

# Explicit Ensemble Prediction of Hail in 19 May 2013 Oklahoma City Thunderstorms and Analysis of Hail Growth Processes with Several Multimoment Microphysics Schemes

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## ABSTRACT

Hail forecast evaluations provide important insight into microphysical treatment of rimed ice. In this study we evaluate explicit 0–90-min EnKF-based storm-scale (500-m horizontal grid spacing) hail forecasts for a severe weather event that occurred in Oklahoma on 19 May 2013. Forecast ensembles are run using three different bulk microphysics (MP) schemes: the Milbrandt–Yau double-moment scheme (MY2), the Milbrandt–Yau triple-moment scheme (MY3), and the NSSL variable density-rimed ice double-moment scheme (NSSL). Output from a hydrometeor classification algorithm is used to verify surface hail size forecasts. All three schemes produce forecasts that predict the coverage of severe surface hail with moderate to high skill, but exhibit less skill at predicting significant severe hail coverage. A microphysical budget analysis is conducted to better understand hail growth processes in all three schemes. The NSSL scheme uses two-variable density-rimed ice categories to create large hailstones from dense, wet growth graupel particles; however, it is noted the scheme underestimates the coverage of significant severe hail. Both the MY2 and MY3 schemes produce many small hailstones aloft from unrimed, frozen raindrops; in the melting layer, hailstones become much larger than observations because of the excessive accretion of water. The results of this work highlight the importance of using a MP scheme that realistically models microphysical processes.

## 1. Introduction

Each year, hail causes on average more than \$1.4 billion (U.S. dollars) in property and crop damage in the United States (Changnon et al. 2009). Densely populated regions in particular are susceptible to costly hail damage; for example, the 1995 “Mayfest” hailstorm near Fort Worth, Texas, caused more than \$2 billion in damage and injured 109 people at an outdoor festival (Edwards and Thompson 1998). The cost of hail-related damage is expected to increase as cities expand; most hail events that have caused more than \$1 billion in insured losses have occurred since

2001 and the frequency of occurrence of such events is increasing (Changnon 2009).

Extending the warning lead time for severe hail has the potential to mitigate hail-related damage, though this is challenging, as severe hail events often develop quickly. To extend severe weather warning lead time, the National Weather Service (NWS) is shifting from issuing severe weather warnings based upon detection by, for example, radar (warn-on-detection), to a paradigm where warnings will be issued based upon high-resolution numerical weather prediction (NWP) model guidance (warn-on-forecast; Stensrud et al. 2009, 2013). Many studies analyze the skill of convective-scale forecasts; however, only a handful of studies (e.g., Milbrandt and Yau 2006; Snook et al. 2016, hereafter

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S16; Luo et al. 2017; Gagne et al. 2017; Labriola et al. 2017, hereafter L17; Adams-Selin and Ziegler 2016; Luo et al. 2018) assess the skill of explicit hail prediction (hail size and coverage). Hail prediction using storm-scale NWP forecasts remains understudied because of the difficult and complex nature of hail storm prediction, hail growth processes, and challenges associated with hail forecast verification.

The skill of explicit hail prediction is dependent upon the NWP model's ability to predict the complex processes and the storm environment that support the growth of hail; this requires an accurate representation of the hailstorm structure, dynamic, thermodynamic, and microphysical processes, as well as the surrounding environment. Hail grows where supercooled liquid accretes to the surface of rimed ice; however, growth is limited if a large number of hail embryos depletes the supercooled liquid (e.g., Heymsfield 1983). Hail particle growth is strongly correlated with storm updraft strength (e.g., Rasmussen and Heymsfield 1987). Strong updrafts suspend large rimed ice particles in regions of supercooled liquid; however, if an updraft is too strong the particles are more rapidly advected outside of the optimal growth zone. Strong vertical wind shear also contributes to hail growth (e.g., Nelson 1983; Foote 1984; Dennis and Kumjian 2017). Increased deep layer shear extends the storm updraft horizontally and increases the residence time of hail within the growth zone. Because of the complex nature of riming, a large number of microphysical parameterizations that have different treatments for rimed ice were developed with the goal of improving the representation of rimed ice.

Microphysical processes in convective-scale NWP models are mostly parameterized using bulk microphysics (MP) schemes (e.g., Lin et al. 1983; Milbrandt and Yau 2005a). Bulk MP schemes predict the bulk characteristics for hydrometeor species  $x$  such as mixing ratio  $q_x$ , total number concentration  $N_{0x}$ , or reflectivity  $Z_x$  to define a commonly used three-parameter gamma particle size distribution (PSD):

$$N(D) = N_{0x} D^{\alpha_x} e^{-\lambda_x D}, \quad (1)$$

where  $\lambda_x$  is the slope parameter,  $\alpha_x$  is the shape parameter, and  $N_{0x}$  is the intercept parameter (Ulbrich 1983). A few MP schemes (Mansell et al. 2010; Morrison and Milbrandt 2015; Morrison et al. 2015; Milbrandt and Morrison 2016) also have the capacity to predict hydrometeor volume ( $v_x$ ) and thus explicitly derive density ( $\rho_x$ ); this is primarily done for rimed ice hydrometeor categories, which undergo large fluctuations in density during growth and melting. Variable density

MP schemes (e.g., Mansell et al. 2010) have the ability to update particle fall speeds and rimed ice production.

To save computational time, one or more PSD parameters are often assumed to be constant, though such assumptions can limit the model's ability to realistically represent microphysical processes. For example, most double-moment MP schemes predict the third ( $q_x$ ) and zeroth ( $N_{0x}$ ) moments of a PSD to diagnose  $\lambda_x$  and  $N_{0x}$ , but assume  $\alpha_x$  is constant. Double-moment schemes are able to simulate qualitatively polarimetric signatures (Jung et al. 2012; Johnson et al. 2016; Putnam et al. 2017b), but the schemes often suffer from excessive size sorting (e.g., Milbrandt and Yau 2005b; Dawson et al. 2014; Johnson et al. 2016; Morrison et al. 2015). Triple-moment MP schemes, which additionally predict the sixth moment of the PSD ( $Z_x$ ) can diagnose  $\alpha_x$ . The shape parameter narrows the hydrometeor size spectra and limits size sorting by having the weighted fall speeds of a hydrometeor type converge toward a singular value (Milbrandt and Yau 2005a; Dawson et al. 2014).

Convective-scale forecasts have been shown to predict spatial extent and size of hail at the surface with some skill. Multiple studies (e.g., S16; L17; Luo et al. 2017, 2018) have used predicted hail quantities (mass, number concentration) to identify regions where hail reaches the surface. Simulated radar products such as the maximum estimated size of hail (MESH; Witt et al. 1998a) have also been used to derive surface hail size from model output (S16; L17; Luo et al. 2017, 2018). Another method, the Thompson hail method (Thompson et al. 2018), uses model diagnosed hail PSDs to approximate the largest "observable" hailstone. Variants of this method have been used to diagnose the maximum surface hail size and distinguish regions of hail growth (Milbrandt and Yau 2006; S16; L17; Luo et al. 2018). The skill of hail forecasts derived directly from model output is strongly dependent upon choice of MP scheme; to account for model biases, machine learning algorithms can be applied to CAM output to create storm-based hail forecasts (Gagne et al. 2015, 2017; McGovern et al. 2017).

Convective-scale explicit hail forecasts are subject to significant error associated with the NWP model (e.g., initial conditions, microphysics). Many recent convective-scale studies use cycled ensemble Kalman filter (EnKF; Evensen 1994, 2003) data assimilation (DA) to generate initial conditions (e.g., Dowell et al. 2004; Snook et al. 2011, 2012, 2015; Dawson et al. 2012; Jung et al. 2012; Putnam et al. 2014, 2017a; Yussouf et al. 2013, 2016; Wheatley et al. 2014; Schwartz et al. 2015; Skinner et al. 2018). This technique is preferred because flow-dependent error covariances derived from the forecast ensemble allow the DA system to update unobserved

variables such as temperature, pressure, and microphysical variables from available observations (e.g., Tong and Xue 2005, 2008a). An EnKF DA system can thus use radar observations, which indirectly observe hydrometeor information, to improve the microphysical state variables of hail producing storms (S16; L17).

Hail forecast verification is a substantial challenge. Hail events are underreported in rural areas, away from major highways, as well as in cases when a more extreme severe weather event (e.g., one or more tornadoes) occurs in addition to hail (Doswell et al. 2005; Witt et al. 1998b; Allen and Tippett 2015). Additionally, hail sizes that correspond to familiar circular or spherical objects (e.g., dimes, softballs) are overreported (Sammler 1993), as such objects are commonly used as size references in reports from the public. New observational databases attempt to mitigate observed hail size biases. The meteorological phenomena identification near the ground (mPING; Elmore et al. 2014) phone application requires users to report hail in size increments of 0.25 in., and the severe hail analysis and verification experiment (SHAVE; Ortega et al. 2009) directly interviews the public to create a hail observation database with high spatial resolution. Nevertheless, surface-based reports remain inadequate for objective hail forecast verification.

The next-generation radar (NEXRAD; Crum et al. 1993) system is the only observational platform that performs full-volume scans of the atmosphere at a temporal frequency (approximately 5 min) sufficient to capture the rapid evolution of convective storms. Radar-derived hail products, such as MESH, serve as a proxy for hail size and can be used for forecast verification (e.g., S16). Although MESH varies in skill throughout the United States, it is not subject to population bias and thus superior to ground-based reports (Cintineo et al. 2012).

Polarimetric radars provide observations such as differential reflectivity  $Z_{dr}$ , copolar correlation coefficient  $\rho_{hv}$ , and differential phase  $\Phi_{dp}$ , in addition to  $Z$  and radial velocity  $V_r$ . With these additional observations, one is able to infer hydrometeor properties such as shape, size, orientation, and phase (Kumjian and Ryzhkov 2008). Hydrometeor classification algorithms (HCAs; Park et al. 2009) apply a fuzzy logic algorithm to both single- and dual-polarization radar data to classify the dominant hydrometeor type. This study implements the Putnam et al. (2017b) HCA, along with hail size discrimination algorithm membership functions from Ortega et al. (2016) to classify the dominant hail size into one of three bins: nonsevere (5 mm), severe (25 mm), and significant severe (50 mm). HCAs often classify hail size with more skill than MESH (Ortega et al. 2016),

although biased observations degrade the skill of the classifications.

Previous studies (i.e., S16; L17; Luo et al. 2017, 2018) evaluated hail forecasts produced using single-moment, double-moment, and triple-moment MP schemes; however, no studies have evaluated hail forecasts produced using a variable density MP scheme, which has the ability to represent a spectrum of rimed ice particle characteristics. In this study double-moment, triple-moment, and variable-density double-moment MP schemes are used to produce explicit ensemble hail forecasts for a severe hailstorm event that occurred in the Oklahoma City, Oklahoma, metropolitan area. Surface hail size forecasts are subjectively and objectively verified against HCA output (Ortega et al. 2016) using the Brier skill score (Brier 1950), reliability diagrams, and relative operating characteristic curves. In addition to forecast evaluations, a microphysical budget analysis is conducted. This study analyzes hail growth and decay processes predicted by each MP scheme to identify sources of forecast error, and determine the impact of microphysical processes on the prediction of hail.

The rest of the paper is organized as follows. In section 2 we provide a brief overview of the 19 May 2013 Oklahoma City severe hail event and describe the experimental design, including the DA and forecast model settings. Hail forecast verification statistics and an in-depth analysis of hail growth and decay processes are presented and discussed in section 3. A summary and some further discussion on the results are given in section 4.

## 2. Case, experiment configuration, and verification procedures

### a. Case overview

On 19 May 2013 an upper-level low pressure system developed in the northern plains of the United States as a negatively tilted trough moved over the southern plains. An 1800 UTC (1300 LST) sounding launched from the Norman Weather Forecast Office (WFO) indicated the environment was both highly unstable ( $CAPE \sim 4878 \text{ J kg}^{-1}$ ) and strongly sheared [0–6-km shear  $\sim 49 \text{ kt}$  ( $\sim 25 \text{ m s}^{-1}$ )], ideal for the development of supercell thunderstorms. Thunderstorms initiated along a dryline boundary west of Oklahoma City starting at approximately 1945 UTC. Three supercell thunderstorms (Fig. 1a) produced large regions of severe hail, and localized regions of significant-severe hail, in the southwestern Oklahoma City metropolitan area (Figs. 1b,c). Aside from producing large hail throughout

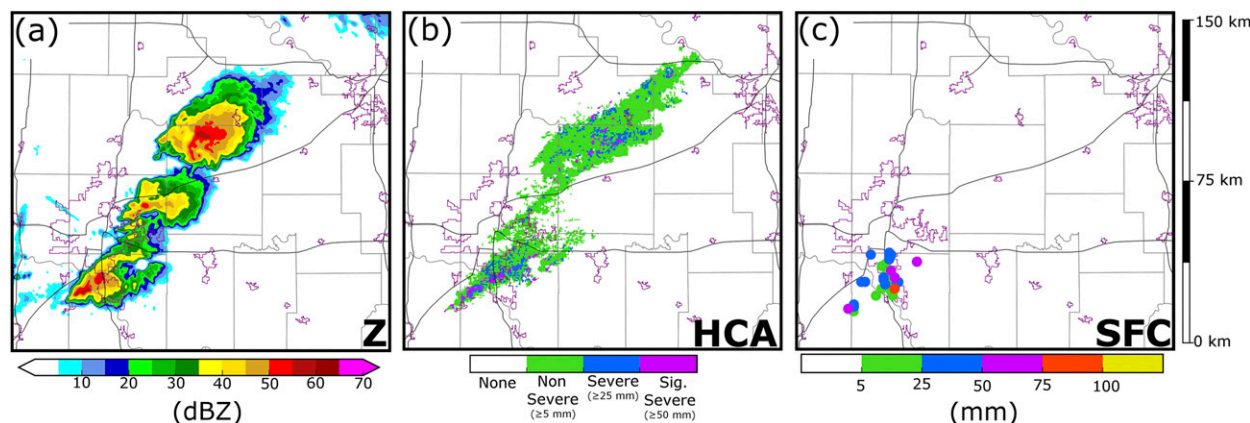


FIG. 1. (a) KTLX observed Z at the lowest radar tilt ( $0.5^\circ$ ), valid for 2220 UTC. (b) A swath of the largest observed hail size between 2155 and 2320 UTC. Hail size is determined by applying an HCA to the lowest radar tilt of KTLX and KINX ( $0.5^\circ$ ) and merging the subsequent output. (c) Surface-based hail size reports (SPC and mPING) between 2155 and 2320 UTC. On the background maps thin black lines are highways and purple lines are urban boundaries.

the region, these storms also produced two long-track tornadoes that produced EF-3 damage [on the enhanced Fujita (EF) scale] in Carney, Oklahoma, and EF-4 damage in Shawnee, Oklahoma. For further discussion on this event we refer the reader to Carlin et al. (2017) and Wienhoff et al. (2018).

#### b. Prediction model configurations

All experiments use the Advanced Regional Prediction System (ARPS; Xue et al. 2000, 2001) as the prediction model. The model domain consists of  $573 \times 515 \times 53$  grid points with a 500-m horizontal grid spacing. The model grid is stretched in the vertical, with a minimum vertical grid spacing of 50 m at the surface and an average vertical grid spacing of 425 m. The ARPS model settings follow those of S16 and L17. Model physics (Xue et al. 2000, 2001) include NASA Goddard Flight Center shortwave and longwave radiation, surface fluxes calculated from surface drag coefficients, surface temperature and volumetric water, a two-layer soil model, and a 1.5-order turbulent kinetic energy-based subgrid-scale turbulence parameterization (Deardorff 1980; Klemp and Wilhelmson 1978; Moeng 1984).

Hail forecasts are produced using three different MP schemes: the Milbrandt and Yau (2005a) double-moment MP scheme (MY2), the Milbrandt and Yau (2005a) triple-moment MP scheme (MY3), and the National Severe Storms Laboratory (NSSL) double-moment variable-density-rimed ice MP scheme (Mansell et al. 2010). Forecast maximum surface hail size ( $D_{\max}$ ) is calculated using a variant of the Thompson hail algorithm introduced in S16.  $D_{\max}$  is calculated at the first model level above the surface ( $\sim 25$  m AGL) and is

defined as the largest diameter for which the hail PSD has at least 1 stone per  $10\,000\text{ m}^4$  (Milbrandt and Yau 2006).

The MY2 scheme is selected for this experiment because it has been used in multiple recent hail prediction studies (e.g., S16, L17, Luo et al. 2017, 2018). The MY2 and MY3 schemes use the same prognostic equations for hydrometeor mass mixing ratio and number concentration; however, MY3 also predicts hydrometeor Z; Z is used to diagnose the PSD shape parameter. This parameter narrows the size distribution and limits further size sorting by causing weighted fall speeds to converge toward a singular value (Dawson et al. 2014). The NSSL scheme includes mass mixing ratio and number concentration, as well as prognostic equations for both hail and graupel volume (and thus density). Variable density is used to update particle fall speeds and conversion rates between hydrometeor species (e.g., dense wet-growth graupel is converted to hail; Mansell et al. 2010). The minimum number concentration threshold is set to  $10^{-8}\text{ m}^{-3}$  for the MY2 and MY3 schemes; the default setting ( $10^{-3}\text{ m}^{-3}$ ) led to excessive removal of hail, including some large hail near the surface. Unlike the MY2 and MY3 schemes, which threshold via number concentration and mass, the NSSL scheme only thresholds via mixing ratio, and thus does not remove as many hail grid points near the surface.

Vertical cross sections taken through the hail core of a thunderstorm, after a 60-min forecast, are used to demonstrate the impact of reducing the minimum number threshold. The default MY2 scheme (Figs. 2a,c,e) removes near-surface hail that is relatively sparse in both mass ( $< 5 \times 10^{-3}\text{ g m}^{-3}$ ) and number ( $< 5 \times 10^{-5}\text{ m}^{-3}$ ); however, the updated scheme (Figs. 2b,d,f)



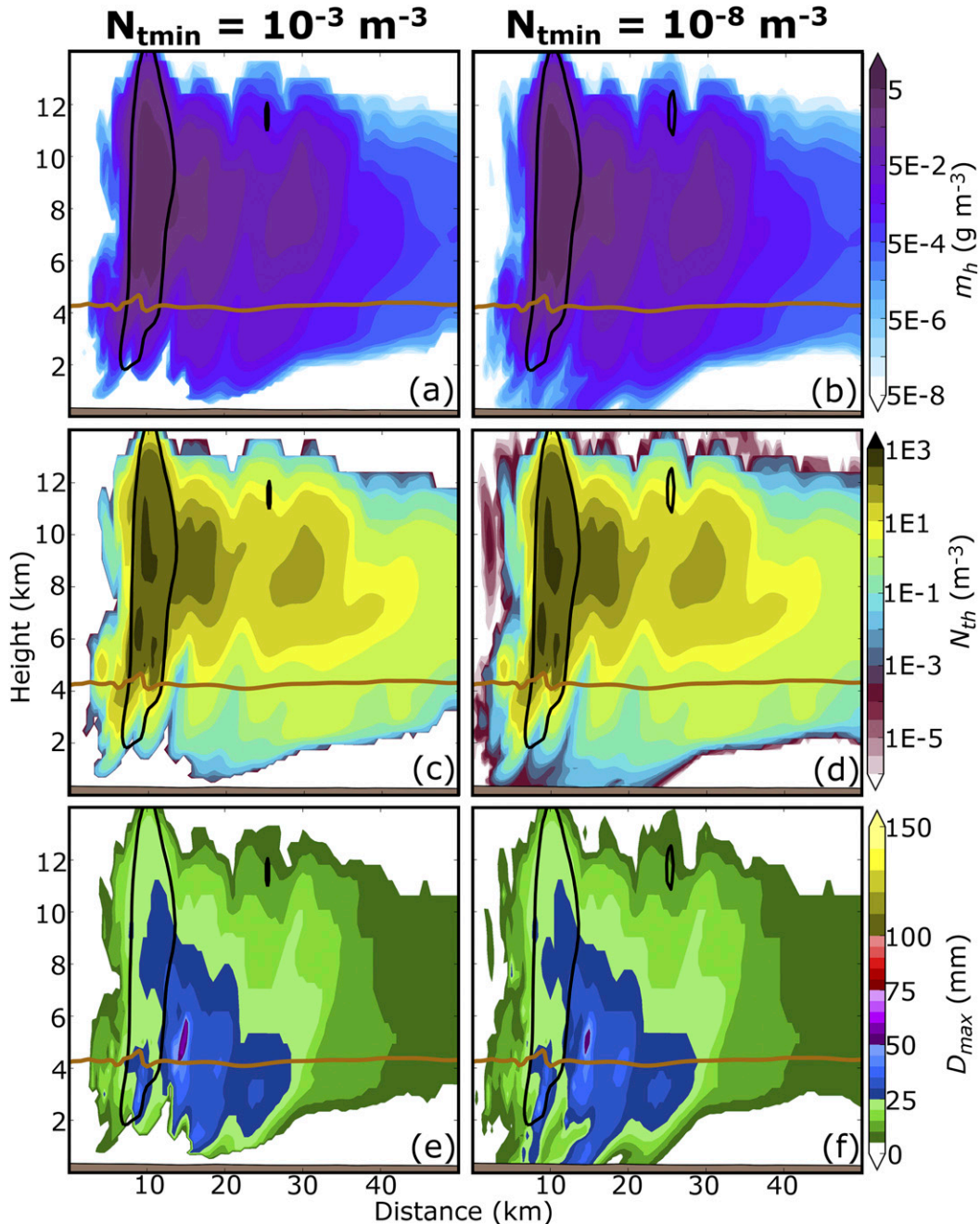


FIG. 2. Vertical cross sections taken through the hail core of a supercell thunderstorm after a 60-min forecast. Forecasts are run using the MY2 scheme where the minimum number concentration is set to (a),(c),(e)  $10^{-3} \text{ m}^{-3}$  and (b),(d),(f)  $10^{-8} \text{ m}^{-3}$ . Cross sections are of (a),(b)  $m_h$ , (c),(d)  $N_{th}$ , and (e),(f)  $D_{max}$ . In each plot a brown horizontal line represents the  $0^\circ\text{C}$  isotherm and black contours indicate regions where updraft velocity exceeds  $5 \text{ m s}^{-1}$ .

predicts more hail to reach the surface that is classified as observable by the  $D_{max}$  algorithm (Fig. 2f). Decreasing the minimum number threshold in the MY2 and MY3 schemes substantially increases surface hail size and coverage (Figs. 3c,d) compared to the default scheme (Figs. 3a,b) because the lower

threshold limits the removal of near-surface hail. Altering the minimum number threshold affects mostly near-surface hail size forecasts (Fig. 3), while model state variables (e.g.,  $w$ ,  $q_h$ ,  $N_{th}$ ) remain relatively similar between the forecasts (Figs. 2a–d) in the rest of the storm. Reducing the minimum number concentration

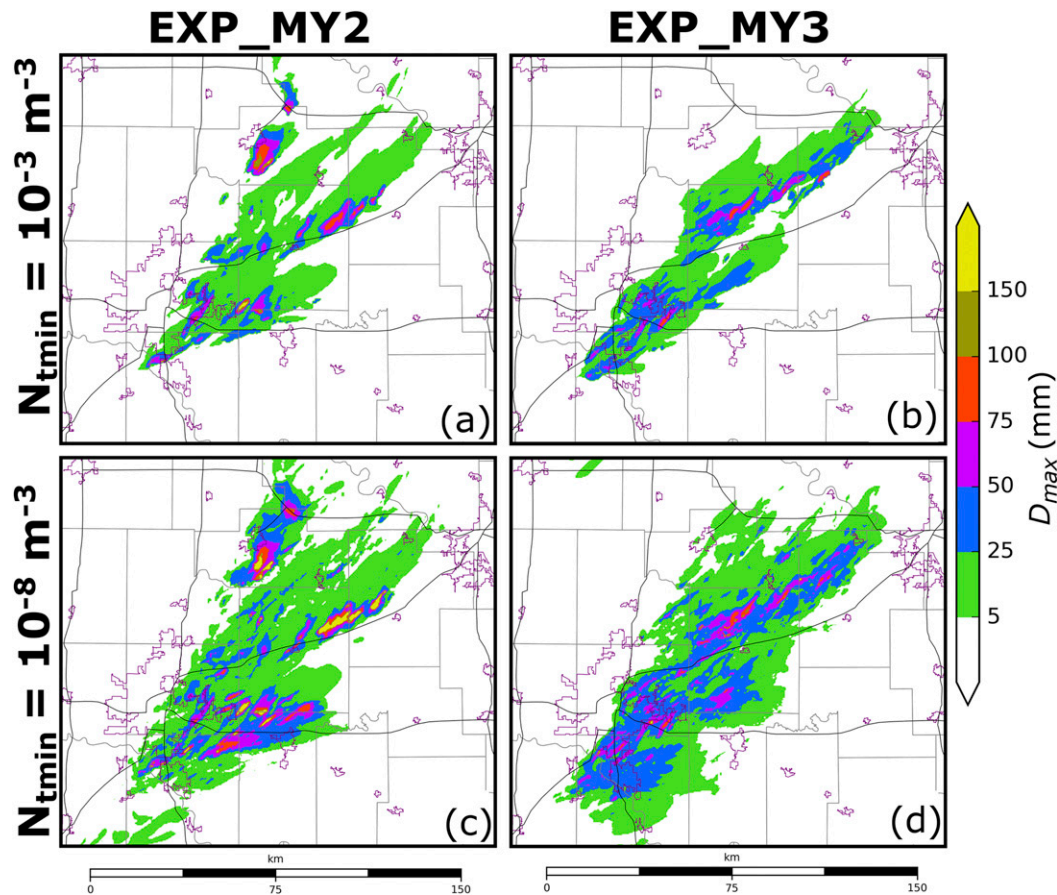


FIG. 3. Swaths of forecast maximum surface  $D_{\text{max}}$  calculated at every model time step between 2155 and 2320 UTC. Forecasts are run using the (a),(c) MY2 and (b),(d) MY3 schemes, where the minimum number concentration threshold is set to (a),(b)  $10^{-3} \text{ m}^{-3}$  or (c),(d)  $10^{-8} \text{ m}^{-3}$ . Background maps are the same as in Fig. 1.

threshold does not increase the number of spurious storms for this study.

### c. Radar observation operators and EnKF DA configurations

The EnKF DA requires accurate and efficient radar reflectivity observational operators consistent with the MP schemes used. The Rayleigh scattering approximation (Jung et al. 2008a,b) is typically used to simulate the  $Z$  of ice hydrometeors in our prior EnKF studies (Dawson et al. 2010; Snook et al. 2011, 2015, S16; Jung et al. 2012; Putnam et al. 2014; L17). While this observation operator is computationally efficient, it frequently overestimates the  $Z$  of large ice particles that cause Mie scattering (e.g., hail). The T-matrix radar simulator/observation operator (Jung et al. 2010) accounts for Mie scattering; it uses the T-matrix calculations (Vivekanandan et al. 1991; Bringi and Chandrasekar 2001) to derive the scattering amplitude of hydrometeors for different size bins and  $Z$  is then calculated by

integrating the scattering amplitude over a given PSD. Recent updates have optimized the computational efficiency of the T-matrix method; for example, Putnam et al. (2018, manuscript submitted to *Mon. Wea. Rev.*) calculates scattering amplitude for select PSD parameters and uses lookup tables and interpolation to derive  $Z$ . This study uses the Putnam et al. (2018, manuscript submitted to *Mon. Wea. Rev.*) T-matrix-based operator, which accounts for Mie scattering at computational speeds comparable to the Rayleigh scattering approximation.

Adjustments are made to the radar forward observation operator to accommodate the different MP schemes. To reduce the size of scattering amplitude lookup tables, the Rayleigh scattering approximation is assumed for the NSSL snow category, since snow particles are small enough that we can assume Rayleigh scattering with little to no error. To simulate the melting layer, a variant of the Jung et al. (2008a) melting model is used to create ice–water mixtures for both graupel and

hail. The ice–water hybrid category (which is generated offline) is determined via a ratio between ice and rain–water mixing ratios. For example, the hail–water mixing ratio ( $f_{wh}$ ) is:

$$f_{wh} = \frac{q_r}{q_r + q_h}. \quad (2)$$

While computationally efficient, this calculation does not account for shedding and thus allows too much liquid to accumulate on the surface of an ice particle. In regions where  $q_r$  is large, this can cause the diameter of melting ice to overinflate. To maintain computational efficiency but to produce more realistic  $Z$  we modify the ice–water mixture number concentration so that the mean mass diameter of the melting ice species is preserved during melting.

A 40-member ensemble is used for both the EnKF DA and forecast periods during this experiment, and one ensemble experiment is run using each of the microphysical schemes listed above (MY2, MY3, NSSL). Hereafter, the ensembles run using the MY2, MY3, and NSSL schemes are referred to as EXP\_MY2, EXP\_MY3, and EXP\_NSSL, respectively. Initial and boundary conditions are provided by the Center for Analysis and Prediction of Storms (CAPS) EnKF-initialized storm-scale ensemble forecast (SSEF; Jung et al. 2014). The EnKF-initialized SSEF is a 40-member ensemble of convection allowing (4-km horizontal grid spacing) forecasts that were run using the Advanced Research version of the Weather Research and Forecasting (WRF) Model (ARW; Skamarock et al. 2008) for the NOAA Hazardous Weather Testbed Spring Forecasting Experiment. The SSEF model domain spans the contiguous United States, forecasts are initialized at 1800 UTC 19 May 2013 and are run until 0000 UTC 20 May 2013. Additional information about the SSEF is provided in Kong (2013).

At 1900 UTC, our 500-m domain ensemble is initialized by interpolating from the SSEF. To introduce additional storm-scale perturbations, a 2-dimensional recursive filter is used to create smoothed Gaussian perturbations with horizontal and vertical correlation scales of 6 km and 3 km, respectively, which are added to the interpolated 1900 UTC SSEF ensemble. The storm-scale perturbations have standard deviations of  $2 \text{ m s}^{-1}$ , 2 K, and  $0.5 \text{ g kg}^{-1}$  for horizontal wind components ( $u$ ,  $v$ ), potential temperature  $\theta$ , and water vapor mixing ratio  $q_v$ , respectively. Sixty-minute spin-up ensemble forecasts are launched from the perturbed initial conditions and run until 2000 UTC when EnKF DA is first performed. The CAPS EnKF system (Xue et al. 2006; Tong and Xue 2008b), which is based upon the

TABLE 1. A list of the assimilated observations along with the corresponding model state variables that are updated by each observation.

Observation	Updated state variables
Surface	$u$ , $v$ , $\theta$ , $q_v$
Profiler	$u$ , $v$
Sounding	$u$ , $v$ , $\theta$ , $q_v$
KTLX $Z$	$u$ , $v$ , $w$ , $\theta$ , $q_v$ , $q_x$
KTLX $V_r$	$u$ , $v$ , $w$ , $q_v$

Whitaker and Hamill (2002) ensemble square root filter (EnSRF) algorithm, is used to assimilate observations every 10 min from 2000 to 2150 UTC, the time of final EnKF analysis. This configuration resulted in the improved suppression of spurious storms.

Assimilated conventional observations include surface (Oklahoma Mesonet, ASOS, and AWOS) and profiler observations. Conventional observations are not used to update vertical velocity ( $w$ ) because in preliminary tests, unreliable covariances led to the development of spurious updrafts. A list of the model state variables updated by each observation type is provided in Table 1. Observation error settings largely follow S16: surface observation errors are assumed to be  $1.5 \text{ m s}^{-1}$  for  $u$  and  $v$ , 2.0 K for  $\theta$ , and 2.0 K for dewpoint temperature ( $T_d$ ). The surface observation covariance localization radius is set to 300 km in the horizontal and 6 km in the vertical. Wind profiler observations are used to update background  $u$  and  $v$  winds and assume an error of  $2.5 \text{ m s}^{-1}$  for both variables and a covariance localization radius of 80 km in the horizontal and 6 km in the vertical.

Radar observations can infer important microphysical information and are assimilated. Data from the nearest Oklahoma City NEXRAD radar (KTLX) are interpolated horizontally to the model grid column locations but are preserved in the vertical at the height of the radar beam (Xue et al. 2006);  $Z$  and  $V_r$  observations are also thinned in the horizontal to one observation every 1 km in regions of precipitation (observed  $Z > 5 \text{ dBZ}$ ) and to one observation every 2 km in regions of clean air (observed  $Z < 5 \text{ dBZ}$ ). Radar data are thinned horizontally to control the ensemble spread reduction due to assimilation of very dense observations. KTLX observation errors are assumed to be  $4.0 \text{ m s}^{-1}$  for  $V_r$  and 6.0 dBZ for  $Z$  (Snook et al. 2013), and the covariance localization function (Gaspari and Cohn 1999) cutoff radius is set to 3 km in both the horizontal and vertical.

As multicategory microphysical variables are underconstrained by observed  $Z$ , large uncertainties may exist within analyzed variables. Uncertainties may increase for multimoment schemes because of the larger number

of predicted variables (Xue et al. 2010); for example, the MY2, MY3, and NSSL schemes predict 12, 17, and 15 state variables, respectively. Multimoment schemes better represent the nonmonotonic relationship between  $q_x$  and  $Z$ , which improves the representation of simulated radar variables (e.g., Putnam et al. 2014), and motivates the use of multimoment schemes in an EnKF during radar DA (Jung et al. 2012). To limit the influence of unreliable covariances we update only hydrometeor mixing ratio during DA (Table 1). Updating only hydrometeor mixing ratios may lead to the slower filter convergence rate but we found the system produces fewer spurious storms than updating more variables.

The relaxation-to-prior-spread (Whitaker and Hamill 2012) method is used to inflate the posterior ensemble spread to 95% of the prior ensemble spread for all updated variables. Although  $N_{tx}$ ,  $v_x$ , and  $Z_x$  are not updated (and thus never inflated), the variables are modified to avoid unrealistic PSD behavior during inflation. During inflation,  $N_{tx}$  is updated to preserve hydrometeor mean mass diameter,  $v_x$  is updated to preserve particle density, and  $Z_x$  is updated to preserve PSD shape parameter. Subjective comparisons between MY3 predicted  $Z$  after inflation and radar-observed  $Z$  (not shown) indicate that the diagnostic update can produce the posterior  $Z$  that compares well with observations.

#### d. Verification procedure

HCA output from the two nearest NEXRAD radars (KTLX and KINX) is used to verify hail size forecasts. Observations from KTLX and KINX are interpolated to the 500-m model grid and a 9-point smoother is applied to the  $Z_{dr}$  and  $\rho_{hv}$  fields to reduce noise. The HCA is performed for radar data on the lowest elevation angle within 120 km of a given radar—120 km is the maximum distance Ortega et al. (2016) used to evaluate hail size. Although the HCA is applied to the lowest radar tilt, the membership functions derived by Ortega et al. (2016) are tuned to detect surface hail size. Next, the data are merged (the largest hail size is selected where radar volumes overlap). Because of small-scale noise, a final smoothing filter is applied to the merged HCA output. The smoothing filter decreases (increases) hail size detections when the four closest grid points are smaller (larger); when a detection is updated, it is set to match the four surrounding grid points. Merged HCA output (Fig. 1b) is used for both subjective and objective forecast verification in this study.

Hail forecast verification is challenging in part because large surface hail is a highly localized phenomenon. When using a traditional neighborhood ensemble

probability (NEP; Schwartz et al. 2010) approach, small position errors between ensemble members can lead to mostly nonzero probabilities that are difficult to verify. Neighborhood maximum ensemble probability (NMEP; Ben Bouallègue and Theis 2014; Schwartz and Sobash 2017) considers the probability of an event occurrence within distance  $i$ ; this expands both the number of probabilistic and observed events and increases the sharpness of the forecast. For objective verification the NMEP method defined in Schwartz and Sobash (2017) is used to verify the probability of an event occurrence within 10 km of a grid point with a smoothing Gaussian filter of 10 km. This distance was selected because it accounts for small forecast storm displacement errors, but maintains the high spatial resolution necessary for warn-on-forecast applications.

Probabilistic hail size forecasts are evaluated using the Brier skill score (Brier 1950), reliability diagrams, and relative operating characteristic (ROC; Mason 1982) curves. The Brier skill score is defined by three separate components: forecast reliability, resolution, and uncertainty. Forecast reliability compares predicted probability of occurrence to observed frequency, which is used to infer model bias and is most commonly displayed in terms of a reliability diagram. In an unbiased system, observed frequency and predicted probability are similar; if there is a mismatch, the model suffers an overprediction (underprediction) bias when the predicted probability of occurrence is larger (smaller) than the observed frequency. ROC curves determine the model's ability to discern between events and nonevents. ROC curves are generated by plotting probability of detection (POD) against probability of false detection (POFD) for increasing probability thresholds. Forecast skill is usually determined by the area under the ROC curve (AUC), if the AUC exceeds 0.7 then the forecast is skilled at predicting the event (Buizza et al. 1999).

### 3. Results

#### a. Forecast evaluation and verification

The 90-min-long hail size forecasts starting from the final EnKF analyses at 2150 UTC are verified between 2155 and 2320 UTC, a period during which multiple supercell thunderstorms produce severe and significant severe hail over the Oklahoma City metropolitan area (Figs. 1b,c). Accumulated swaths of the forecast maximum  $D_{max}$  (calculated at every model time step) are verified against accumulated swaths of observed maximum hail size as indicated by HCA output (calculated every radar volume scan). We note that radars observe the atmosphere far less frequently ( $\sim 5$  min) than



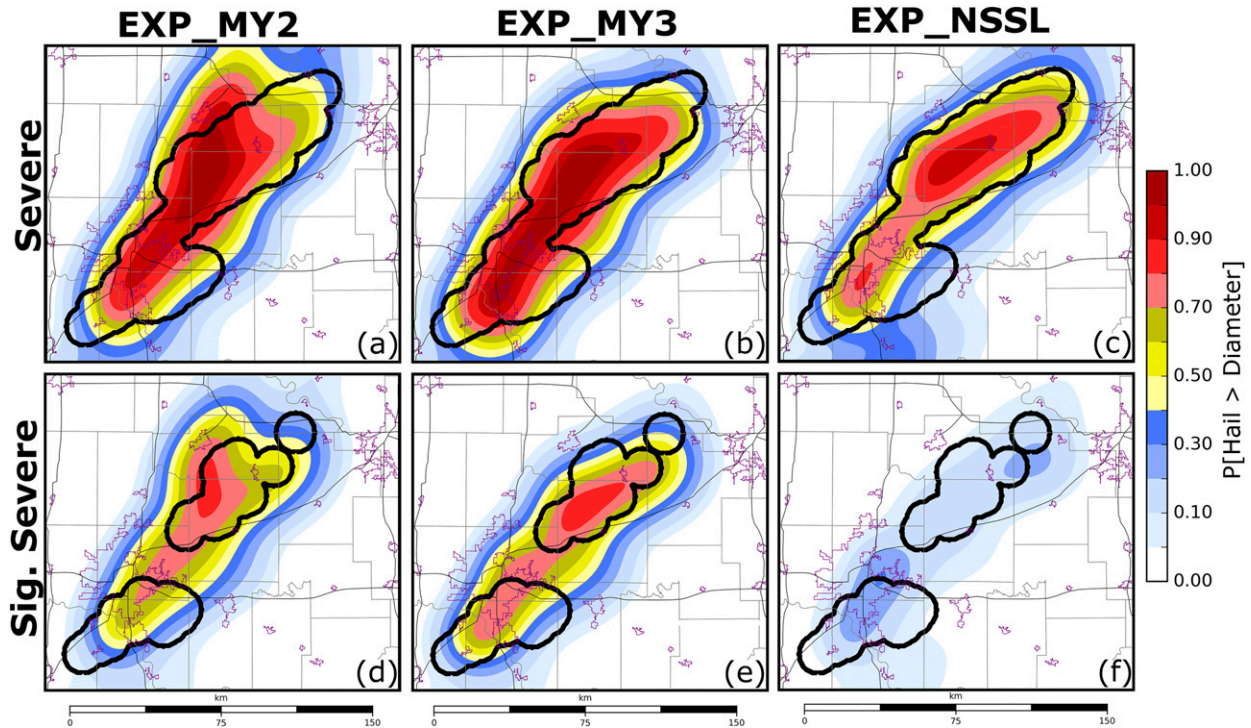


FIG. 4. The NMEP of (a)–(c) severe hail and (d)–(f) significant severe hail predicted using the (a),(d) MY2, (b),(e) MY3, and (c),(f) NSSL MP schemes between 2155 and 2320 UTC. Thick black contours represent the locations of observed severe or significant severe hail according to HCA output. Background maps are the same as in Fig. 1.

a typical storm-scale model time step ( $\sim 1$  s); however, we believe that the higher temporal resolution of the model output better captures the extent and evolution of hail in the model storms, and thus choose to perform verification with the highest temporal frequency data available. Hail size forecasts are verified for two separate diameter thresholds including severe hail (diameter  $> 25$  mm) and significant severe hail (diameter  $> 50$  mm).

For this study, we first subjectively evaluate the NMEP of severe and significant severe hail and note forecast biases. The predicted probability of severe hail ( $P[\text{severe hail}]$ ) (Figs. 4a–c) is generally high ( $> 0.7$ ) in regions where severe hail is observed. Both EXP\_MY3 (Fig. 4b) and EXP\_NSSL (Fig. 4c) predict the northernmost hail swath to be oriented toward the northeast, suggesting these ensembles predict the rightward deviation in storm motion commonly observed in supercell thunderstorms (Bunkers et al. 2000). Unlike the above two ensembles, EXP\_MY2 (Fig. 4a) predicts a bifurcation in maximum  $P[\text{severe hail}]$  values for the northernmost storm. This bifurcation is due to the production of spurious storms in EXP\_MY2 forecasts. EXP\_MY2 forecasts produce many spurious storms, this increases the overall coverage of severe hail and the total area where  $P[\text{severe hail}] > 0.4$  (Fig. 4a). Both EXP\_MY2

(Fig. 4d) and EXP\_MY3 (Fig. 4e) predict  $P[\text{significant severe hail}]$  to exceed 0.5 where HCA output (Fig. 1b) indicates the occurrence of significant severe hail.

Reliability diagrams indicate that ensembles predict the spatial coverage of severe hail (Fig. 5a) with relatively little bias. Although Brier skill scores (Table 2) vary slightly (0.532–0.633), all three ensembles produce skillful severe hail forecasts. For significant severe hail forecasts, the Brier skill score (Table 2) decreases for all three ensembles (0.132–0.463), indicating the ensembles predict more extreme events with less skill. EXP\_NSSL predicts significant severe hail with the least skill (Table 2), in part because the ensemble produces under confident significant severe hail size forecasts (Fig. 5b) and rarely predicts  $P[\text{significant severe}] > 0.2$  (Fig. 4f). Despite lower Brier skill scores (Table 2), EXP\_MY2 and EXP\_MY3 produce reliable significant severe hail forecasts (Fig. 5b) and predict significant severe hail with higher confidence than EXP\_NSSL.

Contrary to the other verification metrics, ROC curves (Fig. 6) indicate the ensembles predict both severe and significant severe hail with a high level of skill ( $\text{AUC} > 0.92$ ). Large variations in  $P[\text{significant severe}]$  between the ensembles does not appear to have a large impact on the AUC score. It is noted that verifications

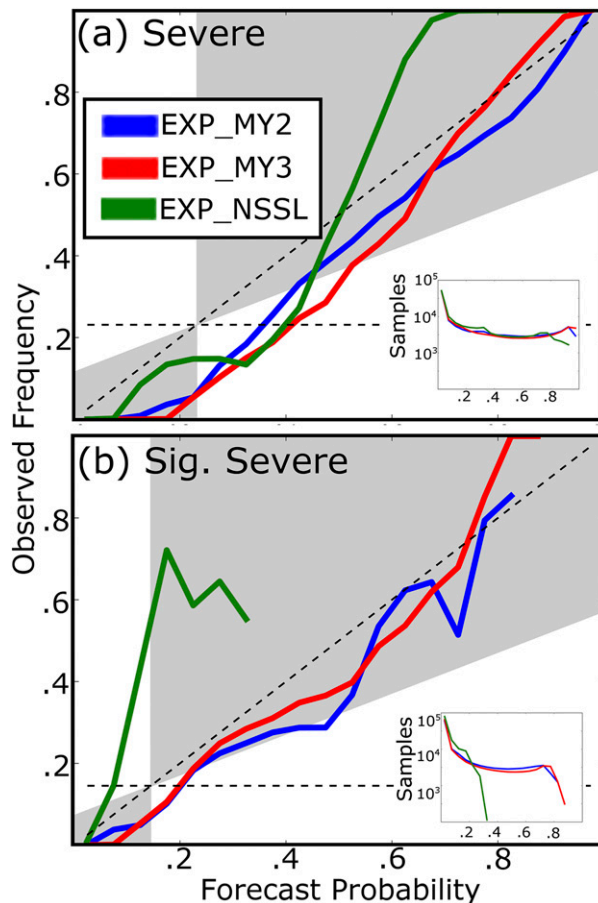


FIG. 5. Reliability diagrams for the probabilistic (a) severe hail forecasts and (b) significant severe hail forecasts shown in Fig. 4. In the diagrams blue lines correspond to EXP\_MY2, red lines to EXP\_MY3, and green lines to EXP\_NSSL. Forecasts exhibit skill when the forecast reliability falls within the gray shading. The number of samples per forecast probability bin is provided for both (a) severe hail and (b) significant severe hail.

are performed for a single event and additional case studies are needed; however, findings from this study are believed to be meaningful.

Swaths of  $D_{\max}$  produced by single ensemble members (Fig. 7) provide insight into hail predictions. EXP\_MY2 overpredicts the size of hail at the surface,

TABLE 2. The Brier skill score for EXP\_MY2, EXP\_MY3, and EXP\_NSSL. Skill scores are calculated for severe hail and significant severe hail forecasts between 2155 and 2320 UTC. The corresponding NMEP used to derive the Brier skill score is provided in Fig. 4.

	Severe	Significant severe
EXP_MY2	0.532	0.352
EXP_MY3	0.633	0.463
EXP_NSSL	0.629	0.132

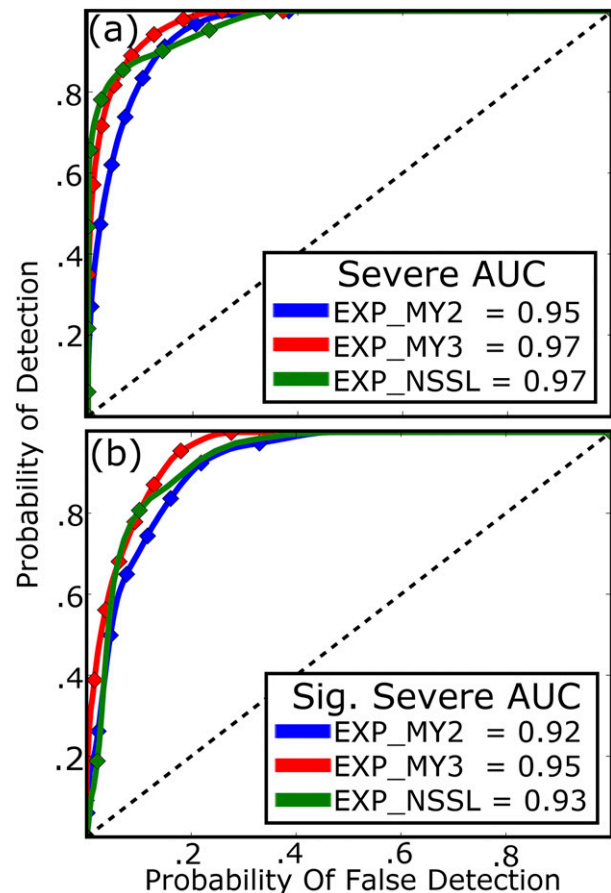


FIG. 6. ROC curves for the probabilistic (a) severe and (b) significant severe hail forecasts shown in Fig. 4. In the diagrams blue lines correspond to EXP\_MY2, red lines to EXP\_MY3, and green lines to EXP\_NSSL. Forecast AUC scores are provided, scores that are greater than 0.7 predict severe or significant severe hail with skill.

with both members 20 (Fig. 7a) and 40 (Fig. 7d) predicting hail to exceed 150 mm (~6 in.) in diameter. The HCA output (Fig. 1b) is unable to determine the upper limit on observed hail size for hail exceeding 50 mm because the algorithm classifies all hail 50 mm and larger as significant severe; however, mPING reports suggest that hail, which occurred during this event, did not exceed 3 in. (~75 mm) (Fig. 1c) in diameter. The MY2 scheme surface hail size overprediction bias is noted in several previous studies; both L17 and Luo et al. (2018) note the scheme produces excessively large surface hail. The MY3 scheme employs the same mass and number concentration tendency equations as the MY2 scheme; as such, EXP\_MY3 also overpredicts hail size (Figs. 7b,e). Despite an overprediction bias, EXP\_MY3 members (Figs. 7b,e) predict smaller  $D_{\max}$  values than EXP\_MY2 members (Figs. 7a,d). Forecast differences are in part because microphysical processes (e.g., sedimentation)

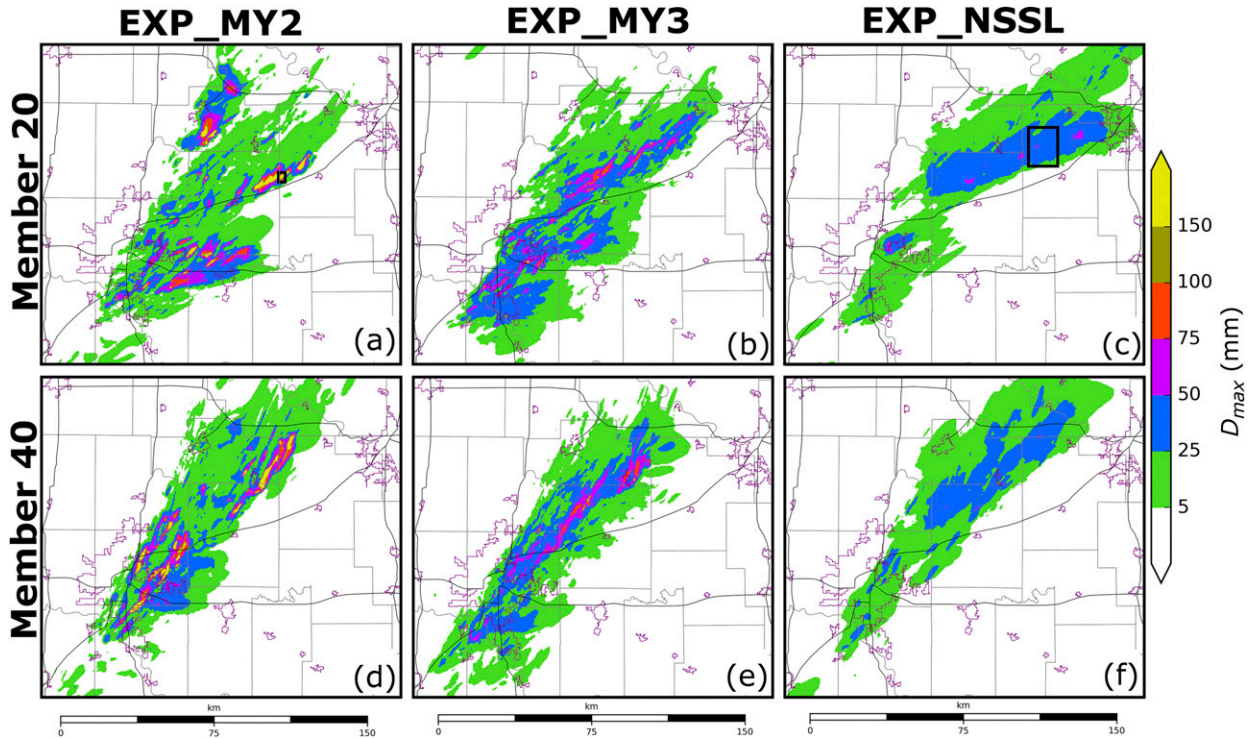


FIG. 7. Swaths of forecast maximum surface  $D_{max}$  calculated at every model time step between 2155 and 2320 UTC. Swaths are created for ensemble members (a)–(c) 20 and (d)–(f) 40, which are run using the (a),(d) MY2, (b),(e) MY3, and (c),(f) NSSL MP schemes. These members are representative of the other ensemble member forecasts. Black polygons in (a) and (c) highlight the locations where hail processes are analyzed later in the study. Background maps are the same as in Fig. 1.

are sensitive to variations in the shape parameter (Milbrandt and Yau 2005a).

EXP\_NSSL forecasts (Figs. 7c,f) qualitatively most closely match HCA output in terms of maximum hail size. EXP\_NSSL members predict the northernmost storm to produce more severe hail than what is observed in HCA output (Fig. 1b), though this is partly because we compare low temporal frequency ( $\sim 5$  min) radar observations to high temporal frequency ( $\sim 1$  s) NWP output. Additionally, the NSSL scheme predicts the storms to produce hail cores that are wider than observations.

### b. Hail production processes

Similar to the microphysical budget analysis conducted in L17, this study analyzes hail growth and decay process predicted by each MP scheme to identify sources of forecast error. Hail microphysical tendency terms (e.g., accretion, sublimation, melting) from a single time step are integrated along a horizontal plane that spans the verification domain (Fig. 4). Terms are plotted as a function of average model level height in order to illustrate the model's treatment of hail both above and below the  $0^{\circ}\text{C}$  isotherm. Vertical cross sections are taken

through the hail core of a supercell thunderstorm during the microphysical budget analysis to illustrate the impact of microphysical processes on the representation of hail. We conduct the microphysical analysis at 2250 UTC; this is one hour into the forecast period when multiple supercell thunderstorms are producing hail.

Hail growth assumptions in MP schemes impact the representation of hail within the model. In EXP\_MY2 (Fig. 8b) and EXP\_MY3 (Fig. 8d), hailstones are almost exclusively created by the three-component accretion of rain; this was first reported in Johnson et al. (2016). Three-component accretion of rain occurs when rainwater freezes to the surface of a frozen hydrometeor species to produce a subsequent third ice hydrometeor type (e.g., rain freezing onto snow to create hail). A static threshold is used to determine the resulting hydrometeor category. The NSSL scheme hail category does not contain frozen raindrops (Mansell et al. 2010), instead hail is created from dense graupel ( $\rho_g > 800 \text{ kg m}^{-3}$ ) in wet growth conditions (Fig. 8f). EXP\_NSSL  $N_{th}$  tendencies (Fig. 8f) are approximately one to two orders of magnitude smaller than EXP\_MY2 (Fig. 8b) or EXP\_MY3 (Fig. 8d)  $N_{th}$



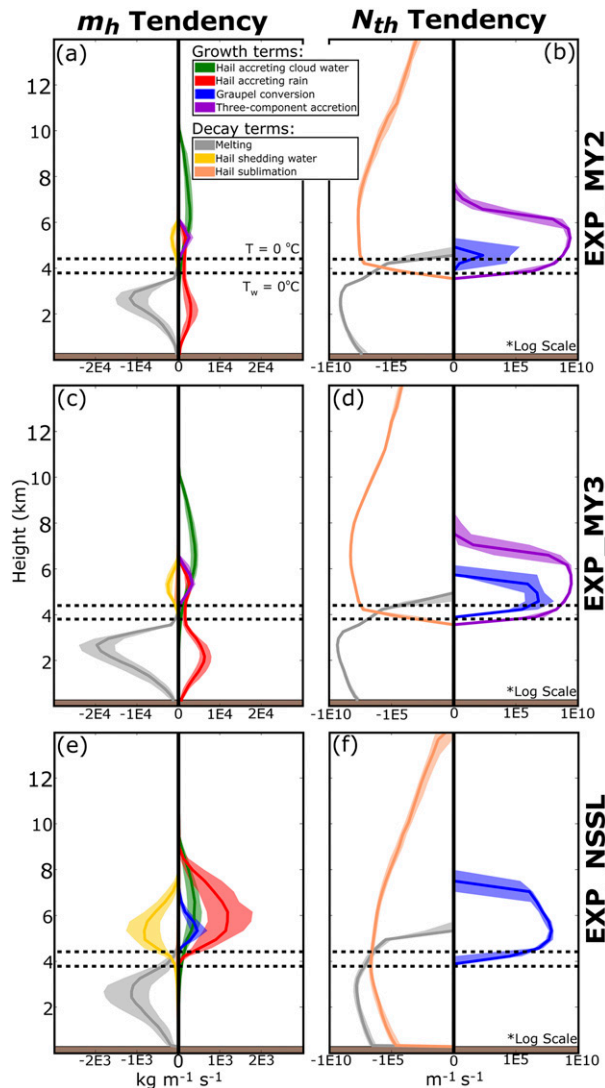


FIG. 8. Microphysical source and sink terms for (a),(c),(e)  $m_h$  and (b),(d),(f)  $N_{th}$  predicted by (a),(b) EXP\_MY2, (c),(d) EXP\_MY3, and (e),(f) EXP\_NSSL at 2250 UTC. Microphysical terms are integrated over each model vertical level spanning the verification domain shown in Fig. 4. The terms are plotted as a function of average model height per vertical level. Terms are summed for each ensemble member, the interquartile range of the ensemble is shaded and median value is denoted as a dark line. Only terms that contribute greater than  $10 \text{ kg m}^{-1}$  for  $m_h$  and  $10 \text{ m}^{-1}$  for  $N_{th}$  are plotted. In each plot the upper horizontal dashed black line is the average  $0^\circ\text{C}$  isotherm height and the lower horizontal dashed line is the average wet-bulb temperature  $0^\circ\text{C}$  isotherm height.

tendencies because dense, wet-growth graupel particles are much less numerous than frozen raindrops.

The three-component accretion of rain increases the number of hailstones above the  $0^\circ\text{C}$  isotherm in the EXP\_MY2 (Fig. 8b) and EXP\_MY3 (Fig. 8d) forecasts; however this process contributes relatively little to the total mass of hail aloft because frozen raindrops are

relatively small (Figs. 8a,c). The two largest  $m_h$  growth processes for the MY2 (Fig. 8a) and MY3 (Fig. 8c) schemes are the accretion of rain and cloud water. Similar to L17, the MY2 (Fig. 8a) and MY3 (Fig. 8c) schemes both predict hail to accrete rainwater near the surface despite the wet-bulb temperature exceeding  $0^\circ\text{C}$ . The consequences of this process are discussed later in section c of this chapter. Unlike MY2 and MY3, the NSSL scheme assumes all accreted water is shed in and beneath the melting layer for hail, eliminating the possibility of  $m_h$  growth. EXP\_NSSL  $m_h$  tendencies (Fig. 8e) are smaller in magnitude than either EXP\_MY2 (Fig. 8a) or EXP\_MY3 (Fig. 8c) because the rate of accretion and melting is proportional to hail number concentration. The NSSL scheme predicts fewer hailstones than the MY2 and MY3 schemes, and thus predicts less rain and cloud water droplets are accreted.

Vertical cross sections of hail-related microphysical variables (Fig. 9) are evaluated to understand the impact of microphysical processes on the behavior of hail. Although the EXP\_MY2 (Figs. 7a,d) and EXP\_MY3 (Figs. 7b,e) predict storms to produce  $D_{\max} > 100 \text{ mm}$  at the surface, cross sections indicate hail is much smaller above the  $0^\circ\text{C}$  isotherm (Figs. 9g,h). The NSSL scheme predicts storms to produce smaller  $D_{\max}$  values at the surface ( $\leq 75 \text{ mm}$ ) (Figs. 7c,f); however,  $D_{m_h}$  remains large throughout the vertical depth of the updraft (Fig. 9i). Differences in microphysical assumptions cause EXP\_MY2 and EXP\_MY3 to predict larger hail than EXP\_NSSL beneath the  $0^\circ\text{C}$  isotherm, but smaller hail aloft.

Both the MY2 (Fig. 9d) and MY3 (Fig. 9e) schemes predict  $N_{th}$  to be large above the  $0^\circ\text{C}$  isotherm at 5 km above mean sea level; this is the approximate level where the three-component accretion of rain is maximized (Figs. 8b,d). Since frozen raindrops are relatively small in diameter, the three-component accretion of rain decreases the  $D_{m_h}$  above the  $0^\circ\text{C}$  isotherm for both EXP\_MY2 (Fig. 9g) and EXP\_MY3 (Fig. 9h) members. The small hail particles aloft are advected downwind of the updraft, creating a plume of hail that extends more than 40 km. Although the MY2 and MY3 schemes produce large plumes of hail aloft, the small hail particles in the plumes mostly melt before reaching the surface. S16 previously noted the MY2 scheme melts hail too quickly and underpredicts the spatial coverage of hail at the surface. The underprediction bias noted in S16 is mitigated when a less stringent minimum  $N_{th}$  threshold of  $10^{-8} \text{ m}^{-3}$  is used, this allows hailstones that are more sparse in number concentration to reach the surface.

Above the  $0^\circ\text{C}$  isotherm, the NSSL scheme produces fewer but larger hailstones than either the MY2 or MY3



schemes because hail is created from wet-growth graupel. Large hail behaves differently from small rimed ice particles: first, upper-level winds do not advect the NSSL scheme hail as far downstream, and most large hail ( $D_{m_h} > 4.5$  mm) thus remains confined to the boundaries of the storm updraft region (Fig. 9i). Second, the larger hailstones have an increased terminal velocity, and thus higher fall speeds. An increased fall speed causes a higher percentage of hailstones to reach the surface before melting and produces relatively wide swaths of severe hail (Figs. 7c,f) that closely match observations (Figs. 1b,c).

### c. Hail melting processes

In this study, EXP\_MY2 forecasts (Figs. 7a,d) predict  $D_{\max}$  to frequently exceed 150 mm (approximately 6 in.), despite HCA output (Fig. 1b) indicating that surface hail is mostly nonsevere or severe (<50 mm in diameter). Much of the hail growth in EXP\_MY2 forecasts occurs in the melting layer; EXP\_MY2 member 20, which is representative of most EXP\_MY2 members, predicts  $D_{\max}$  to increase from 25 mm to more than 150 mm while traversing the 0°C isotherm (Fig. 10a). L17 suggests much of this growth is because hail in the MY2 scheme accretes rain and cloud water in and beneath the melting layer.

To better understand MY2 scheme hail growth, we integrate the total microphysical tendency of  $m_h$  and  $N_{\text{th}}$  over a subdomain that encompasses only the hail core of EXP\_MY2 member 20 (Fig. 7a). Microphysical tendencies vary substantially throughout a single storm, integrating the tendencies over the hail core (Figs. 10b,c) provides more insight into hail growth processes than integrating over the verification domain (Figs. 8a,b). Above the 0°C isotherm  $N_{\text{th}}$  increases rapidly due to the three-component accretion of rain (Fig. 10c). The introduction of many small, frozen raindrops decreases the mean size of hail aloft. Beneath the 0°C isotherm,  $m_h$  increases (Fig. 10b) because hail accretes more rain and cloud water than it melts and sheds. At approximately the same layer below the freezing level,  $N_{\text{th}}$  slowly decreases due to melting (Fig. 10c), this is because accretional growth does not increase  $N_{\text{th}}$ . When  $m_h$  increases but  $N_{\text{th}}$  decreases the slope parameter of the hail PSD  $\lambda_h$  becomes shallow (Fig. 11b) and causes  $D_{\max}$  to exceed 150 mm (Fig. 10a). Accretional growth maintains or expands hail size, allowing more hail to reach the surface.

EXP\_MY3 predicts  $D_{\max}$  (Figs. 7b,e) to exceed 150 mm less frequently than EXP\_MY2 (Figs. 7a,d), even though both MP schemes use the same  $m_h$  and  $N_{\text{th}}$  tendency equations. Hail PSDs selected from three grid points within a hail core of EXP\_MY3 member

5 (Fig. 11a) indicate that when  $\alpha_h > 0$  the PSD follows a gamma distribution which narrows the PSD and moderates  $D_{\max}$  (Fig. 11b). Multiple studies (e.g., Milbrandt and Yau 2005b; Kumjian and Ryzhkov 2012; Dawson et al. 2014) have documented that  $\alpha_h$  increases due to sedimentation. Although the MY3 scheme moderates  $D_{\max}$ , hail continues to accrete liquid in and beneath the melting layer, which causes EXP\_MY3 to predict hail greater than 100 mm in diameter (Figs. 7b,e).

EXP\_MY2 and EXP\_MY3 member 20 forecasts are rerun without the accretion of rain and cloud water beneath the 0°C isotherm (EXP\_MY2\_NOAC and EXP\_MY3\_NOAC, respectively) to demonstrate the sensitivity of surface hail size to the accretion of water beneath the 0°C isotherm. EXP\_MY2\_NOAC (Fig. 12a) and EXP\_MY3\_NOAC (Fig. 12b) predict the maximum surface hail size to be smaller in diameter than EXP\_MY2 (Fig. 7a) and EXP\_MY3 (Fig. 7b) forecasts, and more closely resemble observations (Figs. 1b,c). It is noted that EXP\_MY2\_NOAC (Fig. 12a) predicts hail to exceed 100 mm; however, most of this hail is attributed to initialization from the EXP\_MY2 member 20 analysis, and within 10 min of forecast initialization most of the storms cease to produce hail greater than 100 mm in diameter. Forecast experiments EXP\_MY2\_NOAC (Fig. 12a) and EXP\_MY3\_NOAC (Fig. 12b) are in agreement with the conclusions of this study and suggest the accretion of liquid water beneath the 0°C isotherm is important to the production of large hailstones ( $D_{\max} > 100$  mm).

Coverage of severe hail increases at the surface for EXP\_MY2\_NOAC (Fig. 12a) and EXP\_MY3\_NOAC (Fig. 12b) when compared to EXP\_MY2 (Fig. 7a) and EXP\_MY3 (Fig. 7b) forecasts. In all three MP schemes the  $N_{\text{th}}$  melting tendency term is used to preserve  $D_{m_h}$  during melting; however, both the MY2 and MY3 schemes reduce the  $N_{\text{th}}$  melting term by 90% when accretional growth occurs. Except in updrafts where large quantities of liquid water are accreted, when the  $N_{\text{th}}$  melting term is reduced  $D_{m_h}$  decreases because  $m_h$  decreases more quickly than  $N_{\text{th}}$ . When accretion of liquid water is eliminated beneath the 0°C isotherm the  $N_{\text{th}}$  melting tendency is not reduced, and thus EXP\_MY2\_NOAC and EXP\_MY3\_NOAC predict hail to more frequently exceed 25 mm in diameter (Fig. 12).

EXP\_NSSL exhibits limited hail growth across the 0°C isotherm. EXP\_NSSL member 20, for example, predicts a change of only 0–5 mm across the melting layer (Fig. 13a), compared to a change of 150 mm across the melting layer in EXP\_MY2 member 20 (Fig. 10a). Microphysical tendency terms in the hail core of EXP\_NSSL member 20 (Fig. 7c) suggest  $m_h$  (Fig. 13b) and  $N_{\text{th}}$  (Fig. 13c) decrease at approximately the same

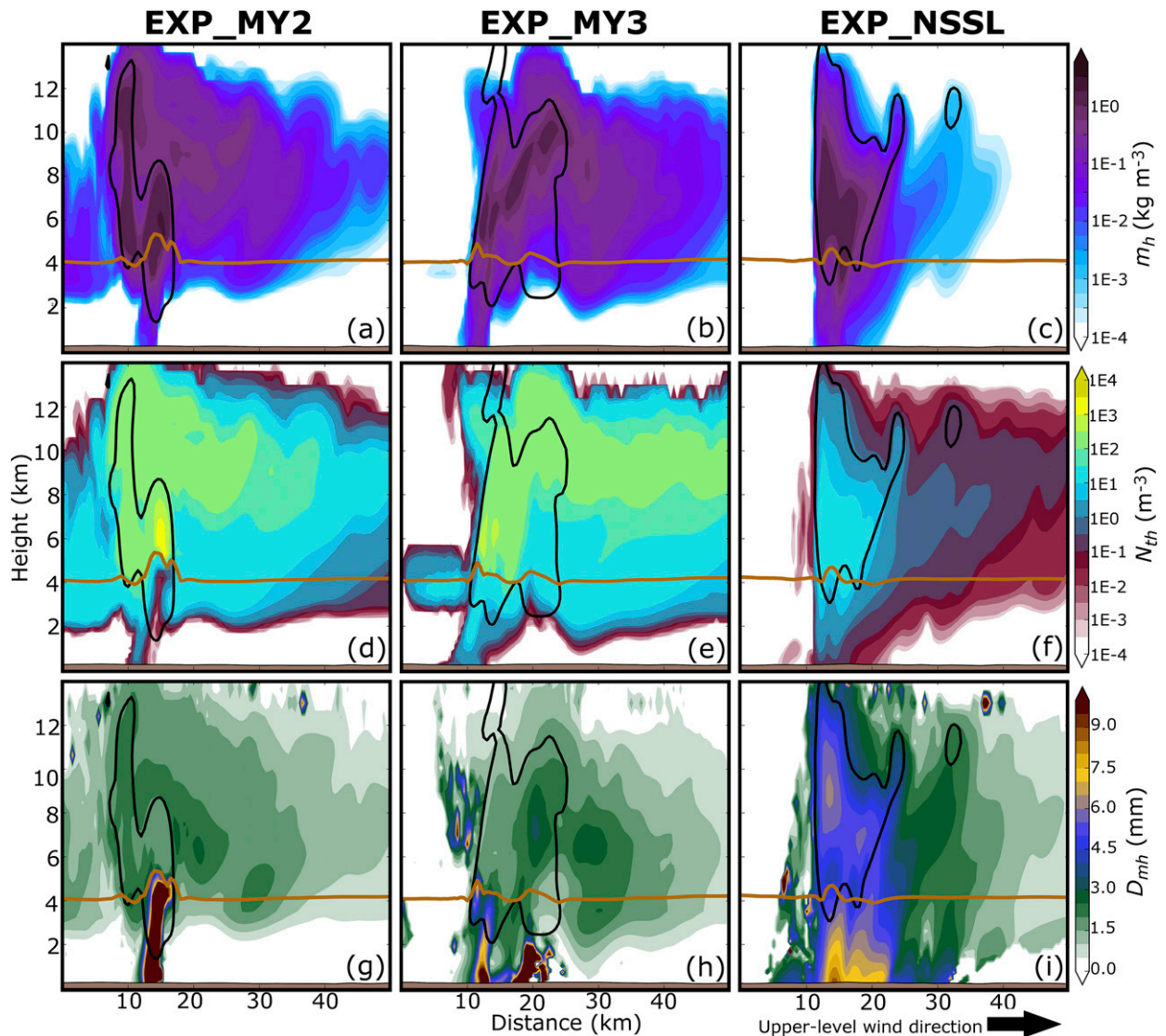


FIG. 9. A vertical cross section taken through the hail core of the northernmost supercell thunderstorm predicted by member 20 of (a),(d),(g) EXP\_MY2, (b),(e),(h) EXP\_MY3, and (c),(f),(i) EXP\_NSSL at 2250 UTC. Cross sections are of (a)–(c)  $m_h$ , (d)–(f)  $N_{th}$ , and (g)–(i)  $D_{mh}$ . In each plot a brown horizontal line represents the  $0^{\circ}\text{C}$  isotherm and black contours indicate regions where updraft velocity exceeds  $5\text{ m s}^{-1}$ .

rate beneath the  $0^{\circ}\text{C}$  and isotherm. Both terms decrease at the same rate because hail sheds all accreted liquid when the air temperature is warmer than  $0^{\circ}\text{C}$ . Although the NSSL scheme limits hail growth within the melting layer, EXP\_NSSL forecasts severe and significant severe hail (Figs. 4c,f) because large hailstones are created from wet-growth graupel above the  $0^{\circ}\text{C}$  isotherm.

Most MP schemes assume ice is dry; this assumption has important consequences on the treatment of hail beneath the  $0^{\circ}\text{C}$  isotherm. Collection of water beneath the  $0^{\circ}\text{C}$  isotherm will increase the ice water fraction and potentially hail size (Lesins and List 1986); however, if too much liquid water is collected than the hailstone

will shed and evaporate excess liquid (Rasmussen and Heymsfield 1987; Chong and Chen 1974). Instead of representing the collection of water, the MY2 and MY3 schemes freeze rain and cloud water onto the surface of hail (i.e., accretion). Since the accreted water is frozen, excess liquid cannot be removed via shedding, and instead  $m_h$  increases when the air temperature is well above freezing (Fig. 10b). In contrast, the NSSL scheme does not allow any liquid water collection beneath the  $0^{\circ}\text{C}$  isotherm (Fig. 13b); this limits a potential hail growth process within the scheme and potentially limits the maximum surface hail size. Without the ability to represent ice–water hybrid categories these

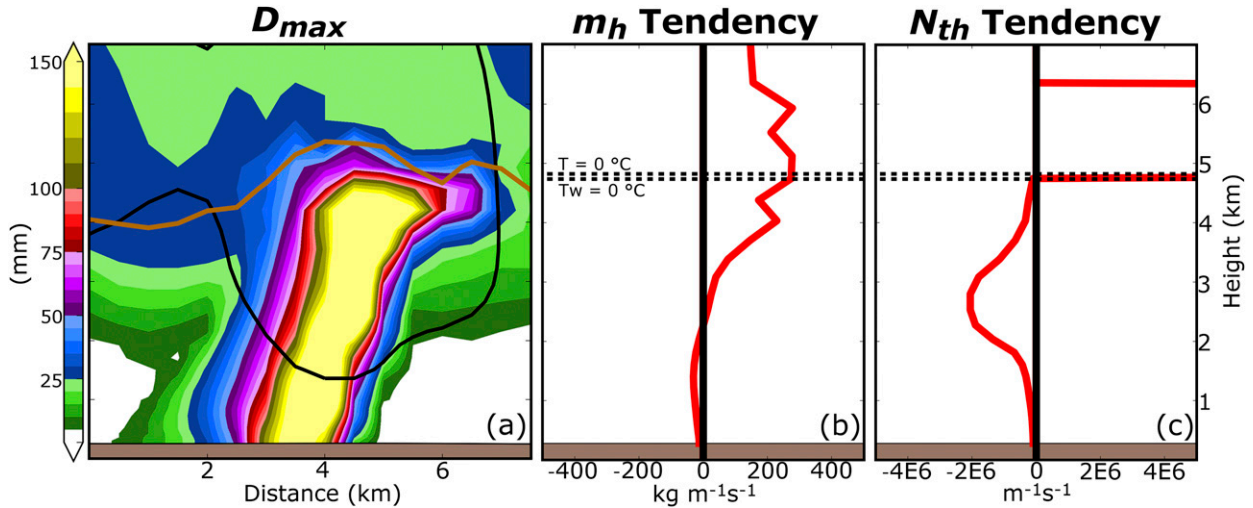


FIG. 10. (a) A vertical cross section of  $D_{max}$  taken through the hail core of the northernmost storm predicted by EXP\_MY2 member 20 at 2250 UTC. (b),(c) The total microphysical tendency for  $m_h$  and  $N_{th}$ , respectively, are also shown. The total microphysical tendency is integrated over each model vertical level in the storm hail core, which is highlighted by the black polygon in Fig. 7a. The tendencies are plotted as a function of average model level height. In the vertical cross section in (a) the horizontal brown line represents  $0^\circ\text{C}$  isotherm and the black contour represents where updraft velocity exceeds  $5\text{ m s}^{-1}$ . In (b) and (c) the upper horizontal dashed black line is the  $0^\circ\text{C}$  isotherm and the bottom dashed line is the wet-bulb  $0^\circ\text{C}$  isotherm.

MP schemes must continue to make assumptions on the treatment of hail during melting.

#### 4. Summary and discussion

In this study we evaluate 90-min-long EnKF-initialized explicit hail ensemble forecasts at 500-m grid spacing for the 19 May 2013 Oklahoma City, Oklahoma, supercell hailstorm event. For the initial

conditions, we assimilate all available surface, profiler, and radar (reflectivity and radial velocity) observations every 10 min, over a 110-min period ending at the final analysis time of 2150 UTC. Three ensembles, consisting of 40 members each, are run for both the data assimilation and forecast periods. Each ensemble is similarly configured but uses one of three microphysics (MP) schemes: the Milbrandt and Yau (2005a) double-moment MP scheme (MY2), the Milbrandt and Yau (2005a)

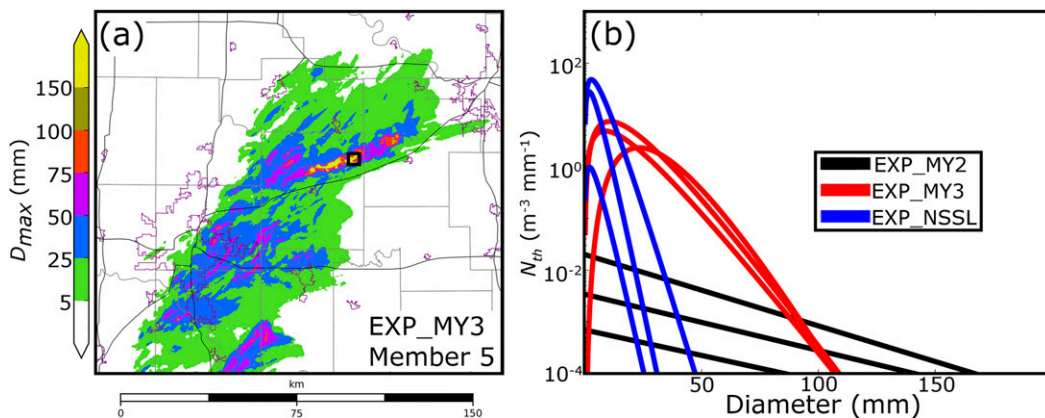


FIG. 11. (a) A Swath of forecast maximum surface  $D_{max}$  calculated at every model time step between 2155 and 2320 UTC for EXP\_MY3 member 5. A black polygon indicates where hail PSDs are sampled for the MY3 scheme. Background maps are the same as in Fig. 1. (b) A comparison of hail core PSDs for EXP\_MY2 member 20 (Fig. 7a), EXP\_MY3 member 5 (Fig. 11a), and EXP\_NSSL member 20 (Fig. 7c) at 2250 UTC. PSDs are sampled from within the black polygons in Fig. 7a, Fig. 11a, and Fig. 7c, respectively. Because the minimum y-axis value is  $10^{-4}\text{ m}^{-3}\text{ mm}^{-1}$  the x intercept is considered to be the maximum observable hail size  $D_{max}$ . For EXP\_MY3, a rightward shift in the PSD peak is caused by an increased shape parameter.



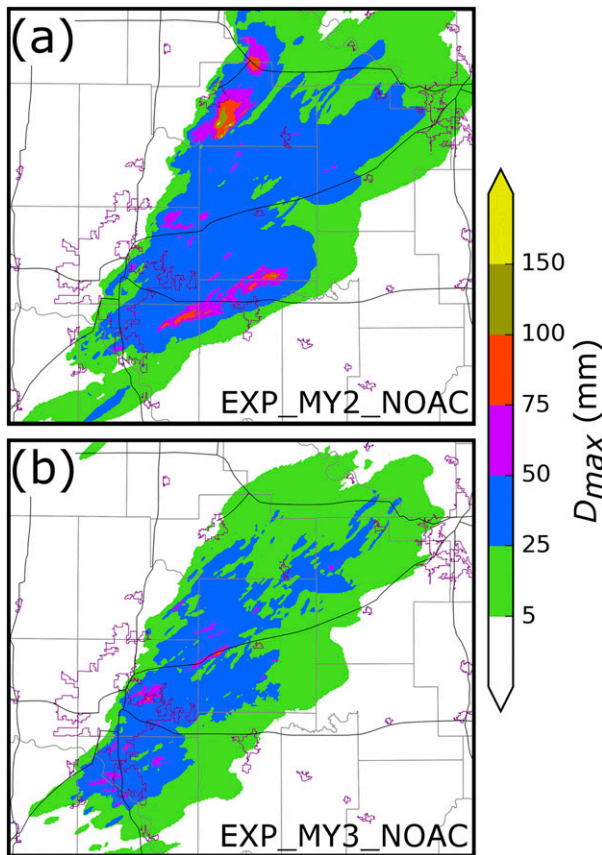


FIG. 12. Swaths of forecast maximum surface  $D_{\max}$  calculated at every model time step between 2155 and 2320 UTC. Forecasts are initialized from member 20 analyses of (a) EXP\_MY2 and (b) EXP\_MY3. Forecasts (a) EXP\_MY2\_NOAC and (b) EXP\_MY3\_NOAC are run with modified MY2 and MY3 schemes, respectively, where the accretion of rainwater and cloud water does not occur beneath the  $0^{\circ}\text{C}$  isotherm. Background maps are the same as in Fig. 1.

triple-moment MP scheme (MY3), or the NSSL double-moment variable-density-rimed ice MP scheme (Mansell et al. 2010). Ensembles run using these schemes are hereafter referred to as EXP\_MY2, EXP\_MY3, and EXP\_NSSL, respectively.

Surface hail size forecasts are verified against output from a hydrometeor classification algorithm (HCA). We determine the forecast maximum hail size ( $D_{\max}$ ) via hail particle size distributions (PSDs), which approximates the largest observable hailstone at each model grid point. Area under the relative operating characteristic (AUC) scores indicate all three MP schemes predict the coverage of severe (diameter  $>25$  mm) and significant severe (diameter  $>50$  mm) hail with a high level of skill ( $\geq 0.92$ ); albeit Brier skill scores suggest the schemes predict significant severe hail coverage with less skill than severe hail coverage. The ensembles

produce reliable hail size forecasts, with the exception of EXP\_NSSL, which produces under confident significant severe hail forecasts. Although EXP\_MY2 and EXP\_MY3 produce skillful severe and significant severe hail forecasts, both ensembles overestimate the maximum surface hail size. Qualitatively, EXP\_NSSL most closely resembles HCA output but the scheme predicts limited significant severe hail coverage.

Hail production tendencies are evaluated during the forecast period to understand hail size forecast differences and biases. Storms predicted by either EXP\_MY2 or EXP\_MY3 generally produce a large number of small hail particles above the  $0^{\circ}\text{C}$  isotherm because the MP schemes create hail via the three-component accretion of rain (i.e., frozen raindrops). The small hail particles are advected downstream of the updraft to produce large plumes of hail aloft that extend more than 40 km downstream of the updraft. EXP\_NSSL predicts storms to produce fewer but relatively larger hail particles above the  $0^{\circ}\text{C}$  isotherm because the NSSL scheme hail category consists of only dense graupel that has undergone wet growth.

Hail size forecast skill is also determined by how MP schemes model hail in and beneath the melting layer. Because water fraction is not predicted by the MY2, MY3, or NSSL schemes, neither scheme is able to represent the collection of liquid water in the melting layer. Both the MY2 and MY3 MP schemes assume all water collected by hail is converted to ice both above and beneath the melting layer; this process causes model predicted hail to exceed 150 mm ( $\sim 6$  in.) in diameter beneath the  $0^{\circ}\text{C}$  isotherm. In regions where hail is exceptionally large the MY3 scheme often increases the shape parameter ( $\alpha_h$ ) of hail. Increasing  $\alpha_h$  narrows the hail PSD and moderates the most extreme hail sizes (diameter  $>150$  mm); however, EXP\_MY3 continues to overpredict hail size. The NSSL scheme sheds all accreted liquid water when the air temperature is warmer than  $0^{\circ}\text{C}$ . While this assumption prevents the rapid growth of hail in the melting layer, it is not realistic when compared to observations (e.g., Rasmussen et al. 1984) and neglects potential hail growth via the collection of water (Lesins and List 1986).

To further improve the hail forecast skill, we must use MP schemes that accurately model hail growth processes. MP schemes such as the predicted particle properties (P3; Morrison and Milbrandt 2015; Morrison et al. 2015; Milbrandt and Morrison 2016) scheme and the Ice-Spheroids Habit Model with Aspect-ratio Evolution (ISHMAEL; Jensen et al. 2017) scheme explicitly predict the mass and volume of accreted rime to avoid converting between predefined hydrometeor species. Hydrometeor conversion between unrimed and rimed



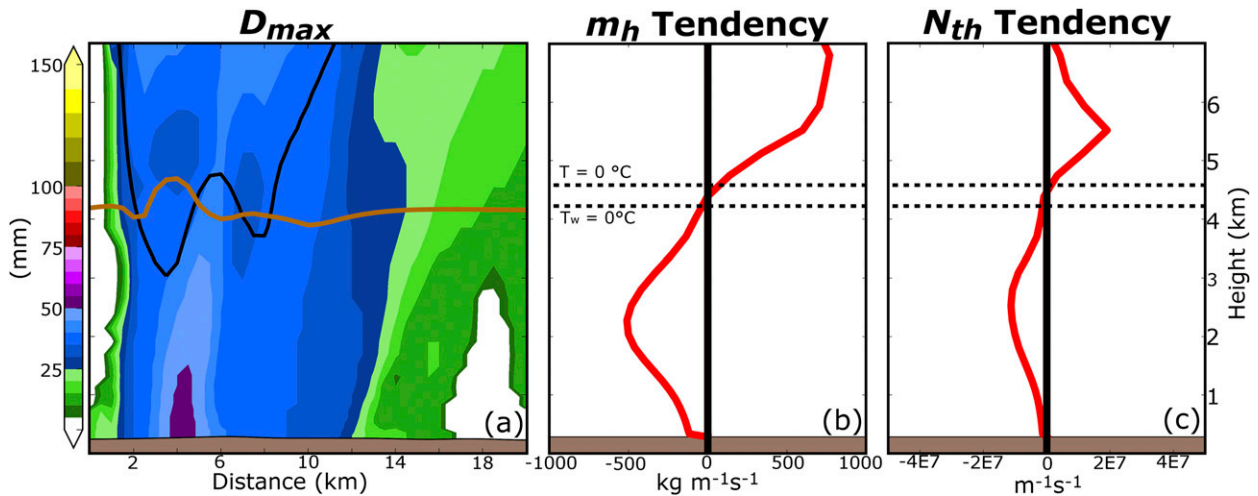


FIG. 13. (a) A vertical cross section of  $D_{max}$  taken through the hail core of the northernmost supercell predicted by EXP\_NSSL member 20 at 2250 UTC. (b),(c) The total microphysical tendency for  $m_h$  and  $N_{th}$ , respectively, are also shown. The total microphysical tendency is integrated over each model vertical level in the storm hail core, which is highlighted by a polygon in Fig. 7c. The total tendency is plotted as a function of average model level height. In the vertical cross section in (a) the horizontal brown line represents  $0^\circ \text{C}$  isotherm and the black contour represents where updraft velocity exceeds  $5 \text{ m s}^{-1}$ . In (b) and (c) the upper horizontal dashed black line is the  $0^\circ \text{C}$  isotherm and the bottom dashed line is the wet-bulb  $0^\circ \text{C}$  isotherm.

ice categories is a large source of model error, this is most evident in EXP\_MY2 and EXP\_MY3 forecasts when frozen rain is converted to hail. ISHMAEL also predicts particle shape, which is shown to impact the spatial distribution of ice due to size sorting (Jensen et al. 2018). Further investigation into the inclusion of prognostic water fraction equations such as those in Ferrier (1994) could also improve representation of ice within the melting layer. Integrating these novel MP schemes into a hail prediction system could be part of an investigation of the benefit of including additional microphysical variables and processes.

Data assimilation can also be used to improve the microphysical state of hailstorms; however, we must determine if microphysical variables are being realistically updated by the DA system. Performing data assimilation when using a multimoment MP scheme is nontrivial. Because of the large number of predicted hydrometeor properties, microphysical variables are typically under constrained by observations (i.e., reflectivity). Microphysical variables also have very nonlinear relations with observations; Xue et al. (2010) notes this introduces large errors into the initial conditions. For this study, we update only hydrometeor mixing ratios to limit the degrees of freedom of the model that are updated. Although this technique produces a stable configuration, the resulting analysis is suboptimal because  $D_{max}$  is often too large in early cycles when forecast errors are large. An optimal data assimilation procedure for multimoment MP schemes should be

developed to improve initial condition estimates and produce a more skilled hail size forecast.

A limited number of studies have analyzed hail forecasts for storm modes other than supercell thunderstorms. Many nonrotating storms produce large hail and cause extensive damage, one such hail event caused more than \$1.4 billion in hail damage in Colorado (Fritz 2017). Studies such as Luo et al. (2017) have evaluated the skill of next-day CAM forecasts for a severe-hail-producing pulse thunderstorm, but to our knowledge no studies have evaluated hail forecasts for nonrotating storms using a frequently updated, high-resolution ensemble forecast framework. In the future we plan to expand the number of case studies to include multiple hail events.

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