

1 **Additional multi-proxy stalagmite evidence from northeast Namibia supports recent models**
2 **of wetter conditions during the 4.2 ka Event in the Southern Hemisphere**

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18

19 **Abstract**

20 The 4.2 ka Event has generally been regarded as a period of decades to at most a few
21 centuries in which comparatively dry conditions existed in the Middle East and more broadly
22 across the mid-latitude Northern Hemisphere. This paper presents new stable-isotopic and
23 petrographic observations from two previously-unreported U-Th-dated stalagmites from Dante
24 Cave in northeastern Namibia. The results are most compatible with wetter conditions during the
25 4.2 ka Event, and wetness during the 4.2 ka Event is the only inference supported by evidence.
26 These new results add to observations previously reported from a third Dante Cave stalagmite
27 suggesting a comparatively wet 4.2 ka Event in which Africa's Tropical Rain Belt migrated
28 southward and rainfall increased along the Congo Air Boundary and/or Kalahari Discontinuity.
29 The new results support findings from three other locations in Namibia and Botswana, from at
30 least seven other locations in the Southern Hemisphere, and at least one in southern China, that
31 suggest a wetter rather than drier 4.2 ka Event in those regions. The pattern emerging from these
32 sites generally agrees with recent modeling results indicating increased moisture over broad areas
33 (but not all) of the Southern Hemisphere. This in turn suggests a 4.2 ka Event that was not a
34 global drought but was instead a set of latitudinally-dependent responses to global-scale
35 southward migration of the Inter-Tropical Convergence Zone (ITCZ), and thus Africa's loosely
36 linked Tropical Rain Belt, as a result of cooling of the Northern Hemisphere, which brought drier
37 conditions to some areas and wetter conditions to others.

38

39 **Keywords**

40 Holocene, paleoclimate, southern Africa, rainfall, Dante Cave, climate change

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43 **1. Introduction: The variability of the 4.2 ka Event**

44 The 4.2 ka Event has been widely investigated both as a significant climatic event and as
45 a driver of societal change (Weiss et al., 1993; Ran and Chen, 2019; Bini et al., 2019). The event
46 was recognized initially in the mid-latitude regions of southwestern Asia (Weiss et al., 1993;
47 Lemcke and Sturm, 1997; Bar-Matthews et al., 1999) but more recently as far north as Iceland
48 (Geirsdóttir et al., 2019) and as far south as southern Patagonia (Ohlendorf et al., 2014).

49 This widespread recognition of the 4.2 ka Event in part hinges on a willingness to accept
50 phenomena that were seemingly not coeval and that were not of the same nature. With regard to
51 the timing of the event, Railsback et al. (2018b) surveyed multiple records and concluded with
52 some skepticism that “if one considers the time-series of data according to their published
53 chronologies, there is significant diachroneity of the event” and that “the duration of the event
54 also differs significantly between records”. With regard to the nature of the event, it was first
55 recognized as a dry phase (e.g., Weiss et al., 1993) but has been reported in other settings as a wet
56 phase (e.g., Li et al., 2018) or, in contrast to both, a time of variable conditions (Kathayat et al.,
57 2017).

58 This range of responses suggests that careful definition of the various expressions of the
59 event is needed. The event was first recognized as a dry phase in paleoclimatological data from
60 the shores of southwestern Asia, specifically in the Gulf of Oman (Cullen et al., 2000) and the
61 Red Sea (Arz et al., 2006), and its chronology has been most firmly established as extending, at
62 least as a single event, from 4.26 ka to 3.97 ka (Carolin et al., 2019). Thus one might speak most
63 strictly and surely of a “Southwest Asian 4.2 ka Dry Event” from 4.26 ka to 3.97 ka. Farther east
64 and west, but in the same latitudinal zone from 20°N to 40°N, the event has been recognized as a
65 dry event from central North American (Forman et al., 1995) to the Mediterranean (Drysedale et
66 al., 2006) to the Himalayas (Nakamura et al., 2016) to Japan (Kawahata, 2019). However, that
67 event is reported across a broader range of time and with some places in that latitudinal zone,
68 such as southeastern China, wetter during that period (Zhang et al., 2018), suggesting a “Mid-

69 latitude Northern Hemisphere ~4.2 ka Generally Dry Event”. Recent work in the southern
70 Hemisphere has suggested wetness around 4.2 ka (e.g., Li et al., 2018; Railsback et al., 2018b),
71 suggesting a still-hypothetical “Southern Hemisphere Variably Wet 4.2 ka event”. Most broadly,
72 it appears that we will be forced to think of a “Globally Variable ~4.2 ka Event”, in contrast to
73 the seeming uniformity of the “global megadrought” of Weiss, (2016).

74 With the evolving context described above in mind, this paper returns to Dante Cave in
75 northeastern Namibia, the site studied by Railsback et al. (2018b), to report further evidence for a
76 wet 4.2 ka Event from two more stalagmites, DAN1 and DP2.

77

78 **2. Setting**

79 *2.1. Physical setting*

80 Dante Cave (19° 24' S; 17° 53' E) is developed in Proterozoic-age dolomitic bedrock of
81 the Otavi Mountainland between the towns of Tsumeb and Grootfontein in the Otjozondjupa
82 region of northeastern Namibia. It is accessed via a single vertical entrance, with a 9 m drop to a
83 debris cone. Steep passages lead to chambers as much as 60 m below the land surface. Sletten et
84 al. (2013), Voarintsoa et al. (2017a), and Railsback et al. (2019) provide more details of the
85 physical setting of the cave.

86

87 *2.2. Climate, moisture sources, and oxygen isotopes*

88 Dante Cave is in southern Africa’s region of summer rainfall, which is greatest in January
89 and February, with about 120 mm each month. In July and August, on the other hand, virtually
90 no rain falls, and June is also typically extremely dry. Temperatures range between an average of
91 26°C in December to 15.2°C in July. In the Köppen classification, the climate of the area is Bsh
92 or hot arid steppe (Beck et al. 2018).

93 Rainfall in northern Namibia in January and February, and more generally in austral
94 summer from November to March, occurs as the Tropical Rain Belt (Nicholson, 2018) migrates

95 southward and/or westward in response to the annual increase in insolation in the Southern
96 Hemisphere. Dante lies consistently south of the Inter-Tropical Convergence Zone (ITCZ), even
97 in summer (Nicholson 2018), and in fact the ITCZ “remains well north of the equator” across
98 Africa all year (Nicholson, 2018).

99 Dante has a more complicated relationship with the Congo Air Boundary (CAB), which
100 marks the convergence of easterly winds from the Indian Ocean and cyclonic winds from the
101 South Atlantic and Congo Basin. Several authors have suggested that the Congo Air Boundary
102 (CAB) has a major impact on rainfall in southern Africa. However, assessing the importance of
103 the CAB in bringing rainfall to Dante is difficult given considerable uncertainty as to its actual
104 location during the austral spring and summer. Figure 1 illustrates the problem because,
105 depending on author, the boundary varies from almost north to south through Botswana
106 (Nicholson 1996, reproduced in Gasse 2000), to largely east to west near or north of the Angola-
107 Namibia border (van Heerden and Taljaard 1998; Chase et al. 2009; Howard and Washington
108 2019).

109 Most recently, the nature of the CAB and its influence on southern African rainfall have
110 been reassessed using the ERA-5 reanalysis dataset (Howard and Washington, 2019) in
111 seemingly the first attempt “to identify the CAB in reanalysis datasets or in global or regional
112 atmospheric models”. In that study, the CAB maintained an ENE to WSW orientation while
113 migrating southwards in austral spring and early summer (Aug to Dec) to a position in southern
114 Angola and southern Zimbabwe. By December (the last month presented in Figs, 2, 4, and 10 of
115 Howard and Washington, 2019), the CAB was still well north of the Angola-Namibia border and
116 was beginning to break down (Fig. 1). A second boundary, the Kalahari discontinuity (KD),
117 marked convergence between the trades and the southerly Namibian nocturnal low-level jet and
118 became more prominent in November and December (Fig. 1). The KD was oriented northwest–
119 southeast from northern Namibia into southwesternmost Botswana and formed the main humidity
120 gradient in November and December as the CAB began to break down. Thus, one understanding

121 of Fig. 1 is that diverse pre-2019 portrayals of the CAB were attempts to represent a synthesis of
122 the CAB (in the strict sense of Howard and Washington, 2019) and the then-unrecognized KD.

123 Because the CAB and KD both divide hotter drier air to the south from moister and often
124 saturated air to the north, both at the surface and aloft, they both have the potential to play a role
125 in interannual variation in rainfall. In addition, breakdown of the CAB and strengthening of the
126 easterlies can trigger the formation of tropical-temperate-troughs (TTTs), which are a major
127 source of rainfall in subtropical southern Africa in summer. Howard and Washington (2019)
128 noted that the early breakdown of the CAB often occurs when the low-level easterlies become
129 more humid. As a result, breakdown of the CAB is often associated with increased rainfall from
130 the Indian Ocean, with the flow of moist air southwards from the Congo, and with a strengthening
131 and westward migration of the KD.

132 Breakdown of the CAB as a result of moister, easterly winds, before the convergence
133 zone reaches Dante Cave, suggests that rainfall at the cave is derived largely from the east, not
134 from the north. Indeed, the isotopic signature of the easterlies has dominated the O-isotope ratios
135 in stalagmite CaCO₃ from northeastern South Africa to northeastern Namibia in the late
136 Quaternary (Fig. 10 of Railsback et al. (2018a), suggesting a large continental effect during
137 transport of vapor westward across southern Africa (see also Section 3.2). Thus, the KD and its
138 weak moisture source from the easterlies may be the best explanation of the source of summer
139 rain in the region around Dante Cave (Fig. 1).

140

141 **3. Methods and criteria**

142 **3.1 Laboratory methods**

143 The two stalagmites from which data are reported here were sectioned and polished along
144 their growth axes, and chips measuring 5.0 cm by 7.5 cm were cut from which thin sections of the
145 intervals near 4.2 kyr in age were prepared. Samples for radiometric dating and stable-isotope
146 analysis were removed using dental burs and a dental handpiece. These samples were positioned

147 so that they did not cross layer-bounding surfaces (Railsback et al., 2013, and Section 3.2 below).
148 For determination of age, the samples were analyzed by use of a multi-collector inductively
149 coupled plasma mass spectrometer (MC-ICPMS, Thermo-Finnigan Neptune) in the University of
150 Minnesota's Department of Earth Sciences using methods described in Shen et al. (2002) and
151 Cheng et al. (2013). The chemical procedure used for separation of U and Th was that of Edwards
152 et al. (1987) and Shen et al. (2002). The initial results were corrected by assuming an initial
153 $^{230}\text{Th}/^{232}\text{Th}$ atomic ratio of $4.4 \pm 2.2 \times 10^{-6}$, which is the value for a material at secular equilibrium
154 with the bulk earth $^{232}\text{Th}/^{238}\text{U}$ value of 3.8. The resulting corrections ranged from 0 to 97 years,
155 with a mean of 13.0 years.

156 For stable isotope analysis, samples of 50 to 100 μg each were analyzed on a GasBench-
157 IRMS system at the Alabama Stable Isotope Laboratory using methods similar to those of Paul
158 and Skrzypek (2007). The results were normalized using NBS-19, which has $\delta^{13}\text{C}=1.95\text{‰}$
159 relative to VPDB and $\delta^{18}\text{O}=-2.20\text{‰}$ relative to VPDB. The resulting isotope ratios are reported
160 relative to the VPDB (Vienna Peedee Belemnite) standard.

161 X-ray diffraction of powdered samples was performed using a Bruker D8 X-ray
162 Diffractometer in the Department of Geology of the University of Georgia by means of scans
163 from 20 to 65° 2 θ with $\text{Co}_{\text{K}\alpha}$ radiation.

164

165 **3.2 Criteria for interpretation**

166 The criteria used to identify the 4.2 ka Event in this project were Type E layer-bounding
167 surfaces, as well as $\delta^{13}\text{C}$ data and $\delta^{18}\text{O}$ data (Figs. 2 and 3). A layer-bounding surface is a surface
168 that cuts across multiple layers of a stalagmite, akin to an unconformity in the stratigraphy of
169 sedimentary rocks. At Type E surfaces (Railsback et al., 2013), the underlying CaCO_3 has been
170 removed, as evidenced by an irregular or jagged surface incised into that underlying CaCO_3 . The
171 origin of these surfaces is most readily understood as a change from dripwater supersaturated
172 with respect to CaCO_3 to dripwater undersaturated with respect to CaCO_3 . An alternate

173 explanation might be physical abrasion, but this would likely lead to a polished smooth surface,
174 rather than a jagged one. Furthermore, Type E surfaces are commonly coated with detrital clay,
175 suggesting flowing or flooding rather than dripping water and thus an increase in water supply.
176 Incision and coating combine to suggest more dilute and more abundant water, both of which
177 would be expected in a change to conditions wetter than those during which the underlying
178 spelean CaCO₃ was deposited. Inference of wetter conditions from the presence of Type E
179 surfaces has been made by Perrin et al. (2014), Shtober-Zisu et al. (2014), Meckler et al. (2015),
180 Martín-Chivelet (2017), Kenny (2018), Liu et al. (2019), López-Martinez (2020), and Fu et al.
181 (2021).

182 Values of $\delta^{13}\text{C}$ in stalagmites are controlled by at least eight different contextual
183 parameters (McDermott, 2004; Table 1 of Voarintsoa et al., 2017b, and sources cited therein).
184 Some of these, such as $\delta^{13}\text{C}$ of the overlying limestone or the extent of Earth's glaciation, are
185 unlikely to change during the deposition of a few layers in one stalagmite. The three contextual
186 parameters most likely to change over short (~decadal) time spans are (1) the extent of vegetation
187 and soil biomass productivity (where lower $\delta^{13}\text{C}$ occurs with more vegetation cover ^[1]_[SEP] in response
188 to wetter conditions) (Hesterberg and Siegenthaler, 1991), (2) the extent of prior precipitation of
189 CaCO₃ from the drip water (where lower $\delta^{13}\text{C}$ occurs with less prior CaCO₃ precipitation typical
190 of wetter conditions) (e.g., Johnson et al., 2006), and (3) the photosynthetic pathway of the
191 overlying plant community. The last of these three is the most confounding, because (a) in wet to
192 semi-arid environments the C3 plants dominant in wetter conditions generate lower values of
193 $\delta^{13}\text{C}$, whereas (b) in semi-arid environments C4 plants typical of hotter, drier or summer-rain
194 climates generate higher values of $\delta^{13}\text{C}$ (Vogel et al., 1978; Brook et al., 2010), but (c) in arid
195 environments C4 grasses do not survive, leaving only C3 scrub as the vegetation, so that lower
196 values of $\delta^{13}\text{C}$ are taken as evidence of drier conditions (e.g., Belz et al., 2020) – although the C3
197 scrub may be so limited that its production of low- $\delta^{13}\text{C}$ soil CO₂ is small and does not generate a
198 noticeably low value of stalagmite $\delta^{13}\text{C}$.

199 Values of $\delta^{18}\text{O}$ in stalagmites are likewise controlled by at least eight different contextual
200 parameters (McDermott, 2004; Lachniet, 2009; Table 1 of Voarintsoa et al., 2017b, and sources
201 cited therein). Some of these, such as altitude of the cave or glacial-interglacial change of $\delta^{18}\text{O}$ of
202 the oceans, are unlikely to change during the deposition of a few layers in one stalagmite. The
203 contextual parameters most likely to change over short (~decadal) time spans are (1) amount of
204 rainfall, where $\delta^{18}\text{O}$ commonly decreases with increasing rainfall (Dansgaard, 1964), (2) extent of
205 evaporation of soil water or drip water, where $\delta^{18}\text{O}$ decreases with lesser evaporation in wetter
206 conditions (Cuthbert et al. 2014), and (3) temperature, where $\delta^{18}\text{O}$ decreases with increasing
207 temperature (McDermott, 2004; Lachniet, 2009). The first two parameters conspire to give lower
208 $\delta^{18}\text{O}$ in wetter conditions, but decreasing temperature can confound that trend. Another
209 contextual parameter relevant at Dante Cave is the distance that water vapor travels to a cave
210 from its source, because greater distance allows more rain-out of ^{18}O and thus a lower $\delta^{18}\text{O}$ of the
211 rainfall (the “continental effect”). However, as noted in Section 2.2, Dante’s location under the
212 shifting CAB and/or KD means that $\delta^{18}\text{O}$ of rainwater can change considerably, further
213 confounding interpretation of $\delta^{18}\text{O}$ from spelean carbonates in the cave.

214 A very general summary of the previous two paragraphs is that lower values of $\delta^{13}\text{C}$ and
215 $\delta^{18}\text{O}$ data are commonly interpreted as evidence of wetter conditions in wet to semi-arid
216 environments (e.g., Section 5.1 of Sletten et al., 2013), but fallibly so. In the best case, a *slow*
217 change to wetter conditions as suggested by the above general model results in lower values of
218 $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ spanning a long time interval and thus across a significant vertical distance in the
219 stalagmite, so that dissolution at peak wetness nonetheless leaves CaCO_3 bearing those lower
220 values of $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ (Fig. 4A). However, a *rapid* shift to wetter conditions leaves only a
221 vertically thin interval of CaCO_3 bearing lower values of $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$, so that dissolution at
222 peak wetness can remove the CaCO_3 that bore the lower values of $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ (Fig. 4B). Thus

223 rapid shifts to wetter conditions may be recorded only by the Type E layer-bounding surfaces
224 described in the first paragraph of this section.

225

226 **4. Results**

227 *4.1. Stalagmite DAN1*

228 Stalagmite DAN1 is 28 cm tall and has a “fencepost” shape that is offset at four points,
229 the uppermost of which is 2.8 cm from the top (Fig. 2A). The interval of DAN1 deposited near
230 4.2 ka BP is, in a coincidence of numbers, 4.2-4.1 cm from the top of the stalagmite. X-ray
231 diffraction of samples drilled 4.3 and 4.1 cm from the top of DAN1 indicates that these samples
232 consist almost entirely of aragonite.

233 Ages from the upper third of Stalagmite DAN1 range from 5.2 to 3.4 ka (Supplementary
234 Document 1). In the same interval of the stalagmite there is petrographic evidence of four
235 surfaces of dissolution (Type E surfaces in the terminology of Railsback et al., 2013 discussed in
236 Section 3.2) to suggest hiatuses that are 5.4, 4.32, 4.15, and 3.89 cm from the top of the
237 stalagmite (Fig. 2C). The closely-spaced hiatal surfaces, between which only one or two ages
238 could be generated, precluded construction of an age model by statistical methods that assume
239 constant deposition (e.g., COPRA or OxCal) in the time range of interest near 4.2 ka (Fig. 2C).
240 The most pronounced of the four hiatuses with regard to both petrographic characteristics and
241 temporal duration is the one 4.15 cm from the top (it is in fact the most visible hiatal surface in
242 the entire stalagmite). Petrographically it is distinguished by an undulating surface of corrosion
243 overlain by a thick layer of fine detrital material, presumably clay, with sand-sized carbonate
244 particles and at least one piece of phosphatic debris that may be a fragment of bone (Fig. 2B). It
245 is by far the most pronounced surface of dissolution and detrital deposition in Stalagmite DAN1.
246 Values of $\delta^{13}\text{C}$ are low around the surface and lowest just below it (Fig. 2D), giving the largest
247 excursion in $\delta^{13}\text{C}$ in the entire stalagmite).

248 Chronologically, the Type E surface at 4.15 cm from the top of Stalagmite DAN1
249 represents a gap of 300 to 400 years that presumably represents (a) time lost from the record
250 because underlying CaCO₃ was dissolved (giving the earlier part of the hiatus) and (b) time when
251 CaCO₃ was not deposited during that dissolution and during deposition of the detrital material
252 (giving the later part of the hiatus) (Fig. 2). The best estimate of the latter, the time of erosion
253 (presumably by dissolution), and of deposition of detrital material (presumably by flowing water),
254 is from 4.1 to 3.9 ka BP (Fig. 2C), which falls within the time range that Railsback et al. (2018b)
255 in their Figs. 2 and 3 found was compatible with eleven previously published chronologies of the
256 4.2 ka Event.

257

258 4.2. Stalagmite DP2

259 Stalagmite DP2 is 37 cm tall and has a nearly perfect fencepost shape (Fig. 3A). The
260 interval of DP2 deposited near 4.2 ka BP is 33-34 cm from the top, and thus near the stalagmite's
261 base. X-ray diffraction of samples drilled 33 and 34 cm from the top of DP2 indicates that those
262 samples consist entirely of aragonite, although a very thin interval (too thin to sample for X-ray
263 diffraction) has the petrographic characteristics of calcite (Fig. 3B).

264 The most prominent feature of Stalagmite DP2 is a thin detrital layer low in the
265 stalagmite, 334 mm from the top. Microscopic examination reveals that this detrital layer
266 overlies a surface of corrosion that is also overlain by spelean calcite, in contrast to the aragonite
267 otherwise present above and below the detrital layer (Fig. 3B). This Type E surface is easily the
268 most visible hiatal surface in the entire stalagmite. Values of both $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ below the
269 detrital layer deviate in the negative direction from the otherwise upward-increasing trend, but the
270 negative anomaly in both cases is only 0.3‰.

271 The exact age of the detrital layer 334 mm from the top of Stalagmite DP2 is uncertain
272 because U-Th ages are inconsistent near the base of the stalagmite. Five ages support an age
273 model readily generated by interpolation yielding an age of 4080 to 4040 yrs BP for the layer

274 (Case α in Fig. 3C), two ages support an age model necessarily generated by interpolation
275 yielding an age of 4340 to 4300 yrs BP (Case β in Fig. 3C), while an age model that is a
276 compromise between all seven ages yields an age of 4210 to 4150 yrs BP (Case γ in Fig. 3C).
277 All three model ages for the detrital layer fall within the time range of anomalies elsewhere that
278 have been ascribed to the 4.2 ka Event (e.g., Zanchetta et al., 2016). More importantly, the age of
279 4080 to 4040 yrs BP suggested by Case α , which is the age model based on the largest number of
280 U-Th ages, falls within the time range that Railsback et al. (2018b) in their Figs. 2 and 3 found
281 was compatible with eleven previously published chronologies of the 4.2 ka Event (Booth et al.,
282 2004; Drysdale et al., 2006; Lemcke and Sturm, 1997; Kaufmann et al., 1988; Berkelhammer et
283 al., 2012; Staubwasser et al., 2003; Cullen et al., 2000; Arz et al., 2006; Sharifi et al., 2015;
284 Kröpelin et al., 2008; Ohlendorf et al., 2014).

285

286 **5. Discussion**

287 *5.1. Inferences from Stalagmites DAN1 and DP2*

288 Stalagmite-based studies of past climate range from those using multiple proxies of past
289 climate (e.g., Scholz et al., 2012; Carolin et al., 2019) to those using just one proxy, typically
290 $\delta^{18}\text{O}$ (e.g., Cheng et al., 2011; Kathayat et al., 2017). Researchers employing multiple proxies
291 presumably do so in hope that meaningful inferences can be made from the synthesis of multiple
292 lines of evidence and/or in hope that, if one proxy fails to record an event or trend, others will
293 nonetheless provide evidence of an otherwise missed paleoclimatological phenomenon. This
294 paper follows that model in reporting five kinds of data or evidence: $\delta^{18}\text{O}$, $\delta^{13}\text{C}$, layer-bounding
295 surfaces, detrital content, and mineralogy of CaCO_3 .

296 The Type E surfaces in Stalagmites DAN1 and DP2 are evidence of dissolution that
297 requires undersaturated waters passing over the stalagmite. Values of $\delta^{18}\text{O}$ show only the
298 slightest of decreases below the Type E surface in Stalagmite DP2 and none in Stalagmite DAN1.

299 However, as Fig. 4 shows, abrupt change to conditions sufficiently wet to cause stalagmite
 300 dissolution (and thus a Type E surface) is likely to remove the $\delta^{18}\text{O}$ signal of that rapid change to
 301 wetter conditions. Thus absence of a low- $\delta^{18}\text{O}$ excursion is not necessarily an argument against a
 302 trend to wetter conditions, and in fact monotonous $\delta^{18}\text{O}$ values below a hiatal Type E surface
 303 suggest *abruptly* wetter conditions (Fig. 4). The expression “abrupt change” is relevant here
 304 because the 4.2 ka Event has been widely regarded as an example of abrupt climate change
 305 (Rashid and Polyak, 2011; Zhang et al., 2018; Carolin et al., 2019; Pleskot et al., 2019; Ön et al.,
 306 2021).

307 In light of the previous paragraph, one logical approach is to use the method of multiple
 308 working hypotheses (Chamberlin, 1890). Here we consider five hypotheses regarding the
 309 paleoclimatological implications of the Type E surfaces and associated geochemical evidence
 310 dating to roughly 4.2 ka from Stalagmites DAN1 and DP2:

311 *Hypothesis 1:* Nothing anomalous took place. In each stalagmite, three things (the hiatus
 312 apparent in the age model, the presence of a corroded surface more pronounced than any other in
 313 the stalagmite, and the accumulation of detrital material in an otherwise much cleaner stalagmite)
 314 combine to document *some* anomalous event. This hypothesis is therefore disproven.

315 *Hypothesis 2:* An unusually wet event or period took place. Three lines of evidence support this
 316 hypothesis. (i) In each stalagmite, accumulation of detrital material in an otherwise much cleaner
 317 stalagmite supports passage of a large flux of water sufficient to transport solids. Such material
 318 may have been deposited either from an exceptionally large flux of falling water or by flooding of
 319 the cave and submergence of the stalagmite. In either case, the fact that the accumulations of
 320 detrital material reported here are the most pronounced such accumulations in their respective
 321 stalagmites suggests that they represent the wettest conditions, either as descending water, flood
 322 water, or both, during those stalagmites’ centuries to millennia of deposition. (ii) In each
 323 stalagmite, presence of a corroded (Type E) surface more pronounced than any other in the

324 stalagmite supports a wet event. The truncation of underlying calcite at irregular surfaces
325 indicates dissolution, rather than the deposition that otherwise generates a stalagmite. Dissolution
326 in turn implies fluxes of undersaturated groundwater through the regolith and bedrock over the
327 cave, and their passage into the cave so rapidly that the water had no time to degas and reach
328 saturation. (iii) In Stalagmite DP2, the change in precipitation from aragonite to calcite is also
329 consistent with wetter conditions (Railsback et al., 1994, and sources cited therein). Thus *three*
330 lines of evidence (detrital content, dissolution, and mineralogy) support an unusually wet event or
331 period.

332 *Hypothesis 3:* An unusually dry event or period took place. Evidence of a dry event in a
333 stalagmite typically consists of a shift of $\delta^{18}\text{O}$ to greater values (see Section 3.2), a shift of $\delta^{13}\text{C}$ to
334 greater values (see Section 3.2), a thinning of layers, and/or a lessening of the area at the top of
335 the stalagmite covered by those layers (a Type L surface in the jargon of Railsback et al., 2013).
336 None of this evidence is present, and so the hypothesis is not supported.

337 *Hypothesis 4:* An unusually wet event was followed by an unusually dry event. As noted with
338 regard to Hypothesis 3, there is no evidence of a dry event, and so the hypothesis is not
339 supported.

340 *Hypothesis 5:* An unusually dry event was followed by an unusually wet event. As noted with
341 regard to Hypothesis 3, there is no evidence of a dry event, and so the hypothesis is not
342 supported.

343 Of these five hypotheses, only Hypothesis 2 (an unusually wet event or period) is
344 supported by the available evidence. Hypotheses 4 and 5 are possible, but there is no evidence to
345 support the “dry” components of these hypotheses.

346 The understanding of detrital material presented above assumes that water routinely
347 moving at very slow rates in the pathways to drips does not have kinetic energy sufficient to carry
348 suspended solids, especially solids as large as the sand-sized grains reported above. In this view,
349 such suspended solids would only accompany exceptionally rapid flow, and the presence of such

350 solids at the crests of stalagmite layers would be evidence of exceptionally wet conditions, either
351 with gushing “drip” water onto the stalagmite or rising standing water around the stalagmite. An
352 alternate understanding might be that cave drips, despite their clarity and slow movement, in fact
353 routinely carry suspended solids (Yadava et al., 2004) as large as fine sand and routinely deliver
354 these to the crests of stalagmites, where greater flow washes the solids off the crests and lesser
355 flow allows the solids to remain. In this view, a drying trend would be indicated by migration of
356 a detritus-rich zone upward through layers/time from the flanks of stalagmite to its crest (Fig. 5).
357 This pattern is not observed in Stalagmites DAN1 and DP2, and thus the drying trend suggested
358 by Hypotheses 3 to 5 above is not supported. Furthermore, Yadava et al. (2004) wrote of “fine
359 detritus”, not sand-sized grains like those shown in Fig. 2B.

360

361 *5.2. The combined data from Stalagmites DAN1, DP2, and DP1*

362 The evidence from Stalagmites DAN1 and DP2 reported here combines with evidence
363 from Stalagmite DP1 reported by Railsback et al. (2018b) to give *three* stalagmites from Dante
364 Cave indicating anomalously wet conditions during the 4.2 ka Event in northeastern Namibia. (A
365 Type E surface and C and O stable isotope data from Stalagmite DP1 were also the evidence of a
366 wet phase in that stalagmite’s record). Paleoclimatological research using stalagmites commonly
367 must be defended from claims of non-replication among the stalagmites themselves (as in Section
368 3.3 of Sletten et al., 2013) or with other data (Betancourt et al., 2002), but in this case the only
369 three stalagmites from Dante Cave known to span the time of the 4.2 ka Event all provide
370 evidence of the same wet paleoclimatological anomaly.

371

372 *5.3. Other proxy evidence of a wet 4.2 ka Event in Namibia and Botswana*

373 There are many proxy records of Holocene climate from northwestern Botswana and
374 northern Namibia, as shown by Fig. 1 of Burrough and Thomas (2013) and Fig. 4 of De Cort et
375 al. (2021), although not all records are of sufficiently high chronological resolution to recognize

376 an abrupt climate event like the 4.2 ka Event. One high-resolution record from the region is that
377 of Chase et al. (2009), who used $\delta^{15}\text{N}$ and $\delta^{13}\text{C}$ data from hyrax middens from the Spitzkoppe
378 inselberg in western-central Namibia (21.8°S, 15.2°E) to estimate change between wetter and
379 drier conditions during the Holocene. Their data (Fig. 6D) show a short-term minimum in $\delta^{15}\text{N}$
380 (and thus a maximum of wetness of climate) and short-term maximum in $\delta^{13}\text{C}$ (and similarly a
381 maximum in wetness in Spitzkoppe's xeric conditions) in which the most negative single $\delta^{15}\text{N}$
382 measurement dates to 4067 BP, at essentially the best estimate of the timing of the 4.2 ka Event
383 (Railsback et al., 2018b, Figs. 2 and 3). That $\delta^{15}\text{N}$ measurement is at the apex of a broader $\delta^{15}\text{N}$
384 minimum from 4223 to 4005 yrs BP matching the time of the 4.2 ka Event in the “precise timing”
385 of the main pulse of the 4.2 ka Event in the Middle East at 4.28 to 3.97 ka BP by Carolin et al.
386 (2019, their Fig. 4). The $\delta^{13}\text{C}$ data from Spitzkoppe similarly have a multi-point maximum that
387 matches the time of the 4.2 ka Event as inferred by Carolin et al. (2019) (Fig. 5D). The data of
388 Chase et al. (2009) thus join the evidence from the three stalagmites from Dante Cave in
389 indicating wetness in Namibia during the 4.2 ka Event.

390 Two records extending from the latest Pleistocene to the Holocene and thus of somewhat
391 lower resolution are those of Cordova et al. (2017) and Belz et al. (2021) (Fig. 7). Cordova et al.
392 (2017) used pollen, spore, and charcoal data to produce a reconstruction of PWetQ (precipitation
393 in the wettest quarter) at Lake Ngami in northwestern Botswana. That reconstruction has a broad
394 maximum at 4.0 ka compatible with a wet 4.2 ka Event at Lake Ngami, although the
395 chronological uncertainty of the record is considerable (Fig. 7A). To the west of that location, at
396 Omongwa Pan in eastern Namibia, Belz et al. (2020) used stable-isotope data to study past
397 climate. Their record of $\delta^{13}\text{C}$ of total organic carbon (TOC) in sediments there has a minimum
398 followed immediately by a maximum (suggesting exceptional dryness followed abruptly by
399 exceptional wetness) in the interval from 4.2 to 3.7 ka. (Fig. 7B) However, the chronological
400 uncertainty of that record is even greater, so that one could infer either exceptional wetness or

401 exceptional dryness in the 4.2 ka event, which most likely occurred around 4.1 ka (Railsback et
402 al., 2018; Carolin et al., 2019). The result is that the records of both Cordova et al. (2017) and
403 Belz et al. (2021) are compatible with (but do not conclusively support) a wet 4.2 ka Event in
404 northwestern Botswana and eastern Namibia and thus are compatible with the inferences of this
405 paper.

406

407 *5.4. Other proxy evidence of a wet 4.2 ka Event in the Southern Hemisphere*

408 Since the first explicit report of evidence of a wet 4.2 ka Event in the Southern
409 Hemisphere (Railsback et al., 2018a), Li et al. (2018) have reported evidence of wetness during
410 the 4.2 ka Event in a stalagmite from Rodrigues Island (19° 42' S, 63° 24' E) in the southern
411 Indian Ocean. Most specifically, both their $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ data suggest a wet pulse at about 4115
412 yrs BP, or more broadly from 4130 to 4020 yrs BP (Fig. 6B).

413 In their Fig. 4, Li et al. (2018) also compiled isotopic evidence of wetness during the 4.2
414 ka Event at five other locations in and around the Indian Ocean. Since their work, Li et al. (2020)
415 have additionally reported evidence of wetness during the 4.2 ka Event on Amsterdam Island in
416 the southern Indian Ocean (Fig. 8). Even farther afield, Yan and Liu (2019) noted that Marchant
417 and Hooghiemstra (2004) had inferred a rise in the level of Lake Titicaca (12.08°S, 69.85°W)
418 around 4.0 ka (Fig. 8). These reports combine to suggest that many regions of the Southern
419 Hemisphere may have had relatively wet climatic excursions during the 4.2 ka Event. In
420 addition, Zhang et al. (2018) reported that $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ data from southern China suggest a wet
421 4.2 ka Event, and Tan et al. (2018) likewise reported evidence of generally wetter conditions in
422 southern China than in northern China during the 4.2 ka Event. These reports combine to suggest
423 that many (but not all) regions, especially in the Southern Hemisphere, may have had relatively
424 wet climatic excursions during the 4.2 ka Event.

425

426 *5.5. Comparison with modeling studies, ITCZ, and AMOC*

427 Yan and Liu (2019) used TraCE-21ka (Simulation of Transient Climate Evolution over
428 the past 21,000 years) simulated using the Community Climate System Model Version 3
429 (CCSM3) to investigate the geographic variability of temperature and rainfall during the 4.2 ka
430 Event. Specifically, Yan and Liu (2019) looked at characteristics between NH warm/wet and
431 cold/dry centennial events found in the 5-3 ka window of a 21ka transient simulation. They found
432 two troughs between 4.4 to 4.0 ka in the annual mean NH surface temperature (down to $-0.18\text{ }^{\circ}\text{C}$
433 from the 5-3 ka mean) and annual mean NH precipitation (down to maximum -0.2 mm/day)
434 simulated timeseries. They also found similar-to-larger amplitude peaks in NH T and P from 4.8-
435 4.5 ka and from 4.0-3.85 ka. To discuss potential spatial characteristics of a 4.2 ka NH cold/dry
436 event, Yan and Liu (2019) plotted spatially the difference in various climate parameters between
437 the 2 centennial NH cold/dry periods and the 2 centennial NH warm/wet periods of the 5-3 ka
438 window.

439 Figure 3B of Yan and Liu (2019) suggests cooling everywhere north of 50°N and nearly
440 everywhere north of 10°S . The modeled precipitation is also latitudinally dependent, but less
441 strikingly so, in suggesting drying everywhere north of 75°N , general drying with lesser regions
442 of increased moisture from 75°N to 5°N , and mixed responses southward from 5°N . The most
443 striking region of increased moisture, with regard to both amount (as much as 70-80 mm/yr) and
444 higher confidence level, was in the zone from the Equator to 35°S (Fig. 8). Yan and Liu
445 concluded that the patterns in temperature and precipitation were consistent with a southward
446 shift in the ITCZ, as had been suggested by Railsback et al. (2018) and Zhang et al. (2018) and
447 more generally is consistent with the findings of Schefuß et al. (2011) regionally and Chiang and
448 Friedman (2012) globally.

449 Figure 8 shows that many of the locations of proxy evidence of a wet 4.2 ka Event
450 discussed in Section 5.4 lie in the regions of wetter conditions at 4.2 ka inferred by Yan and Liu
451 (2019). One should note that this is not just a matter of proxy locations in the Southern

452 Hemisphere and general wetting in the Southern Hemisphere. For example, of two proxies of
453 climate change in South America, the one suggesting increased moisture at Lake Titicaca lies in
454 Yan and Liu's general zone of more rainfall from the Equator to 35°S, whereas evidence of
455 drying at Laguna Cháltel lies in their latitudinal belt of diminished rainfall from 40° to 50°S.
456 Furthermore, this linkage of a wet 4.2 ka Event to ITCZ migration is evident not only in the
457 Southern Hemisphere but also in the Northern Hemisphere: Zhang et al. (2018) and Tan et al.
458 (2018) concluded that the wet 4.2 ka Event in southern China resulted from a southward shift of
459 the ITCZ, so that the East Asian Monsoon did not advance into northern China and instead
460 delivered rainfall to southern China. That short-term observation is consistent with the long-term
461 finding by An et al. (2001) that the region of greatest monsoonal wetness moved from north to
462 south in China across the entire Holocene as the Northern Hemisphere cooled. The abrupt
463 southward shift of the ITCZ around 4.1 ka BP has been attributed to weakening of the Atlantic
464 Meridional Overturning Circulation (AMOC) (Zhang et al., 2018; Yan and Liu, 2019; Geirsdóttir et
465 al., 2019), which brings heat to the Northern Hemisphere.

466

467 **6. Conclusions**

468 New observations from two previously unreported stalagmites from Dante Cave in
469 northeastern Namibia are most compatible with an inference of a wet phase at 4.2-3.9 ka;
470 consideration of multiple hypotheses in fact shows that wetness during the 4.2 ka Event is the
471 only hypothesis supported by the available evidence. These observations combine with previous
472 results from a third Dante stalagmite to indicate wetness, rather than drought, during the 4.2 ka
473 Event. These data from Dante Cave add to a growing body of evidence from eight locations in
474 the Southern Hemisphere and others in southern China suggesting a wet 4.2 ka Event.
475 Supporting the terrestrial data are recent modeling results that suggest zones of increased wetness
476 during the 4.2 ka Event that covered much of the low-to-mid-latitude Southern Hemisphere and
477 more limited regions of the Northern Hemisphere as the ITCZ migrated southward. It thus seems

478 that the 4.2 ka Event cannot be taken to be a global episode of drying as previously suggested but
479 instead must be viewed as a more diverse and complex interval with some areas wetter and some
480 drier with a broad but imperfect latitudinal zonation. Both modeling and proxy data suggest that
481 at least some (but not all) latitudinal zones in the Southern Hemisphere experienced wetter
482 climate.

483

484 **Declaration of competing interest**

485 The authors declare that they have no known competing financial interests or personal
486 relationships that could have appeared to influence the work reported in this paper.

487

488 **Acknowledgements**

489 Funding for this research came from U.S. National Oceanic and Atmospheric Administration
490 (NOAA) grant NA56GP0325 to Brook, Railsback, Jean-Claude Thill and Richard S. Meltzer and
491 from US National Science Foundation (NSF) grant 9908415 to Brook and Railsback. Brook and
492 Eugene Marais collected Stalagmite DAN1 while Ben Hardt and Eugene Marais collected
493 Stalagmite DP2. Dr. Irka Schüller of the Senckenberg am Meer, Marine Research Department in
494 Wilhelmshaven, Germany, provided invaluable help with regard to the age data shown in Fig. 7.
495 P3 Editor Paul Hesse, Reviewer Stacy Carolin, and three anonymous reviewers provided
496 comments that greatly improved the manuscript.

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901

902 Fig. 1. Map of southern Africa showing depictions of the CAB and KD in austral summer by
903 various authors. Red highlights the most recent and detailed of those depictions, as discussed in
904 Section 2.2. Arrows show very generalized wind directions; these winds dictate that humidity is
905 greater on the west or north side of the CAB. The Nicholson (1996) boundary is for January and
906 was reproduced as Fig. 1 in Gasse (2000). The Howard and Washington (2019) CAB and
907 Kalahari Discontinuity are for December, the last month for which they provided a map. The
908 ITCZ is not shown because Nicholson (2018) demonstrated that the ITCZ never comes south of
909 the equator. Gray shading shows the area of January rainfall exceeding 4 mm/day according to
910 the Tropical Rainfall Measuring Mission (TRMM) as shown at
911 https://trmm.gsfc.nasa.gov/trmm_rain/Events/trmm_climatology_3B43.html. Numbers indicate
912 locations of other records discussed in Section 5.3. The underlying base map is from the
913 Alabama Maps website of the Department of Geography of the University of Alabama at
914

915

916 Fig. 2. Images and data from the Type E surface 41.5 mm from the top of Stalagmite DAN1 from
917 Dante Cave in northeastern Namibia. A. Scanned image of the uppermost 10 cm of Stalagmite
918 DAN1. B. Photomicrograph of the Type E surface. The corrosion suggests influx of dilute
919 water and the deposition of detritus suggests influx of much water and possible submergence of
920 the stalagmite. The indexing labels at left are more broadly spaced than one would expect from
921 the scale bar at bottom because the indexing system was generated along the growth axis of the
922 stalagmite, whereas this image is from a location off the growth axis and thus where layers are
923 thinner. C. Age model for uppermost 10 cm of Stalagmite DAN1. D. Plot of C and O stable
924 isotope data along 6 mm of the growth axis of Stalagmite DAN1. The ages and stable-isotope
925 data plotted here are reported in Supplementary Document 1.

926

927

928 Fig. 3. Images and data from the Type E surface 334 mm from the top of Stalagmite DP2 from
929 Dante Cave in northeastern Namibia. A. Scanned image of Stalagmite DP2. B. Photomicrograph
930 of the Type E surface. The corrosion suggests influx of dilute water and the deposition of detritus
931 suggests exceptional influx water. C. Age model for Stalagmite DP2. Three possible inferences
932 for age relationships near the base of the stalagmite are presented. D. Plot of C and O stable
933 isotope data along 9 mm of the growth axis of Stalagmite DP2. The ages and stable-isotope data
934 plotted here are reported in Supplementary Document 1.

935

936

937 Fig. 4. Sketches illustrating that isotopic evidence of rapid climate change to wetter conditions
938 can be removed after deposition. In A, *slow* change to wetter conditions leaves a greater vertical
939 sequence of decreasing $\delta^{18}\text{O}$ or $\delta^{13}\text{C}$, some but not all of which is removed by corrosion. In B,
940 *fast* change to the same value of $\delta^{18}\text{O}$ or $\delta^{13}\text{C}$ leaves a smaller vertical sequence of decreasing
941 $\delta^{18}\text{O}$ or $\delta^{13}\text{C}$, all of which may be removed by the same extent of corrosion. Part B illustrates
942 that failure to recognize surfaces of dissolution (Type E surfaces) may result in failure to
943 recognize rapid climate change to wetter conditions *and* that all isotopic evidence of rapid climate
944 change may be lost to corrosion. The loss of record at such Type E surfaces is analogous to the
945 loss of record at unconformities, the surfaces of erosion, lost strata, and lost history encountered
946 in geology (Hutton, 1788).

947 [Placement of Fig. 4 near the second paragraph of Section 5.1 would be ideal.]

948

949

950 Fig. 5. Pattern resulting from change from wetter to drier conditions if routine transport of solids
951 by drip water occurs wherein greater drip rate washes solids away and smaller drip rate does not
952 (Yadava et al., 2004). This pattern does not occur in Stalagmites DP2 and DAN1.

953

954

955 Fig. 6. Comparison of timing of evidence of the 4.2 ka Event at Dante Cave in northeastern
956 Namibia (C) and nitrogen-isotope and carbon-isotope evidence of a wet period at Spitzkoppe in
957 west-central Namibia (Chase et al., 2009) (D) with other estimates of the timing of the 4.2 ka
958 Event (A and B). The N and C isotope data were detrended by, for each value of $\delta^{15}\text{N}$ or $\delta^{13}\text{C}$,
959 subtracting the result of linear regression of the original data over the period from 6000 to 2000
960 yrs BP. The most negative value of $\delta^{15}\text{N}$ and greatest value of $\delta^{13}\text{C}$ coincide exactly with
961 evidence of wetness in Stalagmite DP1 and DAN2 and with the age of the event suggested by
962 Case α of the age model for Stalagmite DP2, and more generally with the centroid of ages for the
963 4.2 ka Event from eleven sites outside Africa.

964

965

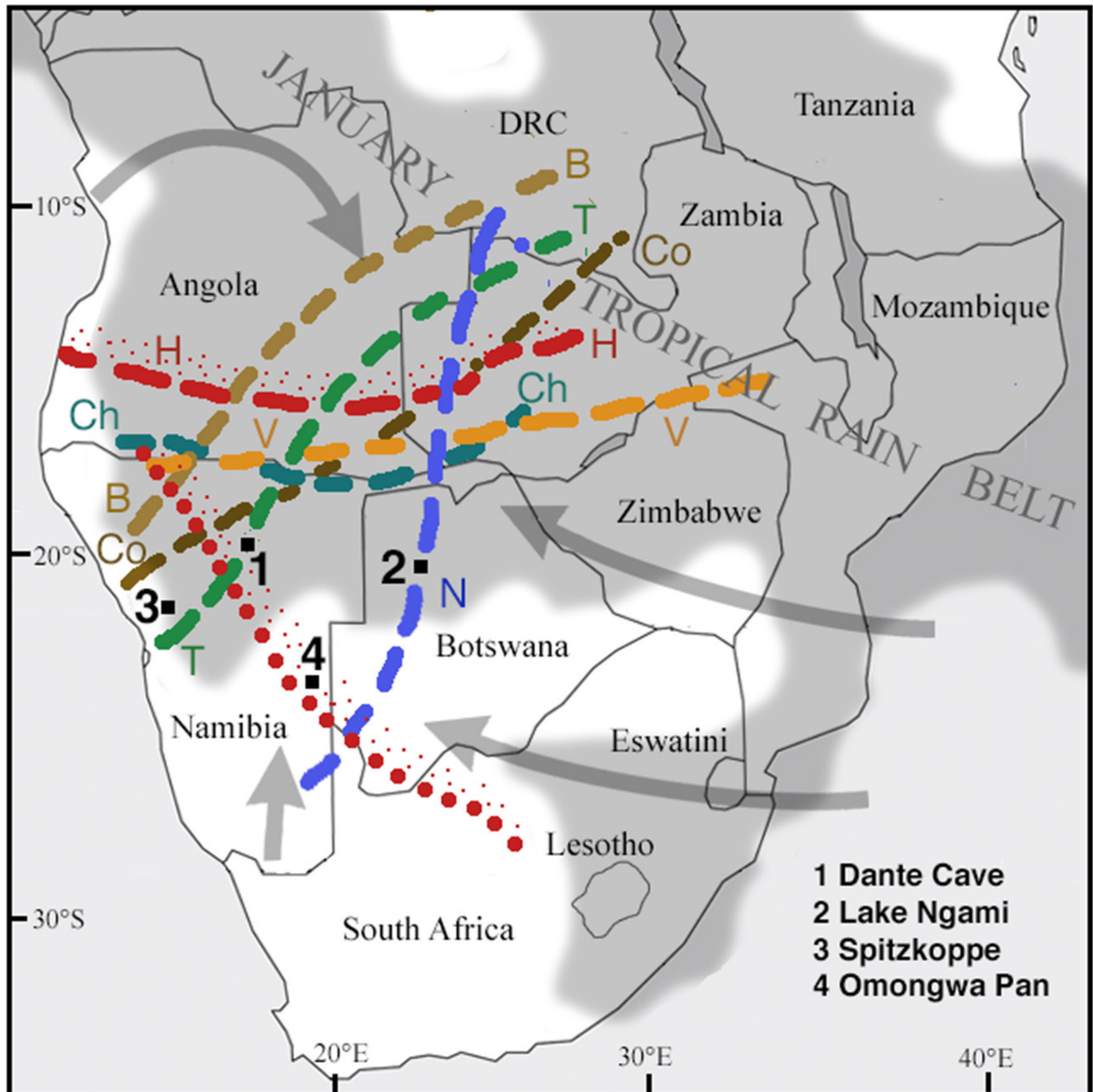
966 Fig. 7. Two proxy records of Holocene climate from near Dante Cave. Upward indicates wetter
967 conditions in both panels A. Reconstruction of precipitation during the wettest quarter of the year
968 of Cordova et al. (2017) from pollen and charcoal data at Lake Ngami in northwestern Botswana.
969 Vertical marks show position in time of pollen samples, but vertical positions of those lines have
970 no meaning. B. Carbon isotope composition of sedimentary organic matter at Omongwa Pan in
971 eastern Namibia (Belz et al., 2020). The age data shown in B are from Schüller et al. (2018)
972 available at <https://doi.pangaea.de/10.1594/PANGAEA.865040> and not from Table 2 of Belz et
973 al. (2020). Both records are compatible with the wet period at about 4.0 ka inferred from
974 stalagmite data from Dante Cave in northeastern Namibia.

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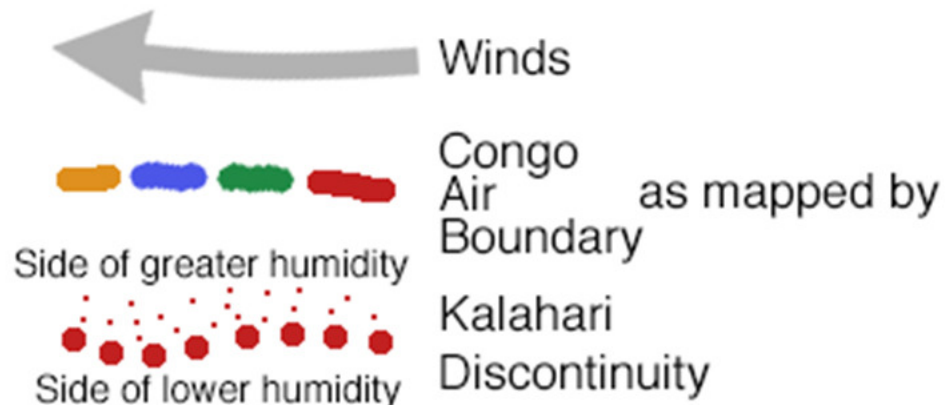
976

977 Fig. 8. Map of the Indo-Pacific region showing locations at which proxy evidence of the 4.2 ka
978 Event has been reported (circles and squares) and showing modeling results of Yan and Liu
979 (2019) (underlying shading). Data from Locations 1 to 4 are shown in Figs. 2, 3, 6 and 7 of this
980 paper. Note the general agreement of the proxy data and model results, both of which suggest
981 that the event was not globally dry, and especially not in mid-latitude regions of the Southern
982 Hemisphere, as discussed in Sections 5.3 and 5.4.

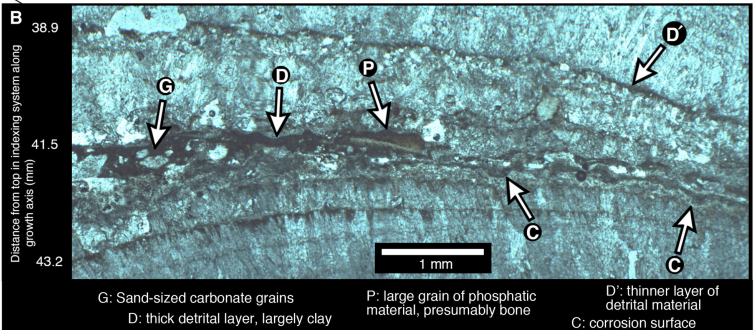
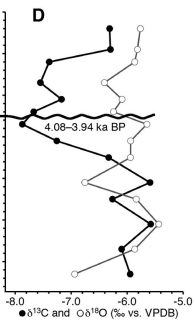
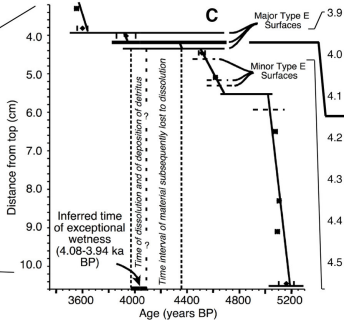
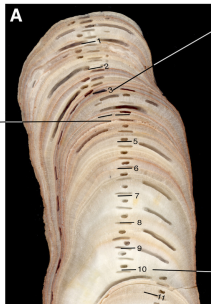
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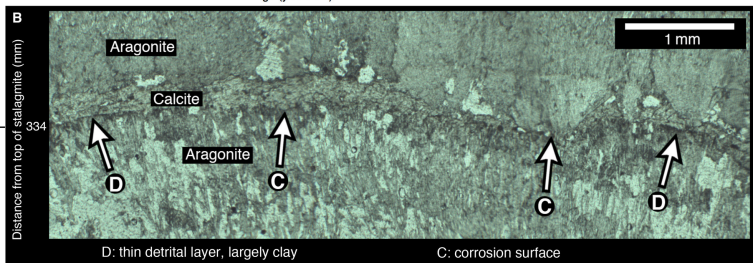
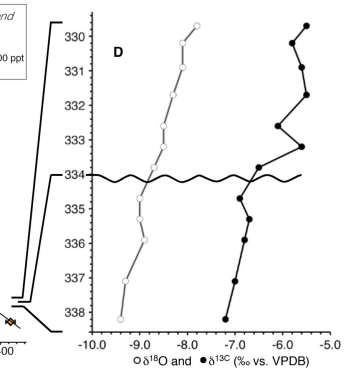
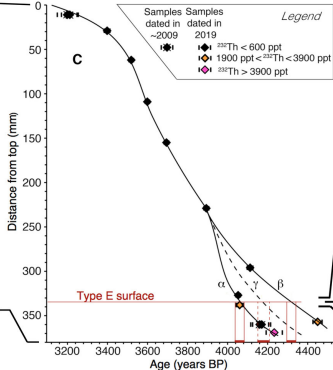


- 1 Dante Cave
- 2 Lake Ngami
- 3 Spitzkoppe
- 4 Omongwa Pan

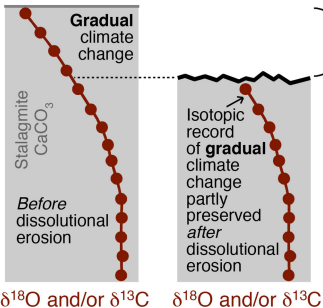


H Howard & Washington (2019)
 V van Heerden & Taljaard (1998)
 Ch Chase et al. (2009)
 Co Cordova et al. (2017)
 B Burrough & Thomas (2013)
 T Tierney et al. (2011)
 N Nicholson (1996)

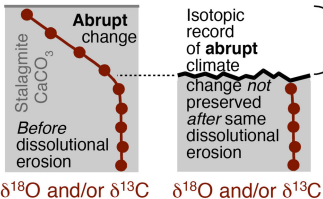




A. Gradual climate change to wetter conditions:

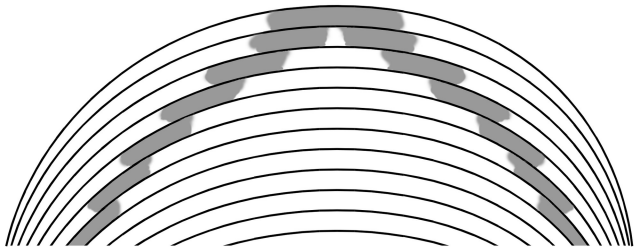


B. Abrupt climate change to wetter conditions:

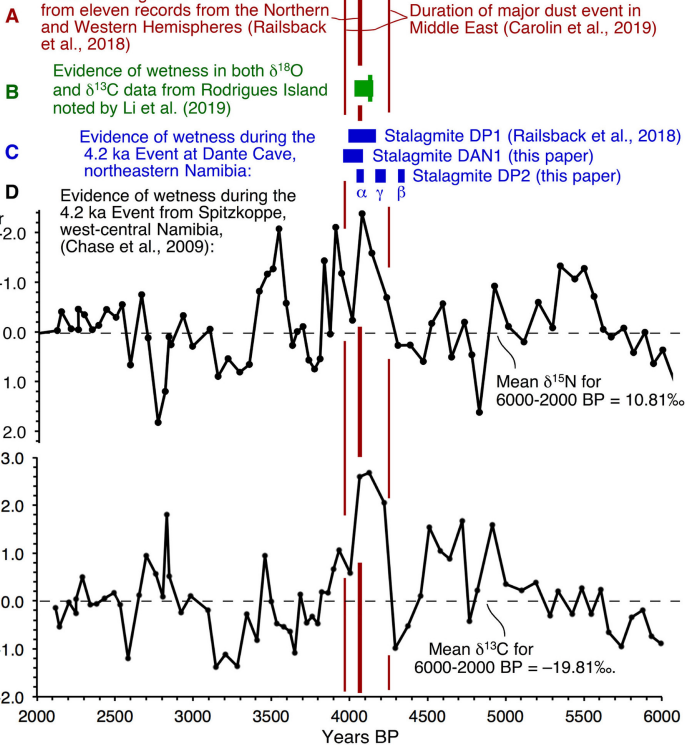


Same vertical extent lost to dissolution

Drier – detritus delivered by drip but not washed from crest

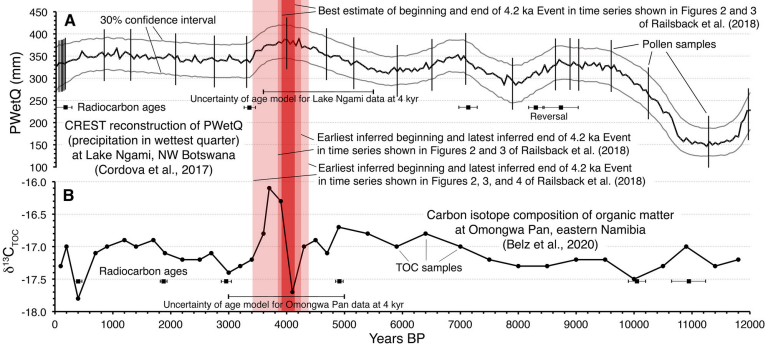


Wetter – detritus delivered by drip but washed from crest by dripwater



Evidence of wetness during the
4.2 ka event at Dante Cave,
northeastern Namibia:

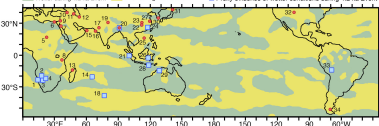
■ Stalagmite DP1 (Railsback et al., 2018)
■ Stalagmite DAN1 (this paper)
■ Stalagmite DP2 (this paper)
 α γ β



Wetter during 4.2 ka Event

Drier during 4.2 ka Event

□ Proxy evidence of wetter conditions during 4.2 ka Event



- 1 Chase et al. (2006) Spitzkoppe inselberg, Namib Desert
- 2 This article, Damie Cave, Otjozondjupa Region, Namibia
- 3 Balz et al. (2020) Omongwa Pan, eastern Namibia
- 4 Cordova et al. (2017) Lake Ngami, northwestern Botswana
- 5 Kröpelin et al. (2008) Lake Yoa, Ouhangja, Chad
- 6 Stanley et al. (2003) Nile Delta, Gharbia, Egypt
- 7 Arz et al. (2006) Core GeoB 5836-2, Sheban Deep, Red Sea
- 8 Thompson et al. (2002) Kilimanjaro ice core, Tanzania
- 9 Bar-Matthews et al. (1999), Soreq Cave, Jerusalem, Israel
- 10 Lemcke and Sturm (1987) Lake Van, Turkey
- 11 Sharifi et al. (2015) Neor Lake, Talesh Mountains, Iran
- 12 Carolin et al. (2019), Stalagmite GZ14-1, northern Iran
- 13 Wang et al. (2019) Stalagmite ANJ94-5, Madagascar
- 14 Li et al. (2018) Rodrigues Island, Indian Ocean*
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* Map symbols represent inference of Li et al. (2018)

** Map symbol represents inferences of original authors and Li et al. (2018)

*** Map symbols represent inference of Yan and Liu (2019)