

# Identifying $Z_{DR}$ Columns in Radar Data with the Hotspot Technique

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**ABSTRACT:** A novel differential reflectivity ( $Z_{DR}$ ) column detection method, the hotspot technique, has been developed. Utilizing constant altitude plan projection indicators (CAPPI) of  $Z_{DR}$ , reflectivity, and a proxy for circular depolarization ratio at the height of the  $-10^{\circ}\text{C}$  isotherm, the method identifies the location of the base of the  $Z_{DR}$  column rather than the entire  $Z_{DR}$  column depth. The new method is compared to two other existing  $Z_{DR}$  column detection methods and shown to be an improvement in regions where there is a  $Z_{DR}$  bias.

**SIGNIFICANCE STATEMENT:** Thunderstorm updrafts are the area of a storm where precipitation grows, electrification is initiated, and tornadoes may form. Therefore, accurate detection and quantification of updraft properties using weather radar data is of great importance for assessing a storm's damage potential in real time. Current methods to automatically detect updraft areas, however, are error-prone due to common deficiencies in radar measurements. We present a novel algorithmic approach to identify storm updrafts that eliminates some of the known shortcomings of existing methods. In the future, our method could be used to develop new hail detection algorithms, or to improve short-term weather forecasting models.

**KEYWORDS:** Convective storms; Algorithms; Radars/Radar observations; Operational forecasting

## 1. Introduction

High  $Z_{DR}$  (e.g.,  $>1$  dB) values found above the environmental freezing level ( $0^{\circ}\text{C}$ ) can be created by, among other things, large raindrops and wet hail suspended in strong convective updrafts (e.g., Kumjian et al. 2014; Snyder et al. 2015). Convective storm updrafts can also sort liquid water drops by size, suspending large drops [e.g.,  $Z_{DR} > 1$  dB corresponding to  $\geq 3$ -mm raindrops Ryzhkov and Zrnić (2019)] while ejecting smaller drops (e.g.,  $Z_{DR} < 1$  dB). Depending on the vertical velocity of the updraft, the region of high  $Z_{DR}$  can extend several kilometers above the environmental freezing level. This polarimetric signature of vertically correlated high  $Z_{DR}$  values found above the melting layer within convective storms is called the  $Z_{DR}$  column.  $Z_{DR}$  columns in polarimetric radar data have been used to study the initiation and development of convective cells in observational studies and in convective storm modeling experiments (e.g., Kumjian et al. 2014; Carlin et al. 2017; Snyder et al. 2017a,b). Used as a proxy for the updraft of the convective storm, a  $Z_{DR}$  column's magnitude, location, cross-sectional area, and development within a particular storm can be associated with other storm-based properties such as storm evolution, hail production, and tornadic potential (e.g.,

Kuster et al. 2019; French and Kingfield 2021; Segall et al. 2022).

The operational utility of  $Z_{DR}$  columns in observations collected by the WSR-88D network has been limited by poor vertical resolution due to scanning strategies that prioritize low-level scans over midlevel scans (e.g., Kingfield and French 2022; Cho et al. 2022). Because of this,  $Z_{DR}$  column detection techniques like the one proposed by Snyder et al. (2015; hereafter referred to as S15) lack consistency when the  $Z_{DR}$  column is in a region of the domain with poor vertical sampling or coverage (which is common at ranges  $> 100$  km). Additionally, the WSR-88D network experiences a  $Z_{DR}$  bias of  $> \pm 0.2$  dB on about 20% of its radars at any given time (D. Zittel 2023, personal communication). Such a bias can create errors in  $Z_{DR}$  column detection and the determination of its depth when a fixed threshold (e.g.,  $Z_{DR} \geq 1$  dB) is used. The high values of  $Z_{DR}$  produced in the melting layer as snow melts into rain can also create errors (false positives) in other  $Z_{DR}$  column identification methods such as the Thunderstorm Risk Estimation from Nowcasting Development via Size Sorting (TRENDSS) algorithm (Kingfield and Picca 2018; French and Kingfield 2021). The TBSS (three-body scatter signature; Zrnić 1987) described by Lemon (1998) is an artifact in the radar data that produces both high  $Z_{DR}$  and low-to-moderate  $Z$ . This artifact is often found adjacent to strong storms and near true  $Z_{DR}$  columns and can be a challenge for all  $Z_{DR}$  column detection algorithms.

Our proposed  $Z_{DR}$  column detection method, the hotspot technique, seeks to identify the base of the  $Z_{DR}$  column on a single level rather than throughout its entire vertical extent. Identifying the base of the  $Z_{DR}$  column on a single level (e.g.,

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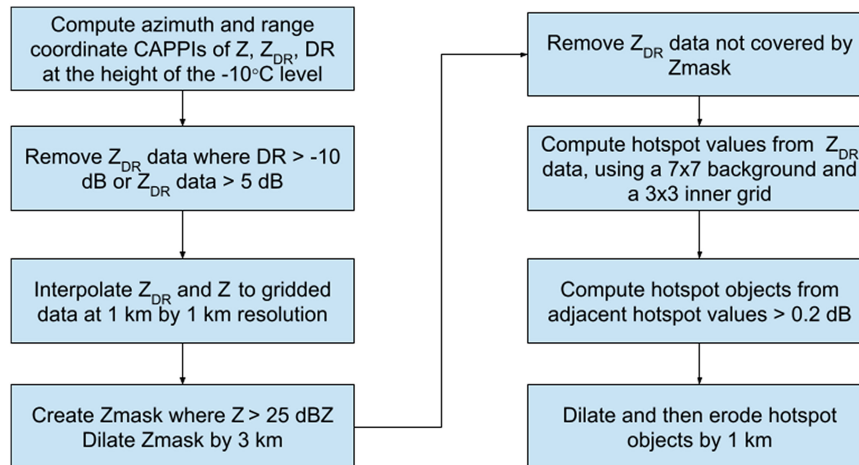


FIG. 1. A flowchart summarizing the computational steps of the  $Z_{DR}$  hotspot technique.

the  $-10^{\circ}\text{C}$  CAPPI) is easier than previous efforts (i.e., S15 and TRENDSS), which try to locate  $Z_{DR}$  columns using data from multiple elevation angles. Our  $Z_{DR}$  hotspot technique can assist forecasters in identifying a storm's updraft location and assessing that storm's strength, structure, and whether it is growing or decaying (e.g., Kuster et al. 2020). The  $Z_{DR}$  hotspot technique identifies areas in the data where the local median of  $Z_{DR}$  is greater than the median of the surrounding data. Due to this local data-to-data comparison, the technique is both  $Z_{DR}$  calibration independent and applicable to radar variables other than  $Z_{DR}$ . Confidence in the identified  $Z_{DR}$  hotspot being related to a storm's updraft is increased by being near a strong VIL (vertically integrated liquid) or reflectivity ( $Z$ ) maximum. Although beyond the scope of this paper, the method could also be applied to correlation coefficient ( $\rho_{hv}$ ),  $K_{DP}$ , azimuthal shear, or other radar fields; for example, if looking for large and giant hail, it can also be used to identify negative  $Z_{DR}$  anomalies (e.g., Kumjian et al. 2010; Witt et al. 2018).

## 2. Algorithm description

This hotspot technique (Fig. 1) begins by computing azimuth and range coordinate CAPPIs of  $Z_{DR}$ ,  $Z$ , and a circular depolarization ratio proxy (Ryzhkov et al. 2017) at the height of the  $-10^{\circ}\text{C}$  isotherm (Figs. 2a–c). Depolarization ratio (DR) is a combination of  $Z_{DR}$  and  $\rho_{hv}$  that serves as a proxy for circular depolarization ratio (CDR), which contains information about scatterers' shape and water (ice) phase composition. Hailstones, wet snowflakes, TBSS, and nonmeteorological scatterers are characterized by large values of DR (i.e., from  $-15$  to  $0$  dB). We chose to remove the  $Z_{DR}$  data where  $\text{DR} > -10$  dB. This DR threshold value specifically aims to remove any TBSS and is less restrictive than the  $-12$ -dB threshold originally proposed in Kilambi et al. (2018), which targets all nonmeteorological data. Figure 2d shows the locations of  $\text{DR} > -10$  dB in orange, which correspond to small TBSS detections. The CAPPI height is set to the height of the  $-10^{\circ}\text{C}$  isotherm (see discussion below) and is constant throughout the domain. It can be obtained from

a model data vertical profile or sounding near the radar of interest. In our experience,  $-10^{\circ}\text{C}$  is the lowest height where reliable detection of the  $Z_{DR}$  column signature can be expected, without contamination by the melting layer. Locations with  $Z_{DR} > 5$  dB, which are not commonly found in meteorological data at S band even when large drops are present (although the biggest raindrops, e.g., those exceeding  $7$  mm in diameter, may be associated with such high  $Z_{DR}$ ), are then removed. We interpolate the  $Z$  and trimmed  $Z_{DR}$  data from azimuth and range coordinates to a  $1$ -km grid with a  $5$  data-point average of the radial data, using the nearest radial point to the target grid point (nearest neighbor) and the surrounding radial points in both azimuth and range ( $5$  gate arithmetic average; Langston et al. 2007; Fig. 2e).

Finally, to focus our method on convection, we create a  $Z$  mask from gridded  $Z$  data on a CAPPI, identifying all locations where  $Z > 25$  dBZ. We dilate this mask by  $3$  km in each (i.e.,  $x$  and  $y$ ) direction, enlarging the detection region and filling gaps (Fig. 2c). The  $Z$  mask is applied to the  $Z_{DR}$  data such that all  $Z_{DR}$  data not covered by the mask are removed. All of the above works to limit the technique's detection to locations with meteorological data near convective storms at the height of the  $-10^{\circ}\text{C}$  isotherm.

To create the  $Z_{DR}$  hotspot field using our trimmed  $Z_{DR}$  data (Fig. 2e), for each grid point we compute the  $3 \times 3$  gridded ( $9 \text{ km}^2$ ) local median and the surrounding  $7 \times 7$  background median; the  $9$  gates of the  $3 \times 3$  local area are *not included* in the median of the  $40$  gates from the surrounding  $7 \times 7$  background area, so the surrounding median covers  $40 \text{ km}^2$  (Fig. 3). The value of the  $Z_{DR}$  hotspot is the difference between the local area median and the background area median:

$$Z_{DR} \text{ hotspot value} = \text{local area } Z_{DR} \text{ median } (9 \text{ km}^2) \\ - \text{background area } Z_{DR} \text{ median } (40 \text{ km}^2).$$

Now that the field of  $Z_{DR}$  hotspots is available, we can highlight those locations in the field that exceed  $0.2$  dB for further review. We selected the value of  $0.2$  dB as our threshold

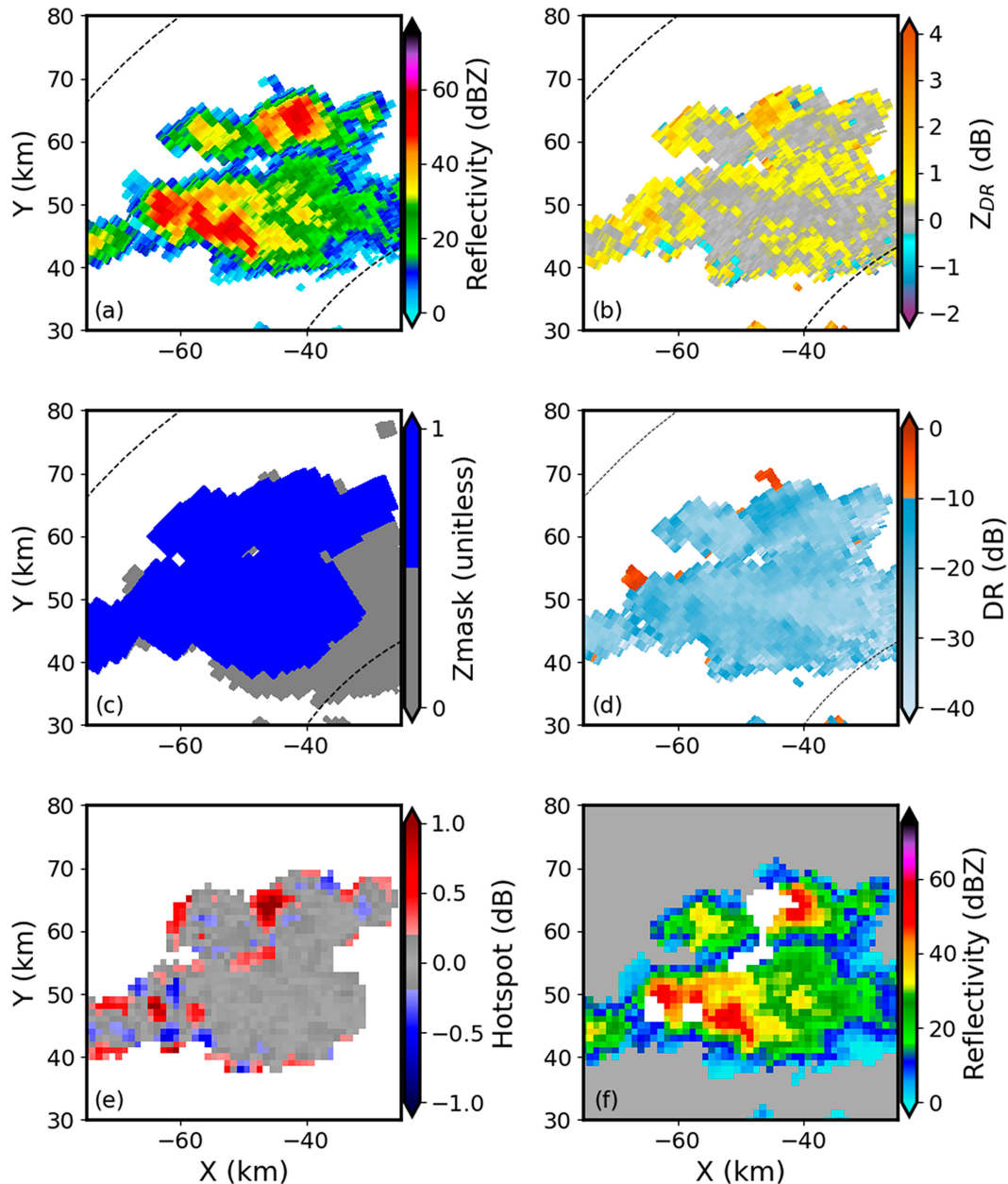


FIG. 2. (a)  $Z_H$  (reflectivity), (b)  $Z_{DR}$ , (c)  $Z$  mask (blue) overlaid on reflectivity (gray), (d) DR, (e)  $Z_{DR}$  hotspots, and (f)  $Z_{DR}$  hotspot objects (white) overlaid on gridded reflectivity. All CAPPI data are from 2057 UTC 12 Jun 2014 at KDYX (Dyess Air Force Base, TX) at the height of the  $-10^{\circ}\text{C}$  temperature level (5.83 km AGL).

because this value would capture almost all points greater than one standard deviation above the mean hotspot value over multiple cases. A list of the cases used to develop the hotspot technique and the number of volumes for each case is listed in Table 1. A histogram of the hotspot values for the KOAX (Omaha, Nebraska) case on 3 June 2014 can be found in Fig. S1 of the online supplemental material and shows the new variable to be normally distributed.

Adjacent points that exceed the 0.2-dB threshold are grouped together and identified as an object. Objects with an

area of  $<5\text{ km}^2$  are removed to reduce the number of objects and to keep only those objects that are likely to exist at the next time step. Each object is required to have at least one point with  $Z > 20\text{ dBZ}$ . Objects then undergo a dilation and erosion procedure that combines objects separated by 2 km or less. Dilation and erosion are morphological image procession operations. In this case, the dilation procedure adds any pixel within 1 km of an object pixel to the object and then the erosion procedure removes any pixel of the object that is within 1 km of the edge of the object. Once the dilation

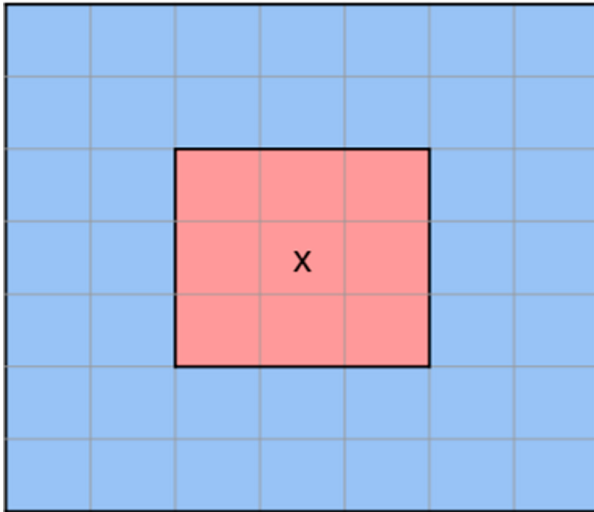


FIG. 3. The hotspot method computed for the location “X” is the difference between the median of the data in the salmon-colored inner box ( $9 \text{ km}^2$ ) and the median of the data in the light blue colored data ( $40 \text{ km}^2$ ). Each box outlined by the light gray and/or black lines is  $1 \text{ km}^2$ .

and erosion processes are complete, the remaining objects are now considered probable updrafts (Fig. 2f).

Now that we have described the hotspot method, it is important to examine the selection of both the height at which we compute the hotspot (i.e., the height of the  $-10^\circ\text{C}$  isotherm) and the size of the grids used to compute the value of the hotspot itself [ $7 \times 7$  background grid (outer) and  $3 \times 3$  inner grid]. Figure 4 shows the hotspot method results for four different sized grids. The hotspot data for Figs. 4 and 5 has been converted into its standard deviation ( $\sigma$  level), where the values of standard deviation greater than 1.0 are colored in red and values less than  $-1.0$  are colored in blue. The conversion into standard deviation allows for a fair comparison between the results from the different grid sizes because the standard deviation of the largest grid ( $11 \times 11$ , Fig. 4f) is much smaller than the standard deviation of the smallest grid ( $5 \times 5$ , Fig. 4c). Finally, a standard

deviation of  $>1.0$  or  $<-1.0$  suggests a location where the value of the hotspot data is of interest.

Figure 4 is presented at a scale of  $150 \text{ km} \times 150 \text{ km}$  so that the reader can evaluate the character of the hotspot method in general. The smallest background grid ( $5 \times 5$ ; Fig. 4c) shows significantly more and separate hotspots (indicated in red) that are smaller in size than that of the largest background grid ( $11 \times 11$ ; Fig. 4f). The inset black box in Fig. 4a shows the location of Fig. 5, where a more detailed view is presented.

Figure 5 shows small, medium, and large hotspots denoted with green, cyan (light blue), and yellow arrows, respectively. The smaller hotspot identified by the green arrow has been broken into two different areas on the smallest background grid ( $5 \times 5$ ; Fig. 5c), but because the signal in the  $Z_{\text{DR}}$  data is weak, this hotspot is not identified in the  $11 \times 11$  background grid (Fig. 5f). The hotspot location pointed to by the cyan arrow is identified on all grids, but we find that the  $7 \times 7$  background grid (Fig. 5d) does the best job of grouping large adjacent hotspot values and identifying the areal extent of the hotspot. Finally, for the largest hotspot (yellow), we note that for the smallest scale ( $5 \times 5$ ; Fig. 5c), the maximum values of the hotspot are not adjacent, and that the areal extent of the hotspot decreases as the background grids increase in size. Our analysis finds that the  $7 \times 7$  background grid with a  $3 \times 3$  inner grid (Fig. 5d) is the best performing combination for  $Z_{\text{DR}}$  data.

When we selected the height at which to run the hotspot method ( $-10^\circ\text{C}$ ) we considered a number of factors. It is our hypothesis that high  $Z_{\text{DR}}$  values found above the environmental freezing level are large raindrops or wet hail held aloft by strong convective updrafts. The creation of CAPPis near  $0^\circ\text{C}$  can contain high  $Z_{\text{DR}}$  values created by melting snow rather than lofted large drops or wet hail. However, weak convective updrafts may not reach much beyond the height of the environmental freezing level. Figure 6 shows the performance of the hotspot method for different heights using the preferred  $7 \times 7$  background grid and  $3 \times 3$  inner grid.

The hotspot CAPPi at  $0^\circ\text{C}$  (Fig. 6c) contains numerous hotspots, while the  $-10^\circ\text{C}$  panel (Fig. 6d) shows a more reasonable number compared to the number of storm cells in the reflectivity data (Fig. 6a). Both the  $-20^\circ\text{C}$  (Fig. 6e) and

TABLE 1. Data used to develop the hotspot algorithm (STD = standard deviation).

Radar ID	Start date	Start time (UTC)	No. of volumes	Avg hotspot value (dB)	STD of hotspot values (dB)
KDGX	25 May 2023	0000	29	-0.0012	0.1896
KDYX	12 Jun 2014	1800	116	0.0304	0.2361
KEWX	4 May 2021	0000	57	0.0121	0.2976
KFDR	22 May 2020	2100	90	0.0134	0.2137
KFWS	12 Apr 2016	0000	83	0.0188	0.2350
KFWS	24 Mar 2019	2100	93	0.0551	0.2742
KICT	24 Jul 2013	0000	98	0.0323	0.2618
KOAX	3 Jun 2014	1800	119	0.0100	0.2059
KSJT	12 Jun 2014	2000	109	0.0295	0.2338
KTLX	27 Feb 2023	0000	66	0.0136	0.2239
KUDX	19 Jun 2015	2300	86	0.0075	0.1810
Total volumes				Avg	Avg STD
946				0.0201	0.2321

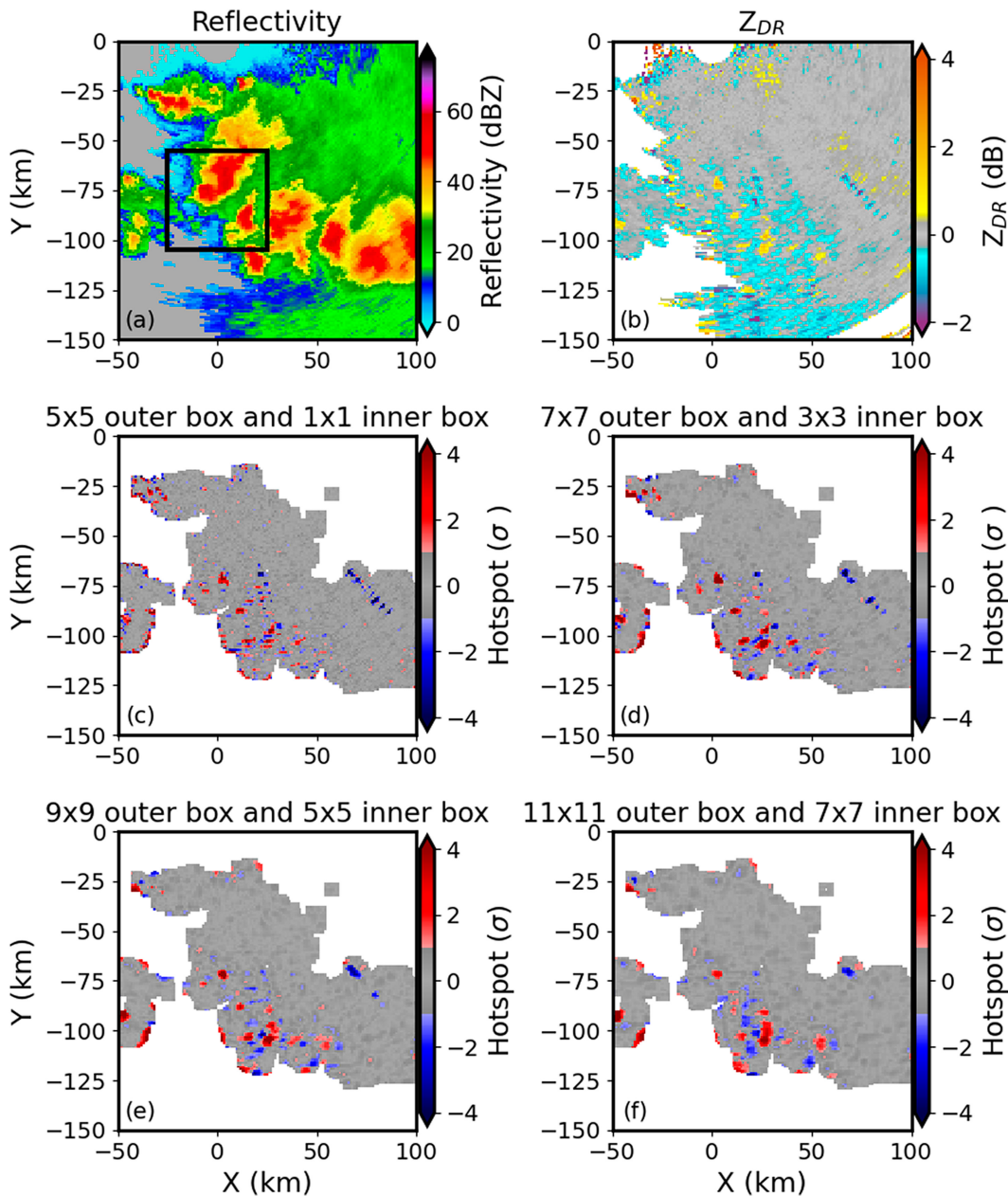


FIG. 4. (a) Reflectivity, (b)  $Z_{DR}$ , and  $Z_{DR}$  hotspots (in  $\sigma$  levels) for a (c)  $5 \times 5$  outer box, (d)  $7 \times 7$  outer box, (e)  $9 \times 9$  outer box, and (f)  $11 \times 11$  outer box. All data are from 0358 UTC 23 May 2020 at KFDR at the height of the  $-10^{\circ}\text{C}$  isotherm (5.547 km AGL). The inset black box in (a) is the region shown in Fig. 5.

$-30^{\circ}\text{C}$  (Fig. 6f) hotspot fields show a distinct lack of hotspot detections, as only the strongest storms can loft large drops and wet hail to these heights.

### 3. Results and discussion

#### a. Results and limitations of the hotspot technique

Results of the hotspot technique are demonstrated in Figs. 7 and 8 where hotspots are computed for a set of severe

storms along the Oklahoma/Texas border on 24 March 2019 as viewed by KFDR (Frederick, Oklahoma). The updrafts of these storms have been identified by the technique and are located adjacent to large areas of higher  $Z$ . Compared to the  $Z_{DR}$  output in Fig. 7b, the hotspot field (Fig. 7c) clarifies the locations of the updrafts, and overlaying all hotspot locations  $> 0.2$  dB as white on the  $Z$  image (Fig. 7d) demonstrates the presence of  $Z_{DR}$  hotspots at plausible updraft locations.

Figure 8 is a zoomed in view of one of the larger and more complex storm cells shown in Fig. 7. A bounded weak echo

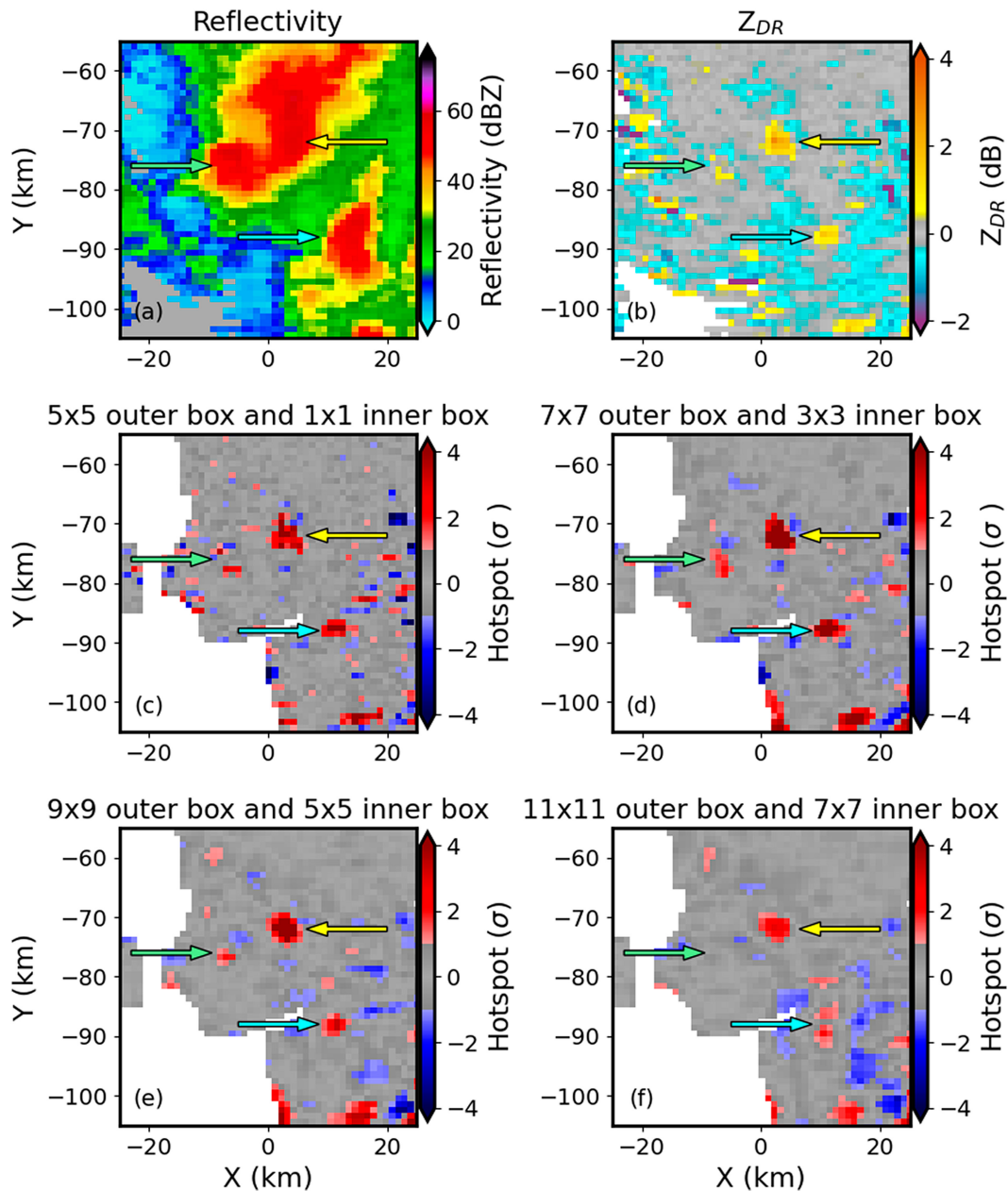


FIG. 5. As in Fig. 4, but for the subsetting area shown in Fig. 4a. The green, yellow, and cyan arrows identify hotspots of different sizes.

region (BWER) (verified by examining multiple elevations not shown) is visible in Fig. 8a at  $x = -25$  km and  $y = -68$  km, which suggests a strong updraft is present. The  $Z_{DR}$  data in Fig. 8b has a number of points where  $Z_{DR} > 1$  dB, suggesting drop size sorting collocated with an updraft and also  $Z_{DR} < -1.0$  dB, suggesting possible large hail. Finally, a region runs from the upper-right corner ( $x = -20$  km,  $y = -50$  km) to the lower-left corner ( $x = -40$  km,  $y = -85$  km) of Fig. 8b, where all of the  $Z_{DR}$  values are lowered, suggesting differential attenuation due to a storm closer to the radar and/or from parts of the storm shown. The  $Z_{DR}$  hotspot technique

(Fig. 8c) identifies three large areas with values  $> 0.2$  dB and several smaller ones with areas less than  $5$  km<sup>2</sup>. The effects of removing the smaller detections and the dilate and erode procedure are evident in the white spots overlaid on the  $Z$  data in Fig. 8d. Only the three larger areas from Fig. 8c remain and those areas have been slightly filled by the dilate and erode procedure. The larger cells were just far enough apart that they were not combined. Overall, the remaining hotspots show two cells associated with the BWER and another cell forming along the southwest edge of the storm. One of the primary benefits of the hotspot technique is the identification

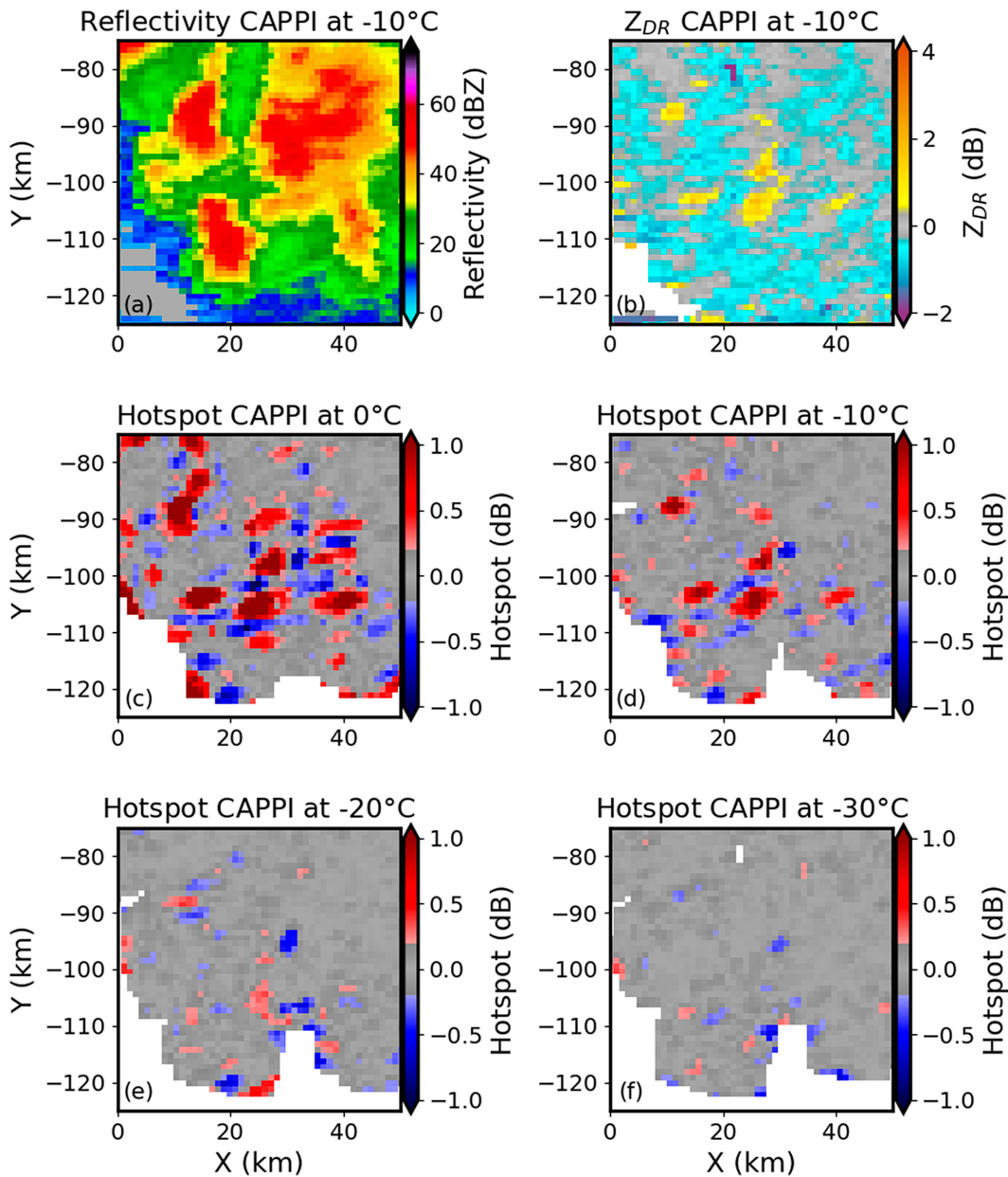


FIG. 6. (a) Reflectivity at  $-10^{\circ}\text{C}$ , (b)  $Z_{DR}$  at  $-10^{\circ}\text{C}$ , and  $Z_{DR}$  hotspots at the heights of the (c)  $0^{\circ}$ , (d)  $-10^{\circ}$ , (e)  $-20^{\circ}$ , and (f)  $-30^{\circ}\text{C}$  isotherms. All data are from 0358 UTC 23 May 2020 at KFDR with a  $7 \times 7$  background grid and a  $3 \times 3$  inner grid.

of newly formed convection, such as the cell on the southwestern edge ( $x = -30 \text{ km}$ ,  $y = -80 \text{ km}$ ).

Figure 9 illustrates an error we observed on multiple occasions. The data are from KEWX (New Braunfels, Texas) at 0010 UTC 4 May 2021 and show three strong storms located from north to south. There is a consistent ribbon of high  $Z_{DR}$  that forms along the western edge of each updraft (yellow colors in Fig. 9b) and down radial from the radar. In the northern storm there is a TBSS evident, which might be one explanation for this region of high

$Z_{DR}$ , but the other two storms do not show evidence of a TBSS and yet contain the elevated  $Z_{DR}$  along the edge of the storm. The hotspot technique identifies this edge region of high  $Z_{DR}$  (Fig. 9, pink arrows) and combines it with nearby updrafts (Fig. 9, yellow arrows) in error. Updrafts embedded in surrounding precipitation do not experience this kind of error. Data sampled with a low signal-to-noise ratio, oblate ice crystals, and TBSS are all possible explanations of this high  $Z_{DR}$  data; regardless, the hotspot technique misidentifies these areas as an updraft region.

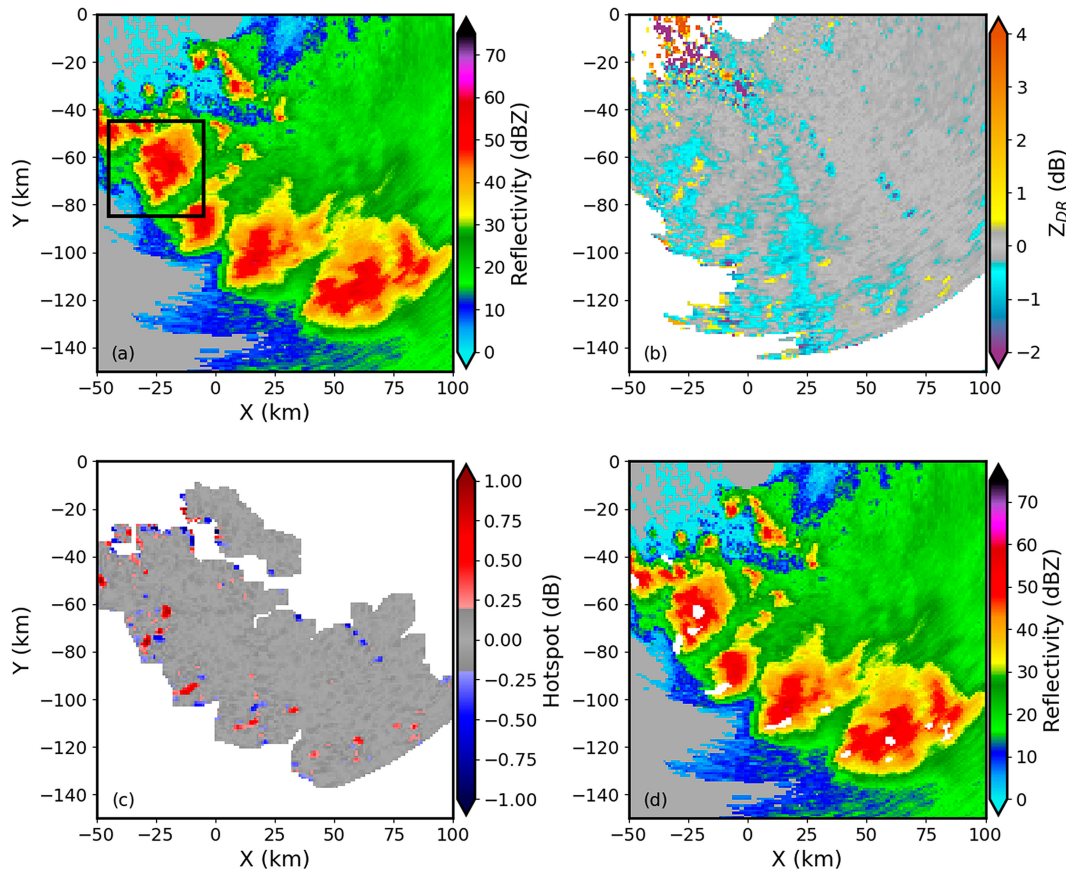


FIG. 7. (a) Reflectivity, (b)  $Z_{DR}$ , (c)  $Z_{DR}$  hotspot, and (d)  $Z_{DR}$  hotspot objects (white) overlaid on reflectivity. All data from 0457 UTC 23 May 2020 at KFDR at the height of the  $-10^{\circ}\text{C}$  isotherm (5.547 km AGL). The inset black box in (a) shows the region shown in Fig. 8.

The technique may also encounter difficulties when applied in regions with poor data quality. Figure 10 shows an example of this in an area located just beyond the maximum unambiguous range. We note that other methods for  $Z_{DR}$  column identification (e.g., TRENDSS, S15) also tend to struggle in regions with poor data quality. In this example, the area located just beyond the maximum unambiguous range ( $\sim 160$  km) impacts the algorithm by producing noisy  $Z_{DR}$  values and several false updraft detections (Fig. 10). The detections inside the black box can be identified as false because they occur away from strong signatures (e.g.,  $Z > 45$  dBZ). The algorithm is usually reliable within the maximum unambiguous range and can be extended to distances beyond it if the poor data quality of the transition zone between the maximum unambiguous range and the unfolded region is taken into account. In our experience the technique produces usable data for storms located  $< 200$  km from the radar. At far ranges (i.e.,  $> 200$  km), the lowest elevation collected by the radar may rise above the height of the  $-10^{\circ}\text{C}$  isotherm, although this is uncommon in convective environments where the height of the  $-10^{\circ}\text{C}$  isotherm almost always exceeds 2 km. At longer ranges, other factors can reduce the

reliability of the technique, such as the size of the beam relative to the size of the updraft, the interpolation of radial data to gridded Cartesian data, and the conversion of data into CAPPI outputs where one elevation may be in the melting layer, or where only one elevation may be used.

The data from KOAX at 0102 UTC 4 June 2014 (Fig. 11) is a complex case with multiple storms in the domain that would require analysis by forecasters. In the center of the image there are two separate storms, each with a developing flanking line. The hotspot technique identifies the updrafts of both the parent storm and storms developing along the flanking line. Forecasters who want to interrogate the data to determine storm strength and issue warnings can quickly find each storm updraft using the results from the hotspot technique.

During our testing and evaluation we noticed strong supercells for which the  $Z_{DR}$  hotspot object would either become too small (i.e., below the  $< 5\text{-km}^2$  detection threshold) or disappear entirely. This lack of detection occurred despite obvious markers of storm strength such as  $Z_H > 50$  dBZ at the height of the  $-10^{\circ}\text{C}$  isotherm. We illustrate this using an example from 0211 UTC 4 May 2021 (Fig. 12), in



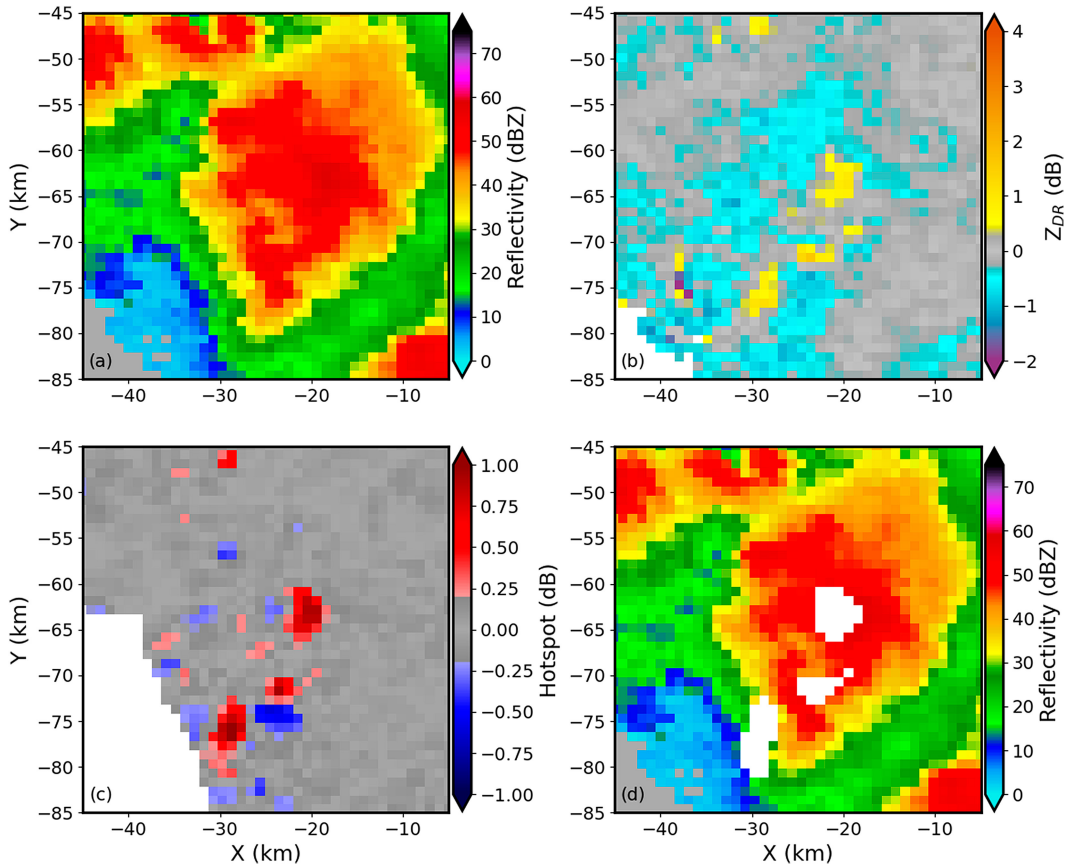


FIG. 8. As in Fig. 7, but for the subsetting area shown in Fig. 7a.

which the  $Z_{DR}$  hotspot became smaller than  $5 \text{ km}^2$  (shown by the yellow arrow in Figs. 12a,c). In this case, we also noticed that raw  $Z_{DR}$  values (Fig. 12b) were lower than the  $Z_{DR}$  from the previous scan (not shown). Evidence of large hail is suggested by an area of low correlation coefficient (CC) where the updraft would be expected to be located (as subjectively assessed using radar training and knowledge of convective storm structure) and a significant TBSS (Fig. 12a, pink arrow). We suspect that the previously observed strongly positive  $Z_{DR}$  values of large liquid water drops are being reduced by the strong  $Z_{DR} \sim 0 \text{ dB}$  contribution of hail, possibly large, inside the updraft (e.g., Kumjian et al. 2014). The disappearance of a  $Z_{DR}$  hotspot for a mature cell could be an indicator of falling hail.

#### b. Comparison with existing updraft detection methods

The  $Z_{DR}$  hotspot technique was compared to both the S15 method and TRENDSS (Kingfield and Picca 2018). In the S15 method, it is assumed that any quality controlled  $Z_{DR}$  value  $> 1 \text{ dB}$  that is located at and above the height of the  $0^\circ\text{C}$  level is due to size sorting in convective updrafts. As such, the output of the S15 algorithm is not a location where size sorting is taking place but rather a  $Z_{DR}$  column depth computed by identifying vertically adjacent  $Z_{DR}$  values  $> 1 \text{ dB}$  in a volume of data. A common

deficiency of the S15 method is its dependence on a constant and arbitrary  $Z_{DR}$  threshold of  $1 \text{ dB}$ . As such, the method can over detect the presence of  $Z_{DR}$  columns due to high  $Z_{DR}$  bias, or under detect due to low  $Z_{DR}$  bias. In practice, however, it is difficult to get consistent results from S15; unreliable detections of  $Z_{DR}$  column depth can be created from volume scan strategies that create large vertical gaps in the data and volume samples that are spaced far apart in time ( $> 5 \text{ min}$ ), tilted or fast-moving storms, TBSS, and areas of poor data quality like those just beyond the maximum unambiguous range. Both the TRENDSS method and the  $Z_{DR}$  hotspot technique can produce  $Z_{DR}$  column depths like those in S15 by applying the method to vertically adjacent data, but the same problems that produce inconsistent results for S15 exist in these approaches.

The TRENDSS method first described in Kingfield and Picca (2018), and then modified by French and Kingfield (2021), was developed to identify drop size sorting in radar data. The method creates histograms of quality-controlled  $Z_{DR}$  for each 5-dBZ range of  $Z$  between 15 and 60 dBZ (e.g., 15–20, 20–25 dBZ, etc.). These histograms of  $Z$ – $Z_{DR}$  are created for three different levels: below the melting layer, within the melting layer, and above the melting layer. The standard deviation of  $Z_{DR}$  is then calculated

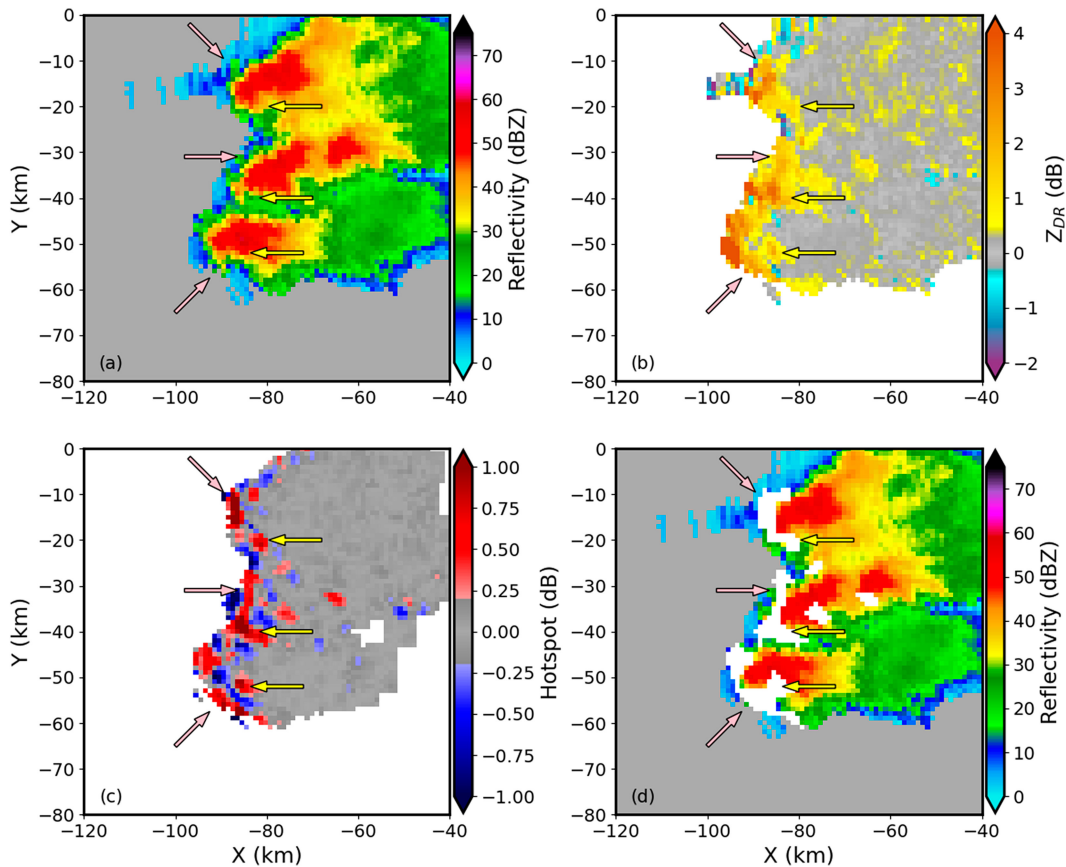


FIG. 9. (a) Reflectivity, (b)  $Z_{DR}$ , (c)  $Z_{DR}$  hotspots, and (d)  $Z_{DR}$  hotspot objects (white) overlaid on reflectivity. All data from 0010 UTC 4 May 2021 at KEWX at the height of the  $-10^{\circ}\text{C}$  isotherm (5.501 km AGL). False detections of updraft regions are identified by pink arrows, and the true updraft regions are identified by the yellow arrows.

for each  $Z$  range at each level. Size sorting locations are identified by computing the standard deviation ( $\sigma$  level) of  $Z_{DR}$  at each location relative to the computed  $Z$ - $Z_{DR}$  histogram at the appropriate height and  $Z$  level. Because the TRENDSS method uses  $\sigma$  as a measure of size sorting it can be directly compared to the  $Z_{DR}$  hotspot technique once the  $\sigma$  of the  $Z_{DR}$  hotspot values has been computed.

To create a fair comparison, we modified the TRENDSS method so that it runs on the trimmed  $Z_{DR}$  data from the  $-10^{\circ}\text{C}$  CAPPI. We then computed the standard deviation of the hotspot data ( $\sigma$  level) as a comparator to the TRENDSS output. Figure 13 compares the three methods for multicellular storms on 23 May 2020 viewed by the KFDR radar. For the group of storms shown, we note that all three methods' maximum values are located in roughly the same place. Also located in this image is an area of data with reduced  $Z_{DR}$ , identified by the yellow box. A strong storm located close to the radar has caused differential attenuation. This is a very common local phenomenon that reduces  $Z_{DR}$  for typically a small number of radials in the volume. Because the S15 algorithm uses a static 1-dB

threshold,  $Z_{DR}$  in this region has been reduced enough that the lowest level does not contribute to the  $Z_{DR}$  column height computation. As a result, S15  $Z_{DR}$  column heights at this location are about 1 km lower than nearby  $Z_{DR}$  columns. The output from TRENDSS is also affected by the differential attenuation because the  $Z_{DR}$  mean and  $\sigma$  values used in TRENDSS are computed from all of the data in the CAPPI. Therefore, the TRENDSS data show a strong negative  $Z_{DR}$  bias in this region (blue). The  $Z_{DR}$  hotspot method, however, is relatively unaffected by this problem due to its reliance on data near the analysis location.

Another comparison between the three  $Z_{DR}$  column methods is presented in Fig. 14. In data from 4 June 2014 at KOAX, the  $Z_{DR}$  field has a positive bias of about 0.42 dB, as determined by computing the average value of  $Z_{DR}$  in dry snow (Richardson et al. 2017). This bias creates a significant over-detection of  $Z_{DR}$  columns by S15 as shown in Fig. 14f. The over-detection of updraft locations by the TRENDSS output (Fig. 14d) is more difficult to explain. We suspect that there are problems in the creation of the  $Z$ - $Z_{DR}$  relationships that underpin the algorithm.

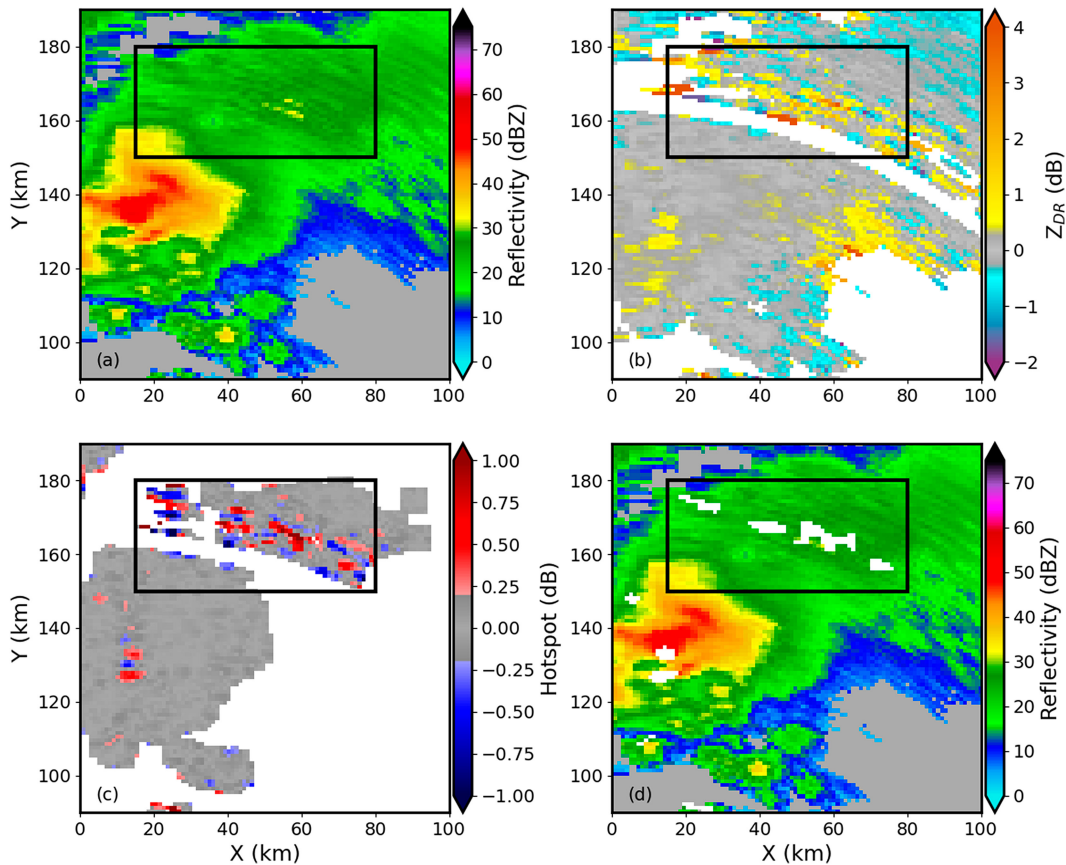


FIG. 10. As in Fig. 9, but for 0224 UTC 25 Mar 2023 at KDGX (Jackson, MS) at the height of the  $-10^{\circ}\text{C}$  isotherm (5.488 km AGL). An area with multiple false detections is indicated by the black box.

Much of the domain, not shown in the image, contains meteorological data in both  $Z$  and  $Z_{\text{DR}}$  at the height of  $-10^{\circ}\text{C}$ . The creation of  $Z$ - $Z_{\text{DR}}$  histograms for this radar at this time at this height could be dominated by ice particles with low and relatively homogeneous  $Z_{\text{DR}}$  values. Such particles would have a much lower standard deviation of  $Z_{\text{DR}}$  than liquid drops. Therefore, when the TRENDSS algorithm computes the  $\sigma$  level of the  $Z_{\text{DR}}$  in locations where liquid drops are present it is always  $>1.0$  (or higher) and as such it may not correspond to size-sorted drops in updraft locations but any type of liquid water or water-coated melting snow with enhanced  $Z_{\text{DR}}$  values. The  $Z_{\text{DR}}$  biases present in the data do not seem to affect the hotspot technique (Fig. 14d), which shows good separation between hotspots and multiple updrafts.

### c. Temporal evolution of updraft objects

Several animations of data were created to demonstrate the  $Z_{\text{DR}}$  hotspot techniques' performance over time. They can be found in the supplemental material. In the first animation, Fig. S2 (KEWX, 1 h 40 min, 20 volumes), we revisit the KEXX case shown in Fig. 9. The detection errors along the edge of the storm echo in this case exist for the first 5–7 time steps ( $\sim 30$  min), but they are not found once

the storm is well organized. By the end of the animation the  $Z_{\text{DR}}$  hotspot technique shows a coherent updraft located on the southwest side of the storm. There are no edge errors in the later outputs even though there is a TBSS present in the data from 0129 UTC onward. Our experience suggests that the issues described in Fig. 9 and present early in Fig. S2 are generally found downradial of newly developed convection.

A much longer animation, Fig. S3 (KFDR, 4 h, 48 volumes), was created for a complex case with numerous hotspot detections present in the volume. This animation begins with a single supercell thunderstorm and then evolves into a multicellular storm. The  $Z_{\text{DR}}$  hotspot technique identifies the initial updraft and the animation shows this updraft's evolution over time. The consistency of the detection improves our confidence in the output of the algorithm. Later in the animation (after 0147 UTC), a multitude of smaller and more short-lived updrafts form and move throughout the domain. Although we have focused on the location of  $Z_{\text{DR}}$  hotspots  $>0.2$  dB as locations of interest, groups of particularly large values ( $>0.6$  dB, approximately  $3\sigma$ ) in our experience suggest a well-organized updraft that forecasters would want to evaluate in more detail.

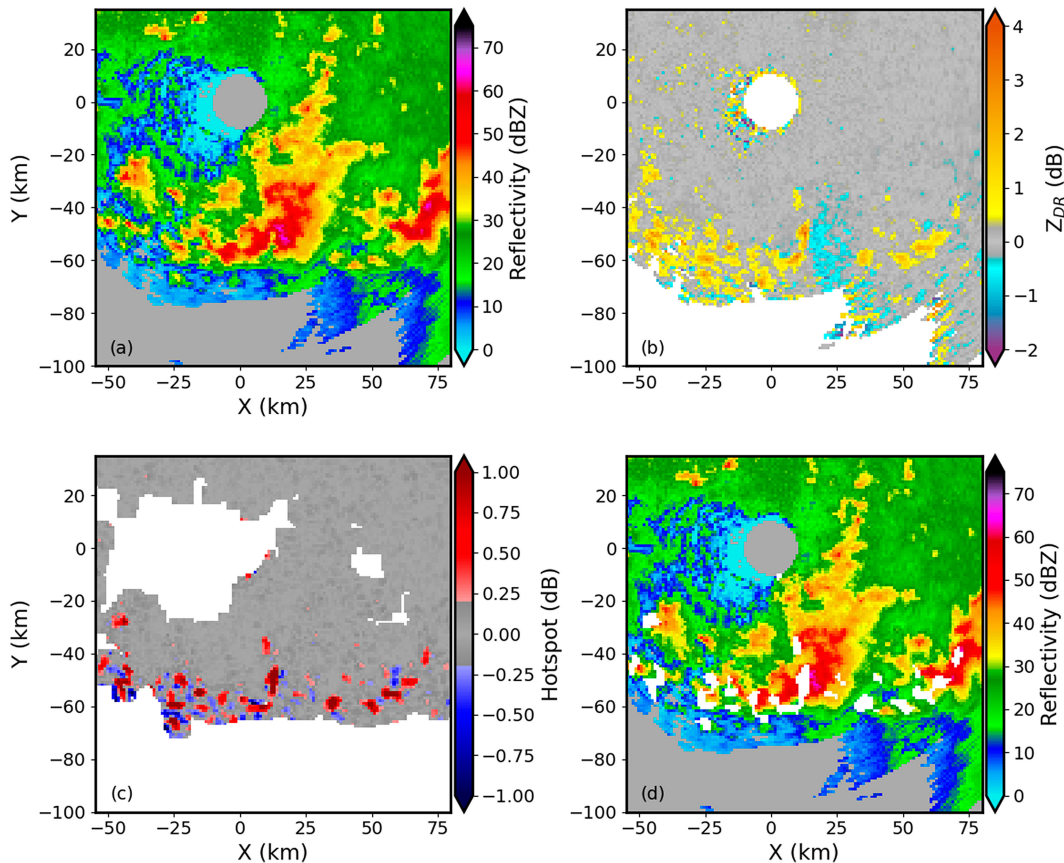


FIG. 11. As in Fig. 9, but for 0102 UTC 4 Jun 2014 at KOAX at the height of the  $-10^{\circ}\text{C}$  isotherm (5.188 km AGL).

Last, we created an animation of a linear system shown in Fig. S4 (KOAX, 2 h, 22 volumes). The hotspot technique identifies the high  $Z_{\text{DR}}$  values along the leading edge of the linear convective system as updraft locations. The edge effects identified in the first animation (Fig. S2) do not seem to be present in this event. As with the previous example, the location of the  $Z_{\text{DR}}$  hotspots with very large values ( $>0.6$  dB) can be used to identify the stronger cells in the linear system.

#### 4. Summary and future work

In this paper we have described and demonstrated the new  $Z_{\text{DR}}$  hotspot technique, which identifies the base of  $Z_{\text{DR}}$  columns in the  $-10^{\circ}\text{C}$  CAPPI data using  $Z_{\text{DR}}$ ,  $Z$ , and DR. We associated locations detected by the technique with liquid water drops that have been size sorted (retaining the large drops) by an updraft. Possible errors in the technique's updraft detection can be created by poor data quality or high  $Z_{\text{DR}}$  values of unknown origin located on the edge of storms near an updraft. A comparison between the  $Z_{\text{DR}}$  hotspot technique and two other  $Z_{\text{DR}}$  column detection methods was shown. The  $Z_{\text{DR}}$  hotspot technique produces improved performance for areas with differential

attenuation and when a  $Z_{\text{DR}}$  bias is present. Forecasters can use the  $Z_{\text{DR}}$  hotspot technique to identify updrafts in the domain and to investigate near them for hazards. In complex nowcasting scenarios when multiple updrafts are present and time is limited, forecasters can use this hotspot technique to triage and focus on locations where the  $Z_{\text{DR}}$  hotspot technique yields very large values ( $>0.6$  dB or  $3\sigma$ ), which we suspect are an indication of a strong updraft in a well-organized storm.

In the future, the  $Z_{\text{DR}}$  hotspot technique could be utilized to study the evolution of storm updrafts over time, and the properties of the identified updraft areas, such as their size and the magnitude of the hotspot value, could be linked to the likelihood of storm hazards. In future studies, we aim to identify precursor signatures for large or giant hail, tornadoes, and downbursts using this approach. The hotspot technique can also be applied to other polarimetric variables such as  $\rho_{\text{hv}}$ ,  $K_{\text{DP}}$ , and/or azimuthal shear, which could enable identification of hail cores,  $K_{\text{DP}}$  columns, and mesocyclones, respectively. Finally, the identification of updrafts might be used by the storm-scale modeling community for polarimetric radar data assimilation (e.g., Carlin et al. 2017) or to evaluate the output of convection-allowing models.

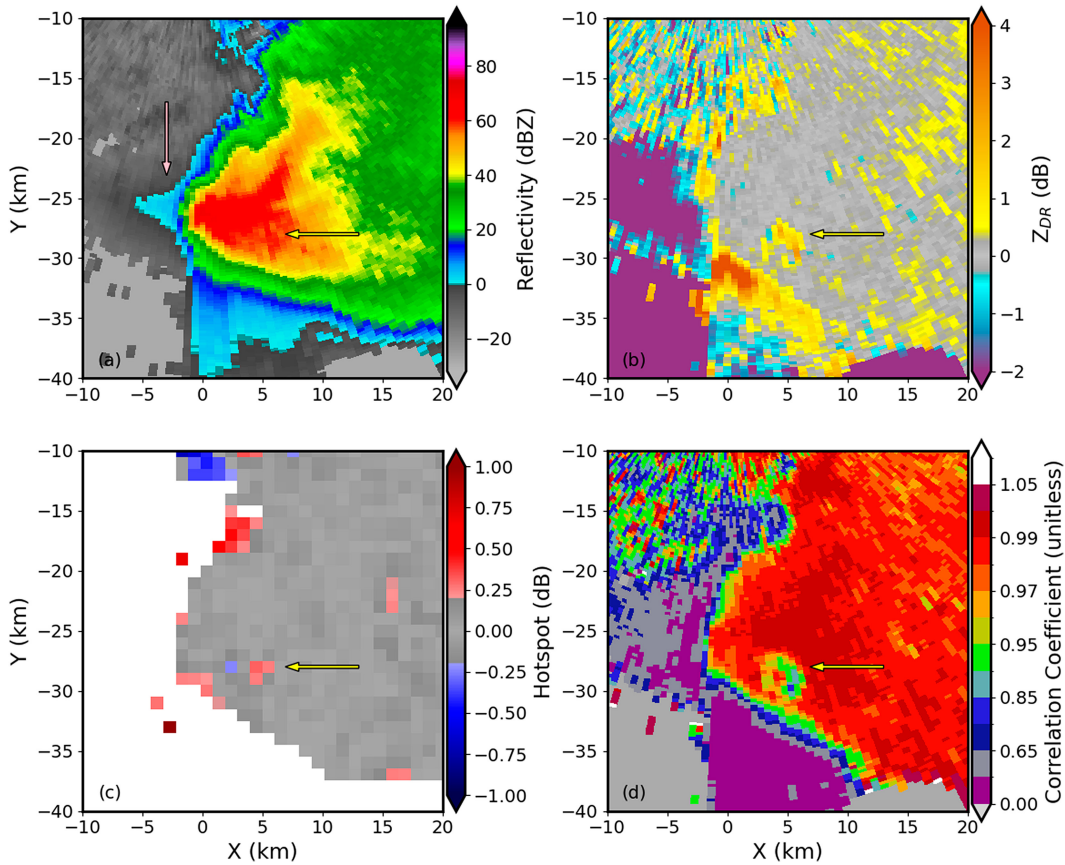


FIG. 12. (a) Reflectivity, (b)  $Z_{DR}$ , (c)  $Z_{DR}$  hotspot result, and (d) CC. All data from 0211 UTC 4 May 2021 at KEWX at the height of the  $-10^{\circ}\text{C}$  isotherm (6.131 km AGL). The yellow arrow identifies the hotspot of interest.

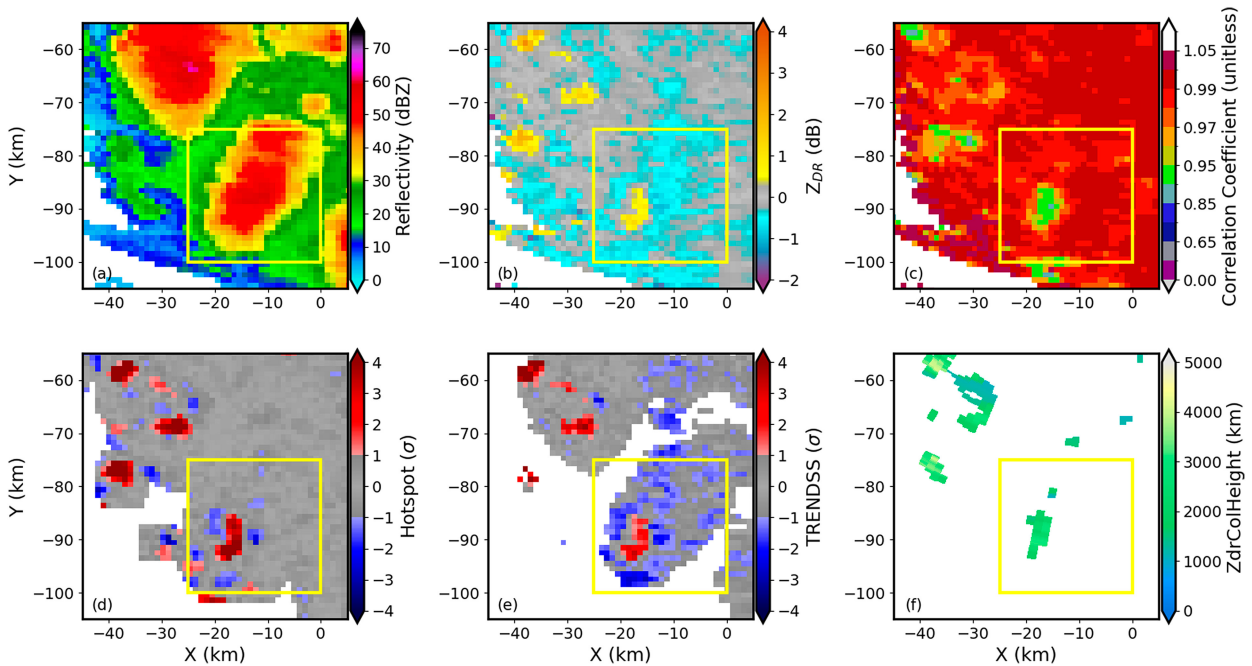


FIG. 13. (a) Reflectivity, (b)  $Z_{DR}$ , (c) CC, (d)  $Z_{DR}$  hotspot result, (e) TRENDSS result, and (f) S15  $Z_{DR}$  column result. All data from 0447 UTC 23 May 2020 at KFDR at the height of the  $-10^{\circ}\text{C}$  isotherm (5.547 km AGL).

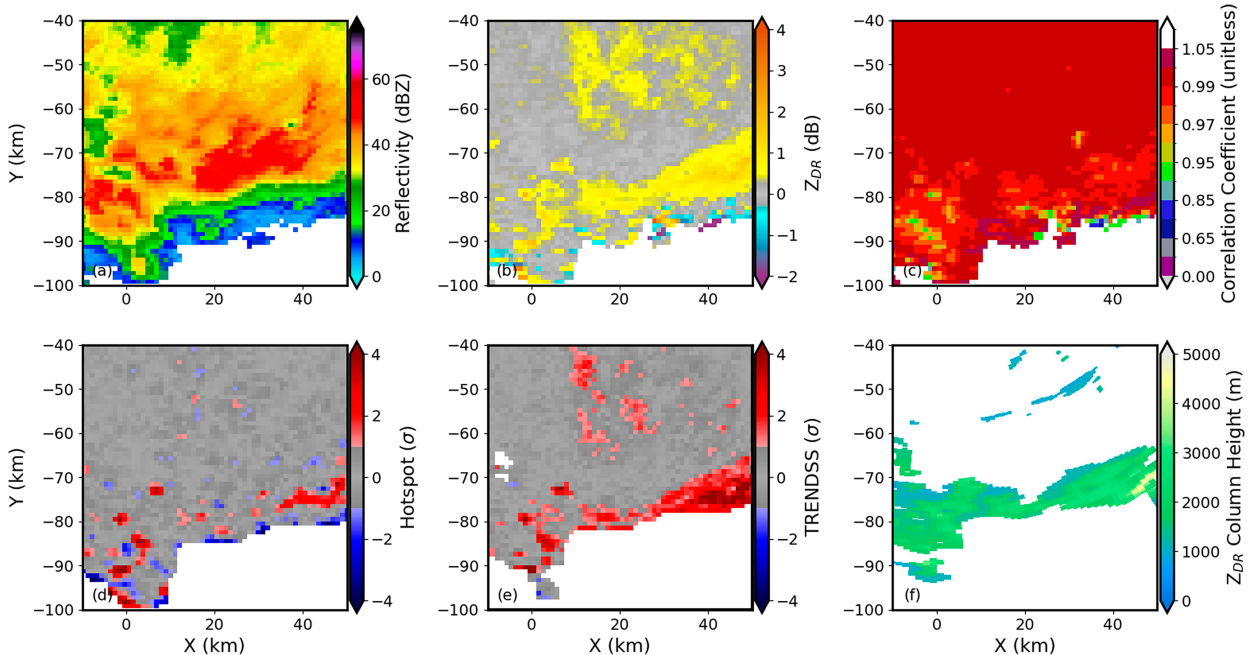


FIG. 14. As in Fig. 13, but from 0455 UTC 4 Jun 2014 at KOAX at the height of the  $-10^{\circ}\text{C}$  isotherm (5.188 km AGL).

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**Data availability statement.** All radar data used in this study are publicly available from National Centers for Environmental Information (NCEI) at <https://www.ncei.noaa.gov/products/radar>.

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