1   # Terrestrial sources as the primary delivery mechanism of

2   # mercury to the oceans across the Toarcian Oceanic Anoxic

3   # Event (Early Jurassic)

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34 **ABSTRACT**

35   This study evaluates the utility of sedimentary mercury (Hg) contents as a proxy for
36 fingerprinting ancient massive volcanism, which is often associated with biogeochemical
37 perturbations. Herein we present new Hg geochemical data from anoxic marine basins across the
38 Toarcian Oceanic Anoxic Event (T-OAE; ~183 Ma) as a test of the complex Hg cycle. The T-
39 OAE was likely initiated by the main eruptive phase of the Karoo-Ferrar large igneous province,
40 which caused a subsequent cascade of environmental perturbations and resulting mass extinction.
41 At present the leading interpretation of sedimentary Hg anomalies has been volcanogenic
42 outgassing as the primary source. Our study and compilation results suggest, however, that
43 Hg/TOC anomalies were restricted to shallow-water, and/or proximal environments, while deep-
44 water, more distal depositional settings document no significant Hg-related anomalies.
45 Furthermore, asynchronous stratigraphic deviations in Hg enrichments favor terrestrially sourced
46 materials and local redox variability, rather than direct volcanogenic emissions, as a primary
47 control mechanism. Additionally, Hg isotope signatures from our only study site documenting an
48 Hg anomaly are also consistent with a terrestrial Hg origin during the T-OAE. Therefore, our
49 results suggest that Hg anomalies in the geological record need to be re-evaluated as a "smoking
50 gun" proxy that only infers volcanogenic inputs.
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52 Keywords: Toarcian Oceanic Anoxic Event; carbon isotope excursion; Hg anomalies; Hg

53 isotopes; large igneous provinces; mass extinction

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1. Introduction

Mercury (Hg) enrichments have recently been identified as a novel proxy to track past events of large-scale volcanism (e.g., Sanei et al., 2012). In the geologic record, the emplacements of large igneous provinces (LIPs) are hypothesized as the lynchpin behind many first- and second-order mass extinction events (e.g., Courtillot, 1999) and other episodes of environmental perturbation. It is thought that volcanogenic outgassing during LIP emplacement catalyzes a cascade of environmental and biogeochemical feedbacks, which have a dramatic impact on the biosphere (e.g., Courtillot, 1999; Bond and Wignall, 2014; Bond and Grasby, 2017). One of the fundamental issues facing the correlation of sedimentological, geochemical, and biological records is unequivocally linking the timing, magnitude, and effects of LIP activity, because converging geochemical processes can often yield non-unique signals in the sedimentary record.

Numerous studies have applied the Hg proxy to Phanerozoic extinctions and oceanic anoxic events (OAEs) associated with documented or hypothesized LIP emplacements, and have noted contemporaneous Hg concentration ([Hg]) anomalies and/or increases relative to total organic carbon (Hg/TOC; e.g., Sanei et al., 2012, Percival et al., 2015; Thibodeau et al., 2016; Grasby et al., 2017; Scaife et al., 2017; Wang et al., 2018; Fantasia et al., 2018; see SI for a list of all ancient Phanerozoic Hg anomalies linked to volcanic activity). The strong affinity of Hg with organic matter (e.g., Ravichandran, 2004) makes it necessary to report Hg/TOC values when identifying Hg anomalies, a standard practice when documenting anthropogenic enrichments in modern sediments (e.g., Fitzgerald et al., 2007). These studies have concluded that LIP injections of CO_2, CH_4, or SO_x into the atmosphere caused the planet to warm and/or cool, depending on the timescale, which then affected the biosphere through environmental and

88 biogeochemical feedbacks (e.g., Courtillot, 1999; Pálfy and Smith, 2000). These suggested

89 feedbacks may represent extinction mechanisms, so understanding the link between LIPs and

90 biological evolution remains critical to resolving the relationship between global environmental

91 change and life.

92   The Early Jurassic Toarcian OAE (T-OAE; ~183 Ma) is a well-documented, transient

93 climate event that was potentially triggered by the initial emplacement of the Karoo-Ferrar LIP,

94 and resulted in global-scale environmental change, culminating in marine mass extinction/s. The

95 T-OAE is characterized by a globally synchronous negative carbon isotope excursion (N-CIE)

96 that represents a major change in the carbon cycle (e.g., Hesselbo et al., 2000; Izumi et al., 2012;

97 Them et al., 2017a; c.f. Them et al., 2018). Other paleoenvironmental changes associated with

98 the T-OAE are an acceleration of the global hydrological cycle and rates of continental

99 weathering (e.g., Percival et al., 2016; Them et al., 2017b; Izumi et al., 2018), increased biomass

100 burning as a result of wildfires (e.g., Baker et al., 2017), carbonate platform demise (e.g., Trecalli

101 et al., 2012; Han et al., 2018), global oceanic deoxygenation (e.g., Them et al., 2018),

102 widespread pyrite burial due to expansion of euxinic (anoxic and sulfidic) conditions (e.g., Gill

103 et al., 2011), and a multi-phased marine mass extinction interval (e.g., Caruthers et al., 2013).

104   Constraining the timing and magnitude of the influence of the Karoo-Ferrar LIP (e.g.

105 Pálfy and Smith, 2000; Sell et al., 2014) is pivotal to our understanding of the effect of massive

106 volcanic eruptions on environmental changes and biological extinction/turnover during this time.

107 Previously published Pliensbachian-Toarcian Hg datasets show variable Hg enrichments (i.e.,

108 enrichment magnitudes and asynchronous stratigraphic enrichment variabilities) during the T-

109 OAE interval (e.g., Percival et al., 2015; Fantasia et al., 2018; Fig. 1); these are concurrent,

110 however, with lithological transitions (i.e., carbonates to organic matter-rich shales, mudstones to

111 clay-rich sandstones, or other similar changes). These Hg enrichments are directly linked to

112 volcanogenic Hg emissions associated with the Karoo-Ferrar LIP (e.g., Percival et al., 2015;

113 Fantasia et al., 2018), although other lines of evidence show increased LIP production and global

114 oceanic deoxygenation before the initiation of the classical defined T-OAE interval (e.g., Moulin

115 et al., 2017; Them et al., 2018).

116

117 **2. Mercury Cycle**

118   Today, Hg is emitted to the atmosphere from several sources including the oceans,

119 biomass burning, rock weathering, soil erosion, volcanism, and anthropogenic activities (Fig. 2).

120 Elemental Hg (Hg^0) is the dominant form of Hg in the atmosphere. The residence time of

121 atmospheric Hg^0 is 0.5 – 1 yr (e.g., Selin, 2009; Driscoll et al., 2013), which is less than the

122 mixing time of the hemispheres (ca. 1–1.5 yr), and therefore it can be globally transported

123 relatively rapidly (e.g., Schuster et al., 2002). Atmospheric Hg^0 is removed through two

124 pathways: [1] transformation to oxidized Hg(II) species, which are particle reactive and soluble,

125 and easily deposited through wet/dry deposition; and [2] direct uptake by vegetation and soils

126 followed by oxidation. Mercury enters the ocean primarily through direct atmospheric deposition

127 and terrestrial riverine runoff (e.g., Holmes et al., 2010; Amos et al., 2014). Mercury in major

128 rivers is particulate-bound, although it becomes decoupled from lithogenic materials after initial

129 deposition (e.g., Zhu et al., 2018); thus, normalizing Hg to lithogenic tracers (i.e., aluminum and

130 titanium; e.g., Scaife et al., 2017; many others) can be problematic. Approximately 25% of

131 riverine-derived Hg might reach the open ocean, with the majority being deposited in shallow

132 proximal locations such as estuaries and nearshore ocean margins (e.g., Amos et al., 2014).

133 Atmospheric Hg also contains more soluble Hg species (e.g., particulate bound Hg(II)), which

134　can be deposited and then re-volatilized to the atmosphere, similar to Hg(0). Importantly, these

135　processes can recycle initial Hg inputs several times through the environment (e.g., Driscoll et

136　al., 2013) before it ultimately enters the ocean where it is scavenged by organic-rich particles,

137　and sinks to the seabed (as reviewed by Fitzgerald and Lamborg, 2014). This results in an

138　average oceanic/atmospheric residence time – from source to ultimate sink – of ~3 thousand

139　years (e.g., Selin et al., 2008).

140　　While volcanoes are the dominant input to the exogenic Hg cycle on geologic timescales,

141　the cycling of Hg through many intermediate reservoirs (Fig. 2) can make it difficult to identify

142　the primary source(s) of bulk sedimentary Hg (e.g., Driscoll et al., 2013). Enrichments alone

143　(i.e., Hg anomalies relative to pre-anomaly values) have proven to be a tenuous fingerprint for

144　the sources of Hg in the modern, thus several studies have used Hg isotopes. Mercury has seven

145　natural stable isotopes (^{196}Hg, ^{198}Hg, ^{199}Hg, ^{200}Hg, ^{201}Hg, ^{202}Hg, and ^{204}Hg), which can be

146　partitioned via mass-dependent (MDF, reported in δ^{202}Hg) and/or mass-independent (MIF,

147　reported in Δ^{199}Hg) fractionations during its cycling (e.g., Bergquist and Blum, 2007). To date,

148　large variations of Hg isotopic ratios (~10‰ for δ^{202}Hg and Δ^{199}Hg values) have been reported

149　for natural samples (e.g., Blum et al., 2014). Although δ^{202}Hg has been used as a source tracer in

150　many case studies, it needs to be cautioned that MDF is ubiquitous and occurs during all

151　biological reactions (e.g., reduction, methylation, demethylation), abiotic chemical reactions

152　(e.g., chemical reduction, photoreduction, oxidation), and physical processes (e.g., volatilization,

153　evaporation, adsorption, and dissolution), which may lead to ambiguous interpretations

154　(reviewed by Blum et al., 2014).

155　　The addition of Δ^{199}Hg signatures greatly increases the usefulness of Hg isotopes because

156　they provide a unique fingerprint of specific pathways. It has been suggested that MDF and MIF

157 signatures can be combined to identify the Hg source(s) in the modern (e.g., Blum et al., 2014)

158 and in the ancient to varying effects (e.g., Thibodeau et al., 2016; Grasby et al., 2017; Gleason et

159 al., 2017; Yin et al., 2017). Plotting Δ^{199}Hg vs. δ^{202}Hg clusters data of unique process(es)

160 together (see Figure 1 in Blum et al., 2014). Although modern Hg isotope studies (e.g., Lefticariu

161 et al., 2011; Das et al., 2013; Bessinger, 2015) show that there is an overlap in the isotope

162 composition of natural sources, the MIF signals could be useful in fingerprinting Hg sources (Fig.

163 2). On geologic timescales, it is suggested that Hg was emitted to the surface environment

164 through volcanic and potentially hydrothermal activities, and these sources have shown Δ^{199}Hg

165 values close to zero (Fig. 2). Δ^{199}Hg anomalies have been predominately observed on Earth's

166 surface environment such as soils, sediments, and atmospheric and biological samples, and Hg

167 photoreactions are thought to play the foremost role in the generation of Δ^{199}Hg anomalies in

168 natural samples (e.g., Blum et al., 2014). It is these Δ^{199}Hg differences that may be useful to

169 distinguish between primary and secondary Hg sources.

170

171 **3. Study sites**

172 Here we investigate the dynamics of the Hg cycle across the T-OAE in two

173 independently constrained anoxic basins from northeastern Panthalassa and the European epeiric

174 seaway (Fig. 1). The first in Alberta, Canada, represents a transect in water depth from proximal

175 (East Tributary) to distal (drill cores 1-35-62-5W6 and 6-32-75-20W6) in the Fernie Formation

176 (Them et al., 2017a, b, 2018). The second represents a semi-restricted locality in the European

177 epeiric sea collected from Dotternhausen Quarry, South German Basin. These regions are ideal

178 localities for testing Hg anomalies, as they both have reducing redox conditions during the study

179 interval (Them et al., 2018). Constraining the local depositional conditions is important to

180   quantify Hg enrichments, as Hg is associated with TOC contents of the sediment and is also

181   redox sensitive (e.g., Ravichandran, 2004; Niessen et al., 2003), which represents the first

182   attempt to do this for the Pliensbachian-Toarcian interval. Additionally, the site with the highest

183   and only Hg/TOC anomaly (East Tributary) was chosen for isotopic analysis to potentially aid in

184   constraining Hg sources as it is the only site suggesting Hg source variability (e.g., Blum et al.,

185   2014; Thibodeau et al., 2016, Grasby et al., 2017; Wang et al., 2018). The sites where no

186   Hg/TOC anomalies were identified were not analyzed for Hg isotopic compositions, as they

187   suggest minimal change in the Hg cycle and confirm the null hypothesis (i.e., there is no change

188   in Hg/TOC across the T-OAE).

189

190   **4. Methods**

191   *4.1 Mercury Concentrations*

192   Splits of samples from previous studies of the Pliensbachian and Toarcian stages of the

193   Early Jurassic (e.g., Them et al., 2017a,b; Gill et al., 2011) were analyzed for Hg mercury

194   concentrations in the School of the Environment at Florida A&M University. Approximately 50-

195   100 mg of sample powder was added to a measuring boat and placed into a DMA-80 (Milestone

196   Inc, Shelton, CT) and heated in stages. Volatilized Hg was collected by gold amalgamation

197   before being measured via spectral analysis. All samples contained quantifiable mercury

198   concentrations. Data were calibrated using international reference standards (SRM Spinach

199   1570a, SRM Mussel 2976, BCR-60 Aquatic Plant, and TORT-2 SRM; see Supplementary

200   Information and SI Table 1 for values), which bracketed the full range of mercury concentrations

201   calculated in these geologic samples. TOC values for these samples were previously published

202   (e.g., Them et al., 2017a,b; Dickson et al., 2017). Mercury data collected by this technique

203   delivered the same values as the samples that were chosen for Hg isotopes and calculated using

204   different chemical processing and analytical instrumentation (see below).

205

206   *4.2 Mercury Isotopes (δ^{202}Hg and Δ^{199}Hg)*

207   Splits of samples from the East Tributary section of Canada were analyzed for Hg

208   isotopes at the Laboratory of Advanced Environmental Analysis, Hong Kong Polytechnic

209   University. Approximately 0.5 g sample was digested in 5 mL of aqua regia (HCl:HNO$_3$ = 3:1,

210   v:v) using water bath (95 °C) for 1 hour. Mercury in the digest solutions were diluted to 0.2 to

211   0.5 ng mL^{-1}, and the acid concentrations were adjusted to 20% (v/v). Mercury and acid matrices

212   of the bracketing Hg standard (NIST SRM 3133) were matched to the sample solutions. Digests

213   of standard reference materials (MESS-1, ocean sediment) were prepared, and measured in the

214   same way as the samples. UM-Almadén secondary Hg standard solution containing 0.2, 0.5, and

215   1.0 ng mL^{-1} of Hg in 20% (v/v) aqua regia was also measured. No statistical differences in Hg

216   isotopic compositions were observed for solutions with different Hg concentrations (see SI Table

217   1), and the results for UM-Almadén and MESS-1 agreed well with previous studies (e.g., Blum

218   et al., 2007; Blum and Bergquist, 2007; Donovan et al., 2013). Uncertainties in the data reported

219   in this study reflect the larger values of either the external precision of the replication of the UM-

220   Almadén or the measurement uncertainty of MESS-1. The Hg concentration and isotope data are

221   included in the SI data.

222   Mercury isotope ratios were measured by a Neptune Plus multi-collector inductively

223   coupled plasma mass spectrometry (MC-ICP-MS). The MC-ICP-MS was equipped with a gas-

224   liquid phase separator and an Apex-Q desolvation unit (Elemental Scientific Inc., USA) for Hg

225   and thallium (Tl) introduction, respectively. Briefly, SnCl$_2$ was continually pumped along with

226 Hg(II) sample solutions and allowed to mix prior to being introduced to a gas/liquid separator,

227 producing gaseous elemental Hg(0). Using the gas/liquid separator, a relatively simple and dry

228 sample stream contain just Ar and Hg(0) was added to the instrument, which minimized any

229 possible matrix effect. It is possible that the wet digestion may cause incomplete reduction of

230 Hg(II) by $SnCl_2$, which may cause Hg isotope fractionation. However, this is not the case for our

231 samples. Hg concentrations of our samples were monitored using the ^{202}Hg signals of MC-ICP-

232 MS, and the total Hg estimated by ^{202}Hg signals were 86 – 107% to that measured by the Hg

233 analyser at the School of Environment at Florida A&M University (see above). Detailed methods

234 for conducting the MC-ICP-MS analysis were described previously (Geng et al., 2018). The

235 signals for ^{202}Hg were <0.02 V for acid blanks. The sensitivities of ^{202}Hg during Hg isotope

236 analysis was 1.4 to 1.7 V per ng mL^{-1} Hg. Mass dependent fractionation of Hg isotopes is

237 expressed in δ^{202}Hg notation in units of per mil (‰) referenced to the NIST-3133 Hg standard

238 (analyzed before and after each sample):

239 $$\delta^{202}Hg\ (‰) = [(^{202}Hg/^{198}Hg_{sample})/(^{202}Hg/^{198}Hg_{standard}) - 1] \times 1000 \quad (1)$$

240 Mass independent fractionation of Hg isotopes is reported in Δ notation (Δ^{xxx}Hg), which describe

241 the difference between the measured δ^{xxx}Hg and the theoretically predicted δ^{xxx}Hg value using

242 the following formula:

243 $$\Delta^{xxx}Hg \approx \delta^{xxx}Hg - \delta^{202}Hg \times \beta \quad\quad (2)$$

244 where β is equal to 0.2520 for ^{199}Hg, 0.5024 for ^{200}Hg, and 0.7520 for ^{201}Hg (e.g., Blum and

245 Bergquist, 2007).

246

247 **5. Results**

248 In northeastern Panthalassa, the most proximal shallow-water site (East Tributary section;

249 n = 36) shows an increase in [Hg] from 0.03 to 0.24 mg/kg (or ppm) and Hg/TOC of 0.02 to 0.05

250 ([mg/kg]/TOC), with the highest values coincident with the N-CIE interval (Fig. 3). During the

251 N-CIE, δ^{202}Hg and Δ^{199}Hg values (n = 8) become more negative (~-0.5 to -1 ‰ and ~0.1 to -0.2

252 ‰, respectively; Fig. 3). The intermediate water depth site (core 1-35-62-20W5; n = 18), shows

253 an [Hg] increase at the base of the N-CIE from one sample, but lacks a corresponding Hg/TOC

254 anomaly due to simultaneous TOC enrichment (Fig. 3). The distal and deepest-water site (core 6-

255 32-78-5W6; n = 17) shows a decrease in [Hg] and Hg/TOC across the N-CIE (Fig. 3).

256 In the European epeiric sea Dotternhausen section (n = 29), a significant increase in [Hg]

257 from ~0.05 to 0.35 mg/kg occurred during the N-CIE, but again lacks a coincident Hg/TOC

258 anomaly as TOC is also enriched (Fig. 4). At this site there are higher Hg/TOC values below the

259 N-CIE, but these reflect extremely low and variable TOC contents (see SI Appendix) as there are

260 no corresponding Hg enrichments (Fig. 4).

261

262 **6. Discussion**

263 *6.1 Evaluating Volcanism as a Direct Driver of Toarcian N-CIE Hg Enrichments*

264 Mercury data from the Panthalassa transect show that only the most proximal location

265 (i.e., East Tributary) documents minor Hg enrichments (as compared to some values from

266 Europe) in both [Hg] and Hg/TOC during the N-CIE (Figs. 3, 5, and 6). As Hg/TOC values

267 increase during the N-CIE, there is a decrease in δ^{202}Hg and Δ^{199}Hg (~-0.5 to -1 and ~0.1 to -0.2,

268 respectively), which suggests that the source(s) or geochemical process(es) controlling local Hg

269 deposition may have changed during the carbon isotope perturbation. As previously stated, there

270 are currently several known mechanisms controlling Hg cycling. These include: volcanism,

271 biomass burning, increased detrital inputs (rock weathering and soil loss), photoreactions, and/or

272   some combination of the above (Fig. 2). The isotope results from the N-CIE interval are

273   inconsistent with purely volcanic signatures which have Δ^{199}Hg near zero or more positive, and

274   therefore, are more suggestive of a trend toward terrestrial sources (e.g., soils, plants, and peat,

275   which have more negative Δ^{199}Hg values but slightly overlap; Fig. 2). Furthermore, the range of

276   δ^{202}Hg and Δ^{199}Hg values, along with the Hg/TOC anomaly, at this location is within the range

277   of modern and pre-anthropogenic marine sediments (e.g., Gehrke et al., 2009; Gleason et al.,

278   2017 and references within). As such, the Hg in our samples suggests the enrichment mechanism

279   is not sourced from direct volcanic emissions during the N-CIE. However, this result does not

280   rule out terrestrial cycling of previously deposited volcanogenic mercury. The recycling of

281   terrestrial Hg is likely related to climatic feedbacks associated with volcanism that could have

282   resulted in increased wildfires (e.g., Baker et al., 2017) and enhanced riverine loading of soil-

283   derived Hg through increased weathering (e.g., Percival et al., 2016; Them et al., 2017b; Izumi et

284   al., 2018). Additionally, at the East Tributary location, the T-OAE N-CIE interval is also

285   associated with the increased deposition of sand-sized grains (e.g., Them et al., 2017a) and

286   higher Al contents (e.g., Them et al., 2017b), which could be due to a local increase in detrital

287   input from an enhanced hydrological cycle (e.g., Izumi et al., 2018). These isotope data are also

288   similar in magnitude with previously published work for other Phanerozoic extinction events that

289   show minor Δ^{199}Hg perturbations (e.g., Grasby et al., 2017; Wang et al., 2018). It must be noted,

290   however, that the Hg isotope values of the multiple Hg sources overlap (Fig. 2); thus, caution is

291   needed when interpreting these data alone to identify specific source(s) as continental shelf to

292   nearshore environments likely record an amalgamation of these sources.

293   The Dotternhausen Quarry Hg/TOC data remain low during the T-OAE, with values only

294   significantly fluctuating below the N-CIE (Fig. 4). The one-point Hg and Hg/TOC increase in the

295 T-OAE interval is intriguing because it may represent an increase in volcanic activity. This

296 sample, however, is associated with elevated pyrite contents and is likely controlled by local Hg

297 cycling (see Supplemental Information). Notably, the fluctuating Hg/TOC values (0.05 to 0.17)

298 prior to the N-CIE are similar in magnitude to several previous studies that attribute such

299 fluctuations to volcanic activity (e.g., Percival et al., 2015; Scaife et al., 2017; Fantasia et al.,

300 2018; many others). There is no change in [Hg], however, and therefore we suggest that the

301 observed Hg/TOC values are driven by low and fluctuating TOC (<0.3%), which are also

302 associated with lithologic changes (organic matter-lean carbonate marls from 0 to 2 meters; Fig.

303 4). It is unlikely these Hg/TOC values are driven by local redox variability, as water column

304 redox conditions are predominantly reducing (anoxic) for the investigated intervals at this

305 location (e.g., Them et al., 2018; see Supplementary Information).

306   Diagenetic alteration of the primary Hg and TOC values could be a possible mechanism

307 to explain the lack of Hg enrichment at Dotternhausen; we suggest, however, that it would be

308 unlikely to have caused the observed Hg/TOC record. While the effects of diagenesis on Hg are

309 relatively unknown, it is probable that Hg and TOC would be affected similarly if Hg is truly

310 associated with organic matter, but could be incorporated in new mineral phases after dissolution

311 in pore fluids. If the Hg and TOC are decoupled, however, then loss of TOC through thermal

312 decomposition or other similar processes during burial may result in spurious Hg anomalies,

313 which is more likely in localities with highly variable TOC. This would be the case if Hg did not

314 also volatilize at temperatures resulting in thermal decomposition of organic matter. The reverse

315 is also true where loss of Hg through volatilization would result in a muted Hg/TOC signal if

316 there was no organic matter loss during burial and heating. Although possible, there is no

317 evidence to suggest that either of these scenarios resulted in the observed Hg/TOC values.

318 Alternatively, the observed increases in Hg during the N-CIE interval are potentially associated

319 with changes in the type of organic matter at this location (e.g., Schwark and Frimmel, 2004).

320 Changing organic matter sources (i.e., prokaryotic vs. eukaryotic, marine vs. terrestrial, etc.) and

321 the subsequent effect on Hg sequestration is not constrained for any ancient deposits, and

322 therefore moving forward will be an important parameter to constrain for the validity of the Hg

323 proxy, especially for events with large, widespread climatic and biological perturbations.

324   Comparable lithologic variability (low TOC to TOC-rich sediments) and ranges of

325 Hg/TOC values (0.14 to 0.4) to those observed at the Dotternhausen Quarry are also observed in

326 a mid-Pleistocene sediment core from the Mediterranean Sea. The Hg and Hg/TOC anomalies

327 are not linked to volcanic activity, but rather local redox variability and low TOC values

328 (~0.25%; e.g., Gehrke et al., 2009; see Fig. 4 and SI Figs. 1 and 2). This suggests that

329 fluctuations of this magnitude are non-unique in the geological record and can be driven by local

330 or regional Hg cycling.

331   All previously reported Toarcian N-CIE Hg/TOC anomalies occur in successions with

332 low and/or variable TOC contents (see Percival et al. (2015) and Fantasia et al. (2018) for these

333 values) and lithological variability (e.g., Percival et al., 2015; Fantasia et al., 2018; Fig. 5). Given

334 this, caution should be used when interpreting Hg data from low TOC settings (i.e., carbonates,

335 low-TOC siliciclastics, etc.) in the geological record (e.g., Grasby et al. (2016) set a lower limit

336 of 0.2%). This was suggested to minimize false positives, and generally this approaches the

337 lower limits of accuracy and precision for TOC measurements and therefore compromises

338 Hg/TOC values.

339   To better constrain the spatiotemporal record of Hg during the Toarcian N-CIE we have

340 compiled all published Toarcian N-CIE data (e.g., Percival et al., 2015, 2016; Fantasia et al.,

341 2018; this study; Figs. 1 and 5). This new compilation documents two observations: [1] there is a

342 clear proximity relationship to landmasses in sections with Hg anomalies (Fig. 6), and [2]

343 enrichments are asynchronous amongst these localities (Figs. 5 and 6). All of the locations

344 recording Hg enrichments (6 out of 7), highlighted in blue, are in close proximity to landmasses

345 (Mochras Borehole, Bornholm, Peniche, Arroyo Lapa, El Peñon, East Tributary; while Yorkshire

346 is the exception with no Hg/TOC anomaly; Figs. 1, 5, and 6). The deepest and more distal sites

347 have no Hg/TOC anomalies (0 out of 5; Sancerre core, Dotternhausen, Velebit core, and Alberta

348 cores). This suggests that proximity to land may control the Hg content of marine sediments

349 during the N-CIE, which is to be expected if the dominant source of Hg was terrestrial. This

350 seems to contradict the idea that the atmosphere was the dominant, direct source of Hg during

351 the N-CIE, which would be expected to deposit Hg more uniformly throughout the ocean as the

352 atmospheric residence time is close to the mixing time for the hemisphere of emission (Fig. 2).

353   The Hg deposited across a landscape directly from the atmosphere via volcanic eruptions

354 would also be expected to be redistributed/cycled during its transport within the intermediate

355 reservoirs (i.e., plants and soils; Fig. 2) before reaching estuarine and marine sediments.

356 Importantly, these reservoirs would have continued to host and cycle Hg long after its initial

357 deposition (e.g., Driscoll et al., 2013; Amos et al., 2013; Zhu et al., 2018; Fig. 2). The increase of

358 Hg from other terrestrial reservoirs (Fig. 2) could therefore mask any direct volcanic Hg signal.

359 Thus, the pattern of the N-CIE Hg enrichments suggests that the observed Hg anomalies most

360 likely reflect changes in regional riverine input. Furthermore, these enrichments would continue

361 well after the initial driver for climatic and environmental change ceased, potentially causing

362 heterogeneity in the timing of Hg burial and enrichment in the sedimentary record (Amos et al.,

363 2014; Fig. 5), an expected outcome of the propagation of some environmental signals to be

364 preserved in marine sedimentary record (e.g., Romans et al., 2016).

365 Increased Hg inputs from rivers (e.g., Amos et al., 2014; Rice et al., 2009) through rock-

366 weathering interactions, soil erosion, and/or biomass burning (e.g., Rice et al., 2009; Grasby et

367 al., 2017) would heterogeneously impact Hg inputs in nearshore environments (e.g., Zhu et al.,

368 2018). The idea of this process exerting an influence on the Toarcian record is supported by

369 evidence that global climates became both warmer and wetter, with increased continental

370 weathering, during the classical T-OAE interval (e.g., Dera and Donnadieu, 2012; Percival et al.,

371 2016; Them et al., 2017b, Izumi et al., 2018). Following this idea, local delivery of Hg through

372 rivers would most directly influence the observed Hg anomalies in proximal locations (e.g.,

373 Percival et al., 2015; Fantasia et al., 2018; this study). Furthermore, modern wildfires can

374 increase the supply of Hg to surface waters and the atmosphere (e.g., Burke et al., 2010). Recent

375 research suggests that a significant increase in sedimentary charcoal abundances is associated

376 with the latter half of the N-CIE interval, indicative of increased wildfires (e.g., Baker et al.,

377 2017), which could also enhance the input of Hg to the oceans on a local to regional scale (Fig.

378 2). Due to their unique accumulation pathways and preservation of Hg and TOC into sediments,

379 one would not necessarily expect to observe increases in charcoal abundances and Hg or

380 Hg/TOC at the same stratigraphic level. A similar relationship with wildfires was suggested for

381 proximal Hg anomalies at the Latest Permian extinction (e.g., Grasby et al., 2017). The increase

382 in Hg deposition from wildfire activity, the magnitude of which intensifies in a warming world

383 (e.g., Pechony and Shindell, 2010), needs to be considered as a probable Hg source during other

384 intervals of climatic upheaval.

385 The proximal locations that do record N-CIE Hg anomalies are not necessarily

386 inconsistent with the interpretation that LIP volcanism resulted in these enrichments, but there

387   are also asynchronous stratigraphic enrichments/depletions in Hg across the N-CIE interval (e.g.,

388   Percival et al., 2015; Fantasia et al., 2018; Fig. 5). It might be expected that an increase in

389   volcanic-derived Hg would result in synchronous stratigraphic Hg enrichments at each location;

390   the observation that Hg enrichments occur at different points in time (using correlations based on

391   carbon isotope stratigraphy and the estimated duration of the 300-500 kyr for the N-CIE interval;

392   e.g. Boulila et al. [2014] and Sell et al. [2014]) in specific locations suggests that local- and

393   regional-scale processes exert a control on the resulting stratigraphic record of Hg (Fig. 5). Thus,

394   the observed asynchronous stratigraphic Hg enrichments (or lack thereof) likely represent the

395   influence of the local/regional cycling of Hg rather than the direct input of Hg from volcanogenic

396   emissions with potential relationships to source-to-sink sedimentary dynamics (e.g., Romans et

397   al., 2016).

398   A striking observation is that there is no Hg/TOC anomaly during the N-CIE interval in

399   the proximal Yorkshire location even though there are small Hg increases (e.g., Percival et al.,

400   2015; Figs. 1, 5, and 6). The absence of a Hg/TOC increase at Yorkshire could be due to many

401   reasons: re-cycling and re-emissions of Hg prior to sedimentation, post-depositional release of

402   Hg due to local redox variability (e.g., Zhu et al., 2018), potential changes in organic matter

403   source affecting Hg bioaccumulation, or water column stratification as is observed at this

404   location (e.g., French et al., 2014). It may also be related to an early Toarcian sea level

405   transgression. Regardless, there will be local deviations to the model put forth to explain the

406   observed Hg/TOC anomalies during the N-CIE interval (Figs. 5 and 6), but it provides the most

407   parsimonious explanation to-date.

408   *6.2 Estimating the amount of mercury release across the T-OAE*

409   It is difficult to quantify the release of mercury from the emplacement of ancient LIPs as

410   there is no equivalent modern analog. A previous estimate by Percival et al. (2015) determined

411   that 150 megatons (Mt) of Hg was released by the K-F LIP based on its size and using the

412   modern Hg/SO_2 emission values from volcanoes. This estimation is much less than the ~3,800

413   Mt release of Hg estimated from the emplacement of the Siberian traps across the Permo/Triassic

414   boundary (e.g., Sanei et al., 2012). These calculations, however, are reasonable first attempts

415   given the number of unconstrained parameters from these ancient LIPs. It is also possible,

416   however, to quantify the increased fluxes of Hg to the ocean/atmosphere system from weathering

417   and wildfires using the modern mass balance (Fig. 2). For example, if global weathering rates

418   increased by approximately two-fold across the T-OAE (e.g., Them et al., 2017b) and global

419   wildfire activity increased by three-fold (c.f., Baker et al., 2017), an additional 210 Mt and 120

420   Mt of Hg, respectively, could be released above pre-perturbation steady state from these two

421   sources assuming a 300-kyr perturbation (see Them et al., 2017b and references within) and a

422   modern mass balance (see Fig. 2). Thus, the potential for increased fluxes should be considered

423   for Hg studies that are associated with other intervals of climatic perturbation and widespread

424   environmental destabilization.

425   *6.3 Mercury isotopes to fingerprint potential source*s

426   As noted above, Hg can be cycled through multiple reservoirs, many of which have

427   isotopic overlap, before it is eventually deposited in the sedimentary record (e.g., Blum et al.,

428   2014; Fig. 2). In near-modern marine sediments, the observed range of paired Hg/TOC and MDF

429   and MIF values (e.g., Blum et al., 2014; Gehrke et al., 2009; Gleason et al., 2017 and references

430   within) is similar to the range observed in the Toarcian record as well as other events in the

431   geologic record (e.g., Blum et al., 2014; Thibodeau et al., 2016; Yin et al., 2017; this study).

432   Notably, these ancient δ^{202}Hg and Δ^{199}Hg data also show less variation than those identified in a

433   modern coastal plain sediment core, where the observed variations were interpreted as a

434   combination of changes in sediment type, grain size, and/or Hg sources (e.g., Das et al., 2013).

435   Although coastal plain biogeochemical processes are not directly analogous to a purely marine

436   system, the potential control of grain size on Hg isotopes should remain ubiquitous in contrasting

437   depositional settings. Furthermore, it has recently been suggested that a negative MIF signal in

438   Mesoproterozoic black shales is indicative of photic zone euxinia (e.g., Zheng et al., 2018),

439   which highlights the need to constrain the local redox when interpreting Hg records. Therefore, it

440   is clear that challenges remain in the use of Hg isotopes to definitively pinpoint the overall global

441   changes in the source(s) of Hg to ancient marine sedimentary rocks if there are significant

442   lithological, grain size, and/or redox variations.

443   Despite these potential complications, compiled Hg isotope values from natural sources

444   of Hg do show negative Δ^{199}Hg values reflecting many terrestrial reservoirs (i.e., plants, soils,

445   rocks), and more positive Δ^{199}Hg values can be sourced from other terrestrial reservoirs (i.e.,

446   plants, soils, rocks) and volcanoes (e.g., Blum et al. (2014) and references within; Fig. 2). As

447   described earlier for the data from the Toarcian, the trends in Hg/TOC anomalies (Figs. 1, 5, and

448   6) also suggest enhanced terrestrial sourcing of Hg in T-OAE Hg cycle. Moreover, it may be

449   possible to more accurately interpret ancient Hg isotope records in geologic settings where local

450   changes in the sedimentary environment are minimal during the interval studied (e.g., Das et al.,

451   2013) as we have attempted to do so with the Toarcian Hg isotope data presented here from the

452   East Tributary section.

453

454   **7. Conclusions**

455   Our new data, when cast in the context of the existing Pliensbachian and Toarcian Hg

456   records, have major implications for the underlying mechanisms driving the Toarcian Hg

457   anomalies. Environmental destabilization during the early Toarcian was most likely driven by

458   multiple feedbacks associated with the emplacement of the Karoo-Ferrar LIP. Our compilation

459   of Hg anomalies along with previous evidence for environmental destabilization during the

460   Toarcian N-CIE suggests the following sequence of events: (1) Main phase emplacement of the

461   Karoo-Ferrar LIP and injection of an unknown quantity of Hg to atmosphere from LIP

462   emplacement; (2) Cycling and redistribution of this Hg in intermediate reservoirs (Fig. 2); (3)

463   Increases in the Hg flux to marine environments from continental weathering, soil erosion, and

464   wildfires; (4) Local changes in water column redox in some locations result in enhanced Hg

465   deposition; (5) LIP emplacement decreases or ceases (or Hg input from volcanism decreases

466   while volcanic activity continues) while enhanced Hg loading in proximal marine environments

467   continues for some time, and biogeochemical cycles and the climate are still perturbed. Events 2

468   – 4 may occur geologically simultaneously but potentially independently from one another.

469   The trends from the Pliensbachian-Toarcian Hg compilation illustrate the need to

470   generate similar records from other events in Earth's history. Warming, weathering, wildfires,

471   redox changes, and volcanism can all potentially affect sedimentary Hg enrichments and isotopic

472   signatures. Despite these complexities in the interpretation of the ancient Hg record, we suggest

473   that LIPs may have been the underlying driver of the observed environmental disturbances that

474   resulted in some Hg anomalies and mass extinctions. Volcanism or LIPs may have introduced

475   significant quantities of Hg to the ocean/atmosphere system, but there are many intermediate

476   steps in the cycling of Hg prior to ultimate burial in the sedimentary record, and other larger

477   reservoirs that respond to climatic perturbations may release Hg on similar timescales (Fig. 2).

478   Therefore, the role of these other mechanisms needs to be considered when interpreting ancient

479   Hg enrichments. To better identify these potential sources and feedbacks in the ancient Hg cycle

480   it is imperative to investigate Hg records from both proximal and distal as well as shallow- and

481   deep-water environments, and from multiple ocean basins with independent evidence for local

482   redox conditions, when possible. Further calibration of the Hg system in modern settings

483   (siliciclastic-dominated, carbonate-dominated, proximal vs distal, fluctuating redox, etc.) will

484   also sharpen our interpretations of the ancient Hg record.

485

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635

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Acknowledgements

649

650   **Competing Interests**

651   The authors declare no competing financial interests.

652

653

654   **Figure Captions**

655   **Figure 1. Global paleogeography of the early Toarcian.** Dark grey represents landmasses,

656   light blue represents shallow seas, and dark blue represents open oceans. Of the 12 locations

657   analyzed across the T-OAE interval (Percival et al., 2015; Fantasia et al., 2018; this study), there

658   are only 6 that have Hg/TOC anomalies. Note that there are three locations within the Alberta

659   study site. This paleogeographic reconstruction was modified from previous versions (e.g., Them

660   et al., 2017a, 2018; Scotese, 2001).

661

662 **Figure 2. Modern Hg cycle and associated Hg-isotope composition of sources to marine**

663 **sediments.** Mercury is recycled in the environment multiple times before being ultimately

664 sequestered in marine sediments. Values in mass balance from Schuster et al. (2018) and

665 references therein. Note the units for Hg reservoirs (Gg Hg) and fluxes (Gg Hg yr^{-1}). Note that

666 δ^{202}Hg and Δ^{199}Hg values of various sources overlap with one another. Mercury isotope values

667 of sources from Blum et al. (2014) and references therein. In the ocean, the sinks of mercury are

668 organic matter and drawdown through sulfide production and preservation. Intermediate sinks

669 that can become sources on short timescales include soils, rocks, and vegetation.

670

671 **Figure 3. Mercury geochemistry of the Alberta sections that represent a basin transect.**

672 Lithostratigraphy, chemostratigraphy, and ammonite biostratigraphy from Them et al. (2017a,

673 2018). Tenuicostat. = Tenuicostatum, carb. = carbonate. East Tributary represents proximal

674 location; core 1-35-62-5W6 represents intermediate location; core 6-32-75-20W5 represents

675 distal location. The proximal East Tributary section is the only section from our Panthalassan

676 study sites that displays an anomaly in Hg/TOC values.

677

678 **Figure 4. Mercury geochemistry of the Dotternhausen section.** Lithostratigraphy, carbon-

679 isotope and TOC chemostratigraphy, and ammonite biostratigraphy from Dickson et al. (2017).

680 Note that the one-point Hg/TOC anomaly during the N-CIE interval is associated with elevated

681 pyrite contents (see discussion in Main Texas and Supplementary Information). Furthermore, the

682 gray circles in the Hg/TOC portion of this section were not interpreted to represent Hg anomalies

683 because their increases are caused by low TOC values.

684

685 **Figure 5. Compilation of sedimentary mercury geochemistry during the Pliensbachian and**

686 **Toarcian.** Lithostratigraphy, carbon-isotope (either organic or inorganic) and Hg/TOC

687 chemostratigraphy across the Pliensbachian-Toarcian interval from Europe, South America, and

688 North America. Blue text represents sites where an Hg/TOC anomaly occurred during the T-

689 OAE interval. Red text represents sites where an Hg/TOC anomaly did not occur during the T-

690 OAE interval.

691

692 **Figure 6. Idealized depositional settings of European, South American, and North**

693 **American locations during Pliensbachian and Toarcian.** (**A**) In the European epeiric sea,

694 multiple semi-restricted to severely restricted basins and emergent landmasses existed. Note that

695 this is not one continuous cross-section, thus it is not depicting the local restriction of the basins

696 as they are connected to other sub-basins (**B**) The two South American locations were directly

697 adjacent to the Laurentian continent and the basins were isolated due to volcanic terranes. (**C**)

698 The North American locations were located in an upwelling zone with strong connections to

699 Panthalassa. The bottom panels depict the study sites with respective max [Hg] and Hg/TOC

700 values during the T-OAE N-CIE interval, as well as their relative distance from any landmasses,

701 small or large. Note that small changes in the relative distance of any study area will not

702 significantly influence the observations. Blue text represents sites where an Hg/TOC anomaly

703 occurred during the T-OAE interval. Red text represents sites where an Hg/TOC anomaly did not

704 occur during the T-OAE interval. A north arrow is shown in each panel. Pan. = Panthalassa,

705 E.E.S. = European epeiric seaway.

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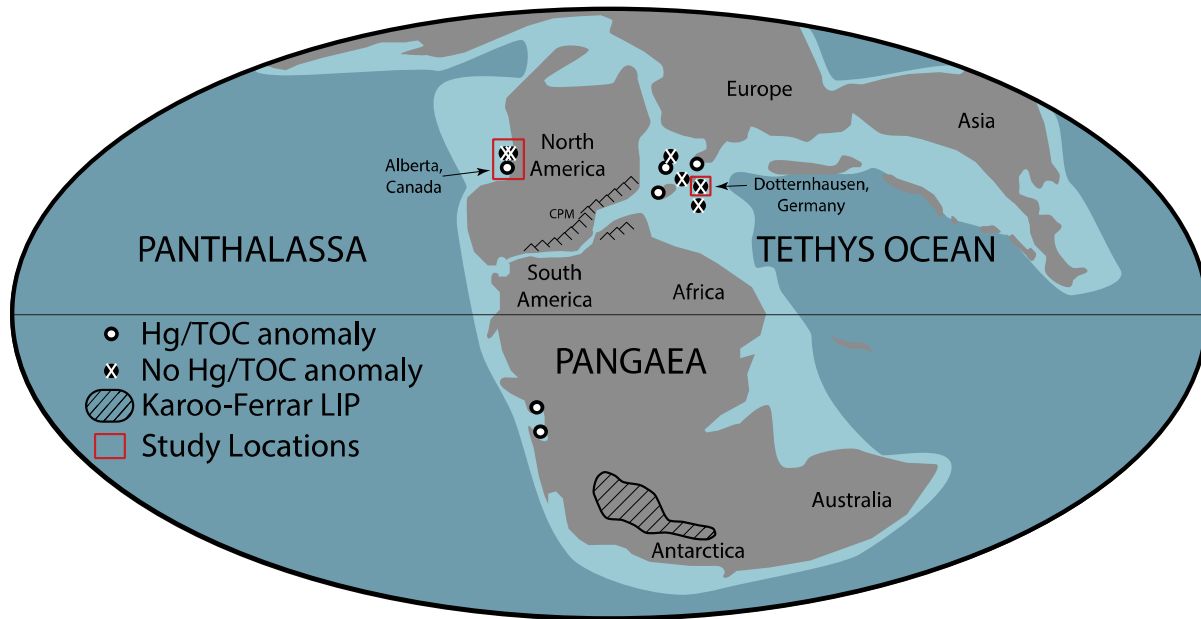
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709   Figure 1.

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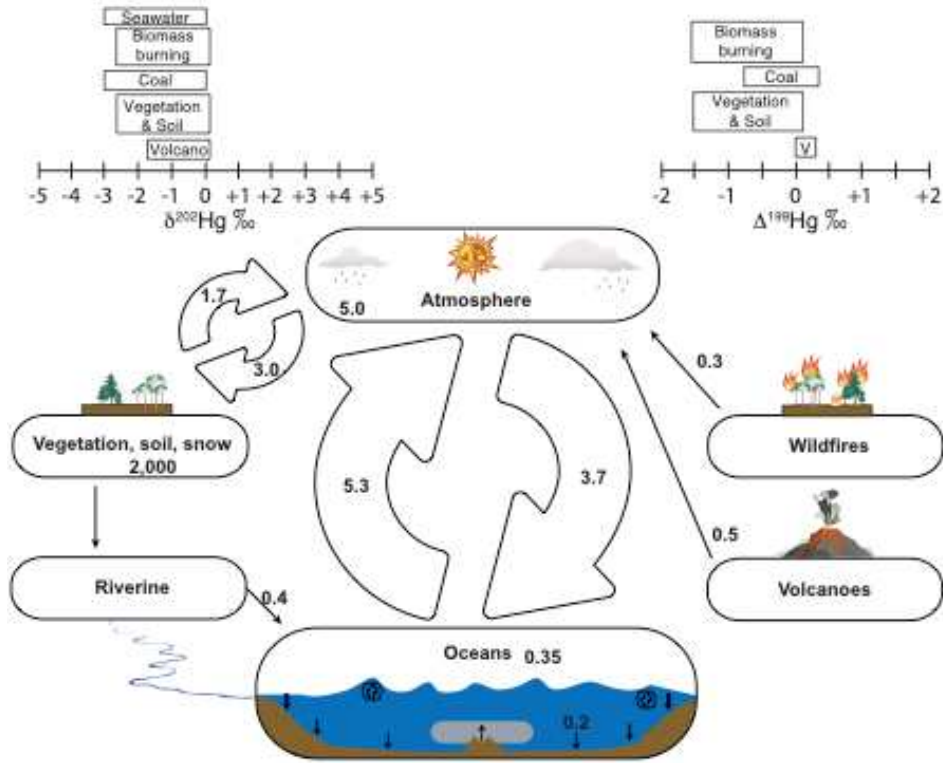
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721   Figure 2.

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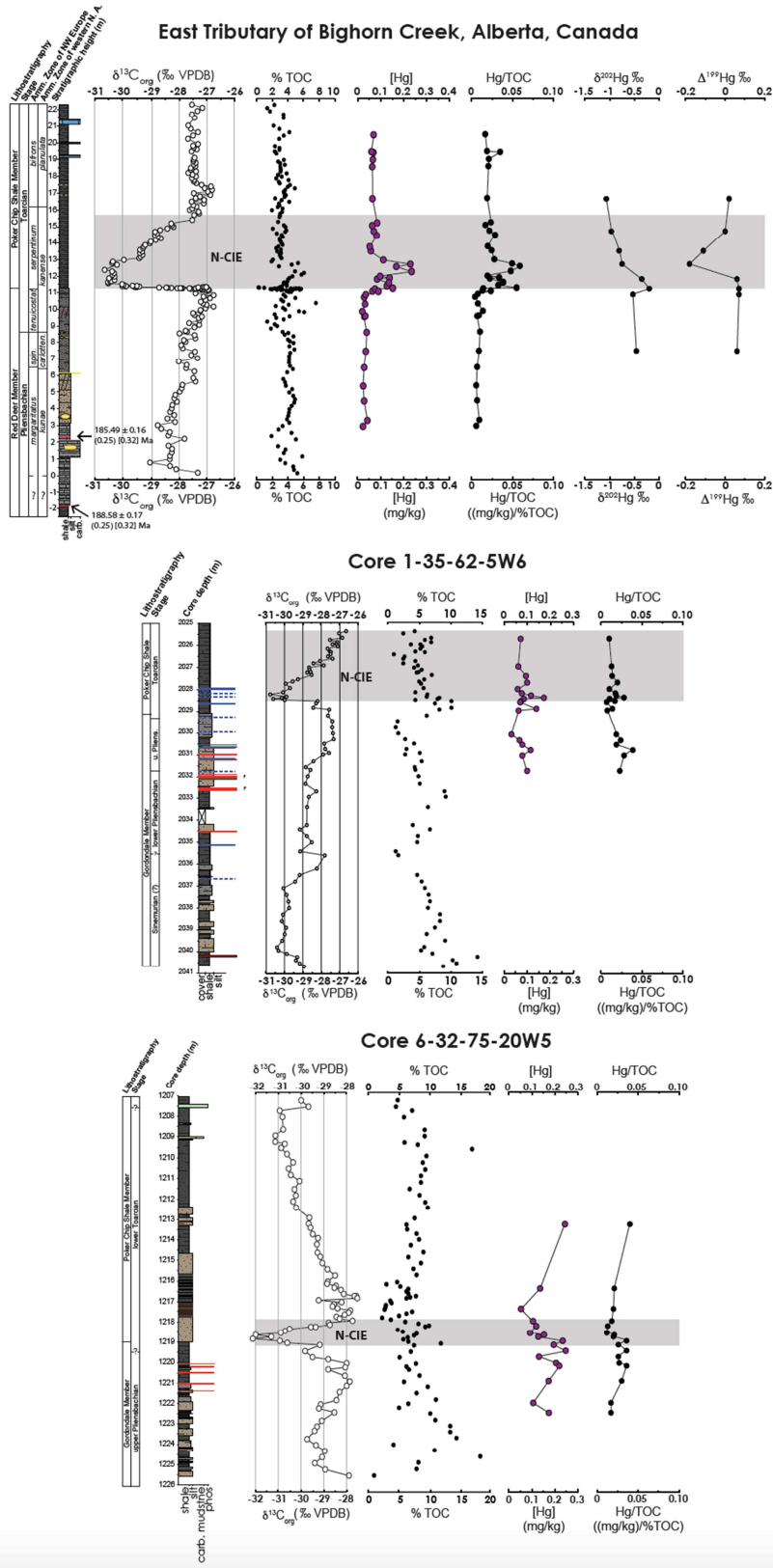
735   Figure 3.
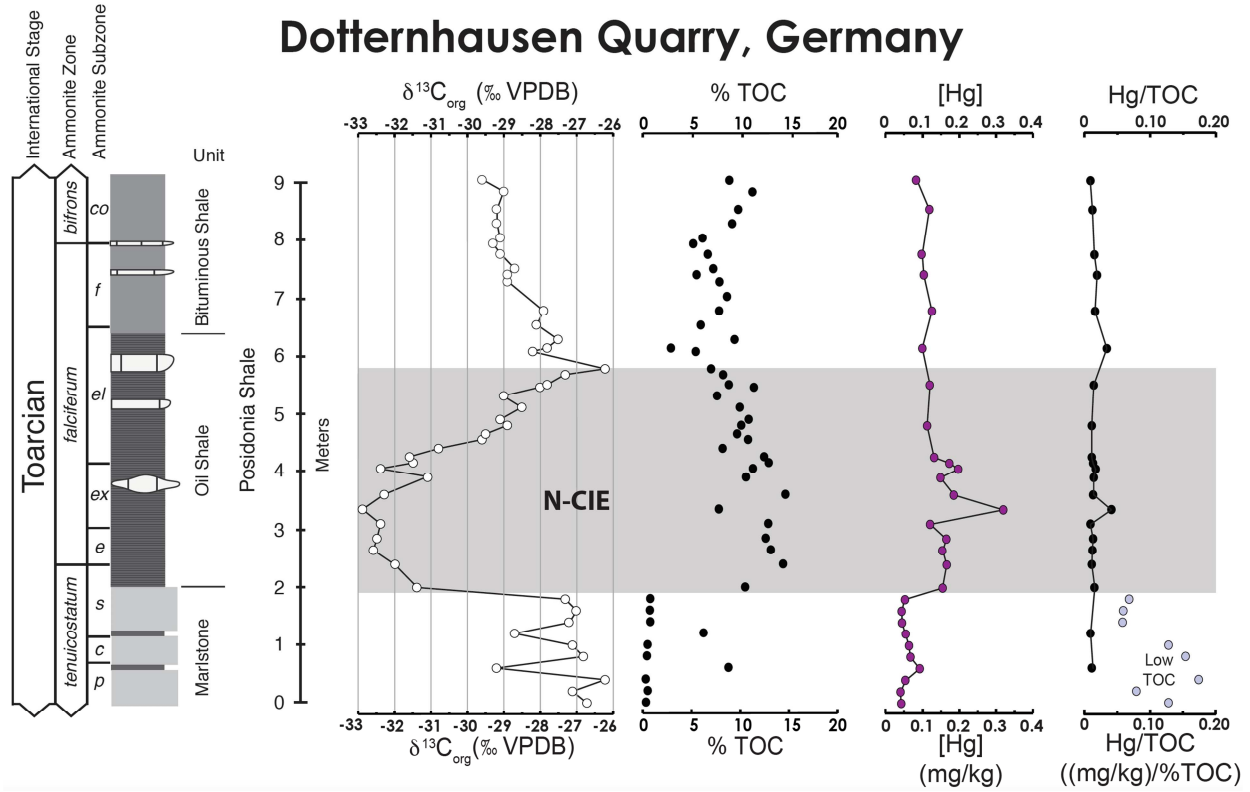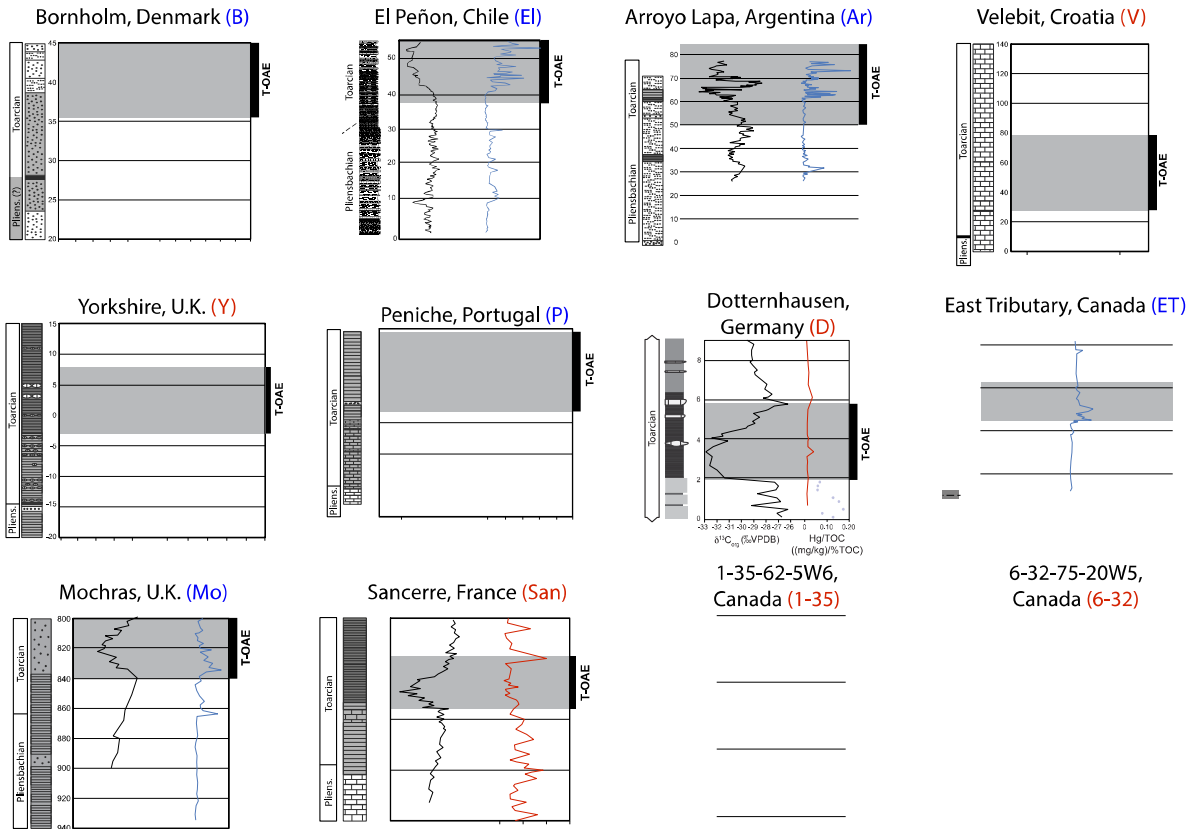
Figure 4.

749

750   Figure 5.

751

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759

760

A — Europe

N←

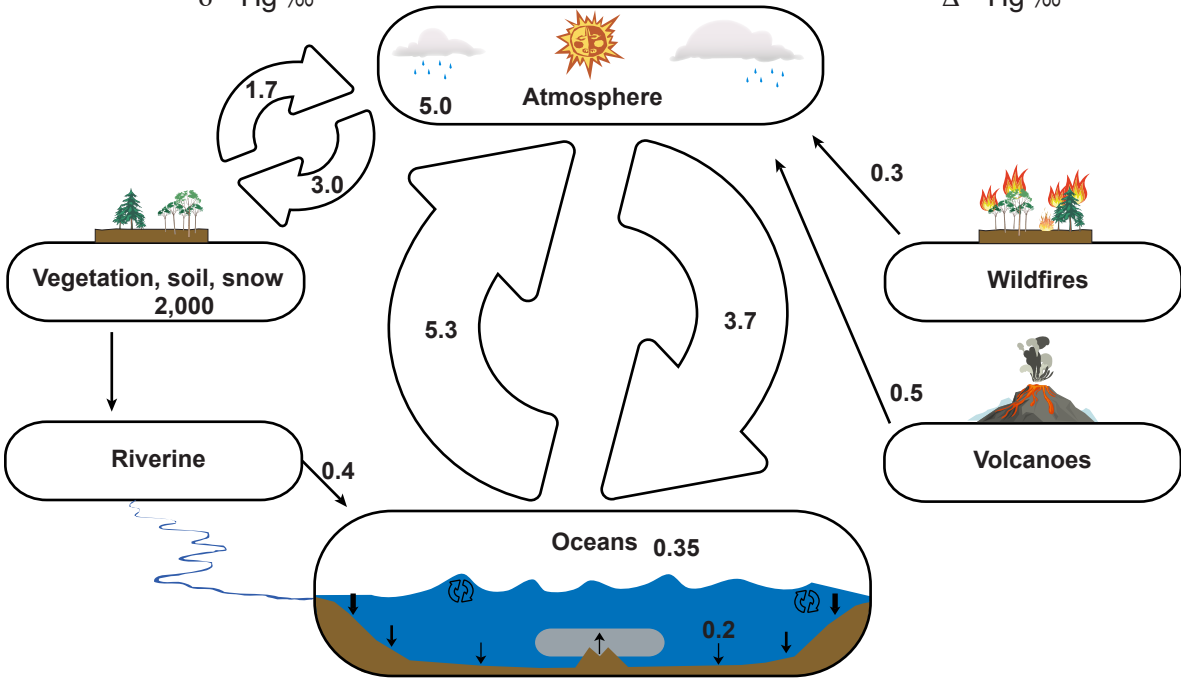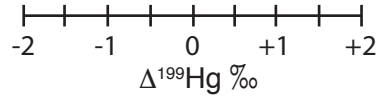*Note that these idealized cross sections do not define whether the basins were restricted.

E.E.S.

Tethys Ocean

— Hg/TOC anomaly during T-OAE   — No Hg/TOC anomaly during T-OAE

B — South America

Ar
El
Pan.
N↑

C — North America

1-35
6-32   ET
Panthalassa
N↑

Max [Hg] (mg/kg) vs Relative distance from landmass

Max Hg/TOC vs Relative distance from landmass

761

Figure 6.

Europe
Asia
Alberta, Canada
North America
Dotternhausen, Germany
CPM
PANTHALASSA
TETHYS OCEAN
South America
Africa
PANGAEA
Australia
• Hg/TOC anomaly
✗ No Hg/TOC anomaly
Karoo-Ferrar LIP
Study Locations
Antarctica

| Seawater | Biomass burning | Coal | Vegetation & Soil | Volcano |

δ^{202}Hg ‰

| Biomass burning | Coal | Vegetation & Soil | V |

Δ^{199}Hg ‰

1.7

3.0

Atmosphere 5.0

Vegetation, soil, snow 2,000

Riverine

Wildfires

Volcanoes

0.3

0.5

0.4

5.3

3.7

Oceans 0.35

0.2

East Tributary of Bighorn Creek, Alberta, Canada
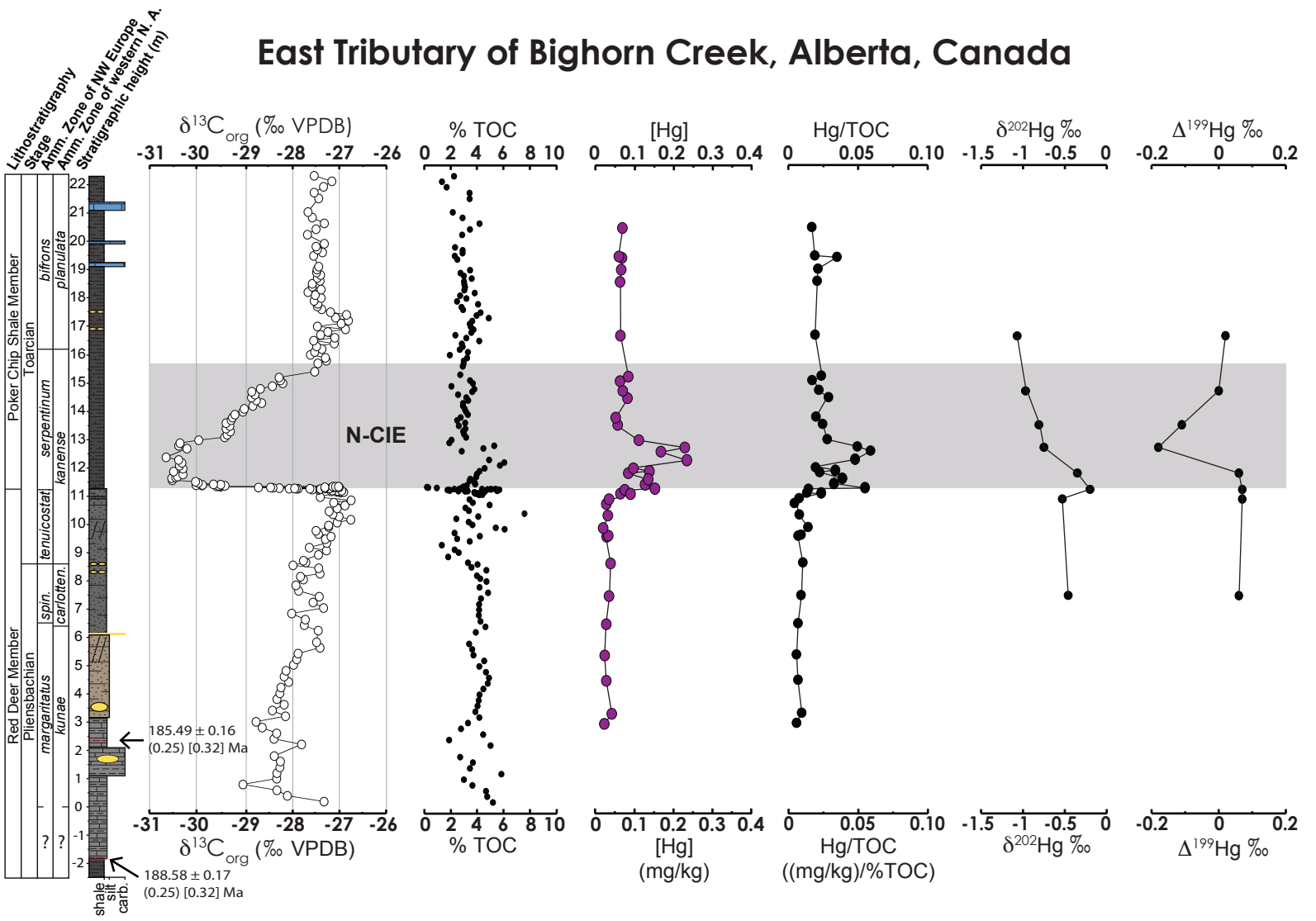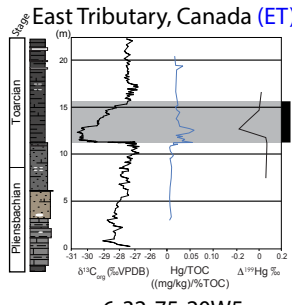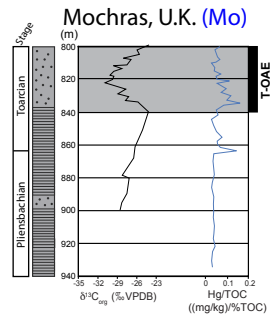

Core 1-35-62-5W6
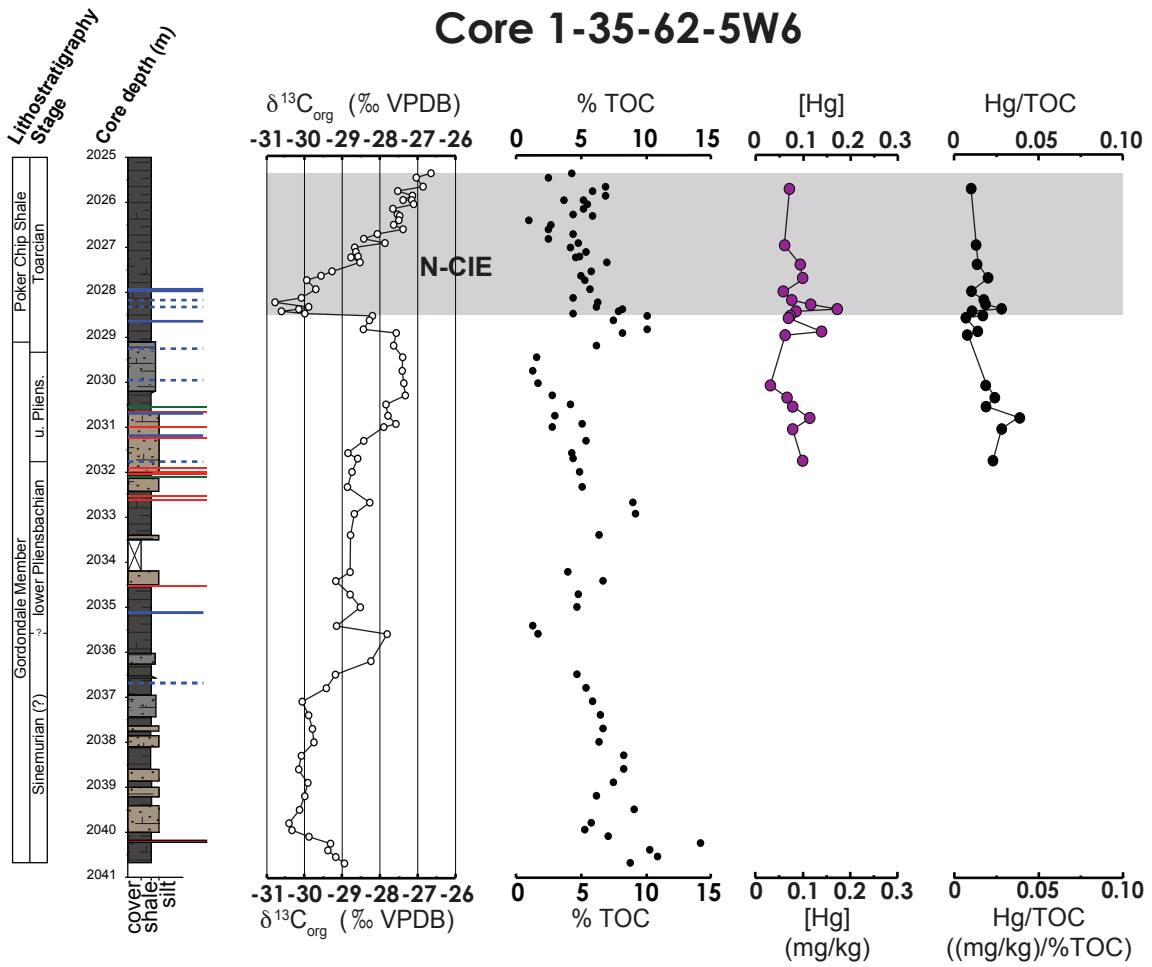
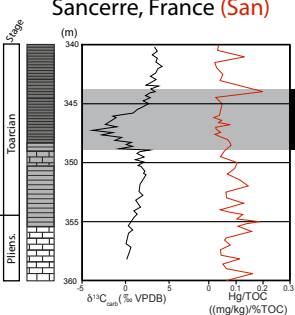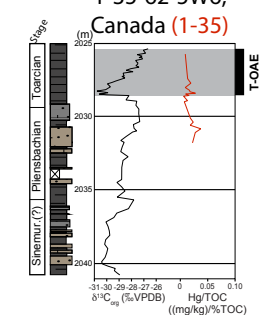## Core 6-32-75-20W5
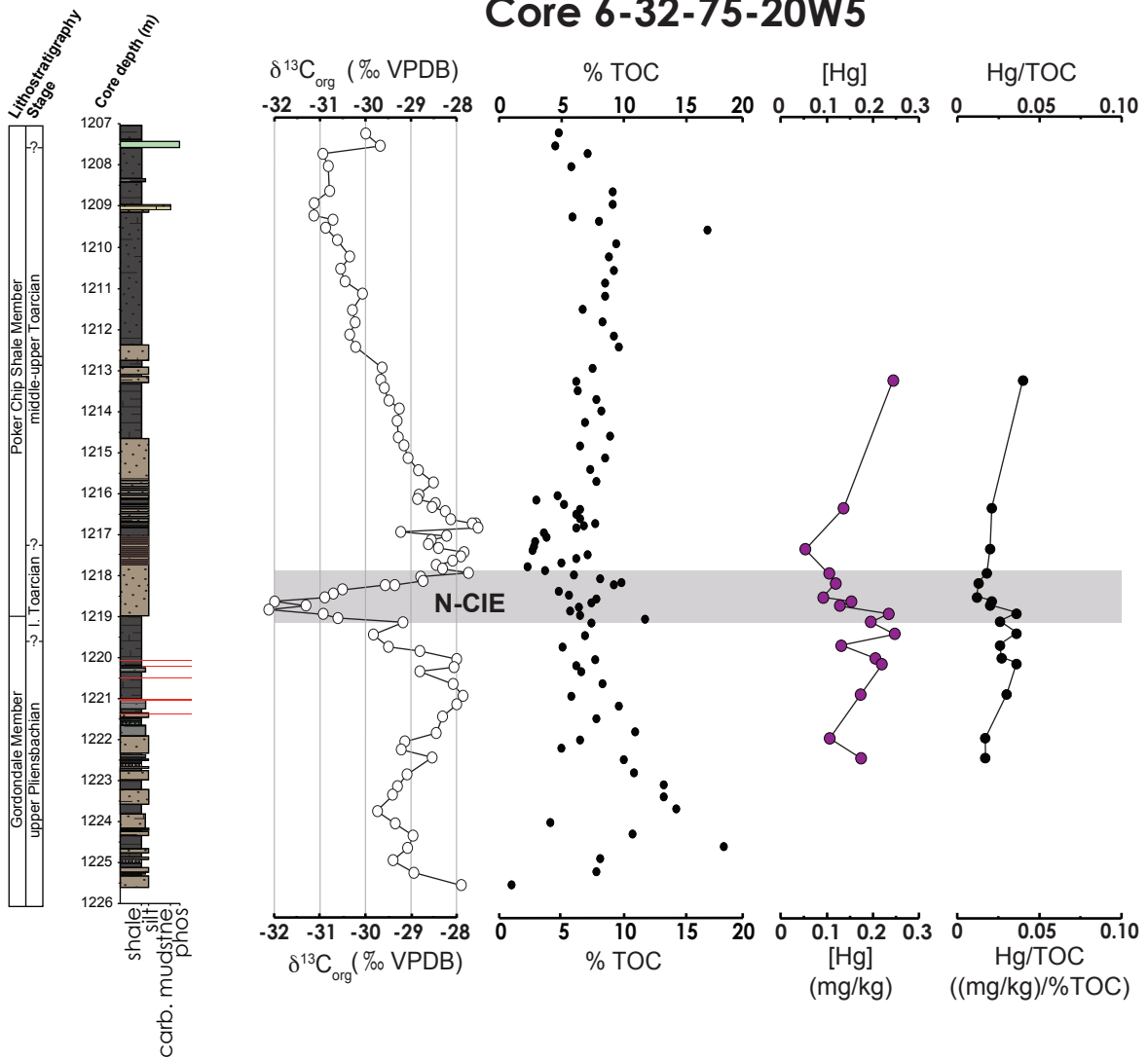
Dotternhausen Quarry, Germany
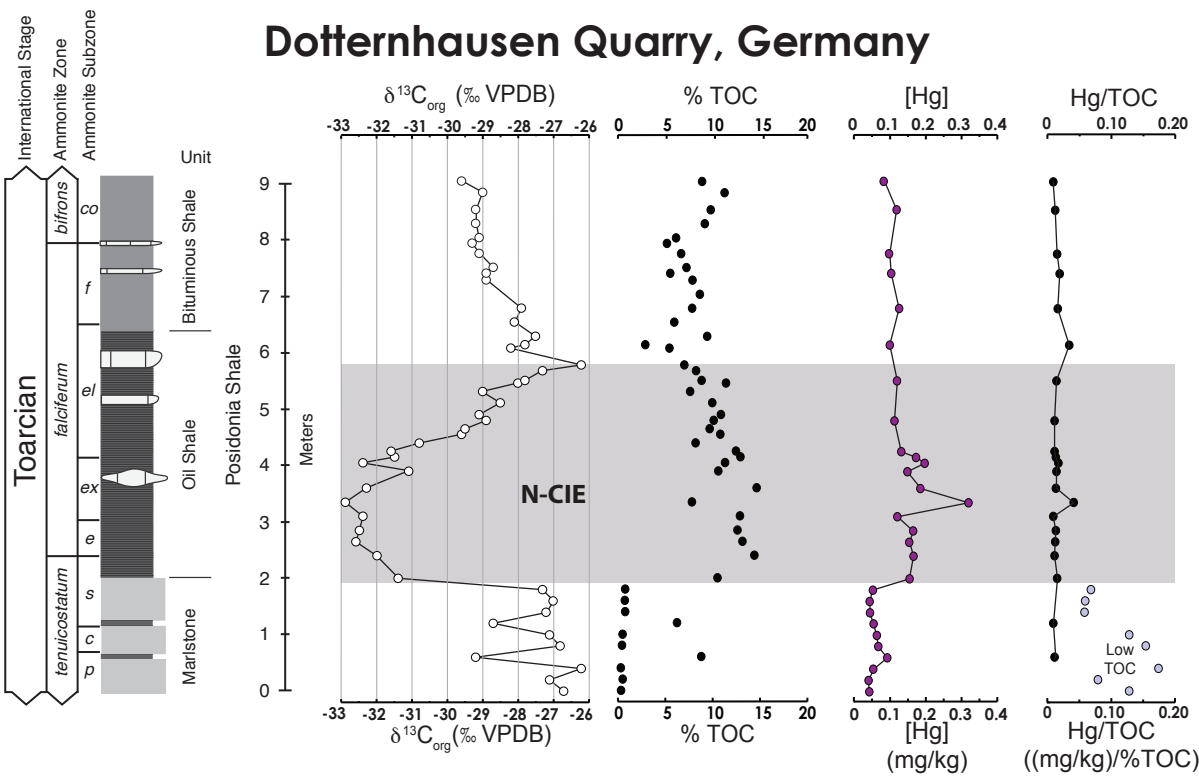
Bornholm, Denmark (B)
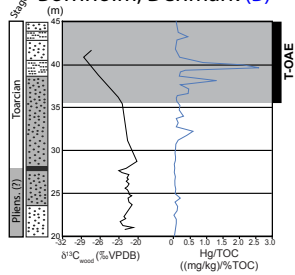
El Peñon, Chile (El)
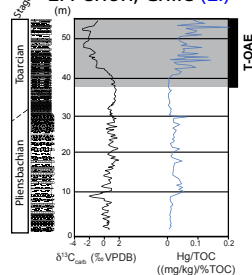
Arroyo Lapa, Argentina (Ar)
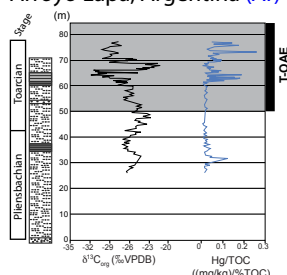
Velebit, Croatia (V)
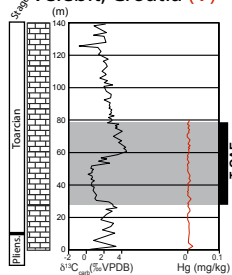
Yorkshire, U.K. (Y)
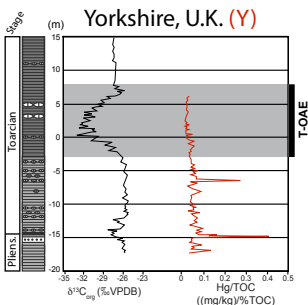
Peniche, Portugal (P)
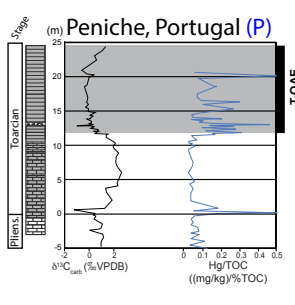
Dotternhausen, Germany (D)
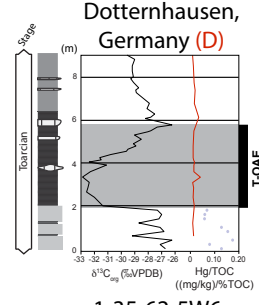
East Tributary, Canada (ET)

Mochras, U.K. (Mo)

Sancerre, France (San)

1-35-62-5W6, Canada (1-35)

6-32-75-20W5, Canada (6-32)

A Europe

N ←

B   P   Mo   San   D   V
         Y

E.E.S.

*Note that these idealized cross sections do not define whether the basins were restricted.

Tethys Ocean

— Hg/TOC anomaly during T-OAE   — No Hg/TOC anomaly during T-OAE

B South America

Ar
El

Pan.

N ↑

C North America

1-35
6-32   ET

Panthalassa

N ↑

Max [Hg] (mg/kg)

0.9 B El
0.8   San
0.7
0.6
0.5   Ar
0.4   D
0.3 Y   ET
0.2 P Mo   1-35   6-32
0.1
0   V

Relative distance from landmass →

Max Hg/TOC

2.5 B
0.5   P
0.4
0.3   Ar
0.2 El   Mo
0.1 Y   ET   D   1-35   6-32 San
0   V

Relative distance from landmass →