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Key Points:

- Local volume solvers are developed in the Joint Effort for Data assimilation Integration (JEDI) software
- Localization operators for various components of the Earth system are developed
- Case studies are presented with various components of the Earth system

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Local Volume Solvers for Earth System Data Assimilation: Implementation in the Framework for Joint Effort for Data Assimilation Integration

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Abstract The Joint Effort for Data assimilation Integration (JEDI) is an international collaboration aimed at developing an open software ecosystem for model agnostic data assimilation. This paper considers implementation of the model-agnostic family of the local volume solvers in the JEDI framework. The implemented solvers include the Local Ensemble Transform Kalman Filter (LETKF), the Gain form of the Ensemble Transform Kalman Filter (GETKF), and the optimal interpolation variant of the LETKF (LETKF-OI). This paper documents the implementation strategy for the family of the local volume solvers within the JEDI framework. We also document an expansive set of localization approaches that includes generic distance-based localization, localization based on modulated ensemble products, and localizations specific to ocean (based on the Rossby radius of deformation), and land (based on the terrain difference between observation and model grid point). Finally, we apply the developed solvers in a limited set of experiments, including single-observation experiments in atmosphere and ocean, and cycling experiments for the atmosphere, ocean, land, and aerosol assimilation. We also illustrate how JEDI Ensemble Kalman Filter solvers can be used in a strongly coupled framework using the interface solver approximation, which provides increments to the ocean based on observations from the ocean and atmosphere.

Plain Language Summary The Joint Effort for Data assimilation Integration (JEDI) is an international collaboration aimed at reducing the time it takes to transition research on initialization of the Earth system models to operation. The JEDI framework is designed to be agnostic of the specific numerical model and, hence, can facilitate collaboration between research institutions and operational centers. This paper documents implementation of the Ensemble Kalman filtering framework within JEDI. The implementation strategy supports a variety of algorithmic approaches to the Ensemble Kalman Filtering and is appropriate for multiple Earth system applications. Specifically, we demonstrate applications for atmosphere, atmospheric composition, ocean, and land data assimilation.

1. Introduction

Data assimilation for multiple Earth system components simultaneously (coupled data assimilation [DA]) is an emerging field, where observations of each Earth system component (such as atmosphere, ocean, ice, land, and aerosols) are combined with a coupled simulation model to produce a statistically consistent estimate of the complete Earth system (Penny et al., 2017). Recent examples of the coupled DA applications build on a rich history of DA in each sub-component of the system. For example, the National Oceanic and Atmospheric Administration (NOAA) (Saha et al., 2006, 2010), European Center for Medium-Range Weather Forecasts (Browne et al., 2019; Laloyaux et al., 2015), and the U.S. Naval Research Laboratory (Barton et al., 2020) operational Earth system forecast models and reanalyzes are composed from legacy DA systems (solvers) for the atmosphere, ocean, and ice. These legacy DA solvers are specific to each Earth-system component and present formidable technical barriers for developing strongly coupled DA systems that treat all Earth system components as a joint statistical estimation problem in which observations of one component can correct the state estimate in the coupled components.



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Figure 1. Separation of concerns in Joint Effort for Data assimilation Integration (JEDI). Image adopted from JEDI academy slides (http://academy.jcsda.org/2021-10/slides/2021-06-21-WhyOOPSJEDI.pptx). Model implementation acronyms correspond to different model implementations and are defined in the main text of the paper. Other acronyms include: OOPS, Object Oriented Prediction System; SABER, System Agnostic Background Error Representation; UFO, Unified Forward Operator; VADER, Variable Derivation Repository; and IODA, Interface for Observational Data Access.

The Joint Effort for Data assimilation Integration (JEDI) is an international collaboration aimed at developing an open software ecosystem for model agnostic DA. Earlier efforts that developed model-agnostic DA infrastructure such as DA Research Testbed (DART; J. Anderson et al., 2009) and Parallel Data Assimilation Framework (PDAF; Nerger & Hiller, 2013) focused on acceleration of research in DA. In contrast, JEDI is developed by a consortium of operational and research weather forecast centers led by the Joint Center for Satellite DA, with the goal of accelerating transition of research results to operations. The JEDI framework achieves this goal by separating the concerns of the developers into generic blocks (see Figure 1) centered on abstract modeling of the DA problem, observational storage, covariance modeling, observation operator, and interfaces to forecast models (grouped inside of the gray box). These basic building blocks in Figure 1 can be reused and specialized by multiple models (orange boxes in Figure 1). Some of these models include atmospheric models like the Finite-Volume Cube-Sphere atmosphere general circulation model—FV3 (Harris et al., 2021), and the Model for Prediction Across Scales-MPAS (Skamarock et al., 2012); ocean models like the Modular Ocean Model version 6-MOM6 (Adcroft et al., 2019) and the Regional Ocean Model-ROMS (Haidvogel et al., 2000); sea ice models like the Los Alamos sea ice model—CICE (Hunke & Lipscomb, 2015); land surface models like the Noah land surface model with Multi-Parameterization options-Noah-MP (Niu et al., 2011), and atmospheric composition models like the FV3-based Global Ensemble Forecast System-Aerosols (GEFS-Aerosols; L. Zhang et al., 2022). Consequently, the generic DA algorithms developed using the JEDI framework can be applied to many of the model implementations currently under development within the JEDI consortium. Having this ability to use the same codebase, data structures, and algorithms for all components of the Earth system model will greatly simplify the transition to strongly coupled DA and testing of the individual components.

This paper focuses on the implementation of local volume solvers within the JEDI framework. By local volume solvers, we mean the large family of algorithms that updates the model forecast at each grid point (e.g., the green square in Figure 2) using a selection of observations in the geographic neighborhood of this grid point (e.g., the orange dots and orange circle in Figure 2). Such updates can be performed independently of each other and can scale well on modern computer architectures. The family of local volume solvers includes the Optimal Interpolation (OI) (Gandin, 1963), the Local Ensemble Transform Kalman Filter (LETKF) (Hunt et al., 2007), the Gain form of the Ensemble Transform Kalman Filter (GETKF) (Bishop et al., 2017; Lei et al., 2018), and LETKF-OI/GETKF-OI (Frolov et al., 2022) algorithms. Local volume updates are in contrast to global updates that are achieved by inverting global covariance matrices through a gradient descent algorithm employed by variational solvers (Daley, 1991).





Figure 2. Local volume solver domain decomposition. Blue grid: grid points stored on the current (this) processing element (PE). Black grid: grid points on the neighboring PE. Green square indicates the updated, analysis point x(i). Dots indicate observations: (orange) observations used in the analysis; (blue) observations that are needed to update all points on this PE; (gray) other observations that are not needed to update grid points on this PE. Circles indicate: (orange) the area used to collect local observations for the current local analysis, (blue) halo region that contains all possible observation locations needed to update grid points on this PE, (gray) observations halo region that contains all possible observation locations needed to update grid points on neighboring PE.

The objective of this paper is to present the implementation strategy for the family of the local volume solvers within the JEDI framework. Specifically, we focus on providing a unified framework that allowed us to implement LETKF and GETKF filters within the JEDI abstract (model-agnostic) layer. We discuss how abstraction of the observation and model containers presents a unique challenge to develop implementations that provide effective load and memory balancing across a variety of abstract JEDI applications such as local volume solvers, variational solvers, forecast, and observational applications. We also note that a skillful application of the DA in multiple Earth System applications (such as atmosphere, ocean, land, and atmospheric composition) requires diverse ensemble localization strategies. To our knowledge the full suite of localization strategies considered in this paper is not yet implemented in other general-purpose Ensemble Kalman Filter (EnKF) packages like DART (J. Anderson et al., 2009) and PDAF (Nerger & Hiller, 2013). Finally, we demonstrate early application of the developed solvers in the context of single observation and cycling experiments in atmosphere, ocean, land, and atmospheric composition.

In writing this paper, we faced a set of unique challenges. The JEDI ecosystem is a set of rapidly evolving, interconnected community packages hosted using the GitHub platform (https://github.com/JCSDA). During the period that this paper was developed, new features were added to the repository stack on a daily basis. At the time of the writing, all operational centers involved in the JEDI development had plans to implement some aspects of the JEDI software in operations within the next 2–5 years. Specifically, at NOAA, the next version of the Global Forecast System (GFS) will transition to JEDI-based initialization for the ocean, ice, aerosols, and land components. During the next 2–5 years, the JEDI software will be configured to be used with the full suite of Earth system observations and modeling components. JEDI will also be hardened and optimized for high-throughput and reliability required in the operational environment. At the same time, we have observed that the number of users of the JEDI system (including work described in this paper) is rapidly growing and so is the demand for a coherent overview of the development strategy for the local volume solvers presented in this paper.

In timing of this publication, we attempted to balance the rapidly evolving nature of the system with the completeness of the system. Specifically, this publication describes a sufficient set of features needed to conduct science and pre-operational trials in a diverse set of Earth system applications such as atmosphere, ocean, ice, land, and atmospheric composition. The system we describe has sufficient throughput for most non-operational



applications with the exception of the most demanding cases in global atmospheric DA. In acknowledgment that the computational performance of the system will evolve rapidly within the next few years, we avoided computational and scalability studies in this paper.

2. Mathematical Notation for the Local Volume Solver

2.1. Ensemble Kalman Filter

A generic implementation of an EnKF consists of the time update and the measurement update steps (Evensen, 2003). In the time update step, ensemble initial conditions are propagated forward in time and the mismatches between actual observations and the ensemble observation equivalents are computed, resulting in a prior estimate (the forecast) of the state and observations:

$$\begin{cases} (a): \mathbf{x}_{j}^{f(t+1)} = M\left(\mathbf{x}_{j}^{a(t)}, \boldsymbol{\eta}_{j}^{(t+1)}\right) \\ (b): \mathbf{y}_{j}^{f(t+1)} = H\left(\mathbf{x}_{j}^{f(t+1)}\right) \end{cases}$$
(1)

where: $\mathbf{x}_{j}^{f(t)}$ and $\mathbf{x}_{j}^{a(t)}$ are the forecast and analysis states for member (*j*) valid for timestep (*t*), *M* and *H* are the forecast and the observation operators; $\eta \leftarrow N(0, \mathbf{Q})$ is the model noise term generated from the zero-mean normal distribution $N(0, \mathbf{Q})$ with the model error covariance matrix \mathbf{Q} ; and $\mathbf{y}_{j}^{f(t)}$ are the observation equivalents of the ensemble member $\mathbf{x}_{j}^{f(t)}$. We note that we will use index (*j*) exclusively in this paper to indicate ensemble members as defined in Equation 1.

Outside of the forecast step in Equation 1 the time indexing (t) is the same for all of the terms in equations and, hence, it will be dropped for clarity through the rest of the paper.

From the prior forecast of observation equivalents \mathbf{y}_{i}^{f} an innovation vector is computed as:

$$\begin{cases} \overline{\mathbf{y}^{f}} = \frac{1}{N_{\text{ens}}} \sum_{j=1}^{N_{\text{ens}}} \mathbf{y}_{j}^{f} \\ \mathbf{y}^{\text{innov}} = \mathbf{y}^{\text{observ}} - \overline{\mathbf{y}^{f}} \end{cases}$$
(2)

where \mathbf{y}_{j}^{f} are the observation equivalents for the forecast of the *j*th ensemble member \mathbf{x}_{j}^{f} ; and $\overline{\mathbf{y}}^{f}$, and $\mathbf{y}^{\text{innov}}$ are the average of ensemble observation equivalents, and the innovations. N_{ens} is the ensemble size. It is common to assume that noisy observations of nature $\mathbf{y}^{\text{observ}}$ are related to the true nature state \mathbf{y}^{true} using $\mathbf{y}^{\text{observ}} = \mathbf{y}^{\text{true}} + \boldsymbol{\xi}$ through observation noise term $\boldsymbol{\xi} \leftarrow N(0, \mathbf{R})$ drawn from a normal distribution with the zero mean and observation error covariance matrix \mathbf{R} .

From Equations 1 and 2, we can also compose matrices of ensemble perturbations in the state and observational space.

$$\begin{cases} \overline{\mathbf{x}^{f}} = \frac{1}{N_{\text{ens}}} \sum_{j=1}^{N_{\text{ens}}} \mathbf{x}_{j}^{f} \\ \mathbf{Z}^{f} = \mathbf{X}^{f} - \overline{\mathbf{x}_{j}^{f}} \end{cases}$$
(3)

$$\mathbf{Y}^{f} = \begin{bmatrix} \mathbf{y}_{1}^{f}, \dots, \mathbf{y}_{N_{\text{ens}}}^{f} \end{bmatrix} - \overline{\mathbf{y}_{j}^{f}}$$

$$\tag{4}$$

where \mathbf{X}^{f} , \mathbf{Z}^{f} , and \mathbf{Y}^{f} are the matrices holding the ensemble of forecasts, ensemble of forecast perturbations, and ensemble of forecast perturbations in the observational space. The average of the forecast ensemble members in the state space and observational space are defined using the overbar notation as \mathbf{x}^{f} and \mathbf{y}^{f} respectively.



In the measurement update step of the EnKF, a posterior estimate of the ensemble members is computed through a linear combination of forecast (prior) ensemble perturbations \mathbf{Z}^{f} :

$$\begin{cases} (a) : \overline{\mathbf{x}^{a}} = \overline{\mathbf{x}^{f}} + \mathbf{Z}^{f} \mathbf{w}_{\{\mathbf{Y}^{f}, \mathbf{R}, \mathbf{y}^{\text{innov}}\}} \\ (b) : \mathbf{X}^{a} = \overline{\mathbf{x}^{a}} + \mathbf{Z}^{f} \mathbf{W}_{\{\mathbf{Y}^{f}, \mathbf{R}\}} \end{cases}$$
(5)

where $\mathbf{w}_{\{\mathbf{Y}^{f}, \mathbf{R}, \mathbf{y}^{innov}\}}$ and $\mathbf{W}_{\{\mathbf{Y}^{f}, \mathbf{R}\}}$ are the ensemble transformation vector and matrix for the mean state and the ensemble perturbations, respectively. Specific forms of the transformation matrices depend on the update algorithm and are discussed in Section 3 below. The subscripts $\{\mathbf{Y}^{f}, \mathbf{R}, \mathbf{y}^{innov}\}$ and $\{\mathbf{Y}^{f}, \mathbf{R}\}$ indicate that the weights are functions of the ensemble of forecast perturbations in the observational space, observation errors, and innovations.

The original implementation of the EnKF filters in JEDI is based on the square-root form of the EnKF, where the ensemble of analyzed states \mathbf{X}^a is obtained using a single transformation matrix ($\mathbf{W}_{\{\mathbf{Y}^f,\mathbf{R}\}}$) and a single mean innovation $\mathbf{y}^{\text{innov}}$. This is in contrast to alternative formulations of the EnKF that use perturbed observations (Evensen, 2004) or use Kalman gain matrices that differ for each ensemble member (Buehner, 2020). Such implementations are possible in the JEDI local volume solver framework; however, they were not implemented at the time this paper was written.

2.2. Local Volume Solver Approximation

In the context of the local volume solver, the measurement update step is performed for each local collection of model states $\mathbf{x}_{(i)}$ as follows. First a local innovation vector is computed by applying the selection operator $\mathbf{S}_{(i)}^{\text{observ}}$ (orange circle in Figure 2) that selects innovations that fall within the localization influence of the grid point (*i*):

$$\mathbf{y}_{(i)}^{\text{innov}} = \mathbf{S}_{(i)}^{\text{observ}} \mathbf{y}^{\text{innov}}$$
(6)

Implementation of the selection operator $S_{(i)}^{observ}$ is further discussed in Section 5.1.

Then, if *R*-localization is used (Hunt et al., 2007), the observation error covariance **R** is inflated as a function of the distance between the location of the update point loc(i) and the observation locations loc(y):

$$\mathbf{R}_{\text{local}} = \left(\mathbf{\Psi}_{\| \text{loc}(i) - \text{loc}(y) \|} \mathbf{S}_{(i)}^{\text{observ}} \right) \mathbf{R} \left(\mathbf{\Psi}_{\| \text{loc}(i) - \text{loc}(y) \|} \mathbf{S}_{(i)}^{\text{observ}} \right)^{T}$$
(7)

where $\Psi_{\||oc(i)-loc(y)\|}$ is the localization value computed as a function of the distance $\||oc(i) - loc(y)\|$ between observation location loc(y) and analysis location loc(i). We discuss different forms of the localization function in Section 5.1. Note that the *R*-localization in Hunt et al. (2007) only applies to diagonal matrix **R** in Equation 7. An alternative to the *R*-localization in Equation 7 is *Z*-localization (Kotsuki & Bishop, 2022), which attenuates columns of the **Y** matrix and can be used in presence of the non-diagonal **R** matrix. Because *Z* localization is not yet implemented in JEDI, we will use observation localization and *R*-localization interchangeably referring to Equation 7 above.

Lastly, the local update is calculated using local versions of the innovations $\mathbf{y}_{(i)}^{\text{innov}}$ and the inflated observational errors $\mathbf{R}_{(i)}$:

$$\begin{cases} \overline{\mathbf{x}_{(i)}^{a}} = \overline{\mathbf{x}_{(i)}^{f}} + \mathbf{Z}_{(i)}^{f} \mathbf{w}_{\{\mathbf{Y}_{(i)}^{f}, \mathbf{R}_{(i)}, \mathbf{y}_{(i)}^{\text{innov}}\}} = \overline{\mathbf{x}_{(i)}^{f}} + \mathbf{x}_{(i)}^{\text{inc}} \\ \mathbf{X}_{(i)}^{a} = \overline{\mathbf{x}_{(i)}^{a}} + \mathbf{Z}_{(i)}^{f} \mathbf{W}_{\{\mathbf{Y}_{(i)}^{f}, \mathbf{R}_{(i)}\}} = \overline{\mathbf{x}_{(i)}^{a}} + \mathbf{X}_{(i)}^{\text{inc}} \end{cases}$$
(8)

where $\mathbf{x}_{(i)} = \mathbf{S}_{(i)}^{\text{state}} \mathbf{x}$ are the variables local to the grid point (*i*) selected using model-state selection operator $\mathbf{S}_{(i)}^{\text{state}}$. Examples of model-state selection operator $\mathbf{S}_{(i)}^{\text{state}}$ include: all model variables associated with a single vertical



column (based on the horizontal distance, i.e., the 2D iterator in JEDI notation); or all model variables associated with a single 3D grid point (based on horizontal and vertical distance, i.e., the 3D iterator in JEDI notation). In principle, one may choose to select local points in more complex ways, for example, only temperature variables located within the oceanic mixed layer.

3. Generic Solvers

At the time of writing, there were two basic solvers implemented in JEDI: the LETKF that uses observation space localization (*R*-localization) and the GETKF that uses a combination of the model space localization (*B*-localization) in the vertical direction and observation space localization (*R*-localization) in the horizontal. LETKF-OI solver introduced in Frolov et al. (2022) discussed in Section 3.3 is a special case of the LETKF solver with artificially created ensemble perturbations.

3.1. LETKF

The LETKF filter (Hunt et al., 2007) can be written (and is implemented in JEDI) as a two-step operation. First, the local update weights $\mathbf{w}_{\{\mathbf{Y}_{(i)}^{f}, \mathbf{R}_{local}, \mathbf{y}_{(i)}^{innov}\}}$ and $\mathbf{W}_{\{\mathbf{Y}_{(i)}^{f}, \mathbf{R}_{local}\}}$ are computed and, second, the update weights are applied to rotate, scale, and translate the local prior perturbations $\mathbf{Z}_{(i)}^{f}$. The weight computation requires an inverse of the $N_{ens} * N_{ens}$ matrix. This is performed using eigen solver as follows:

$$\mathbf{A}_{\text{vec}}\mathbf{A}_{\text{val}}\mathbf{A}_{\text{vec}}^{T} \underset{eigen}{\longleftarrow} \mathbf{A} = \left(\mathbf{Y}_{(i)}^{f}\right)^{T} \mathbf{R}_{(i)}^{-1} \mathbf{Y}_{(i)}^{f} + \frac{(N_{\text{ens}} - 1)}{\rho} \mathbf{I}$$
(9)

where \mathbf{A}_{vec} and \mathbf{A}_{val} are the eigen vectors and eigen values, **I** is the identity matrix, and ρ is the forgetting (inflation) factor (which will be discussed in Section 4.1). We provide two options for this computation: one using a generic C++ Eigen library (Eigen, 2023), and one using direct calls to the LAPACK library (E. Anderson et al., 1999). After the eigen decomposition is performed, update weights are computed as follows:

$$\mathbf{W}_{\text{LETKF}} = \sqrt{N_{\text{ens}} - 1} \mathbf{A}_{\text{vec}} (\mathbf{A}_{\text{val}})^{-1/2} \mathbf{A}_{\text{vec}}^{T}$$

$$\mathbf{w}_{\text{LETKF}} = \mathbf{A}_{\text{vec}} (\mathbf{A}_{\text{val}})^{-1} \mathbf{A}_{\text{vec}}^{T} \left(\mathbf{Y}_{(i)}^{T} \mathbf{R}_{(i)}^{-1} \mathbf{y}_{(i)}^{\text{innov}(k+1)} \right)$$
(10)

The update weights are then applied to the ensemble perturbations and the ensemble mean, giving the ensemble mean increment $\mathbf{x}_{(i)}^{\text{inc}}$ and the ensemble perturbations $\mathbf{X}_{(i)}^{\text{inc}}$:

$$\mathbf{x}_{(i)}^{\text{inc}} = \mathbf{Z}_{(i)}^{f} \mathbf{w}_{\text{LETKF}}$$

$$\mathbf{X}_{(i)}^{\text{inc}} = \mathbf{Z}_{(i)}^{f} \mathbf{W}_{\text{LETKF}}$$
(11)

The full posterior ensembles mean and posterior ensemble members can be computed using Equation 8.

3.2. GETKF

The JEDI implementation of GETKF follows Lei et al. (2018), with some minor changes as described below. The GETKF solver uses the symmetric square-root \mathbf{L}_{vert} (consisting of the N_{eig} columns $\mathbf{l}_{j_{vert}}$) of the vertical localization matrix \mathbf{C}_{vert} .

$$\mathbf{C}_{\text{vert}} = \mathbf{L}_{\text{vert}} \mathbf{L}_{\text{vert}}^{T} \tag{12}$$

where j_{vert} are the indices of the eigen vectors of the vertical localization function. Vectors $\mathbf{l}_{j_{vert}}$ can either be conditioned on the background state or read from an external file.

The localization operator \mathbf{L}_{vert} is used to compute the modulated ensemble \mathbf{Z}^{mod} , by applying the Schur product \circ between each column \mathbf{l}_{ivert} in \mathbf{L}_{vert} , and each column \mathbf{z}_{i}^{f} in \mathbf{Z}^{f} :



$$\mathbf{Z}^{\text{mod}} = \mathbf{L}_{\text{vert}} \Delta \mathbf{Z}^{f} = \left[\mathbf{l}_{1} \circ \mathbf{Z}^{f}_{1...N_{\text{ens}}}, \dots, \mathbf{l}_{N_{\text{eig}}} \circ \mathbf{z}_{N_{\text{ens}}}^{f} \right]$$
(13)

The complete modulated ensemble in Equation 13 is never computed. We either compute N_{eig} modulations of a single original ensemble member $\left[\mathbf{l}_{1} \circ \mathbf{z}_{j}^{f}, ..., \mathbf{l}_{N_{eig}} \circ \mathbf{z}_{j}^{f}\right]$ during the observer step, or we compute all modulated ensemble members corresponding to a local geometry iterator $\mathbf{Z}_{(i)}^{\text{mod}}$ during the measurement-update step.

The GETKF computes two ensembles of forecast perturbations in the observational space. First, identical to Equation 4, we compute perturbations \mathbf{Y}^{f} in the observational space for the original ensemble \mathbf{X}^{f} .

Second, we compute perturbation matrix for $N_{ens} * N_{eig}$ members of the modulated ensemble \mathbf{Z}^{mod} :

$$\mathbf{Y}^{\text{mod}} = \left[H\left(\mathbf{z}_{1}^{\text{mod}} + \overline{\mathbf{x}'}\right), \dots, H\left(\mathbf{z}_{N_{\text{ens}}*N_{\text{cig}}}^{\text{mod}} + \overline{\mathbf{x}'}\right) \right] - \overline{\mathbf{y}'}$$
(14)

We compute update weights using Equations 8–10 in Lei et al. (2018) as follows. First the eigen decomposition is computed using modulated perturbation matrix \mathbf{Y}^{mod} :

$$\mathbf{A}_{\text{vec}}^{\text{GETKF}} \mathbf{A}_{\text{val}}^{\text{GETKF}} \left(\mathbf{A}_{\text{vec}}^{\text{GETKF}} \right)^{T} \underset{eigen}{\leftarrow} \mathbf{A}^{\text{GETKF}} = \frac{1}{N_{\text{ens}} - 1} \left(\mathbf{Y}_{(i)}^{\text{mod}} \right)^{T} \mathbf{R}_{(i)}^{-1} \mathbf{Y}_{(i)}^{\text{mod}}$$
(15)

Then update weights are computed as:

$$\mathbf{W}_{\text{GETKF}} = -\frac{1}{N_{ens} - 1} \mathbf{A}_{\text{vec}}^{\text{GETKF}} \left[\left(\mathbf{I} - \left(\mathbf{A}_{\text{val}}^{\text{GETKF}} + \frac{1}{\rho} \mathbf{I} \right)^{-1/2} \right) \left(\mathbf{A}_{\text{val}}^{\text{GETKF}} \right)^{-1} \right] \left(\mathbf{A}_{\text{vec}}^{\text{GETKF}} \right)^{T} \left(\mathbf{Y}_{\text{local}(i)}^{\text{mod}} \right)^{T} \mathbf{R}^{-1/2} \mathbf{Y}_{\text{local}(i)}$$

$$\mathbf{w}_{\text{GETKF}} = \frac{1}{N_{ens} - 1} \mathbf{A}_{\text{vec}}^{\text{GETKF}} \left(\mathbf{A}_{\text{val}}^{\text{GETKF}} + \frac{1}{\rho} \mathbf{I} \right)^{-1} \left(\mathbf{A}_{\text{vec}}^{\text{GETKF}} \right)^{T} \left(\left(\mathbf{Y}_{\text{local}(i)}^{\text{mod}} \right)^{T} \mathbf{R}_{\text{local}(i)}^{-1} \mathbf{y}_{\text{local}(i)}^{\text{innov}(k+1)} \right)$$

$$(16)$$

The weights are applied to compute analysis perturbations and increments as:

$$\mathbf{x}_{(i)}^{\text{inc}} = \mathbf{Z}_{(i)}^{\text{mod}} \mathbf{w}_{\text{GETKF}}$$

$$\mathbf{X}_{(i)}^{\text{inc}} = \mathbf{Z}_{(i)}^{f} + \mathbf{Z}_{(i)}^{\text{mod}} \mathbf{W}_{\text{GETKF}}$$
(17)

3.3. LETKF-OI

Frolov et al. (2022) introduced a deterministic form of the LETKF algorithm that uses *R*-localization to define a parametric error covariance model for univariate variables. In JEDI, this can be implemented by providing the LETKF with two pseudo ensemble members that are generated from a deterministic forecast step x^{f} , and the standard deviation of the background error σ^{f} as following:

$$x_j = x^f \pm \frac{\sigma^f}{\sqrt{2}} \tag{18}$$

This results in an ensemble with mean of x^{f} , and standard deviation of σ^{f} .

3.4. Coupled Solvers

To support strongly coupled DA, the JEDI framework will need to implement abstract containers for coupled model states, increments, and operators. At the time of this writing these capabilities were in the early development stages and were not yet integrated in the framework of the local volume solver. In this paper, we provide an illustration (or a proof-of-concept) for a strongly coupled DA using the interface solver approximation (Frolov et al., 2016; Sluka et al., 2016). The interface solver uses independent solvers for each component of the Earth system model. However, each of the independent solvers has access to the complete set of Earth system





Figure 3. Example of interface solver implementation in Joint Effort for Data assimilation Integration framework. A pair of Local Ensemble Transform Kalman Filter (LETKF) observers are run first obtaining ensemble observations of all atmospheric and oceanic variables. In this illustration, the near surface SST forecast (NSST) from the diurnal layer model is used as a first guess by the atmospheric observer for observations of the surface-sensitive radiances. Ocean state is used by the ocean observer to estimate the forecast of SST retrievals. Then a pair of LETKF solvers is executed, one producing analysis on the atmospheric and the second on the oceanic grids. Each solver can use a combination of atmospheric and oceanic observations. In this example, the atmospheric solver is tasked to produce the analysis of the NSST variable as it has a more complete picture of atmospheric observations that can potentially influence the NSST estimate.

observations (or at least to the relevant subset of Earth system observations). In other words, the interface solver is coupled in the observational space.

To illustrate the concept of the interface solver, we consider in Figure 3 an example of assimilation into an atmosphere, ocean, and near surface sea surface temperatures (SST) coupled model (similar to Akella et al., 2017; Frolov et al., 2020). Ensembles of coupled states consist of atmospheric, oceanic, and skin SST variables. The EnKF atmospheric observer processes atmosphere-centric observations like radiosondes and radiances (including surface sensitive radiances). The EnKF ocean observer processes ocean-centric observations (such as Argo profiles or retrievals of the ocean SST). We choose to delegate observations of surface sensitive channels to the atmospheric observer, because the atmospheric states usually have all the required information for these observations already interpolated onto the atmospheric grid. We choose to delegate processing of the SST retrievals to the ocean solver because SST retrievals are usually consistent with the ocean temperature observed by the ocean drifters at 1 m depth. Ocean temperature at 1 m depth is more consistent with the gridded oceanic state than with the skin temperature produced by the near-surface SST model. We remind readers that Figure 3 is used for the illustration of the coupled interface solver concept and how it could be implemented by users in their system. We will provide some illustrations of this implementation in Section 7.1. However, we also suggest that other users might choose to configure their interface solver differently (e.g., based on direct radiance assimilation that would include the window channels that are primarily sensitive to the SST instead of relying on the SST retrievals as illustrated in our example presented in Figure 3).

4. Available Inflation Methods

Several options for prior and posterior inflation were implemented in JEDI at the time of writing.

4.1. Prior Inflation

Prior inflation is supported in the form of a forgetting factor (Pham et al., 1998) denoted $\frac{1}{\rho}$ in Equations 9 and 16. For ρ greater than one, the prior ensemble is inflated. For values of ρ between 0 and 1, the prior ensemble is contracted. Values of ρ less or equal to zero are not admissible.

4.2. Posterior Inflation

Two options for posterior inflation are supported: relaxation to the prior perturbations (RTPP; F. Zhang et al., 2004); and relaxation to the prior spread (RTPS; Whitaker & Hamill, 2012). Both options are formulated as operation on the zero-mean ensembles of prior $Z_{(i)}^{f}$ and posterior perturbations $Z_{(i)}^{a}$ defined at a grid point (*i*).



RTPP is defined as:

$$\mathbf{Z}_{(i)}^{\text{inc'}} = (1 - \alpha_{\text{RTPP}}) \mathbf{Z}_{(i)}^{\text{inc}} + \alpha_{\text{RTPP}} \mathbf{Z}_{(i)}^{f}$$
(19)

where $\alpha_{\text{RTPP}} \in (0,1]$ is the scalar inflation coefficient, and $\mathbf{Z}_{(i)}^{\text{inc}}$ and $\mathbf{Z}_{(i)}^{\text{inc}}$ are the analyzed ensemble perturbations at grid point (*i*) before and after RTPP inflation.

RTPS is defined as a grid point operation:

$$\mathbf{Z}_{(i)}^{a'} = \left(\alpha_{\text{RTPS}} \frac{\left(\sigma_{(i)}^{f} - \sigma_{(i)}^{a}\right)}{\sigma_{(i)}^{a}} + 1\right) \mathbf{Z}_{(i)}^{a}$$
(20)

where $\alpha_{\text{RTPS}} \in (0,1]$ is the scalar inflation coefficient, $\sigma_{(i)}^{j}$ and $\sigma_{(i)}^{a}$ are the standard deviations (spread) of the prior and posterior ensemble, and $\mathbf{Z}_{(i)}^{\text{inc}}$ and $\mathbf{Z}_{(i)}^{\text{inc}}$ are the analyzed ensemble perturbations at grid point (*i*) before and after RTPS inflation.

5. Localization

5.1. Observation Space Localization for Earth System Components

Observation space localization $\Psi(\|\log(i) - \log(y)\|, \psi)$ in Equation 7 inflates observations errors for local observations as a function of some distance between the location of the local volume $\log(i)$, the location of the observation $\log(y)$, and the localization length scale ψ . It is customary to use the inverse of a correlation-like function, such as the Gaspari-Cohn (Gaspari & Cohn, 1999), that is equal to 1 at the location of the local volume and decays to zero (infinite inflation) over some support radius ψ .

When localization in more than one dimension is required (e.g., horizontal and vertical), different localization scales (e.g., ψ_{hor} and ψ_{vert}) in each dimension might be required. The total localization function can then be composed as a function of multidimensional distance in one of the two ways.

One could add all normalized distances and then compute the localization function:

$$\Psi(\||\mathrm{loc}(i) - \mathrm{loc}(y)\|, \psi_{\mathrm{hor}}, \psi_{\mathrm{vert}}) = \Psi\left(\sqrt{\frac{\||\mathrm{loc}(i) - \mathrm{loc}(y)\|_{\mathrm{hor}}^2}{\psi_{\mathrm{hor}}^2} + \frac{\||\mathrm{loc}(i) - \mathrm{loc}(y)\|_{\mathrm{vert}}^2}{\psi_{\mathrm{vert}}^2}}\right)$$
(21)

Alternately, one could assume that the localization functions are separable and compose the total localization function as a product of each individual localization:

$$\Psi(\|\text{loc}(i) - \text{loc}(y)\|, \psi_{\text{hor}}, \psi_{\text{vert}}) = \Psi\left(\frac{\|\text{loc}(i) - \text{loc}(y)\|_{\text{hor}}^2}{\psi_{\text{hor}}^2}\right)\Psi\left(\frac{\|\text{loc}(i) - \text{loc}(y)\|_{\text{vert}}^2}{\psi_{\text{vert}}^2}\right)$$
(22)

The JEDI implementation uses the second option for two reasons. First, using separable localizations allows us to keep track of distances in each dimension within the object that implements each of the localization blocks. And second, each localization block can use a different correlation function; for example, horizontal localization could use the Gaspari-Cohn function while vertical localization could use the Gaussian function.

Specifying observation localization for each observation stream independently allows the JEDI local volume solver to be flexible. For example, one could specify wider localization for radiosonde observations than for radar reflectivity observations. In the case of the coupled DA example given earlier, one could specify different localization scales for the SST, the ocean in-situ temperature, and for the atmospheric wind observations.

Finally, the observation space localization (as implemented in JEDI) allows users to limit the number of observations used in a local volume analysis (by selecting those closest to the analysis grid point). Hamrud et al. (2015) showed that selecting the closest observations allows for a significant speed up of the LETKF algorithm with no loss in the fidelity of the LETKF update.





Figure 4. Localization as a function of normalized distance and localization type. For illustration purposes, distance was scaled by a factor of 8 before the second order auto regressive (SOAR) function was calculated to achieve a near zero value for normalized distance of 1 (e.g., the SOAR function value is 0.003 for the distance of 8).

5.1.1. Horizontal, Distance-Based Localization

Several options for horizontal R-localization are supported by JEDI, including Box Car, Gaspari-Cohn, and the second order auto regressive (SOAR) function (see Figure 4). Horizontal observation localization also selects the nearest N observations closest to the updated volume.

5.1.2. Vertical Localization

The Box Car, Gaspari-Cohn, and SOAR functions are also supported for vertical localization. The vertical distance between the local volume and the observation is computed in the vertical coordinate of choice (e.g., pressure, depth, or height). An optional log transformation to the vertical coordinate is also supported before the vertical distance is computed.

To implement *R*-localization in the vertical, the vertical location of each observation must be included in the observation file, and the model interface must then return a matching vertical location for each analysis grid point. For some observation types and model grids this might not be possible. For example, satellite radiances observe integrated vertical properties, which can be difficult to relate to a specific observation height. Alternatively, many ocean-specific implementations of the Kalman filter analyze the entire vertical column of grid points, again making it difficult to assign a specific height to the analysis volume. In such cases, model space localization discussed in Section 5.2 might be a more appropriate vertical localization strategy.

5.1.3. Component-Specific Examples

5.1.3.1. Brasnett Localization for Land DA

To enable localization of land-specific observations like snow, JEDI includes the Brasnett localization (Brasnett, 1999), which attenuates the localization as a function of horizontal distance and the vertical difference in the orography of the observation location and the analyzed grid point:

$$L_{\text{Brasnett}} = (1 + c_x l_x) \exp\left[-(c_x l_x)\right] \exp\left[-\left(\frac{l_z}{c_z}\right)^2\right]$$
(23)

where L_{Brasnett} is the Brasnett correlation function, l_x is the horizontal distance between the two points in meters, l_z is the vertical (terrain) difference between the two points in meters, c_x is the horizontal length scale in m⁻¹ (default value recommended in Brasnett (1999) is 0.000018 m⁻¹, which is equivalent to a 120 km e-folding scale), and c_z is the horizontal coefficient in meters (default is 800 m). Figure 4 in Brasnett (1999) provides a nice illustration of how steep terrain in the Western United States affects correlation structures modeled by Equation 23.

5.1.3.2. Rossby-Radius Based Localization for Ocean DA

Because water has a higher density than air, the first Rossby radius of deformation for baroclinic instabilities varies latitudinally in the ocean from \sim 300 km near the equator to 3 km in the Arctic ocean (Chelton et al., 1998). This variability affects error correlation scales in eddy resolving models. To accommodate this variability, we introduced a variant of the Gaspari-Cohn function with the support radius varying as a function of the local Rossby radius as following:

$$L_{\text{Rossby}} = \min\left\{\max\left[\max\left(L_0 + m_r L_r, m_g \sqrt{a_{(i)}}\right), L_{\min}\right], L_{\max}\right\}$$
(24)

where L_{Rossby} is the support radius provided to the Gaspari-Cohn function; L_{max} and L_{min} are the user-specified maximum and minimum support radii; L_0 is the user-specified base value of the localization radius; L_r is the Rossby radius of first baroclinic deformation from Chelton et al. (1998); $\sqrt{a_{(i)}}$ is the square root of the grid cell





Figure 5. Halo distribution of observations. Shown are (blue grid) the grid points stored on this processing element (PE). (Blue triangle) halo center computed as a center of the grid points. (Blue circle) patch radius R_{patch} that encloses all the grid points on this PE; (Orange circle) halo size radius R_{halo} size that is added to the patch circle. In practice, $R_{\text{halo} size}$ is related to the localization radius. (Green circle) the total halo circle that is computed as a summation of $R_{\text{patch}} + R_{\text{halo} size}$. Observations can then be divided into patch observations that exclusively belong to this PE based on the shortest distance to the patch center (the blue dot). Halo observations (green dots) that are within the green total halo radius. And observations that are not stored on this PE because they are outside of the total halo radius (gray dots).

area at the location of the analysis volume; and m_r and m_g are the userspecified multiples. Equation 24 provides a lot of flexibility to model localization radius L_{Rossby} . The default values for Equation 24 are $L_0 = 0$, $m_r = 1.0$ and $m_g = 1.0$, $L_{\text{max}} = \text{inf}$, and $L_{\text{min}} = 0.0$.

5.2. Model Space Localization

The GETKF solver described in Section 3.2 supports model-space localization in the vertical direction, similar to the original description in Lei et al. (2018). We expect that future implementations might also support model space localization in the horizontal direction, similar to the MLGETKF solver introduced by Wang et al. (2021).

Currently several options are supported for specifying the vertical localization matrix. One option is for the model interface to return the vertical coordinate of the model; then a correlation matrix can be computed based on the localization scale specified in this vertical coordinate. Examples of the vertical coordinates include model level numbers and log pressure. Another option is for the GETKF solver to read the square root of the horizontally-varying vertical localization matrix from disk (one model increment file for each column of the square root). These files can be generated in several ways. For example, one can compute the correlation matrix from a random sample drawn from the parametrized form of the vertical localization operator used in the ensemble variational solver. Alternatively, a user can implement their own vertical localization operator that can output the square root of the vertical correlation as a collection of increment files.

6. Parallel Data Distribution

6.1. Model-Space Data Distribution

Model implementations (lower left block in Figure 1) implement the Message Passing Interface distribution of the model states in JEDI. These distributions can vary depending on the model implementation.

6.2. Observation-Space Distribution

Several options for parallel distribution of observations are available in JEDI at the time of this writing:

- Round robin distribution, which randomly distributes observations across all available processing elements (PEs) to optimize the balance of observations across PEs.
- Inefficient distribution, which replicates all observations on each processor.
- Halo distribution, which stores overlapping sets of observations on each PE. This distribution assumes that the model grid is distributed across PEs in tiles (blue and black grids in Figure 2). The overlapping observation sets are described by a user-specified circle centered on the geometric center of the local model tile (large blue circle centered on the blue tile in Figure 2 or blue triangle in Figure 5). The default halo radius (halo size) is set large enough to encompass all observations on Earth, which reduces the Halo distribution to the Inefficient distribution. Figure 5 shows the details of how Halo distribution is computed. In practice, the Halo size is selected to be just large enough to contain observations that are necessary to update all grid points on a PE.

The current implementation of the solver step of the local volume solvers (Equation 5) assumes that all observations that are required to make an analysis update at a particular model point are available on the PE owning this model point, and no communications are needed. Thus, only Inefficient or Halo distributions can be used for the solver step. Because the solver has to store the ensemble of observations, Inefficient distribution can run out of memory available to each PE on the computational node for large ensembles and observations sizes. Using Halo distribution is highly recommended for global observational counts above 1 million in the GETKF solver.



Tabla	1
Table	1

Summary of the Model and Data Assimilation Configurations Used in Section 7

						Localization	
Figure	Experiment label	Model	N _{ens}	Solver	Inflation	Horizontal	Vertical
Figures 6 and 7	No loc.	UFS p5	80	LETKF	None	N/A	None
Figures 6 and 7	GETKF	UFS p5	80	GETKF	None	N/A	Model-space
Figure 8		UFS p5	80	LETKF	None	N/A	None
Figures 9 and 10	JEDI	GFSv16	80	GETKF	RTPS	R-loc (Gaspari-Cohn)	Model-space
Figures 11 and 12	LETKF	MOM6	30	LETKF	RTPS	R-loc (Rossby)	None
Figure 13	LETKF-OI	NOAH-MP	3	LETKF	None	Brassnett	
Figure 14	LETKF	GEFS-Aerosols	20	LETKF	Multiplicative RTPS RTPP	R-loc (Gaspari-Cohn)	None
Figure 14	GETKF	GEFS-Aerosols	20	GETKF	Multiplicative RTPS RTPP	B-loc (Gaspari-Cohn)	Model-space

In contrast to the solver step, the observation step of the local volume solvers (Equation 1) uses any of the distributions listed above. For the observation step, the round robin distribution provides the best load balancing between different PEs, and has the best time-to-completion because it uses non-overlapping distributions of observations. Inefficient distribution provides the longest time to completion, doesn't scale with the number of PEs, and is impractically slow for observation counts on the order of 1 million. Halo distribution provides a practical balance between the two options.

The current implementation of the local ensemble solver in JEDI also allows the use of different observation distributions during the observer (Equation 1) and the solver (Equation 5) steps. This is accomplished by executing the EnKF application in the observer-only mode (e.g., using the most efficient round robin distribution for the observer step), followed by the execution of the EnKF application in the solver of the solver step). The redistribution of the observations between the two runs is accomplished by writing and reading of the intermediate observer files. Future implementations of JEDI might support this observation space redistribution in memory without the need to dump intermediate files to disk.

7. Case Studies

To illustrate usage of the developed local volume solvers in realistic applications, we provide several case studies that involve assimilation of the atmospheric, oceanic, land, and atmospheric composition measurements. A summary of the experimental configurations is provided in Table 1. These case studies include assimilation of single observations that illustrate how information is spread from observation location to grid points. We also present limited cycling experiments that compare forecasts from the local volume analysis against a combination of other reference systems (e.g., variational or systems with no DA).

7.1. Single Observation Experiments

To demonstrate the impact of localization in the coupled ocean-atmosphere system, we generated an 80-member ensemble using a coupled Unified Forecast System (UFS) model at 1° nominal resolution (UFS, 2023). For this section we used synthetic observations that allowed us to co-locate observation in areas of large ensemble spread and to set the size of the observation error to be equal to the size of the ensemble spread.





Figure 6. Increment to the vertical column of atmospheric temperature (a) and layer pressure (b) from assimilation of single surface pressure observation at 40.5°N, 160.5°E and innovation of 6 hPa. Background ensemble valid for 2015-12-05 18:00Z. Gain form of the Ensemble Transform Kalman Filter (GETKF) vertical localization was used with localization scale of 30 levels and 10 eigen vectors capturing 96% of the variance.

Figures 6 and 7 illustrate the impact of model space localization on spreading of the innovation vertically in the atmosphere from assimilation of a single surface pressure observation (Figure 6) and a lower-troposphere temperature observation (Figure 7).

In Figure 6, the surface pressure observation is localized by modulating the 3D pressure field, instead of prescribing localization directly between surface pressure and the rest of the 3D state. Figure 6 shows that the surface pressure is more correlated with the mid- and high-tropospheric temperatures and layer pressures than with the temperature and layer pressure in the boundary layer directly above the surface. This finding is intuitive as surface pressure measures the integral of the vertical column and is sensitive to the location of the large-scale weather systems that tend to locally displace the tropopause, resulting in areas of high- and low-pressure systems. In contrast, Figure 7 shows that an observation of the near surface temperature has an immediate impact on the



Figure 7. Increment to the vertical column of atmospheric temperature (a) and layer pressure (b) from assimilation of single temperature observation at 40.5°N, 160.5°E, 950 hPa, and innovation of 1 K. Background ensemble valid for 2015-12-05 18:00Z. Gain form of the Ensemble Transform Kalman Filter (GETKF) vertical localization was used with localization scale of 30 levels and 10 eigen vectors capturing 96% of the variance.





Figure 8. Increment to the vertical column of ocean temperature from assimilation of a single atmospheric temperature observation at 27.5°S, 154.5°W, 950 hPa (red) and sea surface temperature observation at the same horizontal location (blue). Both with the innovation of 1 K. Background ensemble valid for 2015-12-05 18:00Z. No vertical localization was used.

temperatures in the atmospheric boundary layer and on the level pressure in the boundary layer and aloft at about 500 hPa.

Applications in the strongly coupled framework—where observations in any fluid can directly increment the initial conditions in the coupled fluid—are still in the early proof-of-concept stage. For example, Figure 8 illustrates the impact on the ocean temperature profile from assimilating a single sea surface temperature (SST—blue line) and a single atmospheric surface temperature (AST—red line) observation. No vertical localization is used for either the within-fluid (SST-ocean) or the across-fluid (AST-ocean) assimilation. Both SST and AST assimilation have a large impact on the ocean temperatures in the top 100 m of the water column. Overall, both increments warm and deepen the ocean mixed layer.

The increments presented in Figure 8 illustrate the cross-fluid assimilation capabilities of the JEDI local volume solver, but we caution readers from over-interpreting the scientific merit of these increments. For example, unlike the SST increment which is confined to the top 100 m, the AST increment has a strong negative increment between 300 and 400 m deep. We speculate that this deep increment is likely due to spurious correlations between the ocean and the atmosphere and should be attenuated with an appropriate cross-fluid localization. We also note that within the mixed

layer, the magnitude of the AST increment is larger than the SST increment. We attribute this to the poorly calibrated ensemble spread in the coupled ensemble, where ocean temperatures are under spread compared to the atmospheric temperatures.

7.2. Cycling Experiments

This section presents limited results that illustrate cycling of the local volume solvers in four distinct application areas: atmosphere, ocean, land, and aerosol assimilation. When appropriate, we compare cycling results for the local volume solvers against other systems (such as variational solvers, an alternative implementation of the same



Figure 9. Root-mean-square error (RMSE) for the 6-hr atmospheric forecast computed for a complete set of in-situ measurements. Statistics were accumulated for a Gridpoint Statistical Interpolation (black line) and Joint Effort for Data assimilation Integration (blue line) cycling system with Gain form of the Ensemble Transform Kalman Filter solver from 2022-01-03 to 2022-01-11. This configuration used 80 ensemble members at 1° atmospheric resolution, assimilating all atmospheric radiosonde observations and radiances from the National Oceanic and Atmospheric Administration 19 satellite, and used relaxation to the prior spread inflation of 0.85.

algorithm, or runs without DA). Several results presented in this section were obtained from early engineering studies focused on demonstrating the feasibility of cycling with a family of JEDI solvers. We expect that, as these systems mature, separate scientific evaluations will be published by the respective research groups fully detailing scientific findings from their cycling experiments.

7.2.1. Atmosphere

Figures 9 and 10 demonstrate that the JEDI implementation of the GETKF solver compares favorably to the reference implementation of the GETKF algorithm in the Gridpoint Statistical Interpolation (GSI; Lei et al., 2018) software. GSI GETKF solver is currently used operationally by NOAA in version 16 of the GFS and version 12 of the GEFS. Specifically, Figure 9 shows that the first guess root-mean-square error (RMSE) for the GSI and JEDI runs are almost identical when compared against in situ measurements of temperature. Figure 10 further illustrates that the ensemble-mean temperature increments produced by both systems are qualitatively similar. It is likely that small differences between the increment from the GSI (top panel) and JEDI (bottom panel) in Figure 10 are due to differences in the analysis grids (Gaussian for the GSI and cube-sphere for JEDI).

Both GSI and JEDI results in Figures 9 and 10 were produced with similar DA configurations. However, the configuration we used in Figures 9 and 10 is different from operations in a number of significant ways. First, we reduced ensemble resolution from $1/4^{\circ}$ (in operations) to 1° (both GSI and JEDI in our





Figure 10. Ensemble-mean temperature increment valid for 2022-01-04 00:00Z for Gridpoint Statistical Interpolation (GSI) (top panel) and Joint Effort for Data assimilation Integration (JEDI) (bottom panel) for model level 76 (~500 hPa).

experiments). Second, we drastically reduced the observing network. In our experiments, both GSI and JEDI only assimilate radiosonde temperature, humidity and wind observations; and AMSU-A radiances from the NOAA 19 satellite. A large group of scientists are currently working on configuring the JEDI system to use the rest of the observing system correctly; however, their effort is outside of the scope of this article. Finally, the GSI system in our experiments was configured to use localization in the 3D pressure space (similar to Section 7.1) instead of direct localization of the surface pressure (used by GSI in operations).

7.2.2. Ocean

In this section, we use the LETKF with 30 members and a small representative subset of marine observations and compare its performance to the 3DVar DA scheme using the JCSDA's Sea-Ice, Ocean and Coupled Assimilation (SOCA) framework for the MOM6 Hurricane North Atlantic domain. This regional study uses the SOCA-





Figure 11. Root-mean-square error for the 24-hr forecast of (a) sea surface temperatures (SST) and (b) absolute dynamic topography (ADT) observations. Colors indicate: (black) no data assimilation, (green) 3DVAR algorithm, and (red) Local Ensemble Transform Kalman Filter (LETKF).

SCIENCE code repository (JCSDA, 2023f) to cycle the experiments. The experiments are cycled every 24 hr. The horizontal resolution of the model is 1/4°, and it has 75 vertical levels. The DA schemes assimilate absolute dynamic topography (ADT) and SST only. The assimilated ADT observations include data from CRYOSAT-2, JASON-3, Saral/ALTIKA, Sentinel-3A, and Sentinel-3B. The level 4 AVISO ADT from Copernicus Marine Services is used to estimate forecast errors but is not assimilated. The SST observations are from the Visible Infrared Imaging Radiometer Suite (VIIRS) instrument onboard the Suomi National Polar-Orbiting Partnership (Suomi NPP) spacecraft. The LETKF here uses Rossby radius-based localization scales, which were found to be better (not shown) compared to the Gaspari-Cohn method.

The results indicate that the LETKF is performing as expected, and even with minimal tuning it is able to outperform the existing 3DVAR (Figure 11). Both DA schemes lead to better performance overall than the no DA case, however, the 3DVAR does show degradations in certain areas near the western boundary currents that the LETKF does not suffer from (Figure 12). This is particularly noticeable in the SST observation departure (O-B) RMSE along the Equatorial currents (Figures 12a and 12b) and in the ADT O-B RMSE along the Gulf Stream (Figures 12c and 12d). This is likely due to the fact that the LETKF is able to apply the velocity increments while 3DVAR currently does not, leading to overall better-balanced state estimation by LETKF. These findings demonstrate that the JEDI-LETKF is both technically and scientifically capable of further development for ensemble ocean DA.

7.2.3. Land

Figure 13 demonstrates the application of the LETKF-OI to assimilate snow depth observations into NOAA's landsurface model, Noah-MP. Point-based snow depth observations from NOAA's Global Historical Climatology Network Daily (GHCN; Menne et al., 2012) are assimilated once daily into an offline (land-only) version of the





Figure 12. Root-mean-square error for the difference between run with no data assimilation and 3DVAR (left column, panels a and c) and Local Ensemble Transform Kalman Filter (LETKF) (right column, panels b and d). Top row shows differences for sea surface temperatures (SST) (panels a and b) and the bottom row shows differences for absolute dynamic topography (ADT) (panels c and d).

Noah-MP model, from 1 September 2019 to 1 May 2020. The model is run at 1° resolution, and forced with archived output from NOAA's GDAS/GFS system. As expected, the LETKF-OI effectively improves the fit between the forecasts and observations, reducing the mean daily O-F standard deviation from 72.0 to 51.3 mm, and reducing the mean daily O-F mean from -14.1 to 0.2 mm. The snow depth LETKF-OI is being prepared for implementation in the NOAA operational system. This system is based closely on the OI-based snow analyses in use at other weather prediction centers (Brasnett, 1999; de Rosnay et al., 2012; Gichamo & Draper, 2022), and will assimilate both station snow depth observations and remotely sensed snow cover observations.

7.2.4. Atmospheric Composition

A global aerosol DA system for the NOAA Global Ensemble Forecast System—Aerosols (GEFS-Aerosols; L. Zhang et al., 2022) was recently developed based on the three-dimensional ensemble-variational (3D-EnVar) framework in JEDI. In this system, the aerosol optical depth (AOD) that represents the total amount of aerosols in a column was assimilated to constrain aerosol concentrations in the GEFS-Aerosols initial conditions. Briefly, the



Figure 13. Root-mean-square error for the 6-hr forecast of the snow depth (GHCN network). Colors indicate: (black) no data assimilation, and (red) LETKF-OI assimilation of snow depth measurements. Line style indicates: (solid) standard deviation of error (dashed) mean error.

one-member control aerosol analysis was obtained by the 3D-EnVar solver in JEDI that used the pure ensemble-estimated background error covariance for aerosols. The resolution of the model was C96L64 with 20 members in the ensemble. The background ensemble was updated by the LETKF or GETKF solver as described in this manuscript. Detailed system descriptions and evaluations were documented in Huang et al. (2023). In this section, we present the initial AOD assimilation application using the 3D-EnVar solver in JEDI in combination of the LETKF and GETKF solvers (hereafter referred to as EnVar-L and EnVar-G, respectively). Two 6-hourly cycled AOD assimilation experiments that employed EnVar-L and EnVar-G, respectively, were performed on 1-30 June 2016 to assimilate within GEFS-Aerosols the VIIRS 550 nm AOD retrievals produced by the National Environmental Satellite, Data, and Information Service at NOAA (Jackson et al., 2013). The AOD assimilation impacts were investigated by comparing with another cycled experiment without AOD assimilation (hereafter referred to as NODA) in the same period. For further details on the experiment configuration, we refer readers to the description of the retrospective experiment in Section 4 of Huang et al. (2023).





Figure 14. (a) Time-series of 550 nm aerosol optical depth (AOD) means from Visible Infrared Imaging Radiometer Suite (VIIRS) (gray), no data assimilation 6-hr forecast (black), the one-member control and ensemble mean analyses in the EnVar-L (blue and green, respectively) and EnVar-G (red and orange, respectively) experiments in a 6-hourly interval in June 2016, and their corresponding (b) biases and (c) root-square-mean errors (RMSEs) against VIIRS AOD retrievals. (d–f) Are the temporally means of the corresponding (a)–(c) over the last 3 weeks in June 2016.

Figure 14 compares simulated AOD from the NODA 6-hr forecast, the one-member control and ensemble mean analyses in the EnVar-L and EnVar-G experiments against assimilated VIIRS AOD retrievals. As expected, compared to the NODA 6-hr forecast, both the control and ensemble mean analyses in the two AOD assimilation experiments showed significantly improved agreement with VIIRS AOD retrievals in terms of AOD value, biases and root-mean-square errors (RMSEs), suggesting the substantial benefits of assimilating VIIRS AOD retrievals within GEFS-Aerosols. In these two AOD assimilation experiments, the one-member control analyses obtained from EnVar showed slightly lower biases and RMSEs than their corresponding ensemble mean analyses produced by LETKF and GETKF, respectively. In addition, the biases and RMSEs in the one-member control analysis in the EnVar-L experiment were marginally lower than those in the EnVar-G experiment, while their ensemble mean analyses were comparable. The differences of the resulting control and ensemble mean analyses in these two AOD assimilation experiment to their specific localization strategies as discussed in Huang et al. (2023). Further tuning of localization length scales and inflation factors in EnVar-L and EnVar-G will be performed to optimize their assimilation performance in the future.

8. Summary, Conclusions, and Future Work

This paper describes the implementation of the local volume solvers in the JEDI framework. We found that EnKF implementations that target Earth system assimilation need to support a wide variety of ensemble localization strategies, as appropriate for the different components of the Earth system. For example, model space localization (GETKF) was requested by the atmospheric and aerosol communities, observation-space localization (LETKF) with localization distance condition on the Rossby radius of deformation was requested by the oceanic community, and terrain-dependent localization was requested by the land-modeling community. Our preliminary demonstration of a strongly coupled DA example (Figure 8) demonstrates that further work is needed to develop, implement, and test methods for cross-fluid localization operators.

Another challenge specific to a generic DA framework such as JEDI, is efficiently distributing computations and memory across PEs. To our knowledge, EnKF-specific implementations like DART have engineered their PE distributions to be optimized specifically for the EnKF computations, by distributing both the model grid points and observations randomly across the PEs (method 4.a of J. L. Anderson and Collins (2007)). This ensures good load balancing at the expense of inter-processor communication. By contrast, the JEDI framework supports a number of applications with different load balancing requirements, such as forward model simulations, observation operators, variational solvers, and EnKF solvers. At the time of writing, our implementation inherited grid point distributions from the model interfaces (e.g., cubed-sphere for the FV3 model and the tri-polar grid for the local volume computation on each PE (Figure 2). This is similar to method 4.b of J. L. Anderson and Collins (2007). We are currently exploring ways to further improve load and memory balancing for local volume computations.

During the development of the framework for local volume solvers, we benefited greatly from the large number of contributors to the JEDI framework, who tested the generic local volume solvers that we developed in this paper using atmospheric, oceanic, land, and aerosol test cases. Our brief summary of the cycling results showed that different variants of the local volume solver (GETKF, LETKF, LETKF-OI) all compared well to a variety of reference solutions including a free running model (land and aerosols) and variational solvers (ocean and aerosols).

Testing of the developed solvers in the cycling system also revealed that further work is needed to improve scalability, computational performance, and memory management of the local volume solver implementation in JEDI. We chose not to highlight specific computational performance results in this paper as the JEDI code optimization is rapidly progressing and any results with regard to the computational performance are likely to quickly become obsolete. Specifically, our group is actively working on the issues related to computational scalability, efficient memory management, and disk input-output optimization that are essential for operational implementation of the JEDI GETKF solver that can use high-resolution global ensembles and a complete set of atmospheric observations.

Finally, we believe that this local volume framework lays the groundwork for future algorithmic developments. Specifically, JEDI can benefit from implementation of the serial solver (J. L. Anderson, 2001), which replaces the matrix inverse in Equations 10 and 16 with a sequence of scalar inverse problems. The serial solver is a pre-requisite to the implementation of non-linear filters, such as the local particle filters (Poterjoy, 2016), or the Gamma, Inverse-Gamma, and Gaussian filter (Bishop, 2016). The most natural implementation of serial filters in JEDI would involve replacing the Ensemble Transform computation for each of the local volumes with the serial EnKF update (similar to method 4.b in J. L. Anderson and Collins (2007)). Another intriguing direction would be to implement iterative Kalman Filters (Bocquet & Sakov, 2014), which can exploit the existing library of linearized observation operators in JEDI. Implementation of the multi-scale localization operators similar to (Wang et al., 2021) would be yet another beneficial development.

Data Availability Statement

The software code for the generic JEDI solvers is available in the OOPS code repository (JCSDA, 2023d). Software implementations for land, atmosphere, and composition are available in the FV3-JEDI code repository (JCSDA, 2023a). Software implementation for the ocean and ice components is available in the SOCA code



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repository (JCSDA, 2023e). The Unified Forecast System model is available in the UFS code repository (UFS, 2023). Training and documentation for the JEDI software are available in JCSDA (2023b, 2023c), respectively.

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