

1 Cool skin signals observed from Advanced Along-Track Scanning
2 Radiometer (AATSR) and *in situ* SST measurements

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8 **Abstract:** Nighttime cool skin sea surface temperature (SST) signals, defined in this study as
9 the differences between the SST_{skin} from the Advanced Along-Track Scanning Radiometer
10 (AATSR) onboard Envisat satellite and *in situ* SSTs from drifting buoys and moorings, $\Delta T =$
11 SST_{skin}-SST_{insitu}, are investigated on a global scale from July 2002 to April 2012. Global
12 mean ΔT , averaged over the full study period, is -0.13 K, with most values falling between -
13 0.1 and -0.2 K. The dominant role of wind speed on the ΔT is shown, with weaker winds
14 usually corresponding to a cooler skin. The effect of air-sea temperature difference is also
15 significant: warm skin ($\Delta T > 0$ K) can be observed under large positive air-sea temperature
16 differences. Other geophysical variables, such as the total column water vapor, *in situ* SST,
17 and net heat flux, also affect ΔT , but to a lesser degree. Significant increase of ΔT size with
18 SST_{insitu} is observed when SST_{insitu} is > 28 °C. Tropical waters, such as the tropical Indian
19 Ocean and the tropical warm pool (western Pacific and eastern Indian Ocean), are more
20 frequently covered with a cool skin, largely due to the calm winds, very warm waters
21 (especially for SST_{insitu} > 28 °C), and other environmental conditions supporting the
22 development of large cool skin events. The ΔT seasonal pattern in the southern hemisphere is
23 more regular, compared to the northern hemisphere. In both hemispheres, larger cool skin
24 signals are seen during the local summer, mainly due to weaker winds. According to several

25 previous cool skin models, higher winds tend to result in thinner cool skin layer depths, and
26 hence in smaller ΔT amplitudes, regardless of stronger evaporation and heat loss. Given that
27 wind is closely coupled with waves and turbulent mixing with wave breaking, the
28 dependencies of ΔT on a few wave parameters are also investigated. A strong (moderate)
29 dependency of ΔT on wave height (wave steepness) is identified, while the dependency of ΔT
30 on wave breaking probability is less discernible.

31 **Key words:** sea surface temperature, cool skin, AATSR, seasonal patterns, wave breaking.

32

33 **1. Introduction**

34 It has been long known that the skin sea surface temperature (SST) is usually slightly cooler
35 than the temperature immediately below, referred to as the cool skin effect or “skin effect”,
36 for short (e.g., Woodcock, 1941, 1947). According to the Group for High Resolution SST
37 (GHRSSST; Donlon et al., 2007) practical definitions, the skin SST refers to the temperature at
38 around 10-20 μm depth, measured by an infrared (IR) radiometer typically operating at 3.7-
39 12 μm wavelengths. Conductive diffusion is dominant in this layer. Cool skin exists because
40 of the combined cooling effects of the longwave radiation and the latent and sensible heat
41 fluxes (e.g., Saunders, 1967; Fairall et al., 1996). Since under most circumstances the net heat
42 flux is from the ocean to the atmosphere, the cool skin is usually present with an amplitude of
43 a few tenths of a degree. In the daytime, when the wind is calm and solar insolation strong,
44 diurnal warming can have amplitudes of several Kelvins, which may offset the cool skin
45 effect (e.g., Fairall et al., 1996; Gentemann et al., 2003). Diurnal warm layer tends to vanish
46 at night, when a near constant temperature profile is restored in the upper few meters up to
47 the bottom of the skin layer, which is referred to as the subskin SST, SST_{subskin} , measured by
48 a microwave (MW) sensor at ~ 1 mm depth. Incorporating a diurnal warm and cool skin
49 layers’ schemes in a numerical weather prediction or a climate model has been shown to
50 improve the model’s accuracy, due to more accurate estimation of air-sea interactions (e.g.,
51 Robertson and Watson, 1992; Zeng and Beljaars, 2005; Brunke et al., 2008; Masson et al.,
52 2012; Clayson & Bogdanoff, 2013; Akella et al., 2017).

53 Numerous studies have been conducted to observe and/or model the cool skin layer (e.g.,
54 Saunders, 1967; Hasse, 1971; Brutsaert, 1975; Liu et al., 1979; Hepplewhite, 1989; Schlüssel
55 et al., 1990; Soloviev and Schlüssel, 1994, 1996; Fairall et al., 1996; Wick et al., 1996; Wick
56 and Jessup, 1998; Artale et al., 2002; Donlon et al., 2002; Castro et al., 2003; Minnett, 2003;
57 Tu and Tsuang, 2005; Ward, 2006; Minnett et al., 2011; Alappattu et al., 2017). Most of these

58 studies used shipborne IR skin SSTs and coincident depth SSTs at a few centimetres to
59 metres depths. The spatial and temporal scales are therefore restricted to the ships' duration
60 and routes. While some studies used both daytime and nighttime data (e.g., Kent et al., 1996;
61 Minnett, 2003; Minnett et al., 2011), it is not uncommon for some authors to adopt nighttime
62 data only, to minimize the complication caused by diurnal warming (e.g., Horrocks et al.,
63 2003). With data of high accuracy but sometimes limited in amount, a series of physical or
64 empirical cool skin models have been developed, which can be generally divided into two
65 groups. The first group, represented by the model proposed in Saunders (1967), considers two
66 essential mechanisms controlling the heat fluxes across the molecular skin layer: free
67 convection, caused by the thermal instability, and the salinity gradient across the cool skin
68 itself under very calm winds ($< 2 \text{ m s}^{-1}$), and forced convection driven by the surface shear
69 stress. Many studies followed with foci on determining the Saunders' proportionality constant
70 λ and then the thickness of the cool skin layer (e.g., Paulson and Simpson, 1981; Robinson et
71 al., 1984; Wu, 1985; Fairall et al., 1996; Artale et al., 2002; Tu and Tsuang, 2005). The other
72 group of parameterizations was developed based on the surface renewal theory, which
73 assumes that a part of the surface layer is removed and replaced by water from beneath (e.g.,
74 Brutsaert, 1975; Liu et al., 1979; Schlüssel et al., 1990; Soloviev and Schlüssel, 1994; Wick
75 et al., 1996; Castro et al., 2003). In addition to the physical models, empirical
76 parameterizations have been also proposed in more recent studies, relating the cool skin layer
77 amplitude to environmental variables such as wind speed (e.g., Donlon et al., 2002; Minnett
78 et al., 2011; Alappattu et al., 2017). Different models have been intercompared with each
79 other in several studies (e.g., Kent et al., 1996; Castro et al., 2003; Horrocks et al., 2003; Tu
80 and Tsuang, 2005).

81 Although spaceborne radiometers have been retrieving the skin SST on a global scale for
82 almost four decades since early 1980s, satellite data have been scarcely used in cool skin

83 investigation, likely due to several reasons. First, satellite IR and *in situ* SSTs represent
84 waters of spatially and temporally different scales. Infrared SSTs are almost instantaneous
85 area averages, whereas *in situ* SSTs are point data, which may be measured either
86 instantaneously or averaged in time. Such systematic differences exist regardless of the sizes
87 of the temporal and spatial windows employed in their collocation. Second, traditionally, the
88 uncertainties in IR SST retrievals are considered too large (with a typical standard deviation,
89 $SD \sim 0.5$ K, when validated against drifting or mooring buoy measurements), making it
90 challenging to analyse the cool skin signal, whose amplitude is typically much smaller.
91 Several factors contribute to this large uncertainty, related to the instrument (spectral
92 response, radiometric noise, in-flight calibration, etc.) and retrieval algorithms (including
93 cloud screening, aerosol detection, correction for the effects of water vapor absorption, etc.)
94 (e.g., Kilpatrick et al., 2015).

95 These issues have been presumably minimized in the (Advanced) Along Track Scanning
96 Radiometers, (A)ATSRs, SST_{skin} data sets produced by the (A)ATSR Reprocessing for
97 Climate (ARC) project. The (A)ATSRs are characterized by more accurate in-flight
98 calibration and dual-view technique, which in turn offer an improved potential for accurate
99 atmospheric correction (Llewellyn-Jones et al., 2001). In particular, the ARC project retrieves
100 SST_{skin} using coefficients based on radiative transfer models (RTM), independently of *in situ*
101 measurements. Recall that many operational SST retrieval algorithms are empirically
102 regressed against *in situ* measurements, and therefore may not be fully independent from
103 those. Although not without its own limitations, the ARC SST_{skin} data may be thus better
104 suited for skin effect studies (Murray et al., 2000; see more detail in section 2.1 and 3).

105 This paper characterizes the cool skin behaviours on a global scale, using nearly ten-year
106 nighttime AATSR SST_{skin} data, in conjunction with collocated *in situ* SSTs, measured by
107 drifting, coastal and tropical mooring buoys. The structure of the paper is as follows. Section

108 2 introduces the data sets and methods. Section 3 briefly describes AATSR and *in situ* data.
109 Section 4 characterizes the cool skin signals, including their statistics and relationships with
110 different environmental variables. Discussion and conclusions are presented in sections 5 and
111 6, respectively.

112 **2. Data and Methods**

113 *2.1. Data sets*

114 *2.1.1 AATSR SST_{skin} data*

115 The AATSR sensor was flown onboard ESA's Envisat satellite, launched in March 2002 as a
116 successor to ATSRs -1 and -2 (launched in July 1991 and April 1995, respectively).
117 Compared to previous missions, several improvements have been made to this family of
118 instruments, including: 1) the dual-view (nadir and forward views $\sim 55^\circ$ from zenith)
119 geometry, within a few minutes of each other, allowing for more effective atmospheric
120 correction; 2) a rigorous pre-launch calibration programme and continuous onboard
121 calibration of the thermal channels against two stable, high-accuracy black-body calibration
122 targets; and 3) the high clarity and sensitivity of the IR image data, mainly due to the special
123 mechanical coolers, which were used for the first time on ATSR-1 (Llewellyn-Jones et al.,
124 2001; Llewellyn-Jones and Remedios, 2012). In addition, from the very onset it was intended
125 that ATSR SSTs should be obtained independently of *in situ* measurements, through the use
126 of RTM to define ATSR retrieval coefficients (Zàvody et al., 1995; Merchant et al., 1999;
127 Embury and Merchant, 2012; Embury et al., 2012a; Embury et al., 2012b).

128 The Envisat crosses the equator at ~ 10 am/pm local time. At night, SST_{skin} is retrieved using
129 bands centred at 3.7, 11, and 12 μm . During the daytime, channel at 3.7 μm is not used, due
130 to solar reflectance and scattering. There are up to four retrieved SSTs in each pixel, referred

131 to as N2 (nadir two channel), N3 (nadir three channel), D2 (dual-view two channel) and D3
132 (dual-view three channel), respectively.

133 The data used in this study, are from the latest third AATSR reprocessing, level 2 pre-
134 collocation (L2P) dataset, spanning more than nine years from late July 2002 to early April
135 2012 (accessed from <ftp://ats-ftp-ds.eo.esa.int>). The third reprocessing products use the same
136 algorithm as the ARC v1.1 data, with the only difference being that the former is GHRSSST
137 format compliant while the latter is not (O. Embury, personal communication). Skin SSTs
138 without sensor-specific error statistics bias (SSES Bias) correction are used. To avoid
139 possible complication caused by diurnal warming, and inconsistent day/night bands and
140 algorithms, only highest quality (quality level, QL = 5) nighttime D3 data are selected.
141 According to prior studies, D3 retrievals are considered the most accurate (e.g., Embury et al.,
142 2012a; Merchant et al., 2012).

143 2.2.2 *In Situ SST measurements*

144 It is known that strong diurnal warming events could persist well into the late night or even
145 the next early morning, especially under very calm conditions (e.g., Gentemann et al., 2003;
146 Gentemann and Minnett, 2008; Zhang et al., 2018). It is extremely challenging to totally
147 exclude the diurnal warming residuals in the upper layer. In this study, it is even more so as
148 the AATSR has an early night local passing time of ~ 10 pm. Nonetheless, previous studies
149 have shown that the diurnal warming residuals in the upper few meters can be minimum
150 under well-mixed conditions, i.e. wind speed $> 2 \text{ m s}^{-1}$ at night (e.g., Donlon et al., 2002;
151 Matthews et al., 2014; Zhang et al., 2016). Hence, SST measurements from *in situ* platforms
152 such as drifting buoys at ~ 20 cm depth, and tropical and coastal moorings at ~ 1 m depth, are
153 expected to well represent $\text{SST}_{\text{subskin}}$. In this study, *in situ* SST data are obtained from NOAA
154 *in situ* SST Quality Monitor version 2 (*iQuam 2*; www.star.nesdis.noaa.gov/sod/sst/iquam/)

155 system (Xu and Ignatov, 2014). Only data of drifting, tropical and coastal moored buoys are
156 used.

157 In addition to applying the appended QL filter, we also looked at the SSTs as a function of
158 latitude. In every 2° latitudinal band, any value falling outside of the mean $\pm 2\times SD$ range,
159 was discarded. Enlarging the window to $3\times SD$ increased the data counts by only 2.3%, while
160 introducing more noise, and therefore was not adopted. In addition, *in situ* SSTs < -5 °C are
161 considered unrealistic and also excluded.

162 2.2.3. Environmental variables

163 All meteorological variables in this study are obtained from the daily European Centre for
164 Medium-Range Weather Forecasts (ECMWF) interim reanalysis (ERA-Interim; Dee et al., 2011),
165 namely the 10 m wind speed (U_{10}), air temperature (T_a), latent heat flux (Q_l), sensible heat
166 flux (Q_s), net longwave thermal radiation (Q_{lw}), and total column water vapour (TCWV). The
167 spatial resolution is by default 0.75° and temporal resolution 3-hrs. For instantaneous
168 variables like U_{10} , T_a , and TCWV, analysis and forecast data are combined: analysis for every
169 6 hours (00:00, 06:00, 12:00, 18:00 UTC) and forecast for the hours in between (03:00, 09:00,
170 15:00, 21:00), while all the heat fluxes are forecast data only. At night, when there is no solar
171 insolation, the net heat flux Q_{net} is calculated as the sum of $Q_l + Q_s + Q_{lw}$, with negative sign
172 indicating heat from the ocean to the atmosphere, and positive the other way around.

173 The ERA-Interim products are among the most widely used reanalysis data sets, thanks to their good
174 quality. The ERA-Interim environmental variables used in this study have been extensively validated
175 against satellite and *in situ* measurements, and inter-compared with other reanalysis data sets
176 (e.g., Brunke et al., 2011; Chaudhuri et al., 2013). For instance, Brunke et al. (2011) found
177 that the ERA-Interim latent heat flux and wind stress fields perform very well when evaluated
178 against direct covariance latent heat flux and inertial-dissipation wind stresses measured from

179 cruises in the tropics, as well as in the mid- and high latitudes. For more detail of the
180 validation and quality of the ERAI meteorological fields, refer to Dee et al. (2011), Brunke et
181 al. (2011), and Chaudhuri et al. (2013) and references therein.

182 2.2. Collocation and quality control

183 The QL = 5 AATSR and QC'd *in situ* SST data are first collocated. The temporal and spatial
184 windows are set to 1 hr and 0.1°, respectively, comparable to (or more conservative than) in
185 other similar studies, which use 3 hr and 10 arc min thresholds (e.g., O'Carroll et al., 2008;
186 Xu and Ignatov, 2016). Any AATSR and *in situ* collocations with absolute difference > 10 K
187 are discarded. Then, meteorological variables at the nearest ERAI pixel are assigned to each
188 collocation, meaning that the temporal and spatial windows are half of the ERAI's resolutions:
189 1.5 hrs and 0.375°. At 10 ± 1.5 pm local time in high latitudes, there may still be positive
190 downward solar insolation, especially in summer. These data, accounting for ~4.3% of all
191 collocations, have been discarded. Finally, the resulting match-up data set includes total N =
192 594,777 collocations.

193 3. Quality of AATSR and *in situ* SSTs

194 The high quality of AATSR SSTs have been illustrated in a number of studies (e.g., Corlett et
195 al., 2006; Noyes et al., 2006; O'Carroll et al., 2006; O'Carroll et al., 2008; Reynolds et al.,
196 2010; Kennedy et al., 2012; Merchant et al., 2012; Merchant et al., 2014; Xu and Ignatov,
197 2016). Several publications used a triple collocation method (TCM) and showed that AATSR
198 SSTs are as precise as *in situ* observations, or better. For instance, O'Carroll et al. (2008)
199 showed that the spatially averaged nighttime AATSR D3 bulk SST observations (converted
200 using the Fairall et al. 1996 model) for 2003 have a SD ~ 0.16 K vs. drifters' SD ~ 0.23 K.
201 Merchant et al. (2012) found the ARC SSTs have a SD ~ 0.14 K during 2003 to 2009,
202 compared to SD ~ 0.15-0.18 K for drifters. Merchant et al. (2014, their Table 2) described the

203 ARC v1.1 data set as the “most accurate and stable SST product available” in comparison
204 with several other data sets. In terms of accuracy, although meeting the < 0.1 K requirement,
205 some authors did observe a small warm bias of ~ 0.04 - 0.06 K when comparing nighttime
206 AATSR data against drifting buoys or radiometers (e.g., Corlett et al., 2006; Noyes et al.,
207 2006).

208 Note that the ARC skin SST data may be also subject to their own limitations. Just like any
209 IR sensor, there is no SST_{skin} retrieved when and where cloud is present. Compared to
210 empirical regression algorithms, an RTM-based retrieval scheme employs numerous
211 environmental data sources (e.g., various atmospheric components, atmospheric and marine
212 aerosols), all of which could contribute to the SST uncertainty to a different extent (Embury
213 et al., 2012b). The RTM may not be fully accurate, either. In addition, the sensitivity of SST
214 retrievals to water vapor or to actual SSTs may vary across different regions (Embury and
215 Merchant, 2012). However, at this moment, this remote sensing skin SST data set is thought
216 to be the most suitable for a global, long-term skin effect study (Merchant et al., 2014).

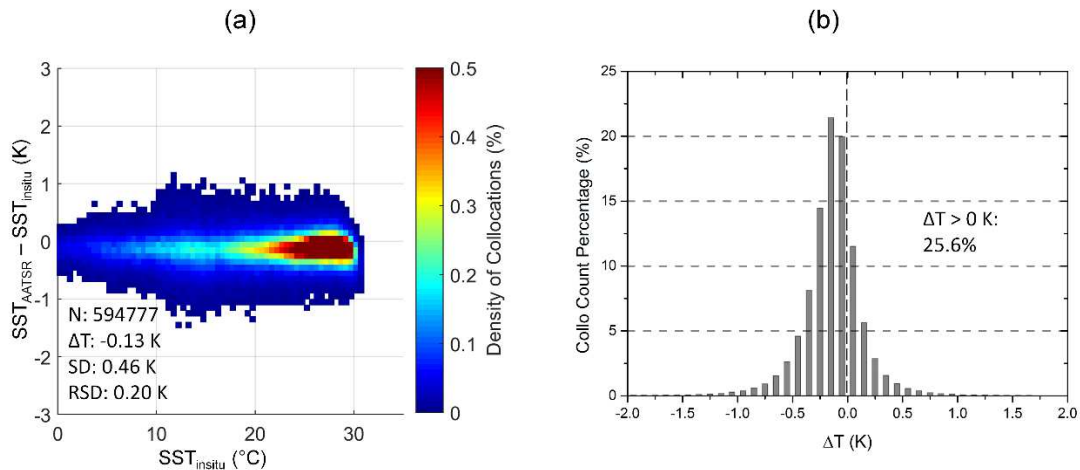
217 Xu and Ignatov (2016) estimated SDs for tropical and coastal moorings as 0.17 K and 0.40 K,
218 respectively. To employ as much data as possible, mooring measurements are also retained in
219 this study, and validation statistics are not stratified by *in situ* data types.

220 **4. Results**

221 *4.1. General statistics*

222 The overall statistics of AATSR satellite SST_{skin} minus SST_{insitu} are shown in Fig. 1. The
223 mean ΔT (defined as $SST_{\text{skin}} - SST_{\text{insitu}}$) is -0.13 K, which is comparable to the findings in
224 previous studies such as Donlon et al. (2002) and Minnett et al. (2011). The 0.46 K SD is
225 obviously larger than the SDs of ~ 0.16 or 0.14 K in O’Carroll et al. (2008) and Merchant et
226 al. (2012), which were based on a TCM, which accounts for the uncertainties in SST_{insitu} . The

227 RSD (robust SD, calculated as 1.5 times the median absolute deviation from the median) of
 228 0.20 K is much smaller, since RSD is less sensitive to outliers. Fig. 1b shows the distribution
 229 of ΔT . The peak is found between -0.2 K and -0.1 K, with 21.4% of data falling in this range.
 230 The spread of the ΔT values is similar to the studies which used shipborne IR skin SST (e.g.,
 231 Hepplewhite et al., 1989; Donlon and Robinson, 1997), except that in this study, the positive
 232 ΔT values form a larger portion (25.6%). These positive ΔT values could be due to many
 233 reasons: residual diurnal warming, satellite and *in situ* data noises, real warm skin signals, etc.
 234 However, the subsequent analyses will show that a large part of those make physical sense
 235 and may be explained by the corresponding environmental conditions.

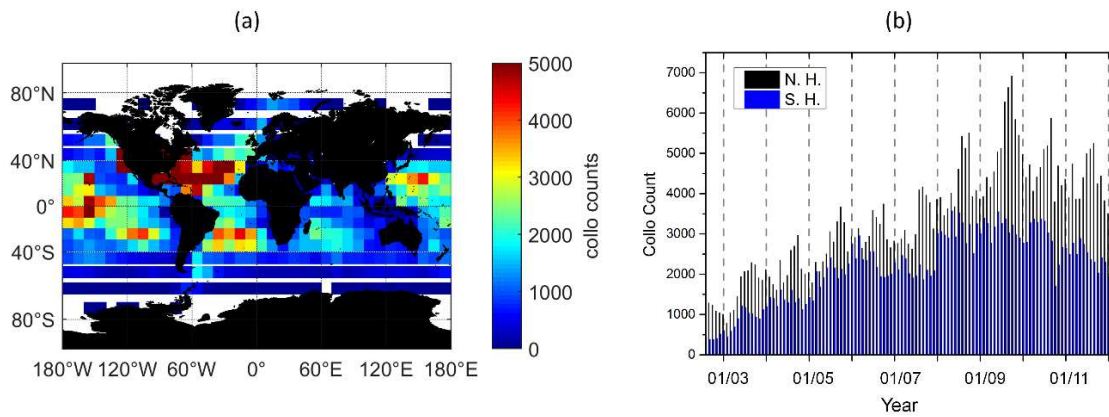


236

237 **Fig. 1.** (a) Density of collocations as a function of $SST_{in situ}$, with overall statistics of ΔT (= SST_{skin} -
 238 $SST_{in situ}$) superimposed. (b) Frequency distribution of ΔT s.

239 The spatial and temporal distributions of the collocations are shown in Fig. 2. The most
 240 populated regions are in the low-to-mid latitudes of the Northern Atlantic (Fig. 2a). The
 241 collocations in the Tropical Pacific are also abundant. Fewer collocations are found for some
 242 areas, particularly in the high latitudes ($> 40^\circ N/S$), largely due to the lack of *in situ*
 243 measurements, and over the eastern Indian Ocean and the tropical warm pool (TWP; tropical
 244 eastern Indian Ocean and western Pacific Ocean) areas, due to a combination of frequent
 245 clouds and sparse *in situ* data. Fig. 2b shows that the number of collocations in both

246 hemispheres has increased with time, as more *in situ* measurements became available. There
 247 are nearly always more collocations in the northern hemisphere than in the southern
 248 hemisphere.

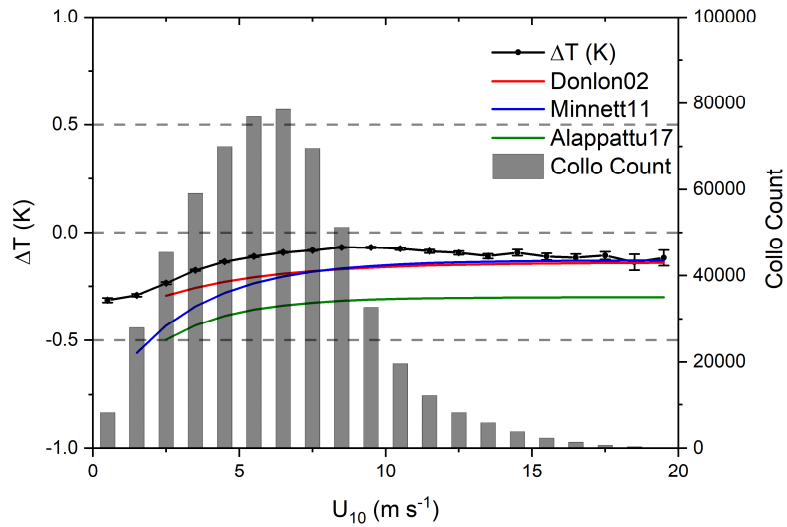


249

250 **Fig. 2.** (a) Spatial distribution of the collocations. Bin size is $10^\circ \times 10^\circ$. (b) Monthly collocation counts
 251 in the northern (black bars) and southern (blue bars) hemispheres.

252 4.2. Dependencies of ΔT on environmental variables

253 The effects of different meteorological factors on ΔT have been investigated, including U_{10} ,
 254 air-sea temperature difference $T_a - T_s$ (here T_s is SST_{insitu}), TCWV, SST_{insitu} , Q_{net} and its
 255 components. Here, QC'd *i*Quam SST_{insitu} observations are chosen to represent T_s over the
 256 ERAI SST reanalysis product. From Table 1 in Dee et al. (2011), since February 2009, the
 257 SST field used in ERAI comes from the Operational Sea Surface Temperature and Sea-Ice
 258 Analysis (OSTIA), a foundation SST product incorporating multiple SST sources, including
 259 ship measurements obtained at a significantly deeper layer than buoys (Donlon et al., 2012).
 260 Traditional ship SSTs, with typically less accurate depth information, are considered less
 261 reliable than buoys (Zhang et al., 2009). Therefore, to adopt the most reliable data and to
 262 keep analyses consistent with the calculation of the ΔT s, QC'd *i*Quam buoy SST observations
 263 are used here.



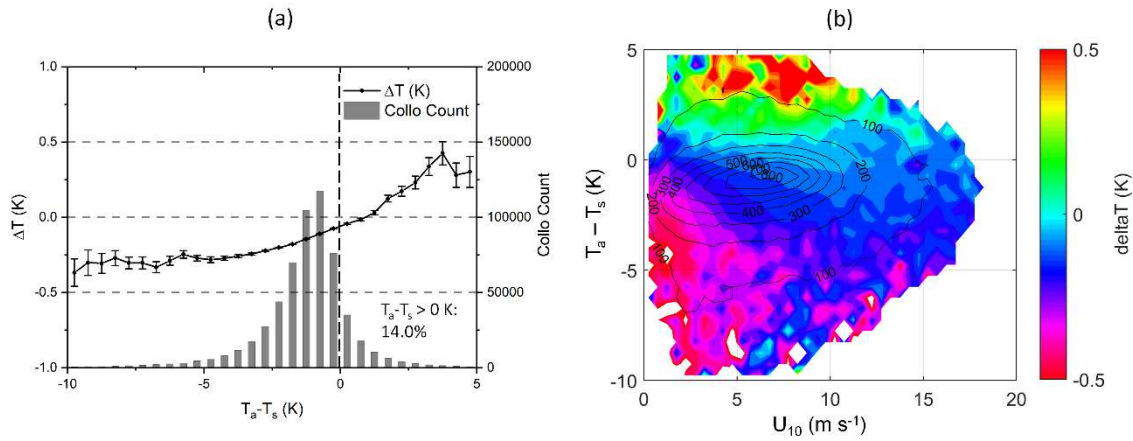
264

265 **Fig. 3.** Solid circles/Black line: dependencies of ΔT on U_{10} . The error bar is the 95% confidence level
 266 margin of error (MoE), i.e. 1.96 times the SD divided by the square root of the collocation number.
 267 The column bars indicate the collocation counts falling within each 1 m s^{-1} U_{10} interval. The red, blue,
 268 and green lines have been calculated using the Donlon et al. (2002), Minnett et al. (2011), and
 269 Alappattu et al., (2017) empirical parameterizations, respectively.

270 The cool skin amplitudes have been long known to strongly depend on U_{10} , as also shown in
 271 Fig. 3. The effect of wind on the cool skin amplitudes is twofold: through turbulent mixing
 272 and net heat flux. Increased winds result in stronger turbulent mixing, which reduces ΔT ;
 273 while larger winds also typically lead to more net heat flux, which is expected to increase ΔT
 274 size. The combined effects of U_{10} and Q_{net} will be discussed later. When the wind is very
 275 calm ($< 2 \text{ m s}^{-1}$ in Fig. 3), the ΔT s have largest amplitudes, with mean values reaching -0.30
 276 K to -0.35 K. As winds get stronger, ΔT climbs up steadily until U_{10} reaches $8\text{-}10 \text{ m s}^{-1}$, when
 277 ΔT starts to level off at -0.09 K to -0.12 K. The trend is very similar to observations made in
 278 previous studies, such as Donlon et al. (2002; hereafter D02), Minnett et al. (2011; hereafter
 279 M11), and Alappattu et al. (2017; hereafter A17), which are also plotted in Fig. 3. The
 280 empirical equations relating ΔT to wind speed are: $\Delta T = -0.14 - 0.30 \cdot \exp(-0.27 \cdot U_{10})$ in D02;

281 $\Delta T = -0.13 - 0.724 \cdot \exp(-0.35 \cdot U_{10})$ in M11; and $\Delta T = -0.30 - 0.55 \cdot \exp(-0.41 \cdot U_{10})$ in A17,
282 respectively. Note that the ΔT range observed in this study, differs from the ones predicted by
283 the models (which also all differ, as defined by the pre-exponent coefficient in the three
284 formulations). The cool skin values reported here are smaller in size than D02 and M11 by
285 0.05-0.07 K for U_{10} between 4-12 m s^{-1} conditions. Model M11 predicts the largest ΔT s $<$ -
286 0.5 K at very calm winds ($< 2 \text{ m s}^{-1}$). The line obtained in this study, and the D02 and M11,
287 all tend to asymptotically converge at high winds ($U_{10} > 13 \text{ m s}^{-1}$). The A17 model estimates
288 significantly larger cool skin sizes for all wind conditions.

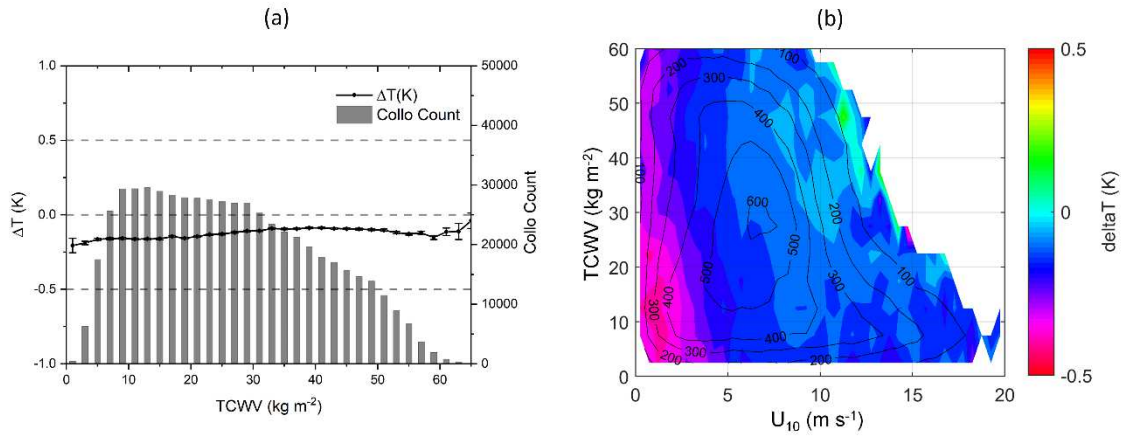
289 Several reasons might account for the ΔT difference between this study and the empirical
290 equations, which all may contain uncertainties of different kinds. In particular, the AATSR
291 skin SSTs are averaged over a relatively large area, in comparison with near-point data of
292 shipborne radiometers and thermistors. Also, the temporal collocation between AATSR
293 SST_{skin} and SST_{insitu} is within one hour, compared to only a few seconds to minutes between
294 shipborne radiometers and thermistors. Finally, there may be a small warm bias residual in
295 the nighttime D3 AATSR data, as found in previous studies (e.g., Corlett et al., 2006; Noyes
296 et al., 2006), although they were validating different versions of AATSR data (see section 3).
297 Diurnal warming residuals could partially be responsible for the smaller ΔT range found in
298 this study when U_{10} is $< 2 \text{ m s}^{-1}$, although they are expected to be progressively less critical
299 for $U_{10} > 2 \text{ m s}^{-1}$. As to the large difference between the A17, on the one hand, and
300 observations and other models; on the other, note that the A17 coefficients were determined
301 from coastal observations, which may have been affected by river discharge (Alappattu et al.,
302 2017). A plot similar to Fig. 3 is also found in Embury et al. (2012a; their Fig. 3) and
303 Alappattu et al. (2017; their Fig. 14).



304

305 **Fig. 4.** (a) Dependency of ΔT on $T_a - T_s$ (with the error bars at 95% confidence level margin of error,
 306 MoE). (b) Dependency of ΔT on $T_a - T_s$ (binned into 0.5 K) and U_{10} (binned into 1 m s⁻¹). Black
 307 contour lines indicate the corresponding collocation counts in each 1 m s⁻¹ × 0.5 K bin. Only bins with
 308 collocation counts ≥ 20 are plotted.

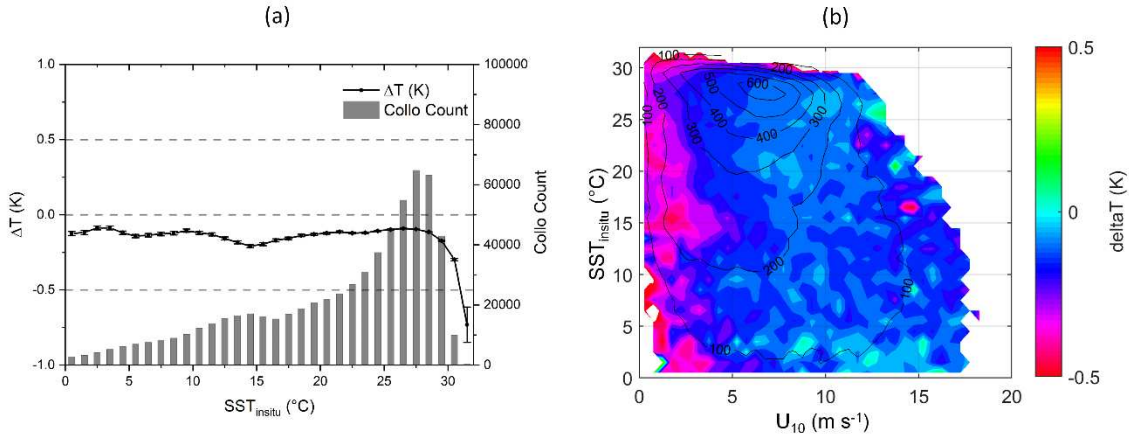
309 The effects of $T_a - T_s$ on ΔT are shown in Fig. 4. Generally, when the air is cooler than the sea
 310 water, heat is lost from the sea to the air, thus favouring the cool skin development. Fig. 4a
 311 shows a steady climb of ΔT as $T_a - T_s$ changes from negative to positive values, reaching ~
 312 +0.4 K when the $T_a - T_s$ is ~ 3-5 K. Since U_{10} is normally considered the most influential
 313 factor, the effect of $T_a - T_s$ is further analysed stratified by U_{10} conditions, in Fig. 4b. As
 314 expected, the skin is coolest when the $T_a - T_s$ reaches its minimal values, in conjunction with
 315 weakest winds. Increasing U_{10} reduces the cool skin amplitude, while increasing $T_a - T_s$ is able
 316 to even reverse the sign of ΔT . For most conditions with positive $T_a - T_s$, ΔT is close to zero or
 317 slightly positive, and sharply increases along with $T_a - T_s$. It is noticed that very large positive
 318 $T_a - T_s$ values typically occur for $U_{10} < 10$ m s⁻¹ conditions. Also, Fig. 4a indicates that positive
 319 $T_a - T_s$ values make up 14.0% of all data, which could explain at least a part of the 25.6%
 320 warm skin values in Fig. 1b.



321

322 **Fig. 5.** Same as in Fig. 4 but for the dependency of ΔT on TCWV (binned into 5 kg m^{-2}).

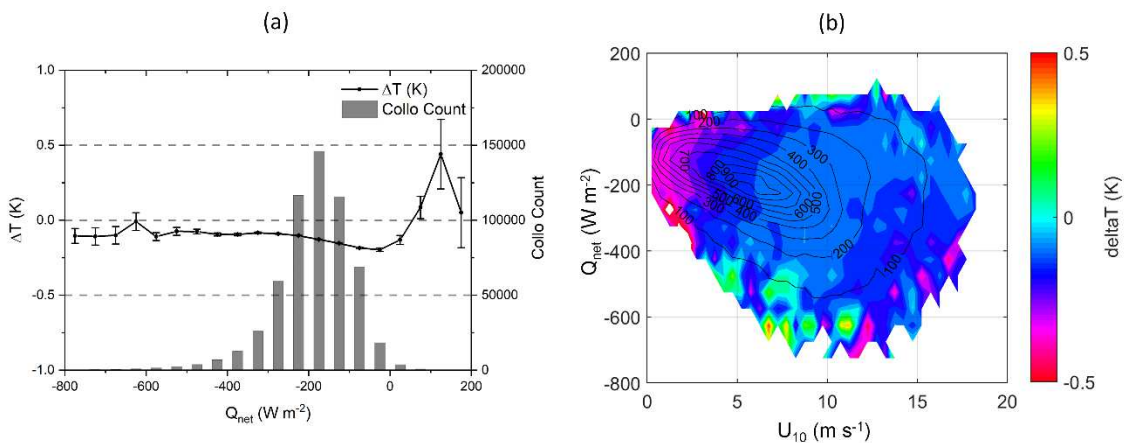
323 The dependency of ΔT on TCWV is shown in Fig. 5. The TCWV variable is selected because
 324 the IR SST_{skin} retrievals mostly aim to correct for its effect as a part of the atmospheric
 325 correction algorithms, and therefore may reveal some residual sensitivity to it. Moreover, the
 326 TCWV field affects the value of the Q_1 . According to Embury and Merchant (2012), the
 327 retrieval scheme of the AATSR data used in this study adopts coefficients banded by TCWV.
 328 The independence of the D3 retrievals on TCWV was illustrated by the near-zero biases over
 329 the whole TCWV range (Embury and Merchant, 2012; their Fig. 7d). In Fig. 5a, ΔT shows a
 330 weak upward trend when the TCWV increases from low ($\sim 5 \text{ kg m}^{-2}$) to middle range (~ 40
 331 kg m^{-2}), but the change of ΔT is small ($\sim 0.11 \text{ K}$). From $\sim 40 \text{ kg m}^{-2}$ to higher TCWV, the ΔT
 332 levels off. When holding U_{10} fixed (Fig. 5b), increasing TCWV can slightly reduce the
 333 amplitude of ΔT . The largest ΔT s appear when the TCWV is low and U_{10} weak. Overall, the
 334 change of ΔT in the full range of the TCWV, under a certain U_{10} conditions, is relatively
 335 minor, except maybe for $U_{10} > 8 \text{ m s}^{-1}$. Near-zero cool skin amplitudes are observed when
 336 U_{10} is strong ($\sim 12 \text{ m s}^{-1}$) and TCWV high ($\sim 50 \text{ kg m}^{-2}$).



337

338 **Fig. 6.** Same as Fig. 4 but for the dependency of ΔT on SST_{insitu} (binned into 1 $^{\circ}C$).

339 In Minnett et al. (2011), a possible temperature dependence of ΔT was reported (ΔT size
 340 increases as bulk SST changes from $\sim 10^{\circ}C$ to $\sim 18^{\circ}C$ when $U_{10} < 3 m s^{-1}$), although no
 341 robust conclusions were drawn due to the small match-up dataset. Here in Fig. 6a, we do not
 342 see a clear dependency of ΔT for $SST_{insitu} < 28^{\circ}C$. However, when SST_{insitu} is $> 28^{\circ}C$
 343 (especially $> 30^{\circ}C$), a sharp drop in ΔT is observed. The validation in Embury et al. (2012a)
 344 indicates little dependency of AATSR D3 SST_{skin} on latitudes, i.e. SST conditions. For fixed
 345 U_{10} in Fig. 6b, there is only little change with SST_{insitu} (except large ΔT amplitudes are seen
 346 for $SST_{insitu} > 30^{\circ}C$). Such warm waters basically only exist in the TWP area under very
 347 calm conditions ($U_{10} < 3 m s^{-1}$). We will investigate the TWP region in more detail in the
 348 next subsection.



349

350 **Fig. 7.** Same as Fig. 4 but for the dependency of ΔT on Q_{net} (binned into 50 W m^{-2}).

351 The effects of Q_{net} on ΔT are complicated by its further coupling with other meteorological
352 conditions, especially with U_{10} . In Fig. 7a, the ΔT dependency on Q_{net} does not show a clear
353 trend for Q_{net} values $< -400 \text{ W m}^{-2}$. A downward trend of ΔT is found for the most frequent
354 Q_{net} values (from -350 to 0 W m^{-2}). This negative relationship is seemingly going against
355 one's expectation of larger Q_{net} leading to larger ΔT amplitude. When plotting Q_{net} together
356 with its accompanying U_{10} , it is noticed that, to a large extent, Q_{net} is determined by U_{10} :
357 strong winds result in larger Q_{net} (Fig. 7b). Mixing induced by strong winds has dominated
358 over the increased Q_{net} , preventing the establishment of large cool skins. This may be the
359 main reason for the downward trend in Fig. 7a. In Fig. 7b, when U_{10} is determined, ΔT s are
360 rather stable regardless of Q_{net} sizes, indicating the weaker role of Q_{net} in ΔT development.
361 The positive ΔT peak for U_{10} between $5\text{-}10 \text{ m s}^{-1}$ and Q_{net} between -400 and -600 W m^{-2} , as
362 unexpected. This signal may be partially due to the sparse collocations under such conditions.
363 It is interesting to note that when Q_{net} is positive, i.e. heat flux is going into the ocean, a sharp
364 upward climbing of ΔT to positive values is seen (Fig. 7a). This makes physical sense in the
365 authors' opinion. In Fig. 7b, we get more near zero and positive ΔT values under the rare
366 positive Q_{net} conditions.

367 The ΔT dependencies on Q_l and Q_s largely follow those of Q_{net} and $T_a - T_s$, therefore not
368 shown here. Q_{lw} has similar effects as Q_{net} on ΔT , but with a smaller amplitude, and also not
369 shown.

370 Table 1 shows the statistics of the environmental variables for non-negative ΔT s and $\text{SST}_{\text{insitu}} >$
371 $30 \text{ }^\circ\text{C}$ situations. Compared to the overall and negative ΔT statistics, non-negative ΔT s
372 correspond to stronger winds and smaller $T_a - T_s$ sizes. This is reasonable since higher U_{10} and
373 smaller air-sea temperature differences both contribute to the development of a near-zero or
374 positive ΔT . In Fig. 6a, we observe a sharp drop in ΔT for $\text{SST}_{\text{insitu}} > 30 \text{ }^\circ\text{C}$ conditions. Such

375 warm waters are normally only located in the tropical areas, especially in the TWP region.
 376 The far lower winds and larger negative $T_a - T_s$ should account for the increased ΔT size.
 377 According to the validation in Embury et al. (2012a; their Fig. 12), AATSR D3 satellite data
 378 over these regions continue to show < 0.1 K biases. Embury and Merchant (2012) showed
 379 that the sensitivity of ARC AATSR D3 SSTs to actual SST is between 0.99 to 1.01 over the
 380 tropical region (their Fig. 14d), and a 10% increase in TCWV causes only a ~ 0.015 K
 381 decrease in the retrieved SST (their Fig. 13d; see also section 4.3 below).

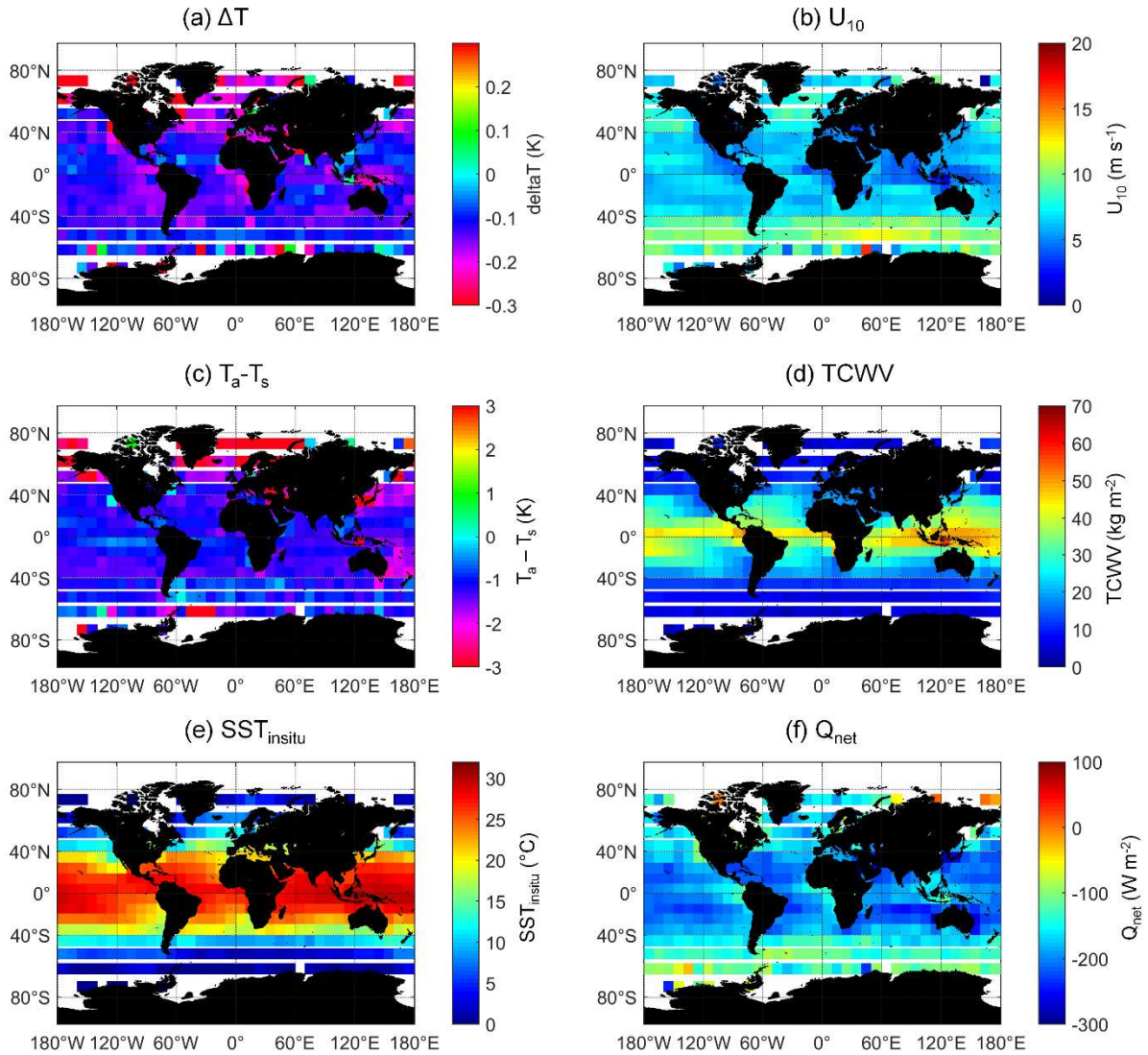
382 **Table 1.** Statistics for non-negative ΔT and $SST_{\text{insitu}} > 30$ °C conditions. The overall and negative ΔT
 383 rows are for comparison.

	Collocation Counts	Mean ΔT (K)	Mean U_{10} (m s^{-1})	Mean $T_a - T_s$ (K)	Mean TCWV (kg m^{-2})	Mean Q_{net} (W m^{-2})
Overall	594,777 (100%)	-0.13	6.20	-1.27	26.32	-187.54
$\Delta T < 0$ K	442,657 (74.6%)	-0.27	6.00	-1.41	26.35	-186.59
$\Delta T \geq 0$ K	152,120 (25.4%)	0.27	6.77	-0.86	26.25	-190.31
$SST_{\text{insitu}} >$ 30 °C	8,331 (1.4%)	-0.36	3.39	-1.90	47.34	-179.57

384

385 4.3. Spatial distribution of ΔT

386 In the previous subsection, we have shown that the major factors regulating ΔT are U_{10} and
 387 $T_a - T_s$. TCWV and Q_{net} play relatively minor roles. SST_{insitu} only has a significant impact
 388 when > 28 °C. In this subsection, we investigate the relationship between ΔT and all the
 389 variables by analysing their spatial distributions.



390

391 **Fig. 8.** Spatial distributions of (a) ΔT ; (b) U_{10} ; (c) $T_a - T_s$; (d) TCWV; (e) SST_{insitu} ; and (f) Q_{net} averaged
 392 over the whole study period. Box size is 10° by 10° .

393 Fig. 8 shows the distribution of ΔT along with other environmental variables averaged over
 394 the study period within 10° by 10° boxes. Several features are quickly spotted. Over most of
 395 the global oceans, average ΔT is around -0.1 to -0.15 K. The distribution pattern of the ΔT
 396 closely follows that of U_{10} : low winds correspond to large ΔT sizes. However, it is also
 397 noticed that large ΔT over different regions can be caused by one or a combination of
 398 different factors. Several regions are selected as examples for illustration.

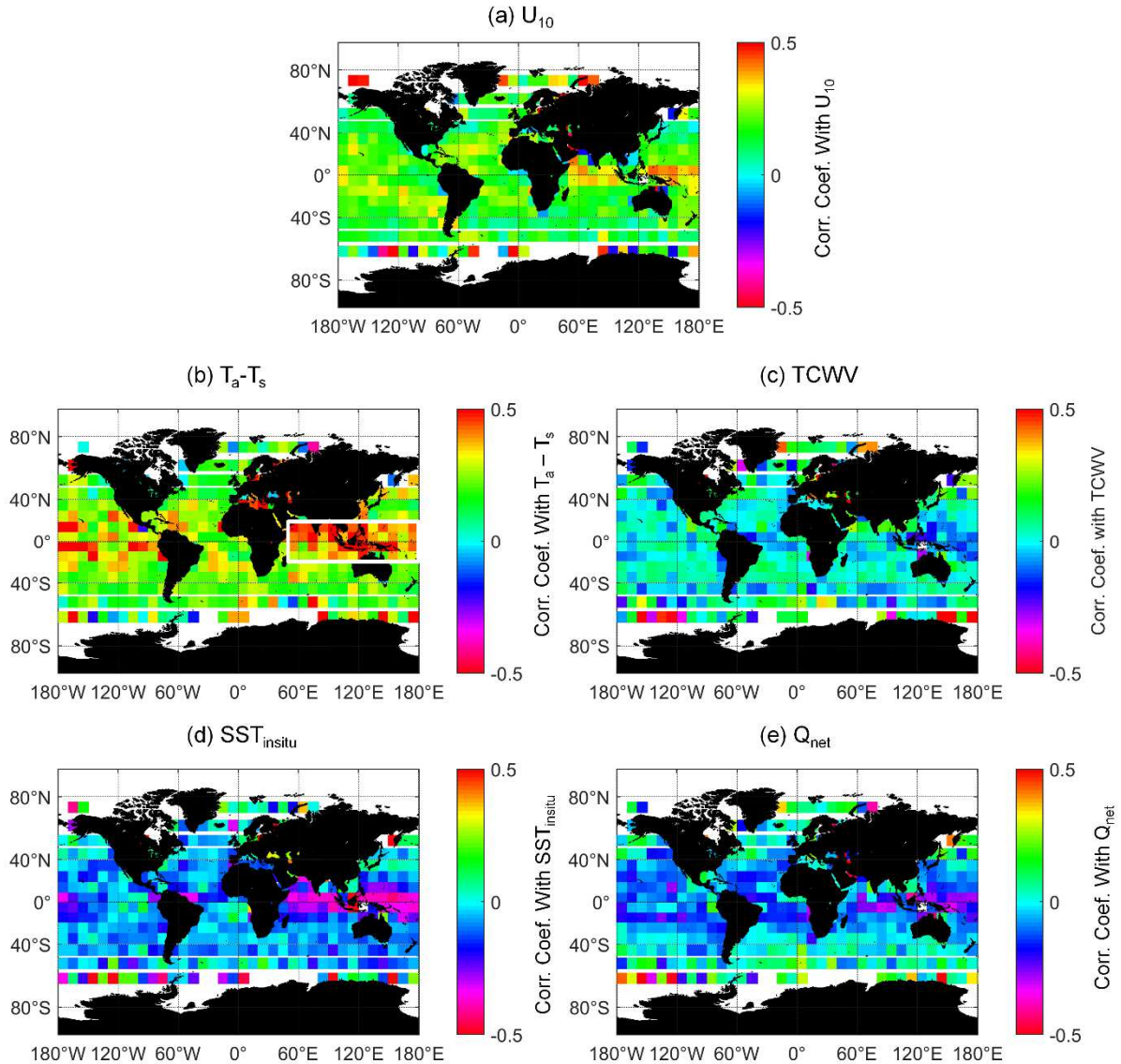
399 (1) Over the tropical areas, namely to the west of the Middle America, Gulf of Guinea, and
400 the TWP region. These areas have nearly year-round low winds ($< 5 \text{ m s}^{-1}$), largely
401 responsible for a very cool skin.

402 (2) Over the west coast of Canada and the USA and to the west of Peru and Chile.
403 Contributing factors to the large ΔT include the low winds (Fig. 8b) and low TCWV (Fig. 8d).
404 The SST_{insitu} values are atypically low for their latitudes, due to the cold California and Peru
405 Currents, respectively, which may also explain the low TCWV here.

406 (3) Over the Mediterranean Sea. Nearly all factors support large ΔT events: low winds, large
407 negative air-sea temperature differences, and low TCWV.

408 (4) Over the high-latitude north Atlantic Oceans. Large ΔT s are observed due to the
409 significantly cooler and dry air, even though U_{10} is relatively large.

410 The ΔT sizes are more irregular in the southern high latitudes, mainly due to the severe
411 sparseness of collocations (see also Fig. 2a).



412

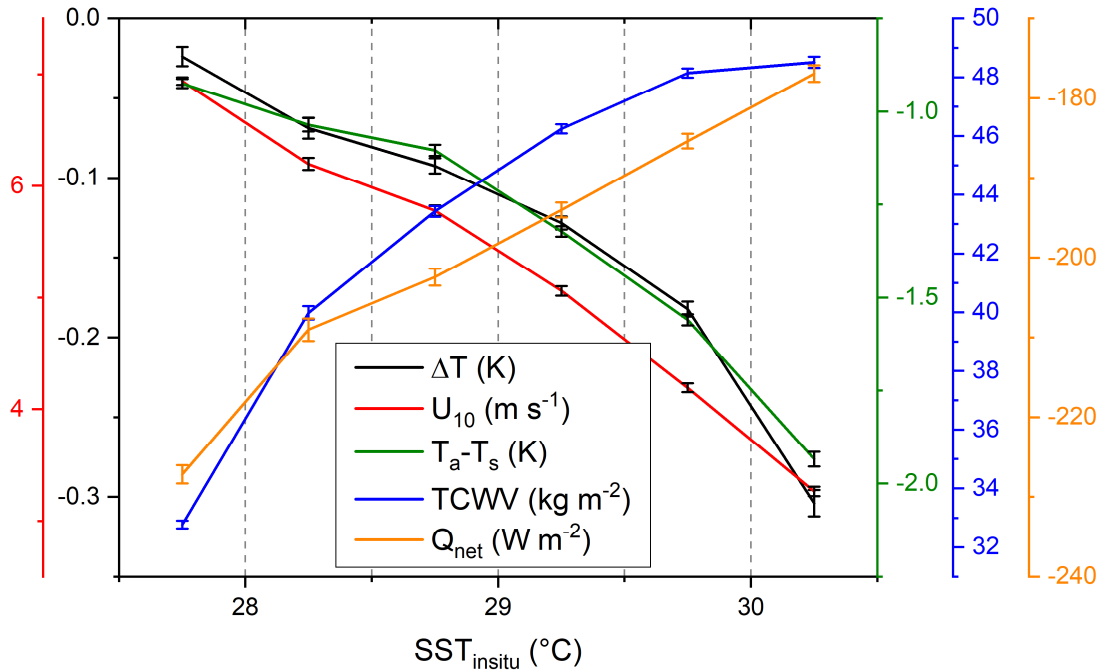
413 **Fig. 9.** Correlation coefficients between ΔT and (a) U_{10} ; (b) $T_a - T_s$; (c) TCWV; (d) SST_{insitu} ; and (e)
 414 Q_{net} . The white box in panel (b) is within (50°E-180°, 20°S-20°N). See text for further illustration.

415 Further to Fig. 8, we calculated the correlation coefficients (R_s) between ΔT and the
 416 environmental variables (Fig. 9). Overall, correlation coefficients are mostly mild (between -
 417 0.5 and 0.5) for all the variables, due to the complicated interactions and noise in the data.
 418 Positive correlations between ΔT and U_{10} and $T_a - T_s$ are observed over most of the global
 419 oceans, as expected (Figs. 9a and b; when interpreting the correlations, please keep in mind
 420 that the majority of ΔT values are negative). Particularly large positive R_s (~ 0.5) are found

421 over the tropics, including the TWP region, tropical Indian Oceans, and the waters adjacent to
422 the Middle America (Figs. 9a and 9b). The Rs between TCWV and ΔT are mixed of near-
423 zero positive and negative values, indicating TCWV's relatively weak influence (see also Fig.
424 5), except over the tropical Indian Ocean and the TWP region, where large negative Rs are
425 found (Fig. 9c). The dependency of ΔT on SST_{insitu} is also weak, and not well defined over
426 most of the oceans (Fig. 9d). However, large negative Rs are seen over the tropical Indian
427 Ocean and TWP domain, consistent with the sharp drop of ΔT for $SST_{insitu} > 28$ °C in Fig. 6a.
428 Over most of the low-to-mid latitude oceans, especially over the tropical Indian Ocean and
429 TWP domain, Q_{net} has a negative effect on the ΔT development, confirming the finding in
430 Fig. 7a.

431 The above results have highlighted the tropical Indian Ocean and the TWP domains, which
432 motivates us to conduct a further regional analysis. The region selected is 50°E to 180°, 20°S
433 to 20°N (see the white box in Fig. 9b). The change of ΔT with U_{10} , $T_a - T_s$, TCWV, and Q_{net} , is
434 now stratified by SST_{insitu} . Focusing only on the $SST_{insitu} > 28$ °C conditions, we divided the
435 SST_{insitu} values into six bands: < 28 °C, every half a degree from 28 °C to 30 °C, and > 30 °C.
436 The results are displayed in Fig. 10. The size of ΔT quickly increases as SST_{insitu} warms up
437 from 28 °C to 30 °C, in line with Fig. 6a. Wind speed drops from around 6.9 m s⁻¹ for
438 $SST_{insitu} < 28$ °C to ~ 3.3 m s⁻¹ when SST_{insitu} is > 30 °C, and $T_a - T_s$ reduces from -0.9 K to -
439 1.9 K, both contributing to a cooler skin. The negative correlation between ΔT and TCWV
440 and Q_{net} is also robust, consistent with Figs. 9c and 9e. Although the impact of TCWV on the
441 cool skin size does exist, yet it is minor compared to the effects of calmer winds and larger
442 $T_a - T_s$. In addition, such warm waters and calm winds, together with the usually high solar
443 insolation in this region, are manifestation of strong diurnal warming events. However, the
444 residual diurnal warming at ~ 10 pm local time, if existing, would have resulted in a warmer

445 skin (compared to the water at drifting or moored buoy depths, 0.2-1 m), rather than
 446 contributing to a cooler skin.



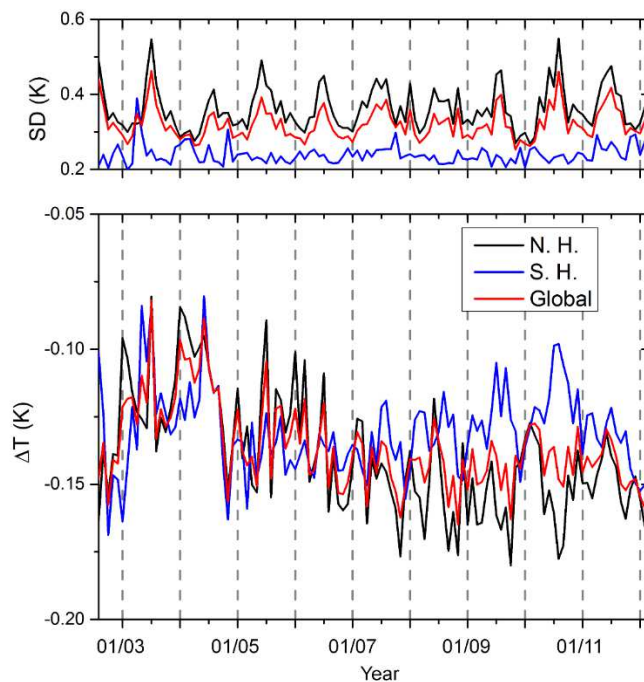
447

448 **Fig. 10.** Change of ΔT (black), U_{10} (red), T_a-T_s (green), TCWV (blue), and Q_{net} (brown), against
 449 $SST_{in situ}$ over the tropical Indian Ocean and the TWP domain (50°E-180°, 20°S-20°N; see the white
 450 box in Fig. 9b). $SST_{in situ}$ are divided into six bands: < 28 °C, every half a degree from 28 °C to 30 °C,
 451 and > 30 °C. The error bar indicates the 95% confidence level MoE.

452 *4.4. Seasonal patterns of ΔT*

453 The time series of monthly ΔT and SDs are shown in Fig. 11, both globally and separately for
 454 two hemispheres. Over the years, global average ΔT basically stays stable between -0.10 to -
 455 0.17 K with variations. The stronger fluctuations in early years are probably due to fewer
 456 collocations. Global ΔT s largely follow those of the northern hemisphere, due to its larger
 457 collocation contribution (Fig. 2b). A seasonal pattern can be easily recognized for most of the
 458 years for the southern hemisphere, with larger ΔT sizes in austral summer (and smaller in

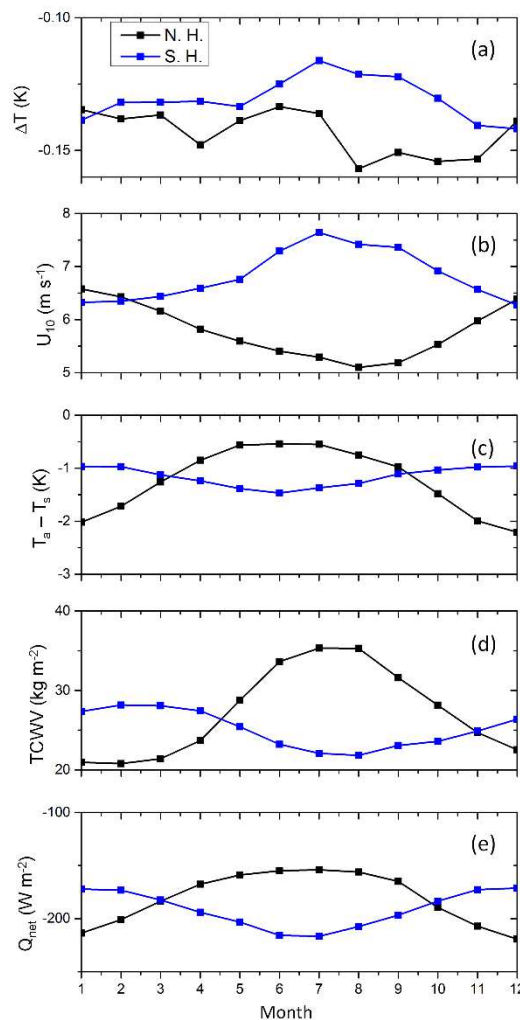
459 austral winter). The seasonal pattern in the northern hemisphere is less regular (except for
 460 year 2010). The SDs also keep steady over the years, but display a clear seasonal pattern,
 461 especially in the northern hemisphere with larger SDs in summer (~ 0.4 K) and lower in
 462 winter (~ 0.3 K). The SDs in the southern hemisphere are much lower and more consistent
 463 being around 0.2-0.3 K and shows no clear seasonal pattern. It is interesting to note smaller
 464 SDs in the southern hemisphere, presumably due (at least partly) to the fewer *in situ*
 465 platforms, and collocations.



466
 467 **Fig. 11.** Monthly time series of ΔT (bottom panel) and SDs (top panel) for the northern hemisphere
 468 (black line), southern hemisphere (blue line), and global coverage (red line).

469 Monthly mean, averaged over the whole study period, ΔT and environmental variables for
 470 both hemispheres are shown in Fig. 12. The seasonal pattern is more regular in the southern
 471 hemisphere with minimum ΔT size (~ -0.11 K) in austral winter and maximum (~ -0.14 K) in
 472 austral summer (Fig. 12a). The ΔT sizes in the northern hemisphere are larger in almost all
 473 the months, reaching ~ -0.16 K in August and -0.13 K in winter. Seasonal pattern of ΔT in

474 the northern hemisphere is less regular (Fig. 12a). The meteorological variables show strong
 475 seasonal pattern in both hemispheres. The wind speeds in the northern hemisphere are smaller
 476 than those in the southern hemisphere for almost all months, largely accounting for the
 477 former's larger ΔT sizes (Fig. 12b). The minimum $T_a - T_s$ in northern winter is lower (-2.2 K)
 478 than its counterpart in austral winter (-1.5 K), which also contributes to the larger ΔT in the
 479 northern hemisphere. In addition, the TCWV in the northern summer is higher than the
 480 maximum TCWV in the austral summer ($\sim 36 \text{ kg m}^{-2}$ compared to $\sim 29 \text{ kg m}^{-2}$), which may
 481 contribute to the irregularity of the seasonal pattern in the northern hemisphere.



482

483 **Fig. 12.** Monthly mean (a) ΔT ; (b) U_{10} ; (c) $T_a - T_s$; (d) TCWV; and (e) Q_{net} for northern (black line) and
 484 southern (blue line) hemisphere.

485 **5. Discussion**

486 The AATSR dual-view three channel (D3) skin SST, in conjunction with high-quality *in situ*
487 SSTs from the NOAA *iQuam* system, are employed here for cool skin analyses, thanks to its
488 high quality (attributable to innovative sensor features, such as the dual-view geometry and
489 stable calibration), and independency from *in situ* measurements (due to the physics-based
490 retrieval algorithm). Several prior studies have touched on the cool skin effect using ATSRs
491 data. For example, Murray et al. (2000) compared four-year ATSR SST_{skin} data with
492 coincident bulk SSTs and found a night (~ 20.30 pm local time) cool skin amplitude of ~ -
493 0.20 K, slightly larger than in this study. Horrocks et al. (2003) calculated nighttime cool skin
494 as the difference between the ATSR-2 and *in situ* SSTs, to test the performance of a skin
495 model. Only the cool skin change against wind speed was shown, and their corresponding D3
496 figure matches our result very well (cf. Fig. 12a in Horrocks et al., 2003). The general
497 statistics shown in section 4.1 of this study, including the mean difference and low RSD
498 between SST_{skin} and SST_{insitu} (Fig. 1a), and the spread pattern of ΔT (Fig. 1b), together with
499 all the following analyses, consistently suggest that the AATSR D3 SST is well suited for
500 skin SST studies. However, the temporal and spatial differences between satellite retrievals
501 and *in situ* observations may be comparable to, or even larger than, those between SST_{skin} and
502 SST_{depth}, which is the focus of this study. These factors, at least in part, may be responsible
503 for the differences between results of this study, and the parameterizations (D02, M11, and
504 A17 models) obtained from the collocated skin and depth data onboard the same ship.

505 The effect of dust on IR SST retrievals was minimized by applying the QL = 5 filter in
506 AATSR data, which restricts the ATSRs Saharan Dust Index (ASDI) to a range of -0.2 to 0.2.
507 ASDI is a new SDI proposed especially for the ATSRs, and it is available in each pixel.
508 Although scaled to produce values comparable with visible aerosol optical depth (AOD), the
509 ASDI differs from the AOD in that it may go negative, with values $> \sim 0.2$ presumably

510 indicating the presence of dust (Good et al., 2012). Stricter ASDI filters, such as retaining
511 data with ASDI from -0.15 to 0.15, have been also tested but no noticeable changes in the
512 results were observed (not shown). Xu and Ignatov (2016) observed large negative
513 differences between nighttime AATSR and *in situ* measurements over the regions such as to
514 the west of tropical Africa and northern Indian Ocean. They argued that this may be due to
515 the Saharan dust outbreaks or Indian aerosol over the Arabian Sea. Note that no dust filter
516 was applied in their work, and findings in this paper should be much less sensitive to the
517 effects of aerosols. For example, large cool skin amplitudes were observed in this study in the
518 Gulf of Guinea, which is just slightly to the south of the areas under strong Saharan dust
519 effects. Also, calm winds typically appear in such areas, which should be the major reason for
520 the cool skin signals. In addition, Good et al. (2012) and Noyes et al. (2006) both showed
521 evidence, suggesting that AATSR D3 data are robust to dust and aerosol effects; in fact, the
522 dual view retrievals can be even slightly positively biased in the presence of Saharan dust.
523 This gives us more confidence in the results obtained.

524 With regards to warm skins ($\Delta T > 0$ K), Minnett et al. (2011) argued that they may be
525 “merely an artefact of the depth at which the bulk temperature measurements are taken”.
526 However, in this study, at least a large portion of the warm skin signals appears real. Warm
527 skins account for 25.6% of all ΔT values, part of which may well correspond to the 14.0%
528 case with positive $T_a - T_s$ conditions. Considering a possible small warm bias in AATSR data,
529 the residual diurnal warming effect (especially when U_{10} is $< 2 \text{ m s}^{-1}$), and noises in both
530 satellite and *in situ* measurements, the true percentage may very likely be $< 25.6\%$. In the
531 daytime, when solar insolation is present (not analysed in this study), it is expected that warm
532 skin events may occur more frequently than at night, as the skin layer can absorb part of the
533 solar insolation (e.g., Saunders, 1967; Fairall et al., 1996), and higher T_a may lead more

534 positive $T_a - T_s$ conditions. Nevertheless, more observations are required to validate this
535 assumption, which may be subject of future work.

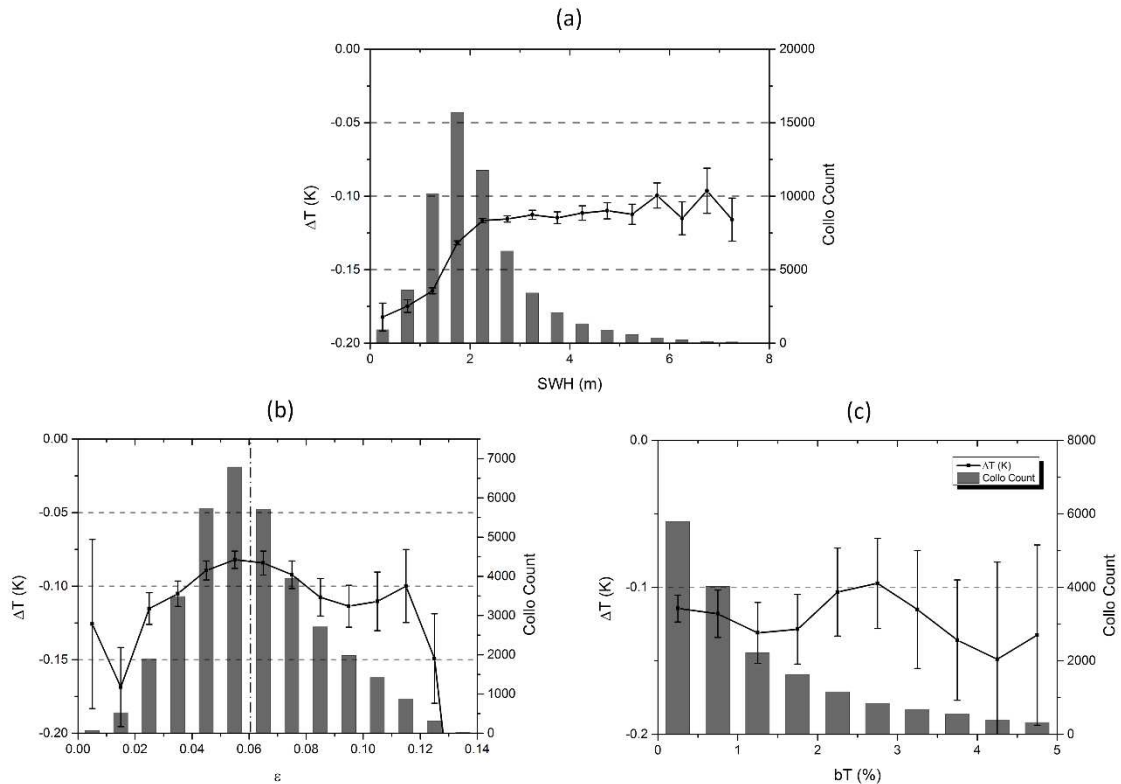
536 According to several previous cool skin models (e.g., Saunders, 1967), higher winds tend to
537 result in much thinner cool skin layer depths, and hence in smaller ΔT amplitudes, regardless
538 of stronger evaporation and heat loss. Since wind is closely coupled with waves and turbulent
539 mixing coupled with wave breaking, one is naturally curious to explore possible links
540 between cool skin and wave variables. As a preliminary check, we used one-year (2011)
541 global wave data produced by running the third-generation spectral wave model
542 WAVEWATCH III (hereafter WW3; WAVE-WATCH III Development Group, 2016). The
543 observation-based physics of WW3 model are described in Rogers et al. (2012), Zieger et al.
544 (2015), and Liu et al. (2019). The wave parameters selected in our analysis include: (1)
545 significant wave height of the full wave spectrum, H_s ; (2) significant wave height and peak
546 wave length of the wind-sea partition, $H_{s,w}$ and $L_{p,w}$, from which wave steepness $\epsilon =$
547 $\frac{1}{2}H_{s,w} \times 2\pi/L_{p,w}$; and (3) dominant wave breaking probability b_T . Among them, the H_s , $H_{s,w}$,
548 and $L_{p,w}$ are formal output parameters of WW3. We calculated b_T based on the
549 parameterization proposed in Babanin et al. (2001), given as:

550
$$b_T = 85.1[(\epsilon_p - 0.055)(1 + H_{s,w}/d)]^{2.33}$$

551 where d is the water depth, $\epsilon_p = 2k_p[\int_{0.7f_p}^{1.3f_p} F_w(f)df]^{1/2}$ is the significant steepness of the
552 spectral peak, $F_w(f)$ is the frequency spectrum of the wind sea after filtration of swell, k_p
553 and f_p are the peak wavenumber and frequency of $F_w(f)$, respectively. The reason why we
554 are running WW3 offline, instead of using wave parameters available in the ERAI reanalysis
555 dataset, is that the unique b_T parameter, estimated from the full wave spectra, is not included
556 in the ERAI. Moreover, our parameterizations for the spectral wave model includes a wave

557 breaking term (which incorporates both the inherent (saturation-based) breaking and
558 cumulative breaking mechanisms; Babanin et al., 2010), a separate swell dissipation term
559 based on the wave-induced turbulence theory (Babanin, 2011), as well as a nonlinear wind
560 input term formularized from field observations (Donelan et al., 2006). Overall, this updated
561 formulation is deemed to better capture the physical mechanisms involved. Wave parameters
562 (e.g., H_s) based on this set of wave physics, have proved more accurate than the ERAI wave
563 product (e.g., Ardhuin et al., 2010; Liu et al., 2019).

564 The results are shown in Fig. 13. Expectedly, ΔT amplitude decreases as H_s increases, since
565 the H_s is typically an exponential function of wind speed, which also explains the climbing
566 pattern of ΔT when H_s grows from 0 to 3 m (Fig. 13a). The function of ϵ is similar to that of
567 U_{10} , although the trend is less regular due probably to a small collocation number and the
568 further dependency of ϵ on wave lengths (Fig. 13b). Finally, b_T is a direct function of wave
569 steepness (according to Babanin et al. (2011), the threshold of ϵ above which waves start to
570 break is $0.055/0.9 = 0.061$, as indicated in Fig. 13b) in the open deep oceans, rather than of
571 U_{10} . Breaking waves only account for a very small proportion of all waves. Of all b_T values,
572 84.7% are zero and discarded. Only the positive b_T values (15.3%) are retained. Fig. 13c
573 indicates that there is no clear dependency of ΔT on the b_T , which is consistent with the trend
574 to the right of the $\epsilon = 0.061$ threshold line in Fig. 13b. Since ΔT is already closely
575 approaching the asymptotic value (~ -0.09 K to -0.12 K) when ϵ is > 0.061 , i.e. waves start to
576 break, there is not much room left for the ΔT to change due to different b_T conditions. In the
577 future, with more data available, further analyses such as the one considered here, could be of
578 use to both SST and wave communities in terms of better understanding the skin effect, and
579 further understand its link with the upper mixing investigation.



580

581 **Fig. 13.** The dependency of ΔT on (a) SWH, H_s , (b) wave steepness, ϵ , and (c) wave breaking
 582 probability (b_T in %). In panel (b), the wave breaking threshold of $\epsilon = 0.061$ is indicated. In panel (c),
 583 only $b_T > 0\%$ (accounting for 15.3% of all values) are retained. In all panels, black line is ΔT and
 584 error bar is the 95% confidence level MoE. Note all Y-axis ranges are only -0.2 K to 0 K.

585 6. Conclusions

586 Cool skin signals revealed from nighttime AATSR skin SST data and NOAA *iQuam in situ*
 587 SST measurements (defined as $\Delta T = SST_{\text{skin}} - SST_{\text{insitu}}$), have been described in detail on a
 588 global scale for nearly ten years from July 2002 to April 2012. Foci are on skin effects'
 589 overall statistics, dependencies on different meteorological variables, spatial distribution, and
 590 seasonal patterns. So far, the cool skin has not been systematically analysed on a global scale
 591 and over a long period, using a combination of high-quality satellite and *in situ* SST data.
 592 Traditional works in this area usually take advantage of the high accuracy shipborne

593 radiometer SST_{skin} data and coincident SST_{depth} , which may have more accurate data but are
594 inevitably limited to the ship's routes and duration.

595 In terms of dependencies of ΔT on environmental variables, wind speed is confirmed to play
596 the most important role in the cool skin development, consistent with the findings in previous
597 studies. The effect of $T_a - T_s$ on the ΔT s is also significant. When the air is warmer than the sea,
598 there can be warm skins. A small(er) variation range of $T_a - T_s$ may be why its effect on ΔT
599 appears secondary to that of U_{10} . The dependency of ΔT on TCWV is relatively weak. No
600 clear dependency of ΔT on SST_{insitu} is observed for $SST_{\text{insitu}} < 28$ °C, above which, however,
601 a dramatic increase of ΔT size is observed. The (partial) effect Q_{net} on ΔT turns out to be mild,
602 since Q_{net} is largely coupled with U_{10} . A strong (moderate) dependency of ΔT on wave height
603 (wave steepness) is identified in this study, while the correlation between ΔT and wave
604 breaking probability is less discernible.

605 Spatially, large ΔT is normally associated with low U_{10} , yet other meteorological variables
606 may or may not contribute. Typically, a combination of weak winds with large negative $T_a - T_s$
607 and dry air profile, can lead to very cool skin, for instance over the Mediterranean Sea. The
608 effects of different variables can sometimes offset (or amplify) each other. The spatial
609 distributions of the correlation coefficients between ΔT and environmental variables further
610 illustrate the more important effects of U_{10} and $T_a - T_s$. The tropical waters such as the tropical
611 Indian Ocean and the TWP region have stood out. The cool skin amplitude gets significantly
612 larger when SST_{insitu} increases from 28 °C to > 30 °C, especially when observed in
613 conjunction with a sharp drop in U_{10} and $T_a - T_s$. The near 1 sensitivity of AATSR D3 data to
614 actual SST and the data's near-independency on TCWV over this region, add to our
615 confidence in the results.

616 The seasonal pattern of ΔT is more identifiable in the southern hemisphere with larger
617 (smaller) average ΔT s in austral summer (winter), which is basically controlled by the wind
618 speed pattern. In the northern hemisphere, ΔT s are larger than those in the southern
619 hemisphere in all months, due to weaker winds. Also, ΔT in the northern hemisphere has a
620 less regular seasonal pattern.

621 The spatial distribution and seasonal patterns of the cool skin can be useful in cool skin
622 modelling as they may be part of the reason why many cool skin models, developed from
623 seasonally and spatially different experiments, behave inconsistently with each other and
624 some may seem less satisfying when intercompared together (e.g., Kent et al., 1996; Castro et
625 al., 2003; Horrocks et al., 2003; Tu and Tsuang, 2005).

626 In the future, it may be worthwhile to explore combined ARC AATSR with microwave (MW)
627 subskin SSTs (instead of *in situ* SSTs), to see if similar results can be obtained. Significantly
628 larger collocation counts can be expected from two satellite products, to achieve improved
629 statistical representativity and robustness. However, that would require highly accurate MW
630 SSTs, which presently are challenging to obtain. For instance, Castro et al. (2008) tried to
631 extract cool skin effect from AVHRR (Advanced Very High Resolution Radiometer) IR skin
632 SST produced by Pathfinder project, and Tropical Rainfall Mapping Mission (TRMM)
633 Microwave Imager (TMI) MW subskin SST retrievals. Although the dependency of the IR-
634 MW differences on wind speed they obtained was consistent with the cool skin effect feature,
635 they concluded that physical analyses of the skin layer process were strongly obscured by
636 retrieval errors in both AVHRR IR and TMI MW SSTs. However, with AATSR SST_{skin} data
637 and possibly more sophisticated MW sensors and/or retrieval algorithms (e.g., Nielsen-
638 Englyst et al., 2018), it may be worth revisiting.

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884

885 Figure captions:

886 **Fig. 1.** (a) Density of collocations as a function of SST_{insitu} , with overall statistics of ΔT ($= SST_{skin} -$
887 SST_{insitu}) superimposed. (b) Frequency distribution of ΔT s.

888 **Fig. 2.** (a) Spatial distribution of the collocations. Bin size is $10^\circ \times 10^\circ$. (b) Monthly collocation counts
889 in the northern (black bars) and southern (blue bars) hemispheres.

890 **Fig. 3.** Solid circles/Black line: dependencies of ΔT on U_{10} . The error bar is the 95% confidence level
891 margin of error (MoE), i.e. 1.96 times the SD divided by the square root of the collocation number.
892 The column bars indicate the collocation counts falling within each $1 \text{ m s}^{-1} U_{10}$ interval. The red, blue,
893 and green lines have been calculated using the Donlon et al. (2002), Minnett et al. (2011), and
894 Alappattu et al., (2017) empirical parameterizations, respectively.

895 **Fig. 4.** (a) Dependency of ΔT on $T_a - T_s$ (with the error bars at 95% confidence level margin of error,
896 MoE). (b) Dependency of ΔT on $T_a - T_s$ (binned into 0.5 K) and U_{10} (binned into 1 m s^{-1}). Black
897 contour lines indicate the corresponding collocation counts in each $1 \text{ m s}^{-1} \times 0.5 \text{ K}$ bin. Only bins with
898 collocation counts ≥ 20 are plotted.

899 **Fig. 5.** Same as in Fig. 4 but for the dependency of ΔT on TCWV (binned into 5 kg m^{-2}).

900 **Fig. 6.** Same as Fig. 4 but for the dependency of ΔT on SST_{insitu} (binned into 1°C).

901 **Fig. 7.** Same as Fig. 4 but for the dependency of ΔT on Q_{net} (binned into 50 W m^{-2}).

902 **Fig. 8.** Spatial distributions of (a) ΔT ; (b) U_{10} ; (c) $T_a - T_s$; (d) TCWV; (e) SST_{insitu} ; and (f) Q_{net} averaged
903 over the whole study period. Box size is 10° by 10° .

904 **Fig. 9.** Correlation coefficients between ΔT and (a) U_{10} ; (b) $T_a - T_s$; (c) TCWV; (d) SST_{insitu} ; and (e)
905 Q_{net} . The white box in panel (b) is within ($50^\circ \text{E} - 180^\circ$, $20^\circ \text{S} - 20^\circ \text{N}$). See text for further illustration.

906 **Fig. 10.** Change of ΔT (black), U_{10} (red), $T_a - T_s$ (green), TCWV (blue), and Q_{net} (brown), against
907 SST_{insitu} over the tropical Indian Ocean and the TWP domain ($50^\circ \text{E} - 180^\circ$, $20^\circ \text{S} - 20^\circ \text{N}$; see the white

908 box in Fig. 9b). SST_{insitu} are divided into six bands: < 28 °C, every half a degree from 28 °C to 30 °C,
909 and > 30 °C. The error bar indicates the 95% confidence level MoE.

910 **Fig. 11.** Monthly time series of ΔT (bottom panel) and SDs (top panel) for the northern hemisphere
911 (black line), southern hemisphere (blue line), and global coverage (red line).

912 **Fig. 12.** Monthly mean (a) ΔT ; (b) U_{10} ; (c) $T_a - T_s$; (d) TCWV; and (e) Q_{net} for northern (black line) and
913 southern (blue line) hemisphere.

914 **Fig. 13.** The dependency of ΔT on (a) SWH, H_s , (b) wave steepness, ϵ , and (c) wave breaking
915 probability (b_T in %). In panel (b), the wave breaking threshold of $\epsilon = 0.061$ is indicated. In panel (c),
916 only $b_T > 0\%$ (accounting for 15.3% of all values) are retained. In all panels, black line is ΔT and
917 error bar is the 95% confidence level MoE. Note all Y-axis ranges are only -0.2 K to 0 K.

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