (molecules  $\text{cm}^{-3} \text{s}^{-1}$ ). Note that the rate-limiting step for ozone production is  
52 formation of NO<sub>2</sub> from NO+HO<sub>2</sub>. We further assume that the loss of NO<sub>x</sub> ( $L_N$ ) and reactive carbon ( $L_C$ ) occur only through  
53 reactions R9 and R4, respectively.

54 To derive the dependence of OH, first we assume O<sub>3</sub>-NO-NO<sub>2</sub> photochemical steady state (photostationary assumption)  
55 from reactions (R6 -R7), i.e.

$$56 \quad J_{\text{R6}} [\text{NO}_2] = k_{\text{R7}} [\text{NO}] [\text{O}_3] \quad [6]$$

We combine equation 6 with equations 1–5, and explicitly solve the linear set of equations to obtain

$$[\text{OH}] = \left( \frac{J_{\text{R6}} k_{\text{R5}} S_N (3[\text{H}_2\text{O}] J_{\text{R1}} k_{\text{R2}} + k_{\text{R3}} k_d [\text{M}])^{3/2}}{\sqrt{2} k_{\text{R3}} k_{\text{R7}} k_{\text{R9}} [\text{M}]^2 (S_C + S_N + F_{\text{STE}})} \right) \times \sqrt{\frac{[\text{H}_2\text{O}] J_{\text{R1}} k_{\text{R2}} (2S_C - S_N + 2F_{\text{STE}}) - k_{\text{R3}} k_d [\text{M}] S_N}{k_{\text{R8}} ([\text{H}_2\text{O}] J_{\text{R1}} k_{\text{R2}} (S_C + S_N - 2F_{\text{STE}}) + k_{\text{R3}} k_d [\text{M}] (S_C + S_N))^2}} \quad [7]$$



57 Using typical tropospheric concentrations, rate constants and fluxes to identify negligible terms, eq. 7 may be reduced to a  
 58 functional form of

$$59 \quad [\text{OH}] \propto \kappa J_{\text{O}(^1D)}[\text{H}_2\text{O}] \frac{S_N}{S_C^{3/2}} \quad [8]$$

60 where  $J_{\text{O}(^1D)} \equiv J_{\text{R1}}$  is the mass-weighted mean tropospheric photolysis frequency for ozone to  $\text{O}(^1D)$  ( $\text{s}^{-1}$ ),  $[\text{H}_2\text{O}]$  is the mean  
 61 tropospheric water vapor concentration (molecules  $\text{cm}^{-3}$ ),  $S_N$  is the total  $\text{NO}_x$  emission rate ( $\text{mol N cm}^{-3} \text{s}^{-1}$ ), and  $S_C$  is the  
 62 total reduced carbon species emission rate ( $\text{mol C cm}^{-3} \text{s}^{-1}$ ), and  $\kappa$  is an effective constant that reflects the relative weighting  
 63 of the second and higher-order terms. Wang et al. (2) were the first to use a similar derivation to relate the changes in global  
 64 mean  $[\text{OH}]$  to  $S_N/S_C^{3/2}$ , in order to describe behavior of simulated OH with and without anthropogenic emissions. Murray et  
 65 al. (1) later expanded the derivation into Eq. (8) to also include the effects of climate-driven variability in water vapor and  
 66 overhead ozone abundances.

## 67 S2. Adjustments to the ACCMIP archive

68 In multi-model inter-comparison studies, native model diagnostics must be converted into standardized fields with common units  
 69 for analysis, and are therefore susceptible to definitional and unit-conversion errors. For example, *eminox* is meant to include  
 70 all anthropogenic and natural  $\text{NO}_x$  sources in units of  $\text{kg N m}^{-2} \text{s}^{-1}$ ; however, many models archived only anthropogenic  
 71  $\text{NO}_x$  and/or in units of  $\text{kg NO m}^{-2} \text{s}^{-1}$ . We have evaluated each archived variable against reasonable physical limits and  
 72 the self-reported values in the supplemental materials of the ACCMIP overview paper (3). Assumptions made to adjust the  
 73 archived values to match the requested units and variable fields or disclude fields from our analysis are outlined below.

**CESM-CAM-superfast.** Ozone photolysis frequencies (*photo1d*) were not archived for this model. Instead, we derive them from the  
 production rate of OH from ozone photolysis (*prodohjo3*),

$$J_{\text{O}(^1D)} = P(\text{OH}) \cdot \frac{3.3 \times 10^{-11} \exp\left(\frac{55}{T}\right) [\text{O}_2] + 2.15 \times 10^{-11} \exp\left(\frac{110}{T}\right) [\text{N}_2]}{2 \left(1.63 \times 10^{-10} \exp\left(\frac{60}{T}\right)\right) [\text{H}_2\text{O}][\text{O}_3]},$$

74 where  $P(\text{OH})$  is *prodohjo3* converted to molecules  $\text{cm}^{-3} \text{s}^{-1}$ ,  $T$  is the local absolute temperature in K,  $[\text{N}_2]$ ,  $[\text{O}_2]$  and  $[\text{H}_2\text{O}]$   
 75 are respectively the number densities of oxygen, nitrogen and water vapor in molecules  $\text{cm}^{-3}$  determined using archived  
 76 temperature, pressure and mixing ratios, and the rate constants are taken from the JPL 2011 recommendations (4).

77 In order to match the emissions reported by Lamarque et al. (3): we consider *eminox* to be actually archived as  $\text{kg NO m}^{-2} \text{s}^{-1}$   
 78 and without lightning  $\text{NO}_x$  or a constant  $2.8 \text{ Tg N yr}^{-1}$  of other natural sources, and for *emivoc* to be actually archived as  
 79  $\text{kg isoprene m}^{-2} \text{s}^{-1}$ .

80 **CICERO-OsloCTM2.** In order to match the emissions reported by Lamarque et al. (3), we consider all archived emissions to  
 81 not include any natural sources. In this study, we add the constant natural emission fluxes reported for these simulations  
 82 by Skeie et al. (5):  $5 \text{ Tg N yr}^{-1}$  of lightning  $\text{NO}_x$  and  $8 \text{ Tg N yr}^{-1}$  of other natural  $\text{NO}_x$ ,  $180 \text{ Tg yr}^{-1}$  of biogenic CO, and  
 83  $397 \text{ Tg C yr}^{-1}$  of biogenic NMVOCs.

84 **CMAM.** Emissions of NMVOCs are zero in this model. The chemical loss of CO for several years of the RCP 8.5 scenario is  
 85 many orders of magnitude too high, and these years are discluded from this analysis.

86 **EMAC.** In order to match the emissions reported by Lamarque et al. (3): *eminox* is considered to be actually archived as  
 87  $\text{kg NO m}^{-2} \text{s}^{-1}$ , and *emivoc* is scaled by the mass ratio of C to total mass of NMVOC assumed by the model (161/210) (6).  
 88 RCP 8.5 is not included in this analysis due to non-physical archived model air masses.

89 **GEOSCCM.** All variables used in this study are correctly archived.

90 **GFDL-AM3.** Archived *eminox* did not include lightning  $\text{NO}_x$ , which we add in this analysis.

91 **GISS-E2-R.** All variables used in this study are correctly archived.

92 **GISS-E2-TOMAS.** In order to match the emissions reported by Lamarque et al. (3), *emivoc* is multiplied by the molar mass ratio  
 93 of isoprene in grams (68).

94 **HadGEM2.** In order to match the emissions reported by Lamarque et al. (3), *eminox* is considered to be actually archived as  
 95  $\text{kg NO}_2 \text{ m}^{-2} \text{s}^{-1}$  and without lightning  $\text{NO}_x$ . One year of RCP 8.5 is not included in this analysis due to error in model level  
 96 height precluding airmass calculation.

97 **LMDzORINCA.** Historical emissions of CO or NMVOC were not archived, so decadal mean *emico* and *emivoc* are prescribed from  
 98 the values reported in Lamarque et al. (3) in this scenario.

99 **MIROC-CHEM.** In order to match the emissions reported by Lamarque et al. (3), *eminox* is considered to be actually archived as  
 100  $\text{kg NO}_2 \text{ m}^{-2} \text{s}^{-1}$  and without lightning  $\text{NO}_x$ . Lightning emissions were not archived, so decadal mean *emilnox* is prescribed  
 101 from the values reported in Lamarque et al. (3).

102 **MOCAGE.** In order to match the emissions reported by Lamarque et al. (3), *eminox* is considered to be actually archived as  
103  $\text{g N m}^{-2} \text{ s}^{-1}$ , and *emico* and *emivoc* are considered to be actually archived as  $\text{g C m}^{-2} \text{ s}^{-1}$ . Lightning emissions in 1983 and  
104 2003 of the historical scenario and 2032 of the RCP 8.5 scenario are unrealistic, and not included in this analysis.

105 **NCAR-CAM3.5.** In order to match the emissions reported by Lamarque et al. (3), *eminox* is considered to be actually archived  
106 without lightning  $\text{NO}_x$ . The *emivoc* field was not archived, so decadal mean values reported by Lamarque et al. (3) are  
107 used instead. The chemical loss rates of CO (*lossco*) and methane (*lossch4*) were treated as if they were archived as  
108  $\text{molecules cm}^{-3} \text{ s}^{-1}$ , and *lossch4* was scaled by an additional factor of 0.01 for all scenarios except RCP 6.0.

109 **STOC-HadAM3.** All variables used in this study are correctly archived.

110 **UM-CAM.** All variables used in this study are correctly archived.

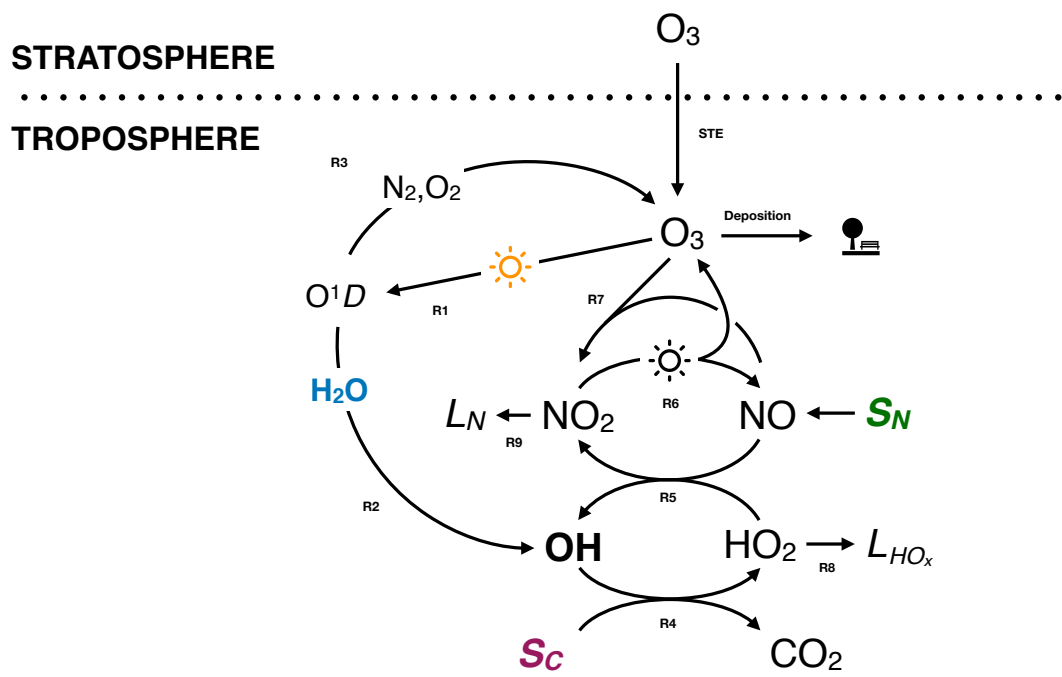
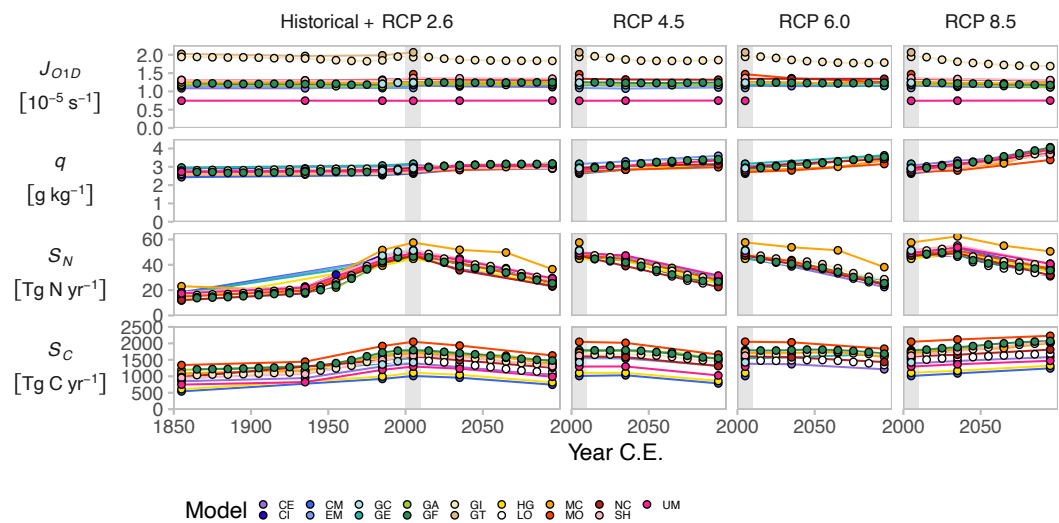
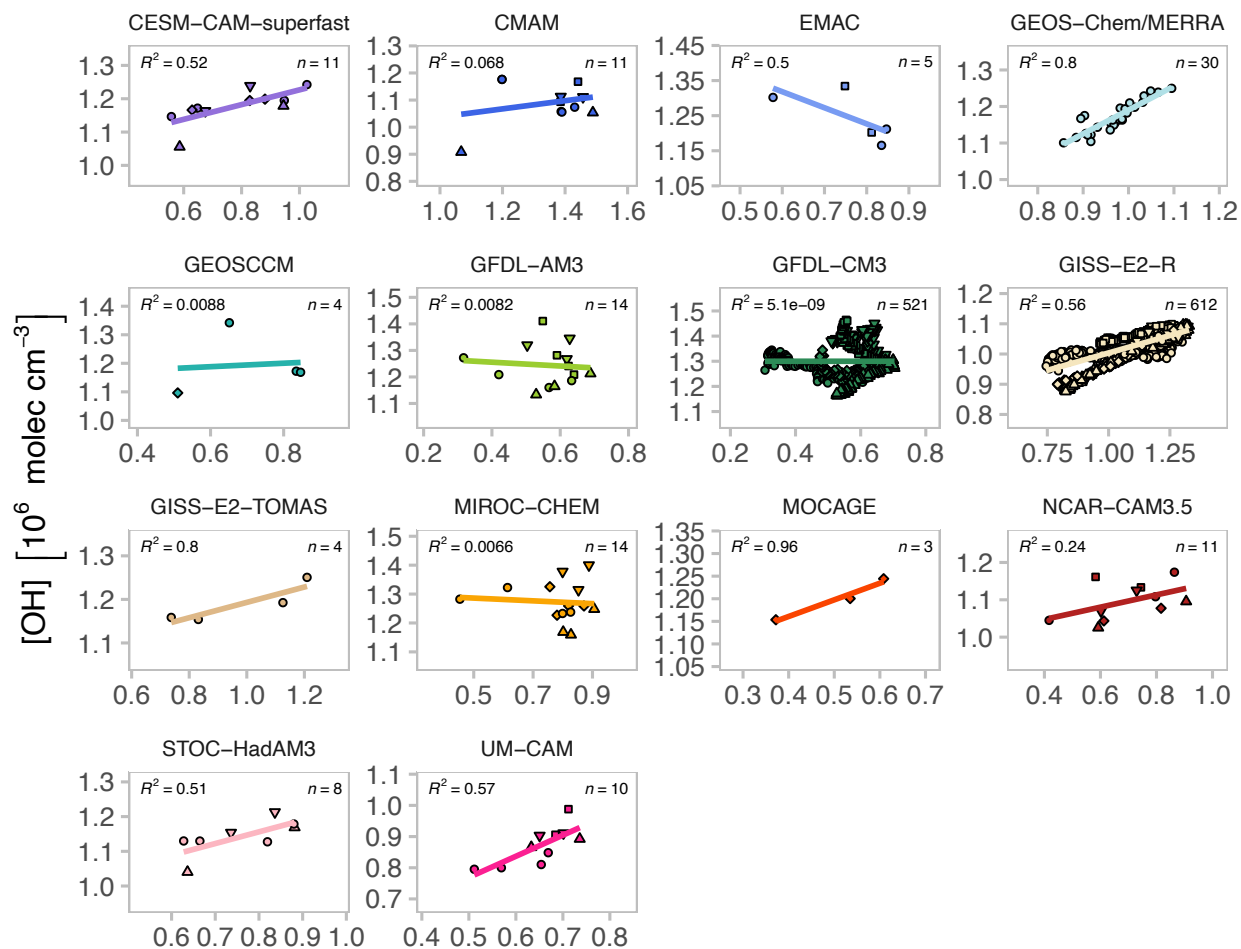


Fig. S1. Diagram of dominant chemical reactions affecting OH. The four key parameters of Eq. [1] are highlighted in colors.



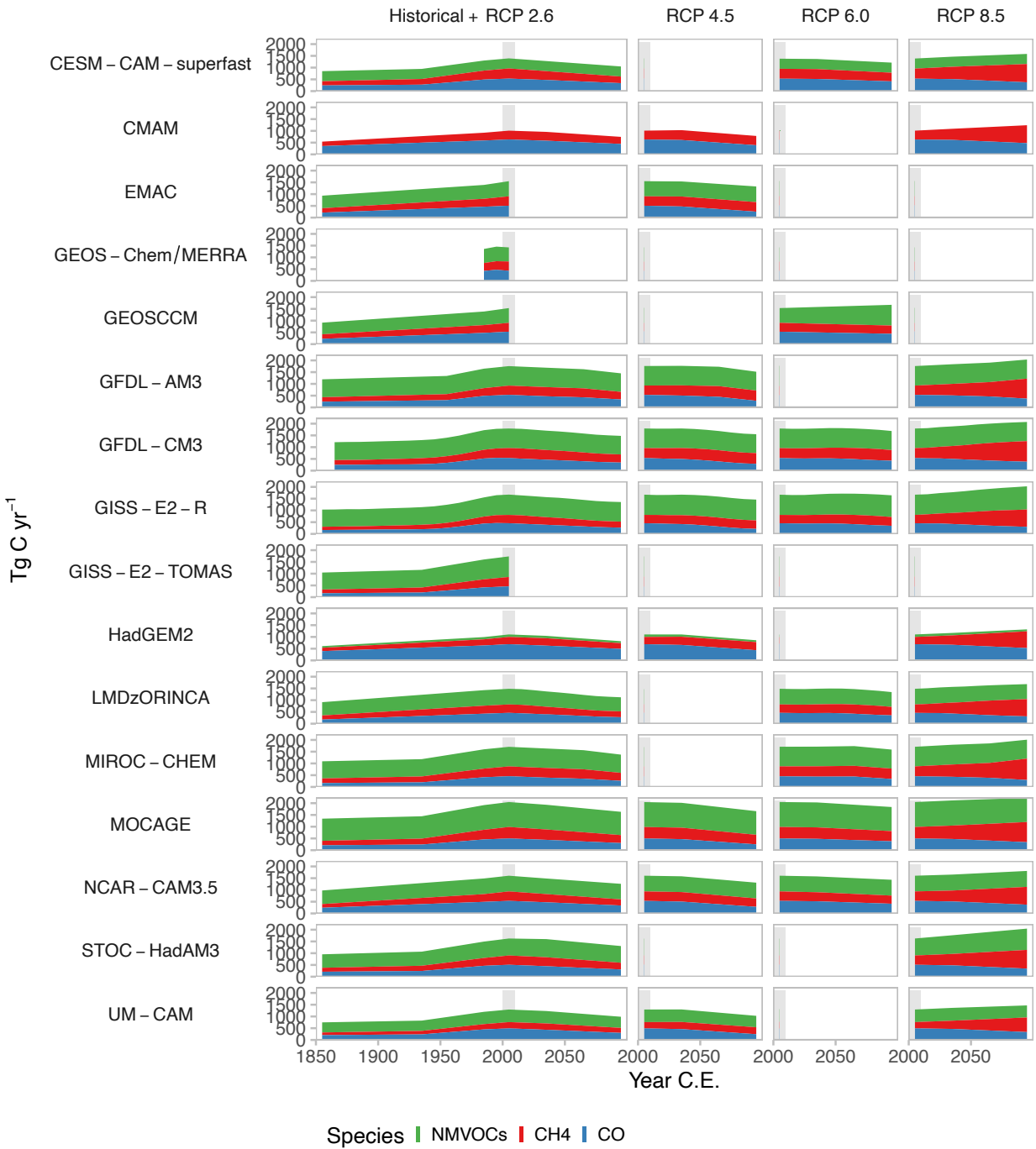
**Fig. S2.** Key parameters for OH photochemistry exhibit large temporal and inter-model variability in the ACCMIP simulations. *Top row:* Tropospheric mass-weighted mean decadal photolysis frequency of ozone to  $O(^1D)$  in  $10^5 \text{ s}^{-1}$ . *Second row:* Tropospheric decadal mean water vapor abundance in  $\text{g H}_2\text{O}$  per  $\text{kg}$  air. *Third row:* Decadal mean reactive nitrogen source in  $\text{Tg N yr}^{-1}$ . *Bottom row:* Decadal mean total reactive carbon source in  $\text{Tg C yr}^{-1}$ . Not all simulations submitted all variables for all experiments.



$$J_{O1D} q \left( \frac{S_N}{S_C^{3/2}} \right) \left[ 10^{-7} \text{ s}^{-1} \text{ g kg}^{-1} \left\{ \frac{\text{Tmol}}{\text{yr}} \right\}^{-1/2} \right]$$

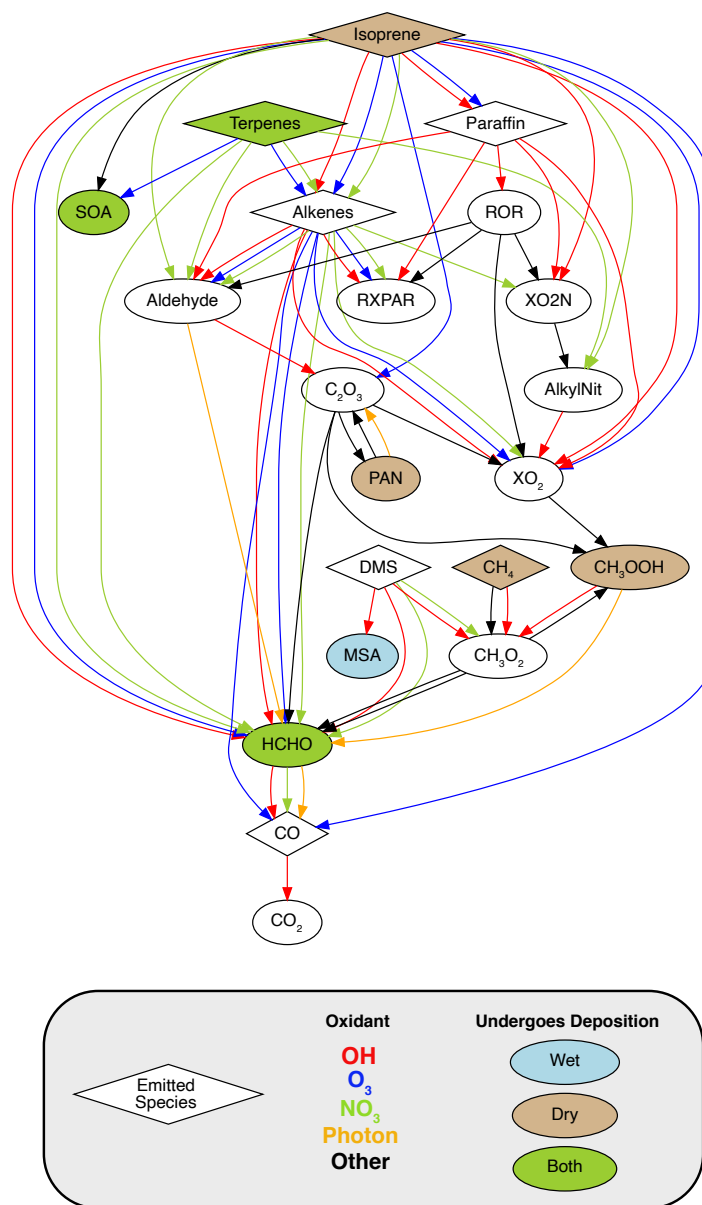


Fig. S3. Same as Fig. 2 in the main text, but with each model that contributed all four key parameters isolated, and showing annual means rather than decadal means.



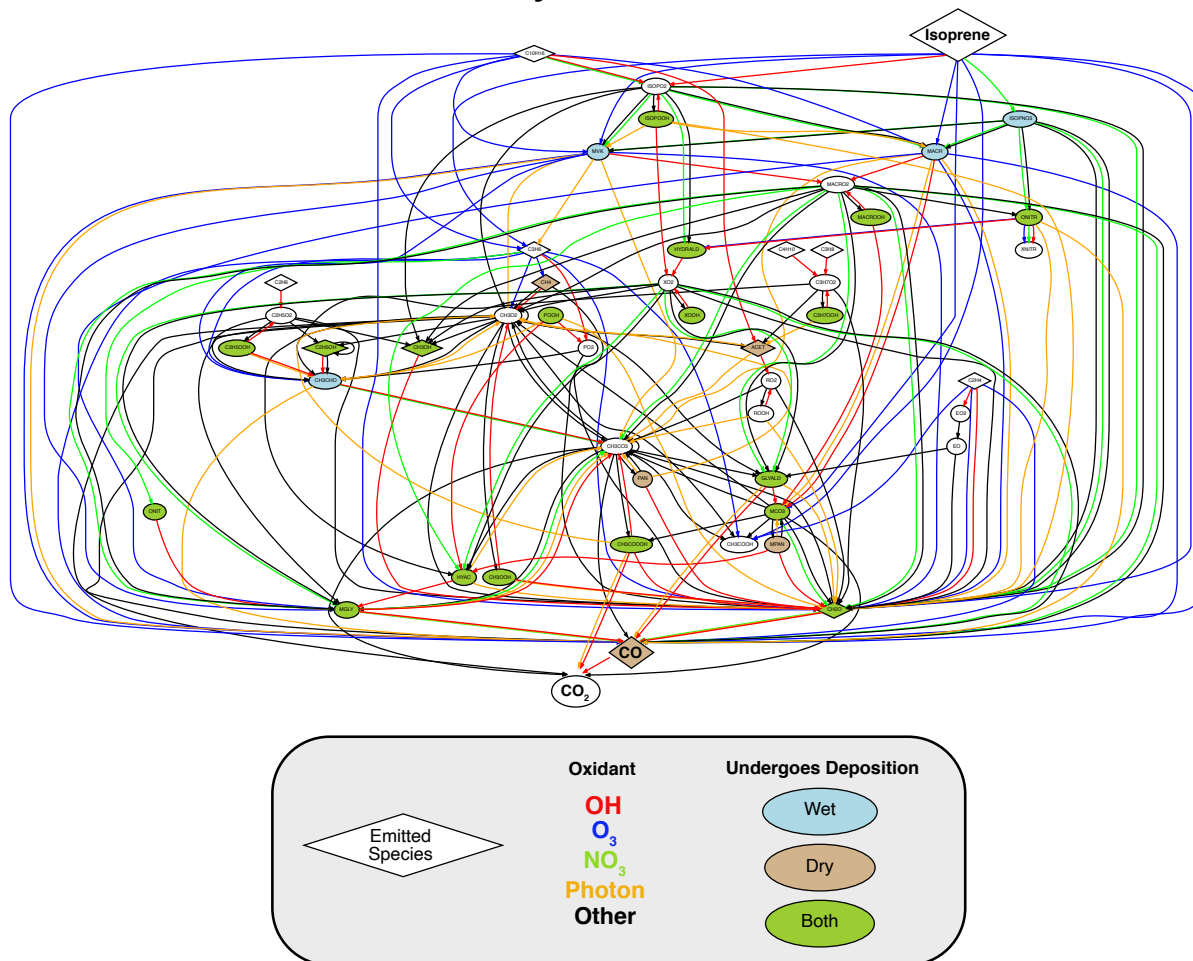
**Fig. S4.** Reactive carbon emission time series broken down by type. Stacked area charts indicate the magnitude of  $S_C$  associated with CO (blue), methane (red), and NMVOCs (green). Each row shows the time series for a different model.

## GISS-E2-R Hydrocarbon Oxidation Mechanism



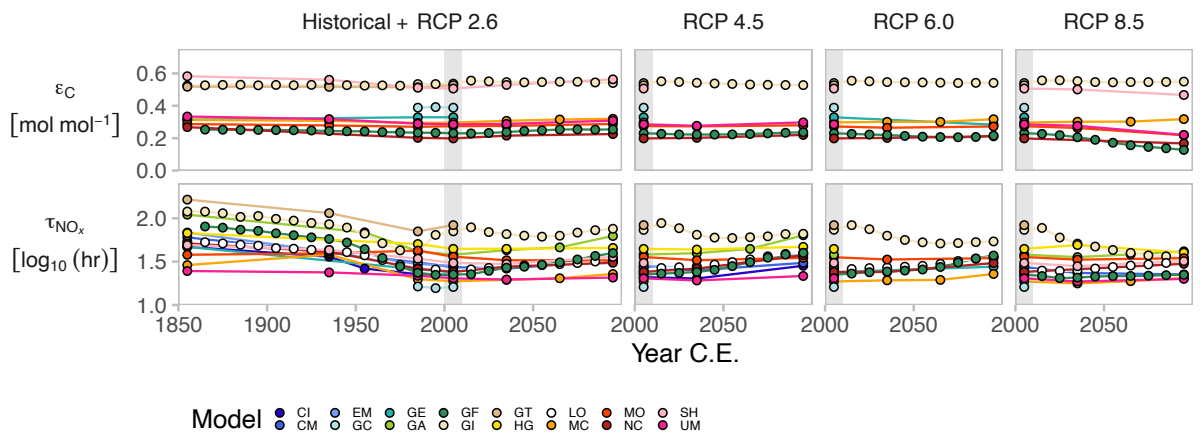
**Fig. S5.** Flow chart from emission to removal of reactive carbon species in the GISS-E2-R oxidation mechanism as described by Shindell et al. (7–10). Emitted reduced carbon species shown as diamonds. Chemical reactions are represented as directional arrows colored by oxidant (red = OH; blue = ozone; green = nitrate radical; orange = photolysis; black = other). Species that may be removed from the atmosphere via wet or dry deposition processes are shaded as light blue or tan, respectively. Those that may be removed by either are shaded as green.

## GFDL-AM3/GFDL-CM3 Hydrocarbon Oxidation Mechanism



**Fig. S6.** Same as Fig. S5, but for the GFDL atmospheric chemical mechanism as described by Horowitz et al. (11, 12). Note, dimethylsulfide (DMS) is also emitted, but only for purposes of the global sulfur budget.





**Fig. S7.** Decadal mean oxidative efficiency ( $\epsilon_{\text{ox}}$ ; top row; unitless) and the  $\log_{10}$ -transformed decadal mean  $\text{NO}_x$  lifetime ( $\tau_{\text{NO}_x}$ ; bottom row;  $\log_{10}(\text{hr})$ ) in the ACCMIP simulations.

$\text{NO}_y \equiv \text{NO}_x + \text{Reservoir Species}$

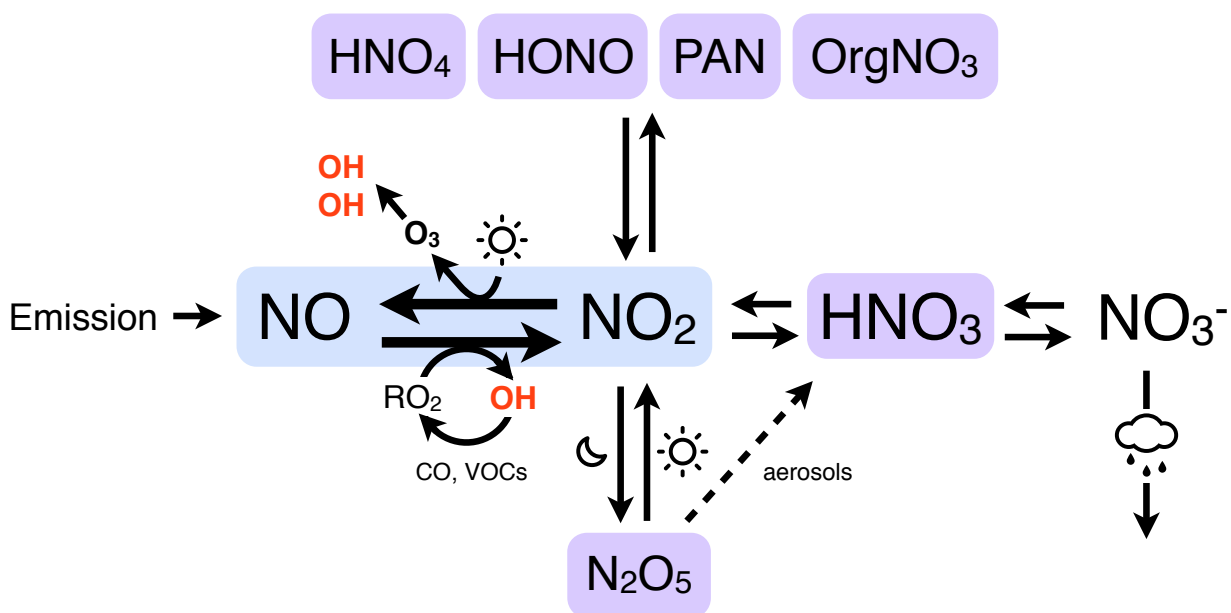


Fig. S8. Schematic of the tropospheric reactive nitrogen budget.

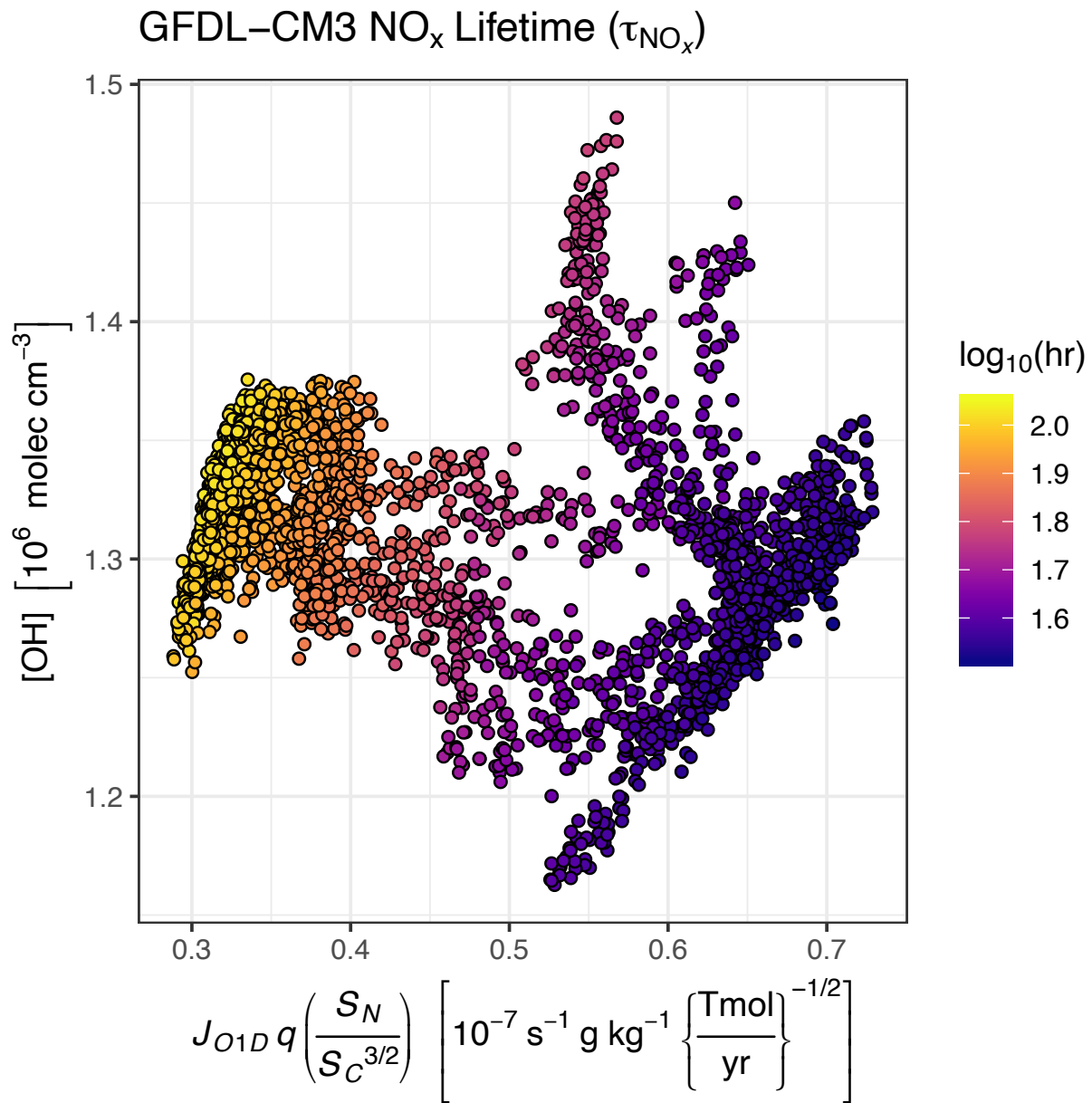


Fig. S9. Annual mean tropospheric  $\text{NO}_x$  lifetime in  $\log_{10}(\text{hr})$  of the entire ensemble of GFDL-CM3 simulations contributed to CMIP5, of which the ACCMIP runs are a subset.

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