

1 Retrieval of macro- and micro-physical properties of oceanic 2 hydrosols from polarimetric observations

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12 Abstract

13 Remote sensing has mainly relied on measurements of scalar radiance and its spectral and
14 angular features to retrieve micro- and macro- physical properties of aerosols/hydrosols.
15 However, it is recognized that measurements that include the polarimetric characteristics of light
16 provide more intrinsic information about particulate scattering. To take advantage of this, we
17 used vector radiative transfer (VRT) simulations and developed an analytical relationship to
18 retrieve the macro and micro- physical properties of the oceanic hydrosols. Specifically, we
19 investigated the relationship between the observed degree of linear polarization (DoLP) and the
20 ratio of attenuation-to-absorption coefficients (c/a) in water, from which the scattering
21 coefficient can be readily computed ($b = c - a$), after retrieving a . This relationship was
22 parameterized for various scattering geometries, including sensor zenith/azimuth angles relative

23 to the Sun's principal plane, and for varying Sun zenith angles. An inversion method was also
24 developed for the retrieval of the microphysical properties of hydrosols, such as the bulk
25 refractive index and the particle size distribution. The DoLP vs c/a relationship was tested and
26 validated against in-situ measurements of underwater light polarization obtained by a custom-
27 built polarimeter and measurements of the coefficients a and c , obtained using an in-water WET
28 Labs ac-s instrument package. These measurements confirmed the validity of the approach, with
29 retrievals of attenuation coefficients showing a high coefficient of determination depending on
30 the wavelength. We also performed a sensitivity analysis of the DoLP at the Top of Atmosphere
31 (TOA) over coastal waters showing the possibility of polarimetric remote sensing application for
32 ocean color.

33 **1. Introduction**

34 The polarization of light is highly sensitive to chemical and physical properties of particles in the
35 atmosphere and oceans (Hansen and Travis 1974, Kattawar and Adams 1989, Mishchenko,
36 Cairns et al. 2004, Chami and Platel 2007, Lotsberg and Stamnes 2010, Knobelspiesse, Cairns et
37 al. 2011, Chowdhary, Cairns et al. 2012). The development of satellite sensors capable of
38 measuring and quantifying the polarization of light emerging from atmosphere-ocean (AO)
39 system is becoming increasingly important for understanding not only the macrophysics, but also
40 the microphysics of particulate matter in AO system. The microphysical properties reveal new
41 details about aerosol/hydrosol characteristics that make it possible to distinguish between
42 different aerosol/hydrosol types from space observations, and to derive precise distributions of
43 size, shape, and concentrations that can help in our understanding of the earth's radiation budget,
44 climate change, and ocean processes (Hansen and Travis 1974, Chowdhary, Cairns et al. 2006,
45 Chowdhary, Cairns et al. 2012). Consequently, NASA is considering the addition of a

46 polarimeter onboard a future ocean color mission: the Plankton, Aerosol, Cloud, and Ocean
47 Ecosystem (PACE) (PACE 2012). The long-term goal of this mission is essentially to assess the
48 carbon cycle and its interrelationship with climate change. It also will extend the ocean climate
49 data records collected since the 1990s and evaluate observed long-term changes. Polarimetric
50 measurements from space will improve atmospheric corrections, especially over bright coastal
51 and shallow waters (Chowdhary 1999, Chowdhary, Cairns et al. 2001, Waquet, Cairns et al.
52 2009, Knobelspiesse, Cairns et al. 2011). While polarization and multi-angle measurements can
53 be used to derive some key properties of aerosols, such measurements can also be useful in
54 retrieving the optical and microphysical oceanic properties (PACE 2012). In addition, the
55 planned Multiviewing, Multichannel, and Multipolarization Imager - 3MI (ESA/Eumetsat) will
56 help with understanding the composition of aerosols and clouds in the atmosphere and possibly
57 the oceanic hydrosols and their impact on climate forcing (Marbach, Riedi et al. 2015). The
58 Second Generation Global Imager (SGLI) on-board the Japanese Global Change Observation
59 Mission (GCOM-C) will provide atmosphere, land, and ocean color products. The two added
60 polarization channels with 250-meter spatial resolution will help to improve aerosol and ocean
61 color retrievals, especially in coastal regions (Honda, Yamamoto et al. 2006). The studies and
62 analyses performed in this work will contribute to the development of comprehensive remote
63 sensing inversion algorithms that utilize the polarimetric signature of the ocean for the retrieval
64 of macro- and micro-physical properties of hydrosols.

65 Several studies have shown the potential of utilizing the polarization characteristics of oceanic
66 light to retrieve Inherent Optical Properties (IOPs) and biogeochemical properties (Chami, Santer
67 et al. 2001, Chami and Mckee 2007, Chami and Platel 2007, Loisel, Duforet et al. 2008,
68 Lotsberg and Stamnes 2010, Tonizzo, Gilerson et al. 2011, Ibrahim, Gilerson et al. 2012). For

69 example, Chami and Platel (2007) studied the use of directional variations and polarization of
70 marine reflectance in the remote sensing retrieval of IOPs. Using an artificial neural network
71 (NN), they demonstrated that adding polarized reflectance to unpolarized reflectance at 490 and
72 665 nm as inputs to the NN improves the retrieval of the scattering coefficient by more than 75%
73 in relative error compared to the use of scalar reflectance alone. The remote sensing reflectance
74 R_{rs} is defined as the water-leaving radiance normalized by the sum of direct and diffuse
75 downwelling irradiances. To a first approximation, R_{rs} is proportional to b_b/a for open ocean
76 waters and to $b_b/(a + b_b)$ for coastal waters, where b_b is the backscattering coefficient. As such,
77 R_{rs} by itself does not contain any information on the light scattered forwardly into the water
78 (Gordon, Brown et al. 1975, Gordon, Morel et al. 1983, Gordon, Brown et al. 1988, Gordon
79 1989). The extraction of coefficient b from coefficient c retrieved from polarization data can
80 provide this information and reduce the retrieval uncertainty of b_b and a from R_{rs} data, especially
81 in optically complex waters.

82 Studies have shown that the particulate attenuation coefficient, c_p , of hydrosols co-varies with
83 the particulate organic carbon concentration (POC) as well as with phytoplankton carbon
84 biomass (Behrenfeld and Boss 2003, Behrenfeld, Boss et al. 2005, Cetinic, Perry et al. 2012,
85 Graff, Westberry et al. 2015). They suggest that there is a first-order relationship between the
86 ratio of c_p to chlorophyll concentration (c_p :Chl), as an index of phytoplankton carbon (C)
87 biomass ratio to chlorophyll concentration (C:Chl) and phytoplankton physiology, which is
88 important for estimating primary production of the oceans. Thus, retrieval of the attenuation
89 coefficient from remote sensing would allow for better understanding of the carbon cycle on the
90 global scale, a primary goal of many ocean color satellite missions (e.g., PACE mission).

91 In coastal waters, hydrosols are composed primarily of two types of particles: algal and non-
92 algal. Algal particles with high water content have a low refractive index (approximately 1.06)
93 relative to that of water and therefore produce only an indistinctive polarization signature similar
94 to that of Rayleigh scattering (Voss and Fry 1984, Tonizzo, Zhou et al. 2009). Their impact on
95 the fraction of polarized light is predominantly by means of their absorption, which reduces the
96 amount of multiple scattered light. Non-algal particles (NAP), such as mineral particles, scatter
97 light more effectively due to their high relative refractive index, typically around 1.18 (Babin,
98 Morel et al. 2003). Through multiple scattering, these particles can significantly decrease the
99 polarization of water-leaving radiance; thus, their concentration should be retrievable using
100 polarization measurements. Although this polarization is highly sensitive to underwater light
101 scattering (Tonizzo, Gilerson et al. 2011), absorbing properties of the water also significantly
102 impact the polarized light field, since the increase in absorption also results in the decrease of the
103 number of the multiple scattering events; in turn this leads to an increase of the fraction of light
104 that is polarized, and therefore, the increase of water absorption in the red part of the spectrum
105 increases the fraction of polarized light. Similarly, the presence of colored dissolved organic
106 matter (CDOM), a strong absorber of blue light, increases the fraction of polarized light in the
107 blue part of the underwater spectrum (Chowdhary, Cairns et al. 2012).

108 In the open ocean, the majority of particles are algae and their by-products whose concentrations
109 co-vary with chlorophyll *a* concentration [Chl]. These types of particles exhibit polarization
110 patterns similar to those of molecular scattering because of their low refractive index (Chami,
111 Santer et al. 2001). Underwater polarization for open oceans is therefore relatively simple (Voss
112 and Fry 1984, You, Tonizzo et al. 2011). It can be reproduced well by vector radiative transfer

113 (RT) computations and persists in the water leaving radiance even when the sea surface is ruffled
114 by strong winds (You, Kattawar et al. 2011, You, Tonizzo et al. 2011).

115 The advantage of using polarization measurements for remote sensing of oceans over scalar R_{rs}
116 measurements is that the polarized components of light intrinsically (mathematically) carry more
117 information about the microphysical properties of scattering particles. R_{rs} is dependent on only
118 one element of the scattering matrix (*i.e.* scattering function) of hydrosols which is an IOP, and
119 the absorption, while the polarized components are affected by all of the elements of the
120 scattering matrix. Therefore, scattering and absorbing hydrosols modulate the polarization of
121 light induced by molecular scattering of water molecules. Changes of light polarization depend
122 on the concentration and composition of these hydrosols and thus can be related to both their
123 absorption and scattering properties.

124 Timofeyeva (1970) found a relationship between the linearly polarized light in artificial milky
125 turbid water and the parameter T, which is equal to the ratio of the attenuation coefficient of the
126 scattered light flux to the direct light flux (Timofeyeva 1970). Inspired by this preliminary
127 relationship, the work presented here seeks to expand that early study and extends it to naturally
128 complex water conditions under natural illumination. It also expands the similar analysis of
129 Ibrahim *et al.* (2012) (Ibrahim, Gilerson et al. 2012) to include a more realistic vector radiative
130 transfer model in order to retrieve the attenuation coefficient, c , and the microphysical properties
131 of hydrosols in the water body. The improvements in the model are focused on the development
132 of a *hybrid* model that relates the microphysics to the macrophysics of hydrosol particles.

133 2. Background

134 2.1. Stokes parameters

135 A full description of light takes the electromagnetic (EM) vector nature of light into account and
136 includes the intensity and the polarization state. The intensity is the energy flux of an EM wave
137 (i.e., the brightness of light), and the polarization state fully describes an oscillating EM wave. In
138 turn, a polarized beam of light can be defined by the 4×1 Stokes vector $\mathbf{I} = \{I, Q, U, V\}'$
139 (Mishchenko). Stokes parameter I is the intensity or radiance, which has the dimension of energy
140 flux per unit solid angle (i.e. the total energy carried by the EM wave). Stokes parameters Q , U ,
141 and V describe the polarization state of the EM wave. The V component, which describes the
142 amount of circular polarization, for most cases of light scattered in AO systems is negligible
143 (Mishchenko). Therefore, one can confine the description of light scattered in AO systems to the
144 intensity and linear polarization of an EM wave, which are given by the first three elements of
145 the Stokes vector. A convenient measure for such cases is the degree of linear polarization
146 (DoLP), which is defined as:

$$147 \quad \text{DoLP} = \frac{\sqrt{Q^2 + U^2}}{I} \quad (1)$$

148 The DoLP describes the fraction of light, which is linearly polarized, so $0 \leq \text{DoLP} \leq 1$. When
149 $\text{DoLP} = 1$, it corresponds to a fully linearly polarized light, and $\text{DoLP} = 0$ corresponds to
150 completely unpolarized light. All other values of DoLP indicate that the light is only partially
151 linearly polarized.

152 2.2. Single scattering matrix

153 In scattering simulations of polarized light, it is important to define the scattering matrix \mathbf{F} . This
154 matrix relates the Stokes vector of light scattered by a particle to the Stokes vector of the incident
155 light, and is therefore a 4×4 matrix, such that

$$156 \quad \mathbf{I}_{sca}(\theta) = \mathbf{F}(\theta) \times \mathbf{I}_{inc}(\theta) \quad (2)$$

157 where θ is the scattering angle between the incident and scattered light. \mathbf{I} is the Stokes vector of
158 a light beam with the subscripts “sca”, and “inc” corresponding to the scattered and incident
159 beam, respectively. For scattering by homogenous spherical particles, \mathbf{F} takes the form of
160 (Kattawar, Hitzfeld.Sj et al. 1973, Hulst 1981):

$$161 \quad \mathbf{F}(\theta) = \begin{bmatrix} F_{11} & F_{12} & \mathbf{0} & \mathbf{0} \\ F_{12} & F_{11} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & F_{33} & F_{34} \\ \mathbf{0} & \mathbf{0} & -F_{34} & F_{33} \end{bmatrix}, \quad (3)$$

162 The (1,1) element of the scattering matrix \mathbf{F} , F_{11} , is the scattering function. The scattering matrix
163 is calculated from the particle refractive index and the particle size distribution (PSD) of a group
164 of particles using the Mie theory detailed in Section 3.2.1.

165 2.3. Multiple scattering in the AO System

166 The following assumptions are made for the AO system: (1) the medium is plane parallel, and
167 (2) only vertical variations of the medium are considered. The RayXP program, developed by
168 Zege *et al.*, is used to calculate the Stokes vector of the light scattered within the AO system
169 (Zege, Katsev et al. 1993, Tynes, Kattawar et al. 2001). The RayXP program performs accurate
170 and extremely fast computations by simulating the Stokes vector of light within any stratified
171 system, achieved by coupling of two methods: (1) the adding-doubling method and (2) the Multi-
172 component approach (MCA). The MCA dramatically reduces the computational time for the

173 cases that involve strong anisotropy in scattering by separating the scattering matrix of the
174 scatterers into two main components. It separates the peaked component from the more diffused
175 remaining component of the scattering matrix. Each of these two separate components is further
176 divided into multiple components that are generally easy to solve for multiple scattering
177 computations.

178 **3. Study cases**

179 **3.1. Atmosphere**

180 The atmosphere is assumed to consist of two homogeneous layers. The top layer is a molecular
181 (Rayleigh scattering) layer. The optical thickness values for this layer are taken from MODIS
182 products to be 0.098 at 550 nm (Ahmad, Franz et al. 2010). The molecular optical properties are
183 obtained from the standard data bank provided by RayXP, and give a molecular depolarization
184 factor of 0.0279 (Hansen and Travis 1974, Young 1980). The bottom layer is assumed to be a
185 continental aerosol layer with optical thickness τ of 0.1 at 400 nm. The continental aerosol model
186 is obtained from a data bank based on simple aerosol models from well-known climatological
187 atmosphere models, which assume a composition made up of 70% dust, 29% water soluble, and
188 1% carbon soot aerosols.

189 **3.2. Ocean**

190 The ocean body is assumed to consist of a single (*i.e.* homogeneous) layer with infinite depth.
191 The benthic optical effects at the bottom of the ocean are not taken into consideration. For the
192 interface between the ocean and the atmosphere (*i.e.* water surface) we use the isotropic
193 distribution of surface slopes given by Cox and Munk (Cox and Munk 1956) for a wind speed of
194 3 m/s. The computation of micro-physical optical properties for the ocean body is discussed in

195 the next section. The macroscopic optical properties for the ocean body are discussed in Section
196 3.2.2.

197 **3.2.1. Microphysical optical properties (Scattering Matrices)**

198 To obtain the scattering matrices as an input for the radiative transfer (RT) program RayXP, Mie
199 calculations were used for both phytoplankton and non-algal particles (NAP) in oceanic bodies.
200 The input parameters for these calculations are the real part of the particle refractive index and
201 the PSD. Open ocean (Case I) waters are assumed to contain only phytoplankton particulate,
202 whereas coastal ocean (Case II) waters are assumed to contain both phytoplankton and NAP
203 particulate. For simplification, it is also assumed that both types of particles are homogeneous
204 spheres with particle radii between 0.1 and 50 μm that follow a Junge-type PSD. The Junge slope
205 parameters ξ_{ph} and ξ_{nap} for the phytoplankton and NAP matter PSD, respectively, vary each
206 between 3.5 and 4.5. The phytoplankton refractive index n_{ph} is fixed at 1.06 (relative to water),
207 whereas the NAP refractive index n_{nap} is allowed to vary from 1.15 to 1.21 (relative to water).
208 The corresponding phytoplankton and NAP scattering matrices are calculated using the Mie code
209 described by Mishchenko *et al.* (Mishchenko, Travis et al. 2002)

210 The RayXP program requires as input the bulk ocean scattering matrix, which consists of a
211 weighted average of the plankton and NAP scattering matrices. For the weights, we use the
212 phytoplankton and NAP scattering coefficients b from Ibrahim *et al.* 2012 (Ibrahim, Gilerson et
213 al. 2012).

214 **3.2.2. Macrophysical optical properties (bulk absorption and scattering coefficients)**

215 The macrophysical optical modeling of bulk oceanic waters is divided into two distinctive cases:
216 open ocean (Case I) and coastal ocean (Case II). Case I oceans contain mainly phytoplankton
217 particles and other co-related, co-variant products such as Color Dissolved Organic Matter

218 (CDOM). Case II waters are impacted by strong interactions with rivers, shores, and
219 anthropogenic activities which make them optically complex because they contain substantial
220 amounts of sediments and minerals whose variability does not depend on the amount of
221 phytoplankton. In what follows, we use chlorophyll *a* concentration [Chl] as a proxy for the
222 amount of phytoplankton. It should be noted that coastal regions can contain high concentrations
223 of phytoplankton (algal blooms), and such eutrophic regions are not considered in this study. The
224 next subsections explain the modeling of the optical properties for Case I and Case II waters
225 separately.

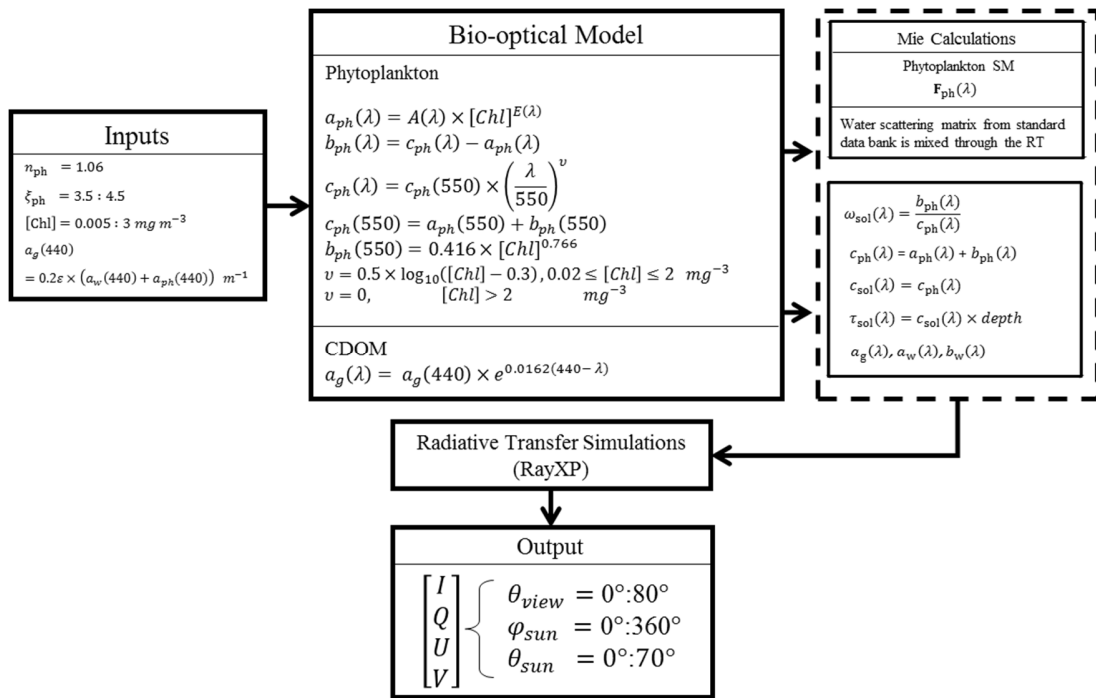
226 *A. Case I waters*

227 For Case I waters, a three-component model is assumed: seawater, phytoplankton particles, and
228 CDOM. The presence of NAP matter is ignored for these waters. The scattering and absorption
229 coefficients b_w and a_w of seawater are taken from Morel 1974 and Pope and Fry (1997),
230 respectively (Morel 1974, Pope and Fry 1997). Since the water molecules and density fluctuations
231 are much smaller in size than the visible wavelength, their scattering matrix can be obtained from
232 the Rayleigh scattering theory. Inelastic scattering, such as Raman scattering, is not taken into
233 consideration in our RT simulations.

234 Following the empirical relationship detailed in Bricaud *et al.* (1998) (Bricaud, Morel et al.
235 1998), the absorption coefficient of phytoplankton, $a_{ph}(\lambda)$, is calculated from the empirical
236 coefficients $A(\lambda)$ and $E(\lambda)$. The phytoplankton scattering coefficient $b_{ph}(\lambda)$ is calculated as the
237 difference between the phytoplankton spectral attenuation coefficient $c_{ph}(\lambda)$ and absorption
238 coefficient $a_{ph}(\lambda)$. The attenuation coefficient at 550 nm, $c(550)$, is obtained from the sum of
239 the absorption coefficient $a_{ph}(550)$ and scattering coefficient $b_{ph}(550)$ at 550 nm, which is
240 obtained from Morel and Maritorena (2001) (Morel and Maritorena 2001). The spectral behavior

241 of $c_{ph}(\lambda)$ is modeled using the power law function given by the Hydrolight RT program (Morel,
 242 Antoine et al. 2002, Mobley and Sundman 2008).

243 The absorption coefficient $a_g(\lambda)$ of CDOM decreases exponentially with increasing wavelength.
 244 CDOM is a poor scatterer of light and its scattering properties can be neglected for the visible
 245 spectrum (Carder, Chen et al. 1999, Babin, Stramski et al. 2003, IOCCG 2006). The absorption
 246 coefficient of CDOM at 400 nm, $a_g(400)$ is correlated randomly within 20% (+/- 10%) to the
 247 phytoplankton and water absorption coefficients at 400 nm as well. Figure 1 shows the block
 248 diagram of the RT modeling for Case I waters.



249

250 **Figure 1. Flow Diagram of the bio-optical and RT models for the generation of the data set for Case I waters.**

251 **B. Case II waters**

252 For Case II waters, a four-component model is assumed that includes seawater, phytoplankton
 253 particles, CDOM, and NAP matter. The specific absorption coefficient of phytoplankton

254 particles $a_{ph}^*(\lambda)$ is modeled as the combination of specific absorption by two dominant
 255 phytoplankton species (*i.e.*, micro- and pico-plankton particles), weighted by their size
 256 parameter S_f (Ciotti, Lewis et al. 2002). The total phytoplankton absorption coefficient $a_{ph}(\lambda)$ is
 257 then calculated as the product of $a_{ph}^*(\lambda)$ and $[Chl]$. The scattering coefficient of phytoplankton
 258 $b_{ph}(\lambda)$ is obtained as the difference between the phytoplankton attenuation c_{ph} and absorption
 259 $a_{ph}(\lambda)$ coefficients (Stramski, Bricaud et al. 2001, IOCCG 2006). The spectral behavior of
 260 $c_{ph}(\lambda)$ is modeled as a power law function with spectral slope Y_{ph} (IOCCG 2006). Note that Y_{ph}
 261 is directly related to the slope ξ_{ph} for the Junge PSD of phytoplankton particles, such that $Y_{ph} =$
 262 $\xi_{ph} - 3$ (Twardowski, Boss et al. 2001). The phytoplankton attenuation coefficient at 550 nm,
 263 $c_{ph}(550)$, is calculated following Gilerson *et al.*, 2007 (Gilerson, Zhou et al. 2007).

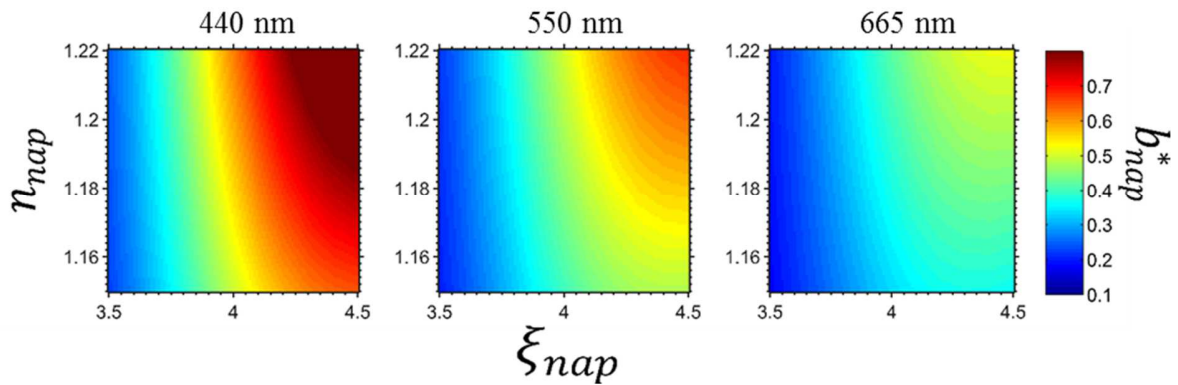
264 The absorption spectrum of NAP is modeled as a decaying exponential function following Babin
 265 *et al.* (2003) (Babin, Stramski et al. 2003). On the other hand, the scattering coefficient is based
 266 on the microphysical properties of NAP matter. This coefficient is modeled as the product of the
 267 specific scattering coefficient $b_{nap}^*(\lambda)$ and the concentration of NAP matter $[NAP]$, where
 268 $b_{nap}^*(\lambda)$ is calculated as follows (Wozniak and Stramski 2004):

$$269 \quad b_{nap}^*(\lambda) = \frac{3\bar{Q}_b(\lambda) \int_{D_{min}}^{D_{max}} N(D)D^2 dD}{2\rho \int_{D_{min}}^{D_{max}} N(D)D^3 dD}, \quad [m^2 g^{-1}] \quad (4)$$

270 In Eq. (4), $\bar{Q}_b(\lambda)$ is the average spectral efficiency factor, ρ is the particle density, D is the
 271 particle diameter, and $N(D)dD$ is the particle size distribution. Following Sec. 3.2.1 we adopt the
 272 Junge PSD with Junge slope parameter $3.5 \leq \xi_{nap} \leq 4.5$ for $N(D)dD$, and the refractive index 1.15
 273 $\leq n_{nap} \leq 1.22$, as input for Mie computations to obtain $\bar{Q}_b(\lambda)$. The density ρ of NAP matter is

274 dependent on the composition of this matter. Assuming an ensemble of 29 mineral species, this
 275 density is calculated as follows (Wozniak and Stramski 2004)

$$276 \quad \rho = 6.779 \times 10^6 \times n_{nap} + 5.232 \times 10^6 \quad [g \ m^3] \quad (5)$$



277
 278 *Figure 2. The specific scattering coefficients b_{nap}^* of NAP according to the Mie theory for 440 nm, 550 nm, and*
 279 *665 nm.*

280 Figure 2 shows the relationship among ξ_{nap} , n_{nap} , and b_{nap}^* . The refractive index n_{nap} is
 281 assumed to be in the range 1.15-1.22 relative to the water, while the values of the Junge slope
 282 ξ_{nap} range from 3.5 to 4.5, which is typical for NAP particles. It is considered a more realistic
 283 approach to have b_{nap}^* vary with ξ_{nap} and with n_{nap} . In our previous study (Ibrahim, Gilerson
 284 et al. 2012), as a first approximation, b_{nap}^* was assumed to be constant over the wavelength and
 285 for different PSD, which is not accurate as shown in Figure 2. It is important to note that the
 286 variation of b_{nap}^* with n_{nap} is weaker than the variation of b_{nap}^* with ξ_{nap} and Y_{nap} , which
 287 validates a part of our previous study. The slope of the NAP spectral scattering coefficient, Y_{nap} ,
 288 is defined as $\xi_{nap} - 3$.

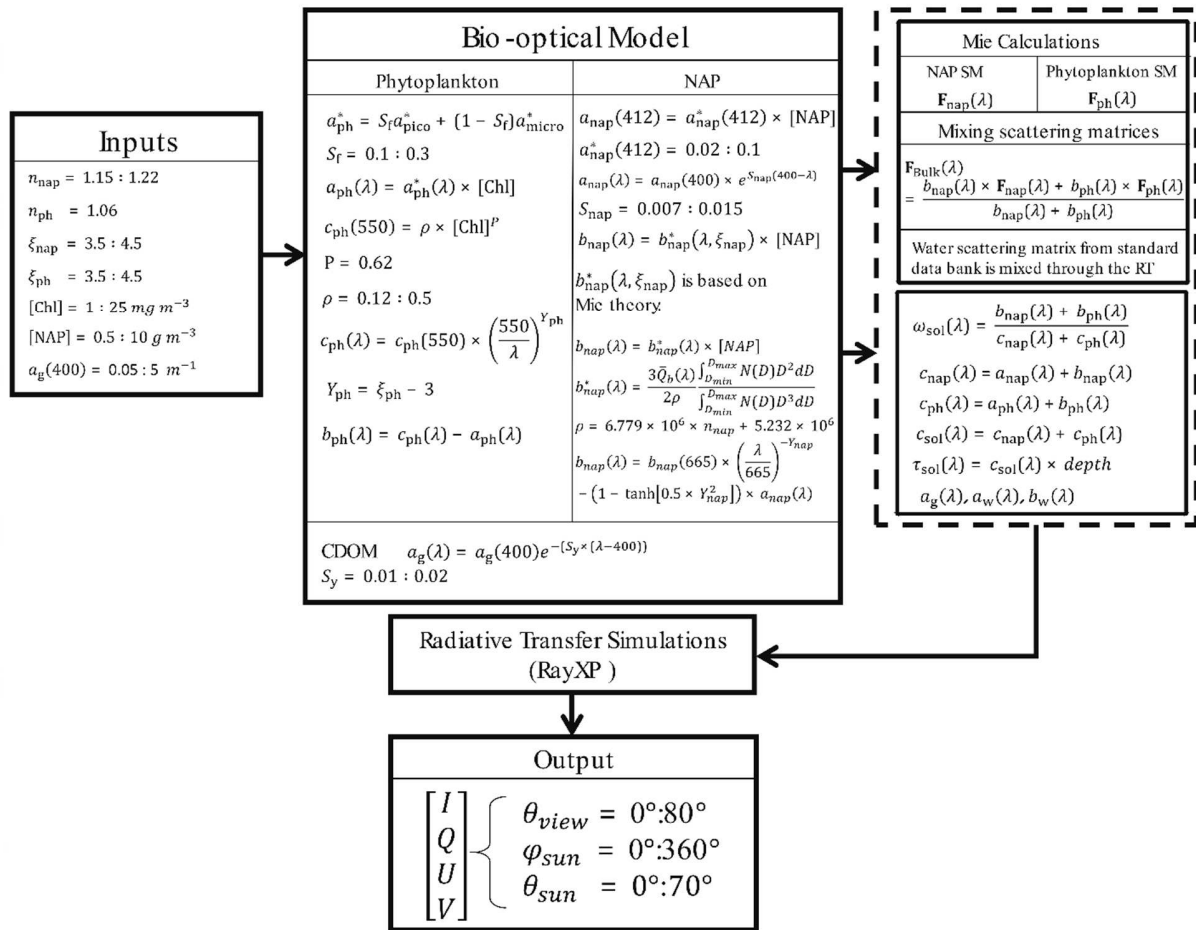
289 The spectral scattering coefficient $b_{nap}(\lambda)$ calculated from Mie theory can have values 5-30%
 290 higher in the blue part of the spectrum when the imaginary part of the refractive index is set to

291 zero in these computations (*i.e.* assuming $a_{nap}(\lambda)$ is zero) (Doxaran, Ruddick et al. 2009). In
292 order to correct for the resulting change in $b_{nap}(\lambda)$ spectra, an empirical model based on
293 theoretical calculations is used to reproduce the actual spectra of $b_{nap}(\lambda)$ from the red channel
294 (665 nm) to the blue channels for non-zero $a_{nap}(\lambda)$ values. This model is given by Doxaran et
295 al. (Doxaran, Ruddick et al. 2009)

$$296 \quad b_{nap}(\lambda) = b_{nap}(665) \times \left(\frac{\lambda}{665}\right)^{-Y_{nap}} - (1 - \tanh[0.5 \times Y_{nap}^2]) \times a_{nap}(\lambda) \quad [m^{-1}]$$

297 (6)

298 The block diagram in Figure 3 shows the flow diagram of our RT computations for Case II
299 waters.



300

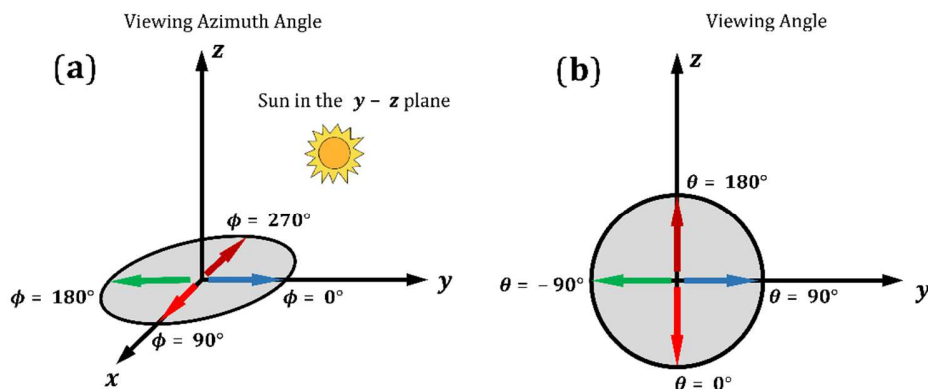
301 **Figure 3. Flow Diagram of the bio-optical and RT models for the generation of the data set for Case II waters.**

302 **4. Results of radiative transfer simulations**

303 Simulated RT values of DoLP are investigated to seek a definitive empirical relationship with the
 304 IOP ratio c/a at three wavelengths: 440, 550, and 665 nm. Such a relationship would provide the
 305 possibility of retrieving the attenuation coefficient, c , of the water constituents from data
 306 obtained by under- or above-water polarization radiometric measurements of upwelling
 307 radiation, since the absorption coefficient, a , is routinely estimated from the remote sensing
 308 reflectance using well-established algorithms (Lee, Carder et al. 2002).

309 **4.1. Geometrical interpretation of the RT results**

310 The RT program, RayXP, provides the Stokes parameters as a function of the viewing and
 311 azimuth angles. The Sun zenith angle is varied from 0° zenith when the Sun is overhead to 70°
 312 when the Sun is near the horizon. The 0° azimuth angle corresponds to the condition when the
 313 Sun and the viewer are in opposition, while 180° azimuth angle corresponds to the condition
 314 when the Sun is behind the sensor (*i.e.* the sensor and the Sun are in the same half plane). The
 315 geometry of the incident, viewing, and azimuth angles is shown in Figure 4.



316

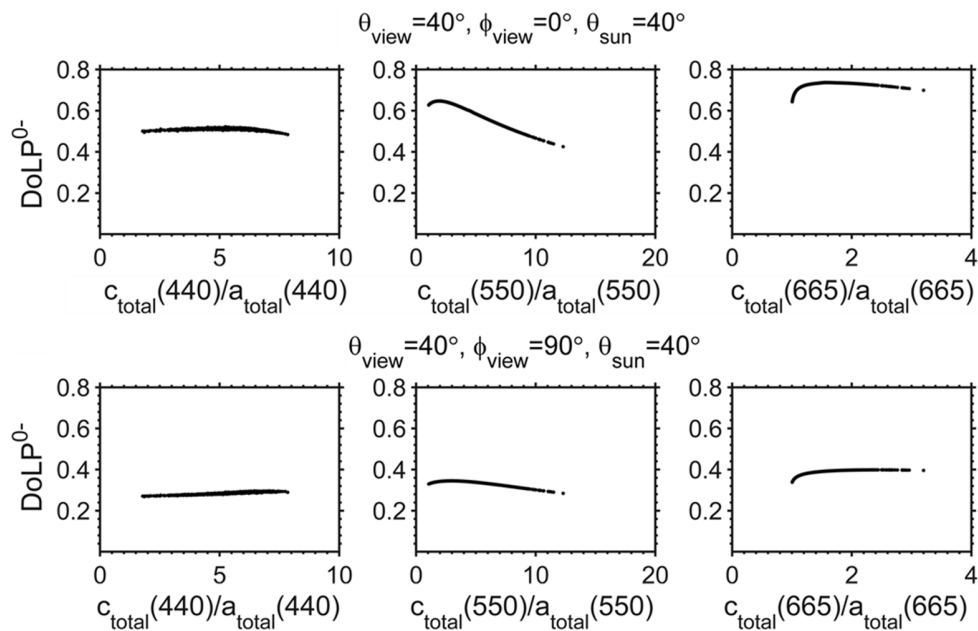
317 *Figure 4. Geometrical interpretation of: (a) the viewing azimuth angle and (b) the viewing angle for RayXP*
 318 *simulations. Arrows point into viewing direction*

319 We also define the scattering angle, θ_{sca} , as the angle between the incident light and the scattered
 320 light (*i.e.* the viewing direction). Based on Ibrahim *et al.* (2012) (Ibrahim, Gilerson et al. 2012),
 321 the best viewing geometries for above water polarization measurements call for a viewing angle
 322 between 40° and 60° from the nadir direction, and a relative azimuth of 90° with respect to the
 323 Sun to avoid Sun glint. Therefore, the results presented here will be at 40° viewing angle and
 324 with two relative azimuth angles of 0° and 90° with respect to the plane containing the Sun.

325 **4.2. Case I simulation results**

326 The RayXP results presented in this section show the variability of the $DoLP^{0-}$ at just below the
 327 air-water interface as a function of the ratio between the attenuation and the absorption
 328 coefficients at 440, 550, and 665 nm. $DoLP$ in-water is mainly affected by the scattering process
 329 of the particulates, which consist only of phytoplankton as assumed in our bio-optical model
 330 (Figure 1). The absorption of light modulates the changes in the $DoLP$ by reducing the number
 331 of scattering events in the water or by reducing the single scattering albedo.

332 Based on the bio-optical model, the largest variability in c/a occurs at 550 nm and 440 nm and
 333 smallest at 665 nm, due to the constant high seawater absorption as compared to the absorption
 334 and attenuation particulate and dissolved components. That in turn drives the dynamic range of
 335 the resultant $DoLP^{0-}$, spectrally. Figure 5 shows the relationship between the $DoLP^{0-}$ and c/a for
 336 40° viewing zenith angles and for azimuth viewing angles of 0° and 90° respectively.



337
 338 **Figure 5. Relationship between $DoLP^{0-}$ just below the sea surface at $\theta_{view} = 40^\circ$ and $\phi_{view} = 0^\circ$ ($\theta_{sca} = 111.2^\circ$) and**
 339 **90° ($\theta_{sca} = 132.2^\circ$) and c/a ratio at three wavelengths.**

340 The variability of the $DoLP^{0-}$ as a function of the c/a ratio is minimal for the 440 and 665 nm for
 341 the 0° azimuth viewing angle. At 550 nm, a noticeable variation exists in the relationship for
 342 most of the viewing geometries since the phytoplankton particles are the dominant constituent in
 343 these water, the $DoLP^{0-}$ becomes sensitive to their scattering coefficient in the green wavelength
 344 channel, where the DoLP is not dominated by Rayleigh scattering (as in the blue channel) and by
 345 the high water absorption (as in the red channel). Therefore, the variability of the c/a ratio due to
 346 the variations at 550 nm is high as compared to the blue and red channels.

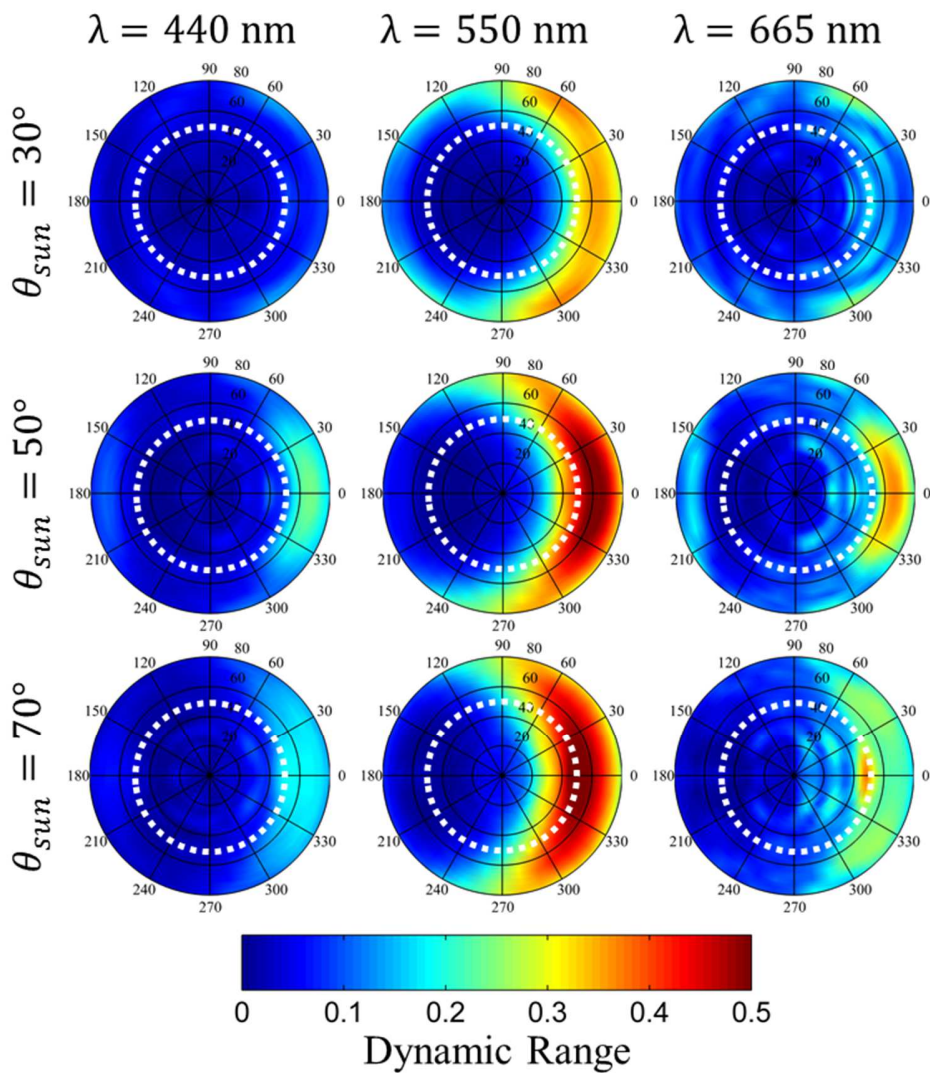
347 The changes of the $DoLP^{0-}$ versus c/a ratio at 90° relative azimuth angle are small. That could be
 348 due to the value of the scattering angle for this geometry (scattering angle is about 132°) which is
 349 higher than that corresponding to the geometry of 0° relative azimuth (scattering angle is about
 350 111°). DoLP is less sensitive to the composition of hydrosols at 132° than at 111° scattering
 351 angle. Despite the 90° relative azimuth angle is a preferred geometry for above water detection
 352 of the polarized water leaving radiance, the invariance of the signal versus the c/a ratio and
 353 therefore the [Chl] does not allow retrieval. These results corroborate with those of Harmel *et al.*
 354 (Harmel and Chami 2008, Harmel and Chami 2012).

355 To understand the geometrical constraints on the variability of $DoLP^{0-}$ versus c/a , we examine in
 356 Figure 6 the Dynamic Range (DR) of DoLP that is defined as follows

$$357 \quad DR(\theta_{view}, \varphi_{view}) = \frac{DoLP^{0-}(\theta_{view}, \varphi_{view})_{max} - DoLP^{0-}(\theta_{view}, \varphi_{view})_{min}}{DoLP^{0-}(\theta_{view}, \varphi_{view})_{theoretical_max} - DoLP^{0-}(\theta_{view}, \varphi_{view})_{theoretical_min}}, \quad (7)$$

358 The DR is calculated for each viewing and azimuth angle as the difference between maximal and
 359 minimal $DoLP^{0-}$ for the same geometry normalized by the difference between theoretical
 360 maximal and minimal $DoLP^{0-}$ simulated for all cases of IOPs at each specific geometry. The
 361 theoretical maximum is calculated from pure water (Rayleigh scattering) conditions, while

362 theoretical minimum is zero. The DR values vary between 0% and 50% of the maximal possible
 363 DoLP⁰ simulated with RayXP at each geometry. Higher DR values explain a large variability in
 364 the relationship and therefore allow for retrievals at specific viewing geometries.



365

366 *Figure 6. Synoptic view of the Dynamic Range (DR) at just below the air-water interface for Sun relative azimuth*
 367 *from 0° to 360° (0° azimuth is for Sun and sensor are in opposition) and viewing angle of upwelling polarized*
 368 *light from 0° to 80° (0° viewing angle is for sensor looking vertically downward) for three Sun zenith angles 30°,*
 369 *50°, and 70° and for three wavelength 440, 550, and 665 nm. Dashed white circles represent the border of Snell's*
 370 *window.*

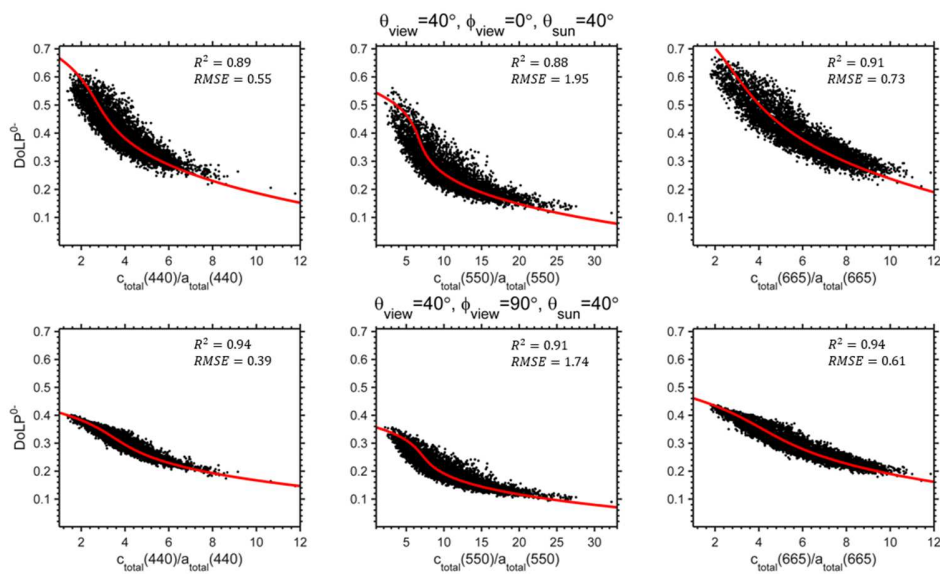
371 Figure 6 shows the DR for all viewing geometries of the upwelling light at just below the air-
372 water interface at three Sun zenith angles and at three wavelengths (440, 550, and 665 nm).
373 Lower values of DR show invariability of the DoLP⁰⁻ versus c/a ratio, thus making retrievals not
374 possible due to the lack of sensitivity in the relationship; on the other hand, higher DR shows
375 more variability in the DoLP⁰⁻ versus c/a for which retrieval is possible. Figure 6 demonstrates
376 that the DR is low at 440 and 665 nm for all of the geometries that corroborates with the scatter
377 plots shown in Figure 5. At 550 nm, the DR exhibits high values mainly in the principal plane of
378 the Sun (*i.e.* $\varphi_{view} = 0^\circ$) and near it. DR values in that scenario can approach 50% of the
379 theoretical DR assumed, showing the high sensitivity of DoLP⁰⁻ to variations in c/a at 550 nm
380 allowing retrievals. It is important to note that such a relationship depends highly on the Sun
381 zenith angle. Lower Sun angles (*i.e.*, larger zenith angle) show an increase in the DR as opposed
382 to a near overhead Sun. This analysis is done for underwater polarization, DoLP⁰⁻ and therefore,
383 measurements of the polarized water-leaving radiance, and will be affected by ocean surface
384 transmission, which limits $\theta_{view} \leq 48^\circ$. Furthermore, the Sun and sky glint corrections will
385 make retrievals even more challenging from above-water measurements. Nevertheless, several
386 options of above-water polarimetric observations in Case I waters should be further studied,
387 including possibilities of measurements in or near the principal plane ($0^\circ \leq \varphi_{view} \leq 60^\circ$) but
388 with different Sun and viewing angles to minimize Sun glint effects. Also of interest for future
389 studies are considerations of more complex (but closer to real) configurations of the
390 phytoplankton particles, with a broader range of refractive indices and hyperspectral
391 measurements of the DoLP or even Stokes components Q/I and U/I, which can reveal more
392 features in polarization signatures and facilitate additional relationships and retrievals.

393

394 **4.3. Case II simulation results**

395 **4.3.1. Assessment of the relationship between DoLP and c/a below the water**

396 Results for Case II water simulations at depths just below the air-water interface are presented
397 next at three wavelengths: 440, 550, and 665 nm. Figure 7 shows the variability of the DoLP⁰-
398 versus c/a ratio at 40° viewing angle and azimuth angles 0° and 90°. For this figure, we fix a
399 single solar zenith angle at 40° and vary the viewing angle to cover the largest range of scattering
400 angles in the backward direction. There is a strong relationship between the DoLP⁰ values
401 and c/a ratios with high variability in the broad range of both parameters for these specific
402 viewing angles in the Sun's principal plane, and at 90° away from it. As a result, it is possible to
403 easily fit the relationship, which allows us to retrieve the attenuation coefficient from the DoLP
404 measurements if the absorption coefficient a is known (note, a can be retrieved with reasonable
405 accuracy using inversion algorithms with good atmospheric correction (AC), although AC can be
406 challenging over turbid waters (Lee, Carder et al. 2002, Ahmad, Franz et al. 2010, Goyens, Jamet
407 et al. 2013)).



409 *Figure 7. Fitted relationship between DoLP⁰ just below the sea surface at $\theta_{view} = 40^\circ$ and $\phi_{view} = 0^\circ$ ($\theta_{sca} =$*
 410 *111.2°) (first row) and 90° ($\theta_{sca} = 132.2^\circ$) (second row) and c/a ratio at three wavelengths.*

411 The red lines in Figure 7 are parameterizations obtained by fitting a third-order polynomial as
 412 follows:

$$413 \quad \left(\frac{c}{a}\right)_{fit} = p_3 \times (DoLP)^3 + p_2 \times (DoLP)^2 + p_1 \times (DoLP)^1 + p_0 \quad (8)$$

414 Here, a and c are the absorption and attenuation coefficients, respectively; p_0 to p_3 are the
 415 polynomial fitting coefficients which vary with the viewing zenith and azimuth angles θ_{view} and
 416 ϕ_{view} . Furthermore, $DoLP \equiv DoLP^0$ for Figure 7.

417 The quality of the fitting of Eq. 8 can be estimated by calculating the root mean squared error
 418 (RMSE) or the coefficient of determination R^2 coefficient defined as

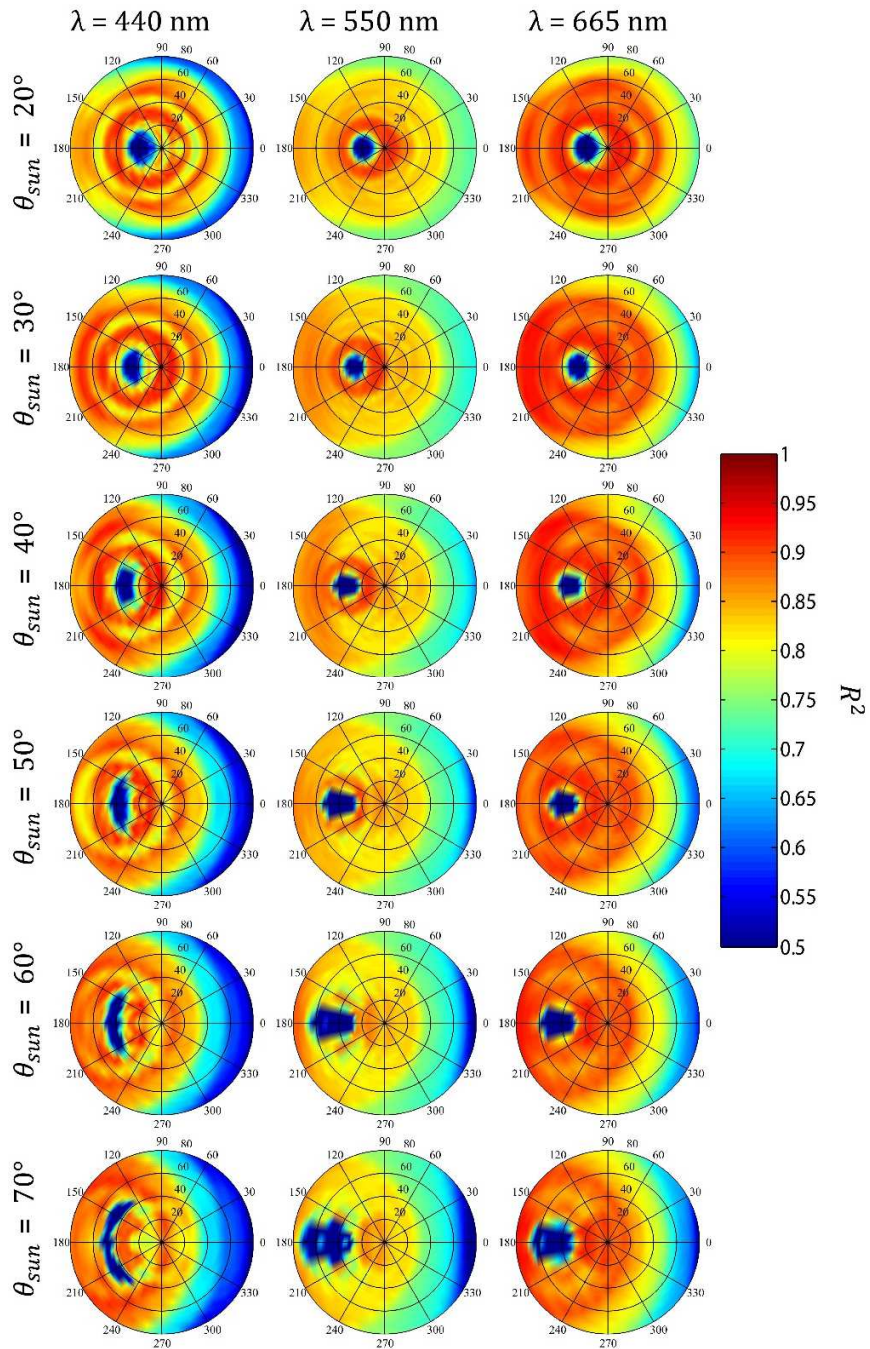
$$419 \quad R^2 = \frac{SSR}{SST} = \frac{\sum_{i=1}^{5000} \left[\left(\frac{c}{a}(DoLP_i) \right)_{fit} - \overline{\left(\frac{c}{a} \right)_{fit}} \right]^2}{\sum_{i=1}^{5000} \left[\left(\frac{c}{a} \right)_i - \overline{\left(\frac{c}{a} \right)_{fit}} \right]^2}, \quad (9)$$

420 In Eq. 9, SSR is the sum of squared differences between the regression fit $((c/a(DoLP_i)_{fit})$ and the
 421 sample mean $\overline{c/a}$, and ‘ i ’ iterates from 1 to 5000 different cases of IOPs in the RT simulations.
 422 SST stands for the sum of squares total, which means the sum of squared deviations of the (c/a)
 423 values around their mean. High values of R^2 indicate good fitting qualities and vice-versa.

424 The average error in the fitting, RMSE, is defined as

$$425 \quad RMSE = \sqrt{\frac{\sum_{i=1}^{5000} \left[\left(\frac{c}{a}(DoLP_i) \right)_{fit} - \overline{\left(\frac{c}{a} \right)_{fit}} \right]^2}{5000}} \quad (10)$$

426 In order to have a synoptic view of the fitting quality, the R^2 values are plotted in Figure 8 for all
 427 viewing geometries available for a given solar angle (*i.e.* $\theta_{sun} = 20^\circ$ to 70°). From this figure, we
 428 can estimate the range of geometries that permit the best retrieval of c from measurements of the
 429 DoLP just below the ocean surface.



431 *Figure 8. Synoptic view of the coefficient of determination R^2 at just below the air-water interface for Sun*
432 *relative azimuth from 0° to 360° (0° azimuth is for Sun and sensor are in opposition) and viewing angle of*
433 *upwelling polarized light from 0° to 80° (0° viewing angle is for sensor looking vertically downward).*

434 In Figure 8, the coefficient of determination (R^2) is higher than 0.9 for 665 nm and higher than
435 0.8 for 440 and 550 nm for most of the viewing geometries below the water surface, which
436 indicates a good and consistent relationship between the simulated data set and parameterization
437 using Eq. 8. It is also noticeable that R^2 degrades in the backscattering direction where the DoLP
438 is minimal. The high R^2 correlation for other viewing geometries in Figure 8 is very promising
439 when considering future air- or space-borne measurements of the polarized water-leaving
440 radiance, since it does not limit the range of viewing angles at which this type of sensor operates.
441 For example, the good correlation that exists at the meridian plane away from the Sun's principal
442 plane (e.g., 90° away and therefore away from Sun glint contaminations) makes the
443 measurements of polarized water-leaving radiance easier and more accurate and creates the
444 possibility of a direct estimation of the attenuation coefficient, which is otherwise physically
445 impossible when using above-water sensors. As already mentioned above, this method to retrieve
446 the attenuation coefficient of hydrosols can be used to study carbon-based ocean productivity
447 and phytoplankton physiology from remote sensing (Behrenfeld, Boss et al. 2005, Cetinic, Perry
448 et al. 2012). Figure 8 shows that excellent R^2 values can be achieved, especially at 440 and 665
449 nm, for a broad range of azimuth and viewing angles. This is most likely due to the moderate
450 absorption (usually dominated by the water absorption at 665 nm and CDOM absorption at 440
451 nm), which reduces the number of the scattering events and increases the DoLP in comparison
452 with 550 nm cases.

453 Table 1 below shows the polynomial fitting coefficients for $\theta_{view} = 40^\circ$ and Sun relative
 454 azimuth angle $\phi_{view} = 90^\circ$ at all three wavelengths for $\theta_{sun} = 30^\circ$ and $\theta_{sun} = 50^\circ$ respectively.

455

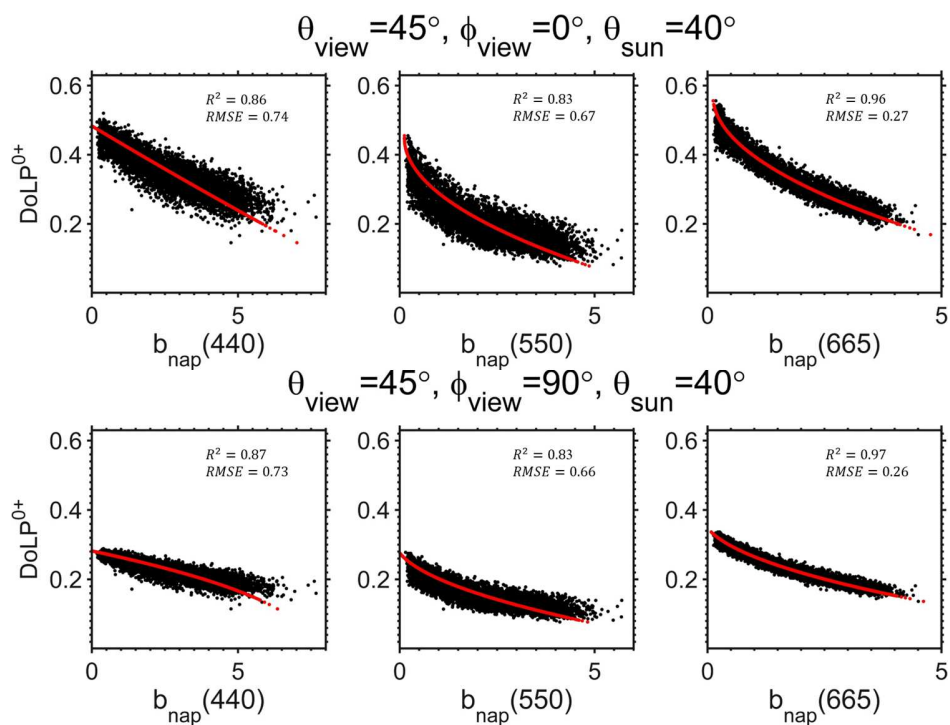
456 **Table 1. Fitting coefficients of the polynomial function for $\theta_{view} = 40^\circ$, $\phi_{view} = 90^\circ$, and $\theta_{sun} = 30^\circ$ ($\theta_{sca} =$
 457 135.3°) and $\theta_{sun} = 50^\circ$ ($\theta_{sca} = 128.8^\circ$).**

Fitting Coefficients	Wavelength (nm)					
	440		550		665	
	$\theta_{sun} = 30^\circ$	$\theta_{sun} = 50^\circ$	$\theta_{sun} = 30^\circ$	$\theta_{sun} = 50^\circ$	$\theta_{sun} = 30^\circ$	$\theta_{sun} = 50^\circ$
P_3	-1540	-783	-4179	-2407	-734	-350
P_2	1291	808	2907	2037	692	414
P_1	-384	-296	-709	-602	-243	-184
P_0	44	41	67	69	35	34

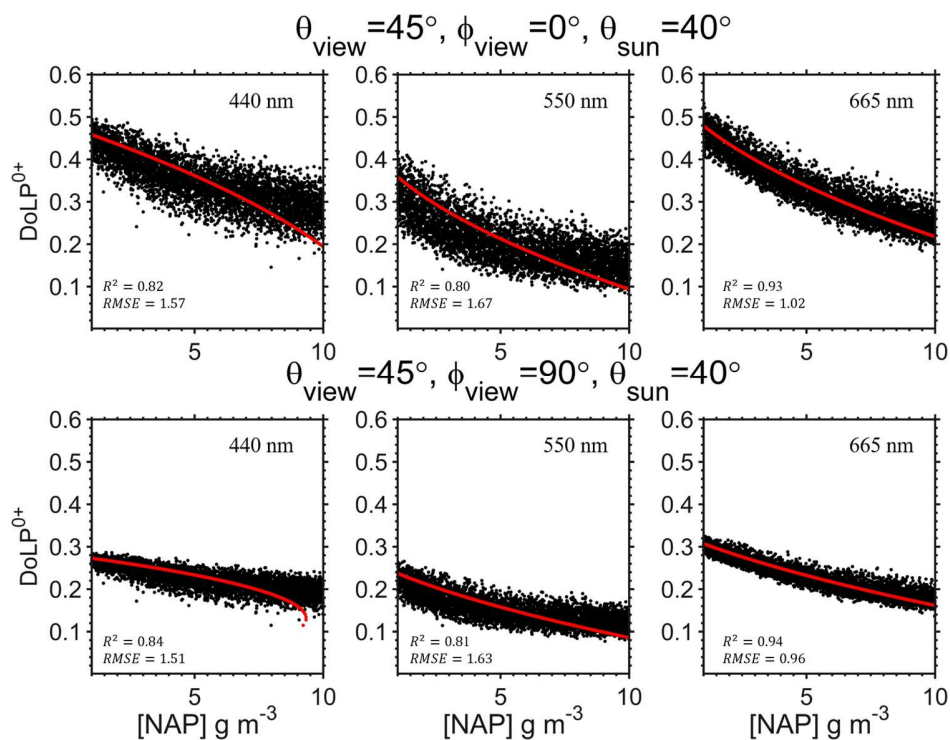
458

459 4.3.2. The retrieval of the concentration of minerals [NAP] from above water

460 The DoLP of upwelling light, either below or above the ocean surface, is highly sensitive to the
 461 scattering properties of high-refractive hydrosols. Our radiative transfer simulations show that
 462 there is a strong relationship between the scattering coefficient of high-refractive NAP matter
 463 and the DoLP. This relationship extends to the concentration [NAP] of high-refractive NAP
 464 matter such as minerals. Based on the conventional understanding of the polarization nature of
 465 light, higher scattering coefficients or high NAP concentrations lead to a higher number of
 466 multiple scattering events, therefore lowering the DoLP and vice versa. Such behavior is
 467 observed in Figure 9 for b_{nap} and in Figure 10 for [NAP]:



469 **Figure 9. Fitted relationship between DoLP^{0+} just above the sea surface at $\theta_{\text{view}} = 45^\circ$ and $\phi_{\text{view}} = 0^\circ$ ($\theta_{\text{sca}} = 95^\circ$)**
 470 **(first row) and 90° ($\theta_{\text{sca}} = 122.8^\circ$) (second row) and b_{nap} at three wavelengths.**



472 *Figure 10. Fitted relationship between DoLP⁰⁺ just above the sea surface at $\theta_{view} = 45^\circ$ and $\phi_{view} = 0^\circ$ ($\theta_{sca} = 95^\circ$)*
473 *(first row) and 90° ($\theta_{sca} = 122.8^\circ$) (second row) and NAP at three wavelengths.*

474 The DoLP⁰⁺ above water surface was calculated from the Stokes vector below the surface using
475 the transmission Mueller matrix for a flat sea-surface give by Kattawar and Adams (1989)
476 (Kattawar and Adams 1989). The co-variation of DoLP⁰⁺ with b_{nap} and with [NAP] is clearly
477 observed, especially at 665 nm, because the polarization of light is mainly driven by the
478 dominant NAP scattering particles in coastal waters. The DoLP becomes less affected by other
479 constituents such as phytoplankton and CDOM absorption or phytoplankton scattering at the red
480 part of the spectrum. Such relationships can be effective in retrievals by themselves and in a
481 combination with typical retrievals of [NAP] from remote sensing reflectance at 665 nm (Babin,
482 Morel et al. 2003, Babin, Stramski et al. 2003).

483 **4.3.3. The retrieval of the microphysical properties from above water**

484 The microphysical properties of hydrosols include but are not limited to the refractive index and
485 the size distribution. The polarization of scattered underwater light, and therefore of water-
486 leaving radiance, is highly sensitive to these properties. One can retrieve such particulate
487 properties from multi-angle polarization measurements by recursive fitting or optimization
488 techniques using RT simulations of polarized light (Chowdhary, Cairns et al. 2001, Waquet,
489 Cairns et al. 2009).

490 In this work, the approach is to use the spectrum of polarized light above the water to retrieve the
491 spectrum of the attenuation coefficient c_{sol} of suspended matter, from which microphysical
492 properties of this matter can be extracted (Twardowski, Boss et al. 2001). From the retrieval of
493 the spectral attenuation coefficient of suspended particulate matter, the slope of the PSD can be
494 approximated as follows (Hulst 1981):

495
$$c_{sol}(\lambda) = A\lambda^{-Y}, \quad (11)$$

496 where Y is the hyperbolic (power law) slope of the hydrosol's spectral attenuation and A is the
497 amplitude. We obtain Y from $c_{sol}(\lambda=440)$ and $c_{sol}(\lambda=665)$ using

498
$$Y = \frac{\log[c_{sol}(440)/c_{sol}(665)]}{\log[665/440]} \quad (12)$$

499 Since the slope of the spectral attenuation Y (assuming it follows a power law) is closely related
500 to the slope ξ of the PSD for a polydisperse hydrosols, it can be approximated as follows (Hulst
501 1981)

502
$$\xi = Y + 3 \quad (13)$$

503 A corrected model of Eq. 13 has been suggested by Boss *et al.* (2001) for cases where larger
504 sized particles are more significant (Boss, Pegau *et al.* 2001, Boss, Twardowski *et al.* 2001).

505 The retrieval of PSD is therefore possible using Eq.s 12 and 13 provided that the hydrosol
506 attenuation coefficient is well estimated. Our algorithms discussed in *Section 4.3.1* allow for the
507 retrieval of the total attenuation coefficient c as long as the total absorption coefficient a is
508 known. The total absorption and attenuation coefficients include contributions from water,
509 CDOM, and hydrosols (phytoplankton and NAP matter). Since the water properties are known in
510 our modeling, they can be easily subtracted from c and a . This leaves CDOM absorption to be
511 subtracted from c and a . In coastal waters that are highly affected by the CDOM, ignoring
512 CDOM absorption can lead to an overestimation in the slope of the PSD (*i.e.* falsely indicating
513 that the water contains higher densities of smaller particles) due to the high absorption at 440
514 nm. Several inversion algorithms that retrieve the CDOM absorption coefficient in coastal waters
515 can be used in this inversion scheme (Lee, Carder *et al.* 2002, Ioannou, Gilerson *et al.* 2011).

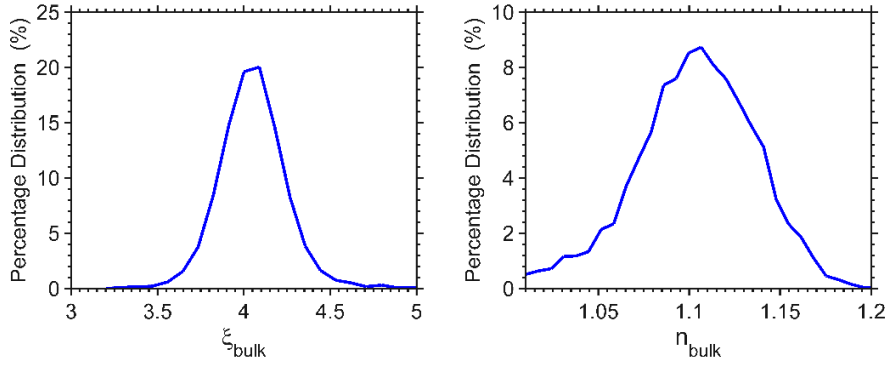
516 To retrieve the bulk refractive index n_{bulk} of the water body, we use an analytical model based
 517 on the Mie theory developed by Twardowski *et al.* (2001) (Twardowski, Boss *et al.* 2001). This
 518 model predicts the bulk ocean particulate refractive index n_{bulk} from Y as follows:

$$519 \quad n_{bulk}(\widetilde{b}_{b_{sol}}, Y) = 1 + \widetilde{b}_{b_{sol}}^{0.5377+0.4867(Y^2)} [1.4676 + 2.2950(Y^2) + 2.3113(Y^4)], \quad (14)$$

520 where $\widetilde{b}_{b_{sol}}$ is the backscattering ratio of hydrosols defined as

$$521 \quad \widetilde{b}_{b_{sol}} = \frac{b_{b_{sol}}}{b_{sol}}. \quad (15)$$

522 In Eq. 15, backscattering coefficient $b_{b_{sol}}$ can be estimated using the QAA (Lee, Carder *et al.*
 523 2002) method and hydrosol scattering coefficient b_{sol} can be obtained from $b_{sol} = c_{sol} - a_{sol}$.



524

525 **Figure 11. Histogram of the retrieval of the bulk refractive index n_{bulk} and the slope of the bulk PSD ξ_{bulk} from**
 526 **simulated DoLP⁰⁺ measurements.**

527 Figure 11 shows the retrieval of the microphysical parameters based on Eq. 11-15. The main
 528 inputs to the inversion was the DoLP⁰⁺ at a 40° viewing angle above water and 90° azimuth angle
 529 relative to the Sun at 440 and 665 nm. After the inversion of the hydrosol attenuation coefficient
 530 and given the total and CDOM absorption coefficients at 440 and 665 nm, the slope of the
 531 particulate attenuation coefficient was calculated as in Eq. 12 and then ξ_{bulk} is calculated as in

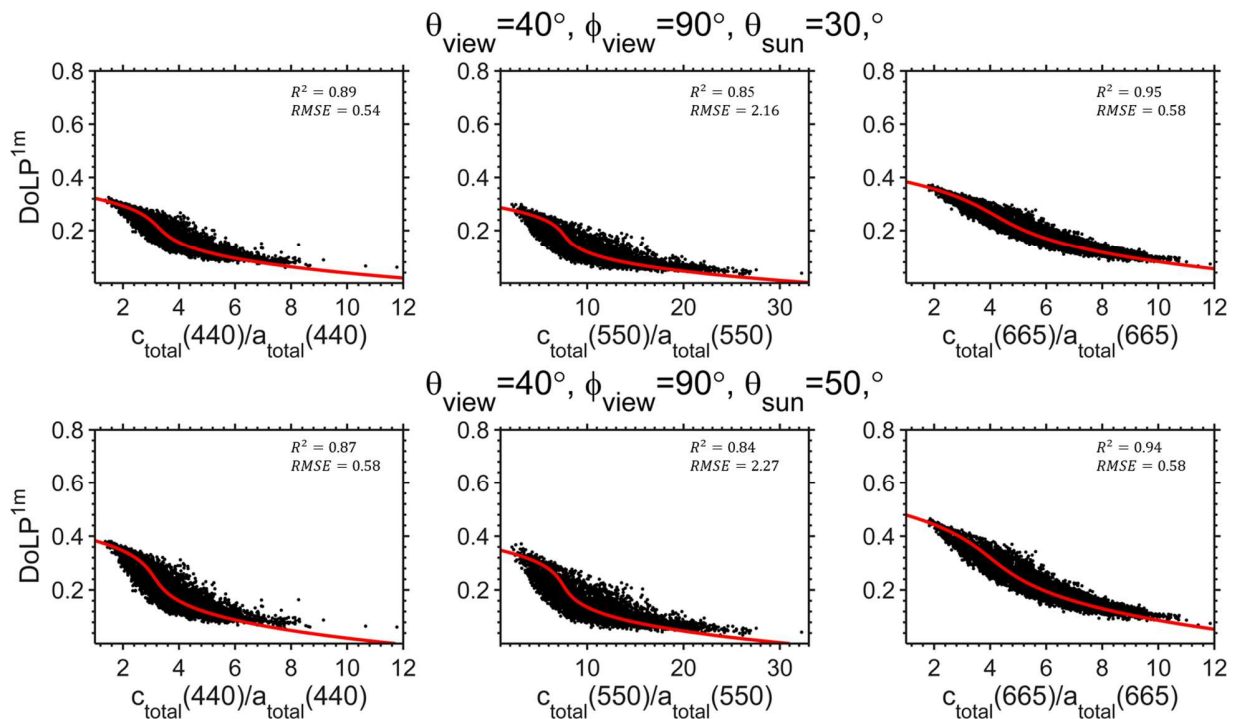
532 Eq. (13). Based on the knowledge of the backscattering ratio given by Eq. 15, the bulk refractive
533 index n_{bulk} is then calculated from Eq. 14. As shown in Figure 11, the ranges of the retrieved
534 microphysical parameters are within the typical ranges for natural oceanic waters (Twardowski,
535 Boss et al. 2001).

536 **5. Validation results for the inversion algorithm against field measurements**

537 **5.1 RT simulations at 1 meter below the water surface**

538 We validate the inversion algorithm discussed in *Section 4.3.1* using field measurements by an
539 underwater polarimeters (Tonizzo, Zhou et al. 2009). In order to carry out such validation,
540 another set of the fitting coefficients for the inversion algorithm is used to retrieve the
541 attenuation coefficient. Based on the RT simulation for the same IOPs described in *Section 3.2*,
542 the DoLPs were obtained at 1 meter below the sea surface. The reason for the 1-meter depth is
543 that the flotation system and geometry of the polarimeter design constrained it to measurements
544 at depths of 1 meter or more below the air-water interface.

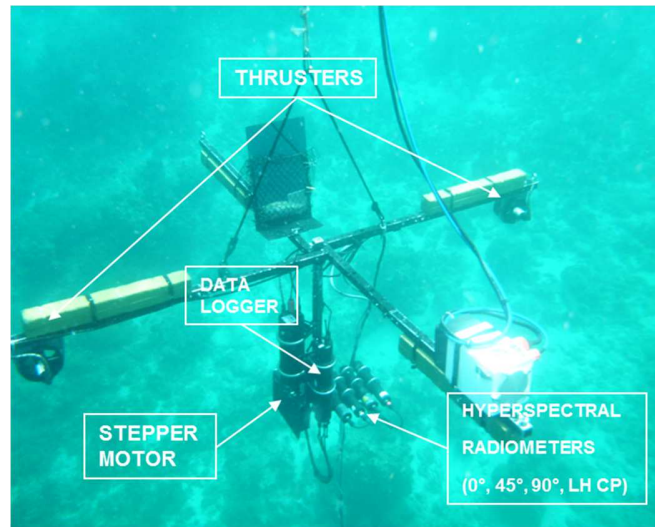
545 Figure 12 shows the relationship between the simulated DoLP at 1 meter below surface
546 ($DoLP^{1m}$) and the c/a ratio for three wavelengths at Sun zenith angles 30° and 50° , at 90° relative
547 azimuth angle, and at 40° in-water viewing zenith angle. That relationship is fitted with the third-
548 order polynomial, shown in the figure in red lines. The variability in the relationship is similar to
549 the $DoLP^{0-}$ just below the sea surface, though $DoLP^{1m}$ exhibits lower values because of the
550 increase in multiple scattered light at 1 me below.



552 **Figure 12.** Fitted relationship between the DoLP at 1 m below the sea surface and c/a ratio at three wavelengths
 553 at $\theta_{\text{view}} = 40^\circ$, $\phi_{\text{view}} = 90^\circ$ and $\theta_{\text{sun}} = 30^\circ$ ($\theta_{\text{sca}} = 135.6^\circ$) (first row) and 50° ($\theta_{\text{sca}} = 128.4^\circ$) (second row).

554 **5.2. Measurements and instrumentation**

555 In order to measure the underwater polarization of light, a hyperspectral, multi-angular
 556 polarimeter was developed by the Optical Remote Sensing Laboratory at the City College of
 557 New York, NY shown underwater in Figure 14 (Tonizzo, Zhou et al. 2009, Gu, Carrizo et al.
 558 2016).

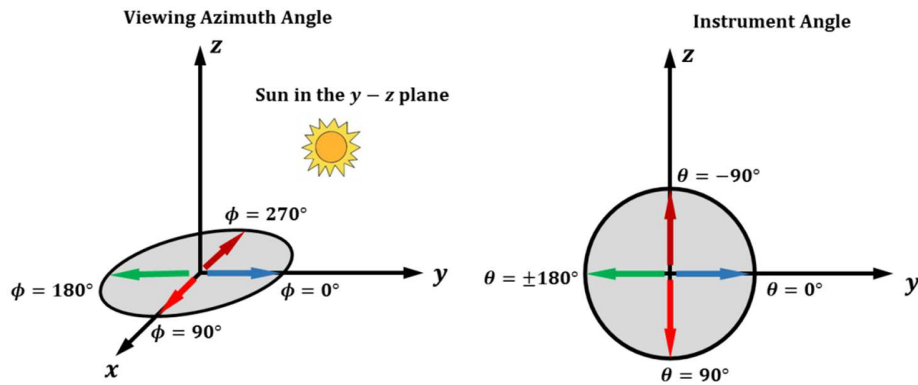


559

560 *Figure 13. The underwater polarimeter deployed to measure the polarization of light in-water.*

561 The instrument consists of three Satlantic Hyperspectral radiance sensors recording intensity
562 over 350 – 800 nm wavelength range, with a 8.5° in-water field of view and mounted on a
563 scanning system controlled by an underwater electric stepper motor (Newmark Systems, Inc.).
564 Its three hyperspectral radiance sensors (HyperOCRs, Satlantic) each receive light through a
565 linear polarizer (Edmund Optics) set to a specific angle: 0°, 45°, and 90° with respect to a
566 reference axis. At this point, one should be careful not to confuse the angles used in describing
567 the geometry of the polarimeter. While the relative angles of the linear polarizers fronting the
568 radiometers are held fixed, the whole radiometer system can be rotated over a full circle of
569 viewing angles. To distinguish these angles from the conventional viewing and viewing azimuth
570 angles, they are henceforth referred to as *instrument angles*. Figure 14 clarifies the distinction.
571 The stepper motor achieves zenith-viewing rotation in the vertical plane for the upwelling and
572 downwelling light. The whole polarimeter assembly uses a pair of remotely controlled thrusters
573 to achieve azimuth rotation and to perform measurements within the Sun principal plane or at

574 angles with respect to it. Additional details about this instrument are available in Tonizzo *et al.*
575 (2011) (Tonizzo, Gilerson *et al.* 2011).



576

577 *Figure 14. Definition of the viewing azimuth angle and the polarimeter's instrument angle.*

578 The IOPs of the water body include the absorption and scattering coefficients of the particulates,
579 and dissolved matter. The absorption and attenuation coefficients were measured hyperspectrally
580 in the visible spectrum using the ac-s (WET Labs) instrument. The ac-s has two operating
581 configurations: one measures the total absorption coefficients of both particulates and dissolved
582 substances, and the other filters the water intake with a 0.2 μm filter pores to pass through and to
583 measure only CDOM. The ac-s samples the IOPs over the depth profile of the water column. The
584 operations of the instrument were directed towards performing sampling at different locations
585 and water conditions.

586 **5.3. Retrieval of the attenuation coefficient c**

587 Since the summer of 2008, dozens of field measurements were collected at different stations
588 using the underwater polarimeter and the WET Labs in-situ instrument package. The data points
589 were collected at different locations near the northeastern coast of the US. The optical properties
590 of the water varied but were within the conditions typical of coastal waters. Table 2 shows the
591 geo-location of the stations used in the validation study.

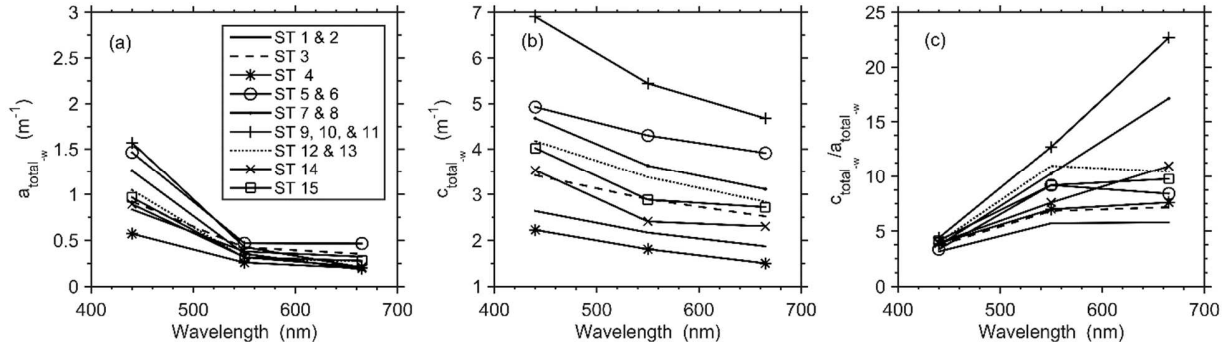
592 **Table 2. The stations used in the validation study.**

STATION NO.	DATE	LOCATION	SUN ZENITH	RELATIVE AZIMUTH	RANGE OF SCATTERING ANGLES (in-water)
ST 1	21-Jul-08	New Jersey Bight	30°	0°	77°-158°
ST 2	21-Jul-08	New Jersey Bight	45°	0°	68°-145°
ST 3	22-Jul-08	New Jersey Bight	40°	10°	72°-151°
ST 4	23-Jul-08	New Jersey Bight	40	0°	71°-151°
ST 5	4-Aug-08	Rockaway Inlet, NY	25°	0°	82°-162°
ST 6	4-Aug-08	Rockaway Inlet, NY	30°	20°	79°-158°
ST 7	9-Jul-09	New York Harbor, NY	20°	0°	85°-165°
ST 8	9-Jul-09	New York Harbor, NY	20°	30°	87°-165°
ST 9	9-Jul-09	New York Harbor, NY	20°	10°	85°-165°
ST 10	9-Jul-09	New York Harbor, NY	20°	30°	87°-165°
ST 11	9-Jul-09	New York Harbor, NY	25°	50°	88°-162°
ST 12	19-Aug-10	Proximate to LISCO*	35°	90°	99°-155°
ST 13	19-Aug-10	Proximate to LISCO	25°	10°	82°-162°
ST 14	19-Aug-10	Proximate to LISCO	30°	90°	99°-158°
ST 15	19-Aug-10	Proximate to LISCO	30°	90°	99°-158°

593 *Long Island Sound Coastal Observatory, part of AERONET-OC network (Harmel, Gilerson et al. 2011)

594 Table 2 shows that in-situ measurements were performed at different Sun illumination and
 595 viewing geometries. These stations passed quality control tests to reduce the environmental
 596 effects that can impact the in-situ optical measurements, such as limiting wind speed to 5 m/s
 597 and choosing stations with clear skies. The underwater polarimeter measured the polarized light
 598 with a full scan of the upwelling light at 1 m below the surface. At this depth, surface effects on
 599 the upwelling light are minimized, but the wind speed is limited to ensure that the apparatus is
 600 stable enough in the upright vertical position to reduce tilt biases.

601 The in-situ measurements of the attenuation and absorption coefficients are shown in Figure 15
 602 and cover a broad range of water conditions.



603

604

Figure 15. Total absorption, (a), attenuation, (b), coefficients and their ratios, (c), without the pure water

605

component at 440, 550, and 665 nm for 15 stations.

606

The absorption and attenuation coefficients displayed in Figure 15 do not include contributions

607

from pure sea water, and are therefore denoted with the subscript “-w” to differentiate them from

608

coefficients that include the sea water contribution. The variability of the coefficients as well as

609

the c/a ratio is for typical moderately turbid coastal waters. The lowest variability of c/a ratio is

610

at 440 nm, while its highest is at 665 nm. By adding the high absorption of the water component,

611

the range of the variability of the ratio becomes smaller. These observations matches the

612

theoretical IOPs modeling used in the VRT simulations.

613

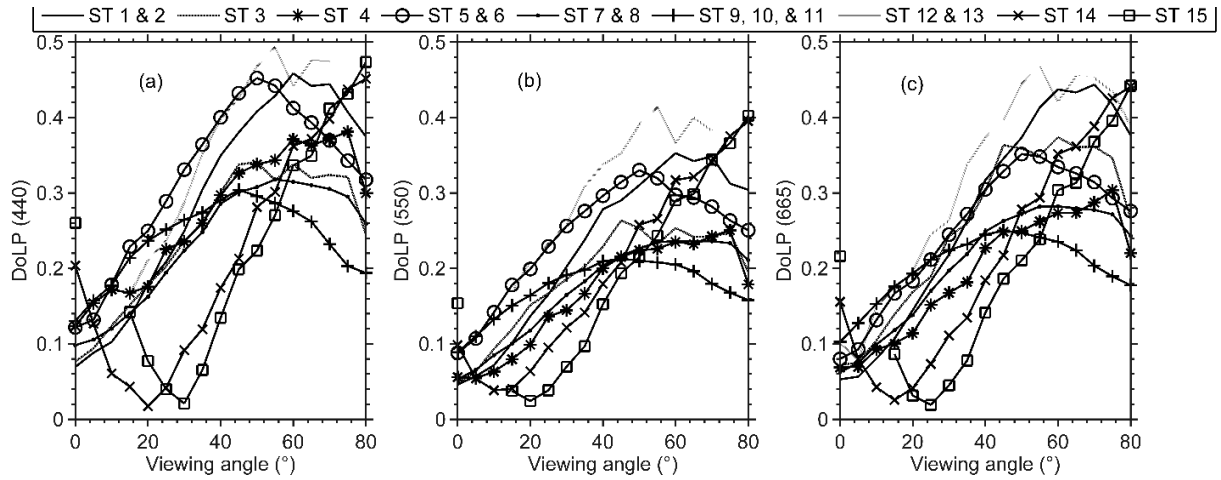
Figure 16 shows the DoLP at 1 m below the sea surface measured using the underwater

614

polarimeter. The DoLP were calculated from the Stokes components of the linearly polarized

615

light measured by the three radiometers as described in *Section 5.2*.

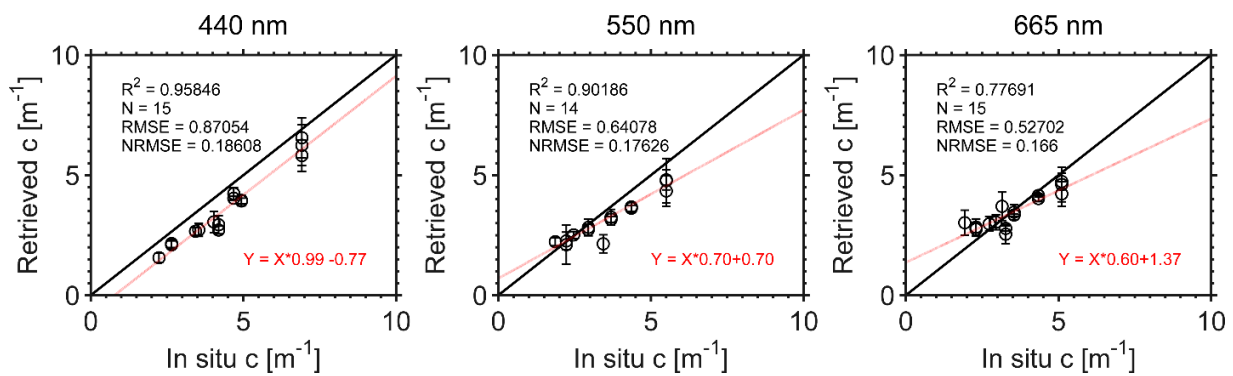


616

617 **Figure 16. DoLP versus viewing angle (as defined in Figure 5) at 1 m below at 440, 550, and 665 nm (a), (b), and**
 618 **(c), respectively for 15 stations. (0° viewing angle is when the detector is looking at the nadir).**

619 The fact that DoLP is not as smooth as the RT simulations predict is due to either radiometric
 620 uncertainties in the polarimeter or to other environmental factors (Tonizzo, Zhou et al. 2009).
 621 The uncertainties of the measurements could be high because of the low upwelling light signal at
 622 1 m below the surface, especially for turbid coastal waters and in the red part of the spectrum,
 623 where the light is highly attenuated (due to water absorption) as shown in Figure 15. The
 624 uncertainties also can increase for specific geometries such as the nadir direction where the
 625 radiances are minimal and the DoLP is low in the near backscattering directions (*i.e.* $\theta \sim 180^\circ$).
 626 The spectral variations of the DoLP are similar to the RayXP simulations predictions, where
 627 DoLP at 440 and 665 nm exhibits higher values than at 550 nm. That is due to the increase of the
 628 total absorption coefficients at the blue and red wavelengths reducing multiple scattering, while
 629 in the green, the absorption is minimal, and increased multiple scattering occurs. Therefore, the
 630 absorption of light by hydrosols has a modulating effect on the polarization of light with respect
 631 to baseline Rayleigh-like scattering and hydrosols scattering themselves.

632 Based on the DoLP measurements in Figure 16, it is possible to retrieve the ratio between the
 633 attenuation and absorption coefficients (c/a) using the polynomial coefficients from Section 4.3.1
 634 for a given geometry of viewing, Sun zenith, and relative azimuth angle. After applying the
 635 polynomial function from Eq. 8 to retrieve c/a , the ratio is then multiplied by the total absorption
 636 coefficient (including water) measured by the ac-s instrument shown in Figure 15(a). The
 637 retrieved total attenuation coefficient from the DoLP is then compared to the total attenuation
 638 coefficient (after adding the water component) measured by the ac-s instrument. A scatter plot of
 639 the retrieval of the attenuation coefficient based on the DoLP measurements is shown in Figure
 640 17 where the coefficients were averaged over the viewing angles from 40° to 80° .



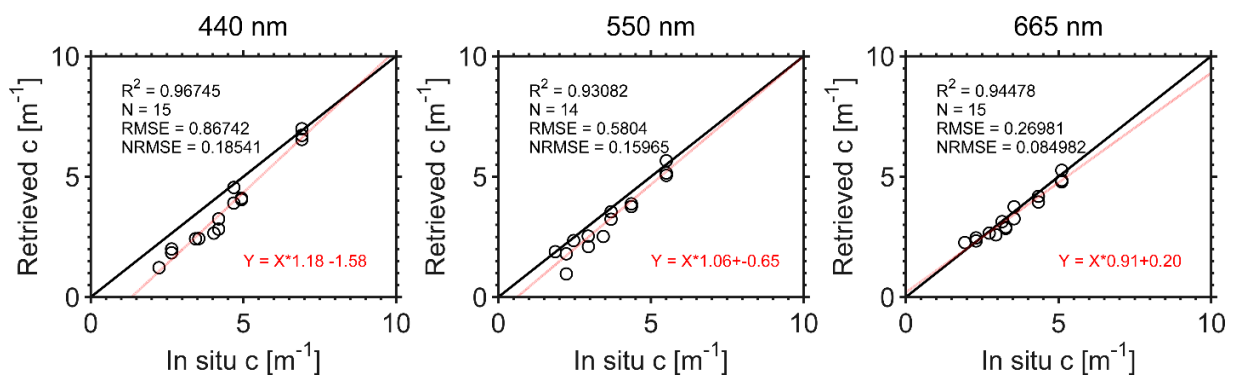
641

642 **Figure 17. The retrieval of the total attenuation coefficient averaged over different viewing angles from 40° to**
 643 **80° and the standard deviation represented as the error bars for three wavelengths.**

644 In Figure 17, the coefficient of determination R^2 is greater than 0.9 for both the 440 and 550 nm,
 645 while it degrades for the 665 nm. The root-mean-square error (RMSE) and normalized root-
 646 mean-square error (NRMSE) values at 665 nm are lower than at 440 and 550 nm, indicating
 647 lower error in the retrieval at the 665 nm, while the decreased correlation is due to the reduction
 648 in the dynamic range of the attenuation coefficient, where it is typically lower at the red part of
 649 the spectrum. Additionally, the number of available stations is limited; therefore, outliers can

650 strongly affect the correlation indicators. At 665 nm, the correlation decreases at geometries
651 where the polarimeter is looking in the nadir direction, as the upwelling radiance is small due to
652 the absorptive nature of the water in that part of the spectrum.

653 Figure 18 shows the retrieval of c from the DoLP measured at 75° viewing angle. The R^2 values
654 are very good for the three wavelengths, where the highest value is at the 440 nm, while the
655 lowest NRMSE is at the 665 nm.



656

657 **Figure 18. The retrieval of the total attenuation coefficient for 75° viewing angle for three wavelengths.**

658 In both Figures 17 and 18, the red lines show the linear regression between the in-situ and the
659 retrieved c . The slope and bias in all the cases might be due to multiple sources of error, such as
660 radiometric uncertainties, which play a major role in introducing differences between retrieved
661 and measured in-situ values. For example, at the cases where the attenuation coefficient
662 increases, the radiometric uncertainties of the polarimeters increase as well due to the decrease in
663 the measured DoLP. Therefore, the retrieval of c in turbid waters could be more challenging due
664 to signal to noise ratio limitations. Meanwhile, the in-situ measurements might also introduce a
665 slope as a result of calibration errors of the ac-s instrument (Mckee, Piskozub et al. 2008). The
666 inverse model itself also might introduce some of the uncertainties in the retrieval because of the
667 inadequacy either of the bio-optical model, or in the polynomial fitting function that does not

668 produce a high quality fit. While investigating of such problems (*i.e.* calibration and modeling) is
669 necessary, due to the lack of in-situ measurements of the DoLP and the IOPs at the appropriate
670 environmental and illumination/viewing geometries, and the lack of microphysical properties of
671 the ocean (*i.e.*, refractive index and PSD) that reproduces the scattering matrix of the polarized
672 light used in the hybrid bio-optical model, this work should be revisited when such data are more
673 available within the ocean color community.

674 **6. Remote sensing application**

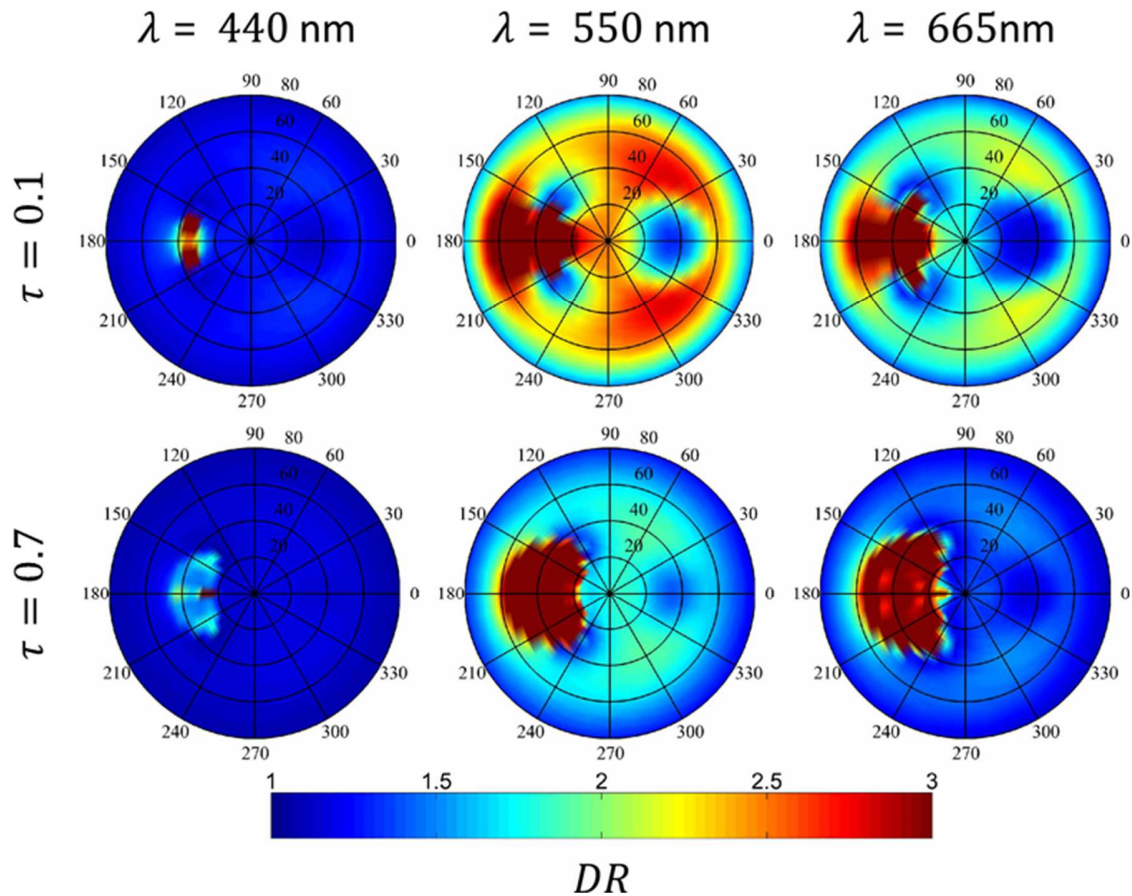
675 Remote sensing algorithms for ocean color retrievals from satellite measurements require
676 atmospheric correction of the top of atmosphere (TOA) radiance. From scalar satellite TOA
677 measurements, AC algorithm removes the contribution of atmospheric path radiance (*i.e.*
678 Rayleigh, and aerosol scattering) and compensates for gaseous absorption in the atmosphere, as
679 well as ocean surface glint. These effects typically are pre-computed using a fully coupled
680 radiative transfer simulations for pre-determined aerosol models, optical thicknesses, and
681 geometries. The residual signal is the ocean contribution to derive ocean color products such as
682 chlorophyll concentration and IOPs. Likewise, polarimetric remote sensing algorithms will
683 require adequate knowledge of the atmospheric effects on the TOA polarimetric signal. The
684 TOA polarized signature is largely attributed to the single scattering of light by the mix of
685 aerosol particles and atmospheric molecules (Hansen and Travis 1974, Mishchenko, Cairns et al.
686 2007, Chowdhary, Cairns et al. 2012). To understand the impact of the atmosphere on the
687 polarized signature of the ocean emerging at the TOA, we simulated 5000 cases of coastal waters
688 and analyzed the sensitivity of the relationship between DoLP^{TOA} and c/a ratio for clear and
689 highly turbid atmosphere (*i.e.*, $\tau(400\text{ nm}) = 0.1$ and 0.7 , respectively). Since the TOA signal's
690 main contribution is from the atmosphere (with fixed optical properties of the continental aerosol

691 model), while the signal from the oceanic layer with varying optical parameters is to be assessed,
692 we performed a sensitivity analysis by calculating the dynamic range (DR^{TOA}) of the $DoLP^{TOA}$
693 for all possible viewing geometry at the TOA. DR for the 5000 cases of oceanic layer is
694 calculated as follows:

$$695 \quad DR^{TOA} = \frac{\max(DoLP^{TOA})}{\min(DoLP^{TOA})}, \quad (16)$$

696 where the DR^{TOA} change is ≥ 1 . Higher DR signifies higher variability in the $DoLP^{TOA}$ versus c/a
697 ratio and therefore a larger impact of the ocean signal at the TOA.

698 Figures 19 is the synoptic polar plots of the DR^{TOA} for three wavelengths and at Sun zenith angle
699 $\theta_{sun} = 30^\circ$ for $\tau = 0.1$ and 0.7 . Of note, the 440 nm channel shows a low DR which can be
700 accounted for by the high contribution of the atmospheric molecular and aerosols scattering.
701 Conversely, DR^{TOA} at 550 and 665 nm is much higher, reaching more than 3, depending on the
702 viewing, Sun geometry and AOT. Higher DR does not necessarily allow a better retrieval of the
703 ocean parameters. Two conditions should be satisfied to allow retrieval at the TOA: 1) the DR is
704 large enough to allow the measurements of Stokes components with low radiometric
705 uncertainties at the TOA, and 2) there must be a definitive relationship between DoLP and c/a .
706 Therefore, the use of a single viewing geometry for the retrieval needs to be carefully assessed
707 for possible illumination scenarios and atmospheric conditions.



708

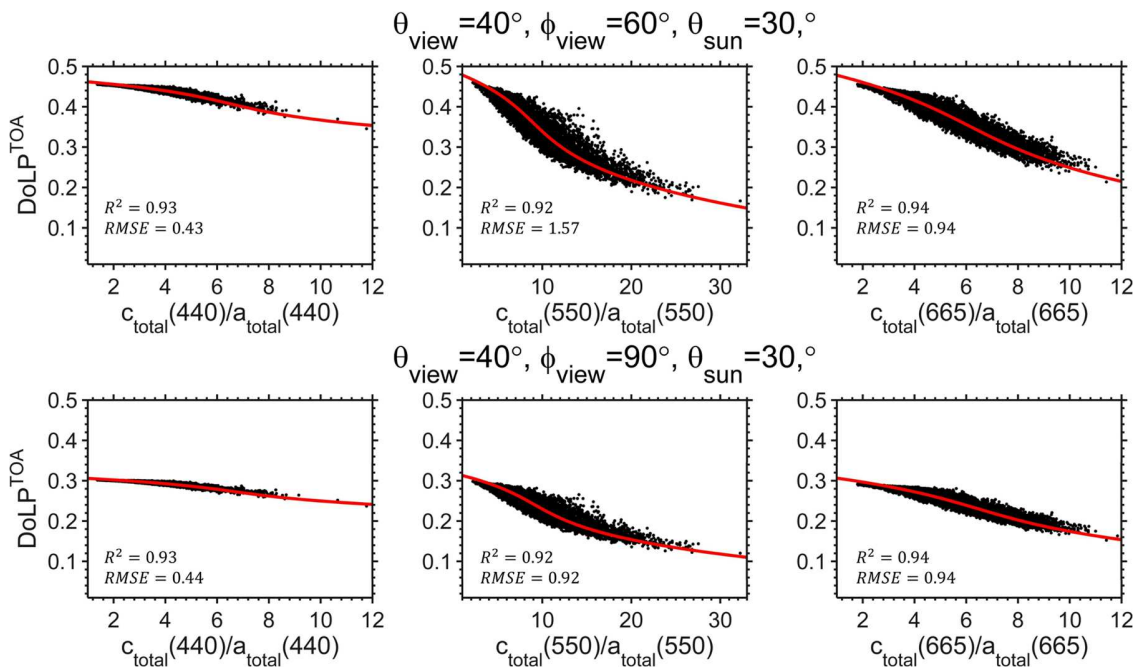
709 *Figure 19. Synoptic polar plot of the Dynamic Range (DR) calculated for $\tau = 0.1$ and 0.7 , and for Sun zenith 30°*
 710 *at 440, 550, and 665 nm.*

711 Increasing AOT leads to degradation of the relationship between DoLP and c/a ratio. The DR
 712 decreases dramatically for all wavelengths and viewing/illumination geometry. The polarized
 713 signature of the atmosphere masks the influence of the hydrosols on the degree of polarization at
 714 the TOA. The highly turbid atmosphere scenario cannot allow the retrieval of the c/a ratio.

715 The high DR regions in the polar plots are in the backscattering direction where the light is
 716 completely or highly unpolarized. At such geometry, there is no relationship between the DoLP
 717 in water and the c/a ratio. Therefore, the high DR is likely due to the fact that both numerator

718 and denominator values of DR are close to zero, which thus induces meaningless values of the
 719 ratio DR.

720 In the next figure, we show the scatter plot between DoLP^{TOA} and c/a at 40° viewing zenith angle
 721 and 90° and 60° viewing azimuth angle. The 60° azimuth angle shows a reasonable high DR of 2
 722 and above.

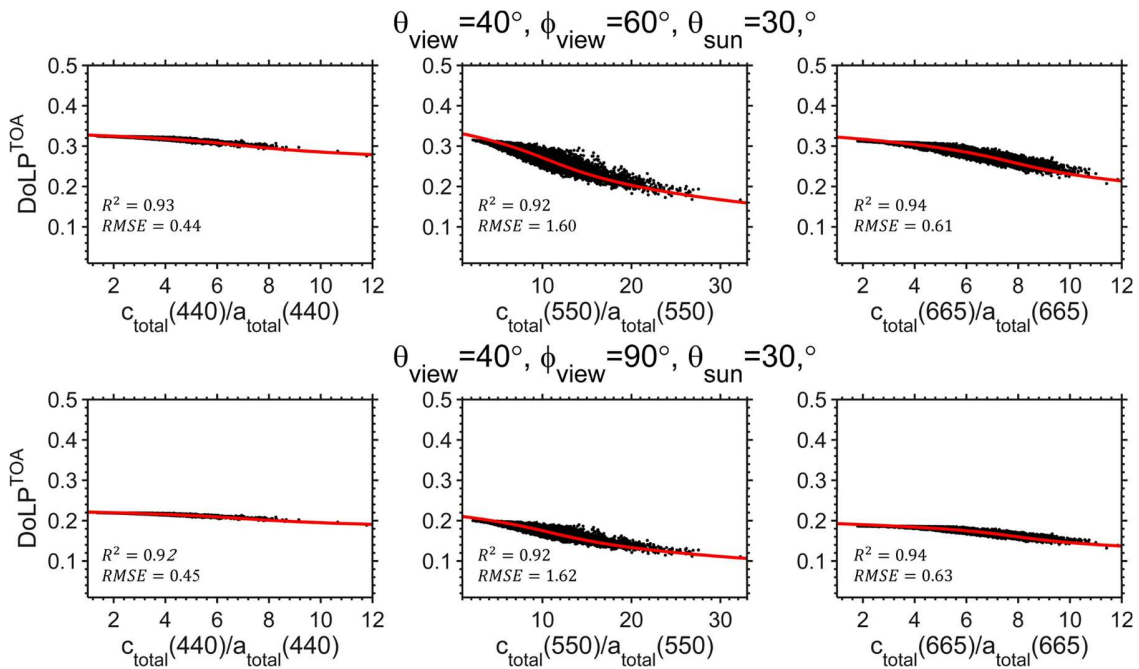


723
 724 **Figure 20.** The relation between DoLP^{TOA} , simulated for $\tau = 0.1$, and c/a at the TOA at three wavelengths, at 40°
 725 viewing zenith angle and for Sun relative azimuth angle of 90° and 60° . Sun zenith angle is fixed at 30° . ($\theta_{\text{sca}} =$
 726 120.2°) (first row) and 50° ($\theta_{\text{sca}} = 131.5^\circ$) (second row).

727 In Figure 20, there is a distinctive relationship at all three wavelengths and at the two viewing
 728 geometry when $\tau = 0.1$. At 440 nm, the DR of the DoLP^{TOA} is very small, making it more
 729 difficult to retrieve ocean's optical properties, whereas at 550 and 665 nm, the DR is much
 730 higher and the relationship between DoLP^{TOA} and c/a is strong. The DR at 60° relative azimuth
 731 viewing angle is higher than at 90° . At 60° azimuth angle, the ocean signal can be contaminated

732 by the Sun glint depending on the wind speed. Higher wind speed can lead to a higher
 733 contribution of the glint and therefore a contamination of the total signal at the TOA.

734 Figure 21 shows the same scatter plot as in Figure 20 but for $\tau = 0.7$. The turbid atmosphere, as
 735 expected, shows a lower DR while it retains the relationship between DoLP and c/a . The 440 nm
 736 channel shows a large degradation in the sensitivity due to the increasing aerosol's impact on the
 737 TOA signal. On the other hand, the relationship at the 550 nm is obvious especially at 60°
 738 azimuth angle. The 665 nm channel shows degradation in sensitivity, which could account for
 739 the weak polarized radiometric quantities that emerges from the ocean at the TOA.



741 **Figure 21. The relation between $DoLP^{TOA}$, simulated for $\tau = 0.7$, and c/a at the TOA at three wavelengths, at 40°**
 742 **viewing zenith angle and for Sun relative azimuth angle of 90° and 60° . Sun zenith angle is fixed at 30° . ($\theta_{sca} =$**
 743 **120.2°) (first row) and 50° ($\theta_{sca} = 131.5^\circ$) (second row).**

744 The synoptic polar plots in Figure 19 show that the 440 nm channel is a good channel for the
 745 retrieval of aerosol parameters due to the low contribution of the ocean polarized signature of

746 Case II water to the total signal at the TOA at all viewing geometry. The low DR is also apparent
747 at the 550 and 665 nm channels in the principal plane and near it.

748 Based on the previous sensitivity analysis, the following recommendation can allow a future
749 development of polarimetric ocean color retrieval from future polarimetric satellite missions:

750 1) TOA polarized light is weakly dependent on the polarized ocean contribution in the
751 shorter wavelengths which corroborates with the results obtained by Chami (2007)
752 (Chami 2007), thus allowing AC of the aerosols contribution in-conjunction with Near-
753 infrared (NIR) AC algorithms.

754 2) Deriving the attenuation coefficient from DoLP^{TOA} is possible at the green-red
755 wavelength over coastal waters under a given specific geometry.

756 3) Based on this study, a single viewing geometry can allow retrieval of the attenuation
757 coefficient for angle range $\theta_{\text{view}} = 40^\circ \sim 70^\circ$ and $\varphi_{\text{view}} = 30^\circ \sim 90^\circ$.

758 **7. Conclusion**

759 This work examines the development of an inversion methodology, which exploits the
760 polarization properties of scattered light to retrieve optical and microphysical properties of
761 oceanic waters. The theoretical modeling presented in this work encompasses a full
762 characterization of the relationship between the DoLP and the ratio of attenuation/absorption
763 coefficients (c/a). It is based on vector RT simulations performed for plane-parallel AO system,
764 and examines the impact of different illumination/viewing geometry on the DoLP. An inverse
765 algorithm is developed to retrieve the attenuation coefficient c at 440, 550, and 665 nm. Based
766 on this retrieval, microphysical properties of hydrosols, such as the bulk refractive index and the
767 PSD, can also be retrieved. The strong dependence of the polarization signature on multiple

768 scattering permits parameterization of the relationship between DoLP and the concentration of
769 in-organic mineral particles, [NAP], from which [NAP] at 665 nm can possibly be retrieved.

770 Two bio-optical models are developed to address, separately, the type and the optical variability
771 of open (Case I) oceans and coastal (Case II) oceans. These models were used as inputs to the
772 VRT program, RayXP, to simulate the Stokes components of light at just below the sea surface,
773 which shows the strong dependency of the DoLP on the ratio of the attenuation/absorption
774 coefficients (c/a) for Case II waters. An analytical third-order polynomial fit is then applied to
775 parameterize this dependency, which, in turn, is used for the retrieval of the attenuation
776 coefficient c at multiple illumination/viewing geometry. Retrievals with the parameterized
777 relationship are assessed for uncertainties at the three wavelengths of the simulations showing R^2
778 values higher than 0.85 for most of the geometries (except in the backscattering direction). For
779 above-water detection, the 90° azimuth angle and 40° viewing zenith angle is minimally affected
780 by Sun glint, permitting easier corrections of surface glint effects in the upwelling polarized
781 reflectance, while a strong relationship between DoLP and c/a still persist allowing the retrieval.

782 For this geometry, we also developed an inversion method to obtain the bulk refractive index and
783 the PSD of hydrosols from the retrieved spectral attenuation coefficient at two wavelengths. This
784 inversion approach obtains the slope of the Junge PSD from the slope of the particulate spectral
785 attenuation coefficient. Then, following the approach developed by Twardowski *et al.*
786 (Twardowski, Boss et al. 2001), the bulk refractive index of hydrosols can be retrieved given the
787 backscattering ratio coefficient and the PSD.

788 Case I waters, as opposed to Case II waters, show a weaker relationship between DoLP and c/a ,
789 especially at 440 and 665 nm, because the polarized signal is masked by Rayleigh scattering in
790 the blue channel and the high water absorption in the red channel. However, at 550 nm, there is a

791 noticeable relationship for most of the viewing geometries, except for those near the
792 backscattering direction. This is so because the phytoplankton particles are the dominant
793 constituent in open water conditions, and DoLP becomes sensitive to their scattering coefficient
794 in the green wavelength. In the solar half-principal plane, the relationship showed its maximum
795 dynamic range or sensitivity. Several options are discussed for further possible applications of
796 polarimetric observations and retrievals in Case I waters.

797 To validate our inversion approach, measurements of the DoLP with a custom-built underwater
798 polarimeter are used for different water conditions and illumination/viewing geometries.
799 Absorption and attenuation coefficients are measured at 1 meter below the sea surface, while a
800 new parameterization for the relationship between DoLP and c/a was used to retrieve the
801 attenuation coefficient c . The results show promising retrievals at the 440, 550, and 665 nm,
802 although additional validation points are required to perform a more thorough validation
803 analysis. The averaged DoLP measured over viewing angles ranging from 40° to 80° lead to
804 retrievals with R^2 higher than 0.9 at 440 and 550 nm. This correlation degrades at 665 nm due to
805 the limited variability of the attenuation coefficient at that wavelength, in addition to the limited
806 number of validation points used, thus increasing impact of outliers. However, retrievals at 665
807 nm show NRMSE values similar to those at other wavelengths. The retrievals of c show
808 improvements at the 75° viewing angle for the three wavelengths, with R^2 higher than 0.93. The
809 NRMSE values are generally lower than 0.18 and can be as low as 0.08 for the 665 nm channel.
810 This improvement in retrievals is due to the reduction in radiometric uncertainties and the
811 physical stability of the polarimeter system when measuring underwater light.

812 The TOA preliminary sensitivity analysis of the DoLP^{TOA} showed that it is possible to retrieve
813 the attenuation coefficient using a single viewing geometry at the green-red wavelengths, while

814 the DoLP^{TOA} at 440 nm is mainly attributed to the atmospheric scattering, which can be used for
815 the development of AC algorithms. A complete remote sensing algorithm needs to be developed
816 for future satellite missions, such as PACE, 3MI, and GCOM-C, that will utilize polarimetric
817 remote sensing. The work presented here can provide a starting point for the development of a
818 pre-launch algorithm.

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