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A Mechanism for the Skew of Ensemble Forecasts

Cécile Penland¹ and Prashant D. Sardeshmukh²

^{1,2}*NOAA/ESRL/Physical Sciences Laboratory, Boulder, CO*

²*Cooperative Institute for Research in Environmental Sciences,
University of Colorado, Boulder, Colorado.*

¹cecile.penland@noaa.gov

²prashant.d.sardeshmukh@noaa.gov

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23 **Abstract**

24

25 It is often not appreciated that forecast ensembles are generally skewed. The skew can
26 arise from the state-dependence of the chaotic system dynamics responsible for the ensemble
27 spread. Generation of skew by this mechanism can be demonstrated in even the simplest dynamical
28 system with state-dependent noise, and even when the initial and the asymptotic (that is, the
29 “climatological”) forecast distributions are both symmetric. Indeed, forecast distributions of
30 systems with state-dependent noise in the dynamical tendencies must in general be both skewed
31 and heavy-tailed, with implications for forecasting extreme anomaly risks. Ensemble forecast
32 systems that misrepresent such state-dependent noise have state-dependent errors in their forecast
33 probability distributions. Because such errors depend on both the initial condition and forecast lead
34 time, they cannot be removed by simple *a posteriori* bias-corrections of the forecast distributions.

35 The ensemble standard deviation is often used as a simple metric of ensemble spread even
36 when the forecast distribution is not Gaussian. In a similar spirit, the ensemble skew S may be used
37 as a simple metric of the difference D between the ensemble-mean and most likely forecast as well
38 as the risk ratio R of extreme positive and negative deviations from the ensemble-mean forecast.
39 This is motivated by the facts that 1) the probability distributions of many geophysical quantities
40 are approximately Stochastically Generated Skewed (SGS) distributions, for which simple
41 analytical relationships exist between these quantities, and 2) Gaussian distributions are a sub-
42 class of SGS distributions. However, S may serve as a useful metric of R and D even when the
43 distributions are not strictly SGS distributions.

44

45

46 **1. Introduction**

47 A basic feature of weather and climate predictions is their sensitivity to small initial errors.
48 The chaotic nature of earth system dynamics renders accurate deterministic predictions impossible
49 even if the initial errors are tiny.

50 Operational weather forecasting centers commonly account for forecast uncertainty by
51 providing an ensemble of forecasts generated from an ensemble of slightly different initial
52 conditions, and more recently at some centers, also by stochastically perturbing the forecast
53 model’s tendencies at each model time step. It is clear that both initial uncertainties and model
54 uncertainties contribute to forecast uncertainty, and both need to be accounted for to give the best
55 probabilistic forecasts.

56 Current ensemble forecasting methods yield a crude probability distribution of possible
57 future states, with the ensemble-mean forecast often interpreted as the most likely future state, and
58 the ensemble spread as the forecast uncertainty. For the ensemble to be “reliable” in a probabilistic
59 sense, the actual future state must generally occur within the ensemble spread, that is, the actual
60 forecast uncertainty must be consistent with the expected forecast uncertainty. One measure of this
61 reliability is the consistency between the root-mean-square error of the ensemble-mean forecast
62 determined over many forecast cases and the root-mean-square of the ensemble spread, also
63 determined over many cases. A longstanding problem with most ensemble forecasting systems is
64 that the ensemble spread is generally smaller than the error of the ensemble-mean forecast, that is,
65 the expected forecast uncertainty is too small. This has consequences, especially for predicting
66 extreme event risks.

67 There are large ongoing efforts in the numerical modeling community to reconcile the
68 forecast error and spread, both by reducing the error and increasing the spread (e.g., Bechtold et al

69 2008, Leutbecher et al 2017). Efforts to reduce forecast error generally focus on improving the
70 representation of physical processes and increasing the spatial resolution of the forecast model.
71 Efforts to increase forecast spread have traditionally focused on using an ensemble of perturbed
72 initial conditions centered on a more accurately estimated “control” initial condition with special
73 perturbations that grow most rapidly, obtained as either singular vectors or “bred” vectors of the
74 perturbation evolution operator (Molteni et al 1996; Toth and Kalnay 1997). An alternative
75 approach is to use an ensemble of initial states consistent with observational uncertainty, generated
76 using an ensemble data-assimilation algorithm such as an Ensemble Kalman Filter (Evensen
77 1994). Neither approach has proven effective in eliminating the gap between the forecast error and
78 spread growth with forecast lead time. More recently, some promising results have been obtained
79 by stochastically perturbing the physics and dynamics of the forecast model itself throughout the
80 forecast (Palmer et al 2009, Berner et al 2009, Leutbecher et al 2017, Palmer 2019). Forecast
81 improvements due to such “stochastic parameterizations” are also generally consistent with those
82 obtained by increasing model resolution (Berner et al. 2012, 2017).

83 Stochastic perturbations are usually justified as accounting for chaotic feedbacks from
84 unresolved model physics and dynamics that are not represented by deterministic
85 parameterizations. Nonlinear interactions, including “deterministic” interactions, that occur on
86 multiple timescales are often described well by a stochastic dynamic “forcing” even when they are
87 generated internally. For example, individual molecular collisions act as state-dependent noise in
88 macroscopic fluid dynamics (Landau and Lifshitz 1959; see García and Penland 1991 for
89 numerical verification). The existence of chaotic interactions at all scales in the geophysical system
90 ensures that stochasticity from various sources is an important contributor to variability in weather
91 and climate. The rich literature showing this to be true has been explored in texts edited by, e.g.,

92 Imkeller and Von Storch (2001) and Palmer and Williams (2010), as well as in classic studies such
93 as that of Hasselmann (1976). The discussion in Gottwald (2021) is particularly interesting for its
94 practical advice on how to ensure consistency with pertinent limit theorems in implementing
95 dynamically-based stochastic parameterizations.

96 Accurate forecasting requires accurate representation of both the resolved aspects of the
97 physical system and the interactions between resolved and unresolved processes. Simply throwing
98 in random numbers to increase the ensemble spread with no regard for the particular physical
99 processes they represent must have limited advantage, although even this is apparently effective
100 in improving estimates of forecast uncertainty (Buizza et al 1999).

101 Meaningful representation of stochastic effects in a multiscale system must also account
102 for the degree of timescale separation between resolved and unresolved processes. If their
103 separation, defined as the ratio of their autocorrelation timescales, is “infinite”, the stochastic
104 effects may be represented rigorously as a forcing of the resolved scales by a Gaussian white noise
105 forcing with vanishing temporal autocorrelation scales (e.g., Khas’minskii 1966; Papanicolaou and
106 Kohler 1974; Hasselmann 1976). If the separation is large but finite, one may employ a noise
107 forcing with “memory”, such as a Gaussian Ornstein-Uhlenbeck (red noise) forcing with small but
108 finite temporal autocorrelation scales. If the separation does not allow for a complete decoupling
109 of resolved and unresolved timescales, such as when the amplitude of the unresolved processes
110 depends on the amplitude of the resolved processes, one may represent the unresolved chaotic
111 feedbacks on the resolved scales as a state-dependent stochastic forcing. In the simplest case of
112 an energy-conserving system with slow (“resolved”) and fast (“unresolved”) system components,
113 one can rigorously show that such a stochastic coupling yields a *Stochastically Generated Skewed*
114 (SGS) probability distribution of the slow variables (Sardeshmukh and Penland 2015), which in

115 the limit of zero coupling reduces to a Gaussian distribution. Unfortunately, this is where
116 analytical tractability of a smooth transition from a white noise stochastic forcing to a more
117 complex stochastic forcing ends. At the next level of complexity, the chaotic nonlinear resolved-
118 unresolved interactions generally cannot be treated as a linear stochastic process, and the particular
119 forms of the nonlinear terms in a forecast model's equations become important.

120 Our purpose here is to highlight the fact that forecast ensembles in systems with state-
121 dependent noise are generally skewed, and that the ensemble standard deviation alone does not
122 provide a reliable quantification of forecast uncertainty. The skew strongly affects forecasts of
123 extreme anomaly risks, but is rarely discussed in the forecasting literature (see e.g., overview
124 articles by Leutbecher and Palmer 2008, Palmer 2012, Gneiting and Katzfuss 2014, Leutbecher et
125 al 2017). We stress that the skew due to state-dependent noise is a fundamental property of multi-
126 scale systems with slow and fast components (Sardeshmukh and Penland 2015), and should be
127 distinguished from the skew generated in forecast ensembles by errors in initial conditions and in
128 the deterministic forecast dynamics, such as errors in propagating waves and fronts (e.g. Miller
129 and Ehret 2002, Miyazawa et al 2005, Hodyss and Reineke 2013, Schulte and Georgas 2018).

130 We investigate the conditional (that is, the forecast ensemble) skew rigorously in the
131 simplest dynamical system with state-dependent noise, specifically a linear scalar system with
132 SGS dynamics, assuming that its sensitivity to initial conditions and forecast lead times will also
133 be illustrative of more complex systems. We do of course recognize that skewness in forecast
134 ensembles can also be generated by deterministic nonlinear processes, not just by state-dependent
135 stochastic processes, and errors in modeling deterministic nonlinear processes as well as
136 observational errors might present additional sources of skew. However, even eliminating such
137 errors would not eliminate the skew in forecast ensembles generated by state-dependent stochastic

138 processes, and we focus on this unavoidable, dynamically-based skew in a simple setting. Further,
139 Sardeshmukh et al (2015) and Sardeshmukh and Sura (2009) provide strong evidence that state-
140 dependent stochastic noise accounts for some basic features of observed large-scale atmospheric
141 non-Gaussianity that are otherwise hard to explain. For example, Sardeshmukh et al. (2015, see
142 their Figs. 2 and 3) used the long-term Twentieth Century Reanalysis dataset to show that for
143 several important meteorological variables, the skew of their probability distributions, the
144 distinctive relationship between the skew and kurtosis, and the approximate equality of their
145 probability densities at standardized positive and negative anomaly magnitudes of $\sqrt{3}$ are all
146 consistent with state-dependent stochastic noise. SGS dynamics may therefore be considered an
147 important generation mechanism for atmospheric non-Gaussianity.

148 In general, the conditional skew S is associated with a difference D between the ensemble-
149 mean forecast (that is, the mean of the forecast distribution for any desired lead time) and the most
150 likely forecast (the mode of the distribution for that lead time). S is also associated with the risk
151 ratio R of extreme positive and negative deviations from the ensemble-mean forecast being
152 different from unity. Less intuitively but no less importantly, S can also cause the rank histograms
153 of ensemble forecasts determined over many forecast cases (sometimes referred to as Talagrand
154 diagrams), which should ideally be flat, to have a *symmetric* U shape. Such histograms are often
155 used to assess forecast reliability (but see Hamill 2001 for a critique). Their U-shape, an
156 undesirable feature of many ensemble forecast systems, is interpreted as a general underprediction
157 of extreme anomalies and is almost always ascribed to a deficiency in the ensemble spread, not in
158 the ensemble skew. The numerical experiments presented here provide a perspective on how large
159 such effects of S are likely to be in real forecasting contexts for them to matter.

160 A rigorous characterization of D and R is possible when their underlying stochastic
161 generator is known. Here we exploit the fact that the probability distributions of many large-scale
162 meteorological variables are approximately SGS distributions (e.g., Sardeshmukh et al. 2015;
163 Sardeshmukh and Sura 2009). Such distributions are generated by a combination of correlated
164 additive (that is, the state-independent) and multiplicative (that is, the linearly state-dependent)
165 noise forcing, often referred to as CAM-noise forcing. To distinguish the impact of the stochastic
166 forcing on the forecast spread from that of initial uncertainty, we focus on situations in which the
167 initial conditions are perfect, i.e., δ -functions. Note that the SGS process, being linear, may
168 superpose an ensemble of δ -function initial conditions to describe the results of the whole.

169 Note also that our focus here is on the skew of the forecast *pdf*, not on the nonlinear
170 evolution of the ensemble-mean forecast *per se*. The ensemble spread and skew are, almost by
171 definition, due to chaotic system dynamics. Our approach of approximating the chaotic dynamics
172 as state-dependent stochastic noise in the simplest model consistent with observed non-Gaussian
173 atmospheric statistics is consistent with the representation of nonlinear chaotic physics by “SPPT”
174 types of stochastic noise parameterizations employed in several state-of-the-art NWP models, and
175 indeed we have found in ongoing research (Sardeshmukh et al 2022, manuscript in preparation)
176 that even our scalar model is useful for understanding the impact of such stochastic
177 parameterizations on forecast skill in the NOAA/GFS model (Wang et al 2019).

178 We also recognize that there are important meteorological variables whose distributions
179 are not SGS distributions. Because SGS distributions fit the data much better than Gaussian
180 distributions (Sardeshmukh et al 2015), we submit that the skew S of a forecast ensemble is a
181 useful metric of both D and R even when the distributions are only approximately SGS
182 distributions, given the simple relationships between S , D , and R in CAM-noise driven systems.

183 We submit this noting that the standard deviation of a forecast ensemble is a useful metric of
184 forecast uncertainty even when the distribution is not Gaussian.

185 The article is arranged as follows. We review the SGS distribution in section 2,
186 summarizing the connection of its parent stochastic differential equation to a nonlinear
187 stochastically forced oscillator, recalling the properties of the unconditional (that is, the stationary)
188 distribution, and revisiting evidence that many important large-scale meteorological variables
189 obey the SGS distribution. In section 3, we note that the conditional (that is, the forecast)
190 probability distribution associated with the stochastic dynamical system obeys the same Fokker-
191 Planck equation obeyed by the stationary distribution, and use it to derive its first four moments
192 for a single initial condition. We numerically generate a large ensemble of realizations of the
193 dynamical system with this initial condition and different realizations of the stochastic noise during
194 system integration. We then repeat these numerical integrations with other initial conditions. The
195 ensemble members from all such integrations are then used to analyze the forecast probabilities as
196 a function of initial condition and forecast lead time. The results presented in sections 3 are based
197 on very large 50000-member ensemble integrations of this simple system, performed to check both
198 numerical accuracy and consistency with theoretical expectations. We also investigate the
199 sampling uncertainties that arise if much smaller but more realistic (100 or 200 member)
200 ensembles sizes are used. Section 4 provides a summary and concluding remarks.

201

202 **2. Review of SGS dynamics**

203 The SGS distribution was introduced (Sardeshmukh and Sura 2009) as a way to reconcile
204 the observed non-Gaussianity of weather and climate variations with their approximately linear
205 predictability and linear response to external forcing. The linear predictability is evident in many

206 contexts, such as the predictability of tropical SSTs (e.g., Penland and Sardeshmukh 1995,
 207 hereafter PS95; Newman and Sardeshmukh 2017). Using Linear Inverse Modeling (LIM) to
 208 estimate the SST evolution operator from observed lag-covariances, PS95 found that the results
 209 did not seem to depend on the training lag used for the estimation (see also Shin et al 2010). That
 210 is, their so-called “tau-test” for linearity was passed. Such a result indicates linear dynamics,
 211 almost by definition. However, although seasonally-averaged SSTs are close to Gaussian, monthly
 212 averaged SSTs are distributed with significant skew (e.g., Martinez-Villalobos et al 2019), which
 213 appears to require nonlinear dynamics. These apparently opposite conclusions regarding linearity
 214 can be reconciled if a linear process is driven by correlated additive and multiplicative stochastic
 215 noise (CAM noise), which occurs naturally in even very simple multivariate nonlinear systems
 216 with slow and fast variables such as a two-dimensional nonlinear oscillator (Sardeshmukh and
 217 Penland 2015). The resulting equation for the slow process $x(t)$, adjusted to have zero-mean, is

$$218 \quad \frac{dx}{dt} = Lx + (Ex + g)\xi_1 + b\xi_2 - \frac{1}{2}Eg. \quad (1)$$

219 In Eq. (1), ξ_1 and ξ_2 are independent Gaussian white noises (to be integrated in the sense of
 220 Stratonovich), while L , E , g , and b are constants. The stationary, that is the climatological,
 221 probability density function (*pdf*: $p(x)$) is evaluated from the stationary Fokker-Planck equation
 222 (see Appendix A) as

$$223 \quad p(x) = \frac{1}{N} [(Ex + g)^2 + b^2]^{-(\nu+1)} \exp \left[q \arctan \left(\frac{Ex+g}{b} \right) \right] \quad , \quad (2)$$

224 where $\nu = - [(L / E^2) + 1/2]$, $q = 2g\nu / b$, and N is the normalization constant. Note that ν is
 225 constrained to be strictly positive for $p(x)$ to exist, and a necessary condition for the n^{th} moment
 226 $\langle x^n \rangle$ to exist is that $2\nu > (n - 1)$. Further, since the arctangent is bounded and constant for large
 227 $|x|$, $p(x)$ has power law tails, with the *pdf* varying as $x^{-2(\nu+1)}$ for large $|x|$.

228 The multiplicative noise $Ex\xi_1$ and additive noise $g\xi_1 + b\xi_2$ are uncorrelated if $g = 0$.
229 However, when the physical processes represented by these noises are correlated, $g \neq 0$, and cause
230 the stationary *pdf* to be skewed, basically because the magnitude of $Ex + g$ is different for positive
231 and negative x . The stationary mean of x is zero by construction, but the peak of the probability
232 distribution, the mode, is at $x_p = E g / (L - E^2/2)$. In fact, since the deterministic parameter L can be
233 absorbed in ν , it is possible to factor L completely out of Eq. (2), and thus out of the location of
234 the mode of the stationary *pdf*, by rescaling E^2 , g , and b^2 by the decay rate λ of the lagged
235 autocorrelation of x (see Eq. A4, and Sardeshmukh et al. 2015).

236 There are other specific properties of $p(x)$ and relationships between moments when $g \neq 0$,
237 which have been used in previous publications (Sardeshmukh and Sura 2009; Sardeshmukh et al.
238 2015) as strong evidence that the unconditional (that is, the climatological) probability
239 distributions of daily anomalies of important meteorological variables are approximately SGS
240 distributions. Sardeshmukh and Sura (2009) justified the relevance of the SGS distribution
241 associated with a univariate CAM-noise process Eq. (1) in the obviously multivariate real
242 atmospheric system by invoking a principle of “diagonal dominance” in the equations for the
243 higher statistical moments (such as skew, kurtosis, and higher moments) in multivariate linear
244 systems. More rigorously, Sardeshmukh and Penland (2022, manuscript in progress) have recently
245 shown the SGS distribution to be the unconditional distribution of any single component, or any
246 linear combination of components, of a multivariate linear system governed by a vector form of
247 Eq. (1) with multivariate CAM noise forcing.

248 In this paper we are not so concerned with the unconditional as with the conditional
249 distribution. In a forecasting context, by conditional distribution we mean a distribution
250 conditioned on an initial condition, that one would obtain if the initial condition were perfect or

251 very nearly perfect. Such a distribution would have a finite spread at any forecast lead time because
252 of the chaotic system dynamics. To highlight the general skewness of such conditional
253 distributions, we consider ensemble forecasts from perfect initial conditions in a simple system in
254 which $g = 0$ in the generating equation and the climatological *pdf* is symmetric. As we shall see,
255 the multiplicative noise interacts with any nonzero initial condition in even such a system to
256 generate a conditional skew that reaches a maximum and then decays to zero with forecast lead
257 time.

258

259 3. Conditional SGS

260 Conditional *pdfs* obey the same Fokker-Planck equation that stationary *pdfs* do, except
261 their time derivative in the equation is not zero. Also, in general the normalization constant in the
262 formula for the conditional *pdf* is time-dependent. This is a serious complication in deriving
263 analytical expressions for conditional *pdfs* in CAM-noise driven systems, and indeed we have not
264 derived them. We have, however, derived analytical expressions for their statistical moments in
265 Appendix A. As shown below, our numerical results match these analytical expressions within
266 sampling uncertainty, and the *pdfs* themselves also appear to be SGS *pdfs*, albeit with parameters
267 conditioned on the initial condition.

268 To demonstrate the general skewness of conditional *pdfs* in even the simplest multiplicative
269 noise driven systems, we numerically generated ensembles of a version of Eq. (1) with $g = 0$. The
270 other parameters were chosen so that the resulting time series $x(t)$ had unit variance and its
271 autocorrelation function decayed exponentially with decay rate $\lambda = 1$. Henceforth, all values of x
272 are presented in units of the climatological standard deviation, and time in units of the decay time
273 $1/\lambda$. Specific values of the parameters are $L = -1.125$, $E^2 = 0.25$, and $b^2 = 1.75$. Each member of

274 an ensemble was integrated from the same initial condition x_o , i.e., $p(x, t = 0) = \delta(x - x_o)$. The
275 different members resulted from differed seeds provided to the random number generator returning
276 different time series of ξ_1 and ξ_2 .

277 An ensemble of 50000 members was generated for each initial condition. Each member
278 was integrated using a stochastic fourth-order Runge-Kutta (RK4) scheme (Rümelin 1982). A
279 Mersenne Twister was used to provide random deviates from a uniform distribution, which were
280 then converted to Gaussian deviates using a Box-Müller scheme. The RK4 scheme has been
281 shown to give results consistent with a Stratonovich Mil'steyn scheme (e.g., Kloeden and Platen
282 1992), and the two schemes have the same formal order of accuracy. Nonetheless, we chose to use
283 the RK4 scheme since it is favored by many scientists.

284 The number of ensemble members (50000) is much larger than can be reasonably expected
285 of either operational or most research ensemble calculations. We chose such a large ensemble size
286 for our control ensemble to verify our analytical results without seriously having to consider
287 sampling uncertainty. Of course, sampling uncertainty is an important issue in real forecasts. We
288 investigated it by subdividing our 50000-member ensemble into 500 100-member and 250 200-
289 member ensembles and recomputing the results using these smaller and more reasonable
290 ensembles sizes.

291 We monitored the forecast ensemble histograms for integrations from initial conditions x_o
292 = 1, 3, and 5. Results for the control ensemble using $x_o = 5$ are shown in **Fig. 1a** at forecast lead
293 times $\tau = 0.25, 1, \text{ and } 7$, starting from the delta function at x_o . For reference, we also show the
294 Gaussian histograms (Fig. 1b) obtained from similar ensemble integrations of an Ornstein-
295 Uhlenbeck process, i.e. with both E and g set equal to 0 in Eq. (1) and with L and b^2 adjusted to
296 give unit variance and unit decay time. All histograms were smoothed with a 5-point smoother.

297 Figure 1 shows that in both the SGS and Gaussian cases the delta-function initial condition evolves
298 into a symmetric stationary *pdf* at long lead times. Note that the mean of the conditional *pdf* (that
299 is, the ensemble-mean forecast) is the same in the two cases at all forecast lead times. However,
300 unlike the Gaussian case, the conditional *pdfs* in the SGS case are skewed and heavy-tailed, and
301 for this initial condition, wider than even the stationary *pdf*!

302 As shown in Appendix A, the conditional moments can be evaluated using an equation
303 such as Eq.(A6) derived from the Fokker-Planck equation. Solving this equation analytically for
304 the conditional first, second, third, and fourth moments allows evaluation of the conditional
305 variance, skew, and excess kurtosis which can then be compared with the corresponding quantities
306 estimated directly from the forecast ensembles. (Note that unlike the stationary moments,
307 analytical evaluation of these conditional moments does require knowledge of the decay rate,
308 which in our examples is unity). This comparison is shown in **Fig. 2** for the control ensemble as
309 a function of forecast lead time τ for the three initial conditions x_o . To assess the robustness of
310 the numerical results with respect to sampling uncertainty, which can be an issue in estimating
311 higher moments, we also show the analytical and empirical skew and excess kurtosis obtained in
312 the control ensemble integrations from negative mirror initial conditions, $x_o = -1, -3, \text{ and } -5$. The
313 numerical conditional skew curves for the positive and negative initial conditions are nearly mirror
314 images of one another, as expected, and in excellent agreement with the analytical curves. The
315 agreement of the conditional excess kurtosis curves is however not as good, even using such a
316 large ensemble. Note that in all cases the maximum skew occurs fairly early in the forecasts, at
317 lead times τ between 0.7 and 1.0 decay times.

318 In the following, we will introduce approximations that rely more on the skew than the
319 excess kurtosis since the skew is more accurately estimated than the kurtosis. In fact, the sample

320 histograms of all conditional moments, including variance, from the 250 sets of 200-member
321 ensembles and from the 500 sets of 100-member ensembles are themselves visibly skewed, and
322 this skew increases with the order of the moment. For this reason, we show median values as well
323 as the 10% confidence levels of the conditional skew as estimated from sets of 100-member and
324 200-member ensembles in **Fig. 3**. That is, 90% (450 values) of the sample skews estimated from
325 the 100-member ensembles lie above the dotted red line in **Fig. 3**. Thus, the uncertainties in sample
326 skew are large, but not so large as to render the skew unusable. Not surprisingly, estimates of
327 kurtosis using even the larger 200-member ensembles are unreliable.

328 Since the conditional *pdfs* in a system with a stationary Gaussian *pdf* are Gaussian, the
329 question arises whether conditional *pdfs* in a system with a stationary SGS *pdf* are also SGS *pdfs*.
330 To address this, we used the conditional moments (variance, skew, and excess kurtosis) in our
331 numerical ensembles to estimate the conditional parameters E , g , and b , assuming that they are
332 related to the conditional moments in the same way as the stationary parameters are related to the
333 stationary moments in Eq.(A4). **Fig. 4** shows the parameters estimated from the control ensemble
334 for three initial conditions, $x_o = 1, 3$ and 5 , as a function of forecast lead time τ . (In practice, we
335 suggest fitting the conditional *pdfs* using a more robust technique, like the maximum likelihood
336 method or a Bayesian procedure (Bianucci and Mannella 2021), than the “method of moments”
337 used here). The conditional parameters estimated from the control ensemble were then specified
338 in Eq.(A2) to determine candidate conditional SGS *pdfs*. **Fig. 5** shows that these candidate
339 conditional distributions are nearly identical to the histograms obtained from the forecast
340 ensembles for the three initial conditions at unit forecast lead time. We have performed similar
341 comparisons for the other lead times and find agreement between histograms and analytically-fit
342 *pdfs* in all cases. Thus, although we have not been able to show analytically that the conditional

343 *pdfs* are SGS *pdfs*, it is clear that SGS *pdfs* are at least an excellent approximation to the conditional
 344 *pdfs*. We are also encouraged to find that these distributions appear to depend only weakly on the
 345 kurtosis, since the kurtosis is generally poorly sampled (Fig. 2d).

346 The generation of skew in the forecast ensemble by multiplicative noise results in a
 347 difference D between the mean and mode of the conditional *pdfs*. This difference, that we call the
 348 forecast distributional bias, is related to the skew S and the multiplicative noise parameter E of the
 349 conditional *pdf* as

$$350 \quad D = x_{mean} - x_{mode} = \frac{\sigma_f}{2} \left(\frac{\lambda - E^2}{\lambda + E^2} \right) S \approx \frac{\sigma_f}{2} \left(\frac{n-3}{n+1} \right) S, \quad (3)$$

351 where σ_f is the standard deviation of the conditional *pdf*, and the last approximation follows from
 352 assuming that $E^2 \approx 2\lambda/(n-1)$ if the highest moment of $(x - x_{mean})$ that exists is
 353 $\langle (x - x_{mean})^n \rangle$. **Fig. 6** shows D estimated from the 100- and 200-member ensembles, but
 354 assuming an accurate estimation of E^2 , in units of the *stationary* standard deviation σ to facilitate
 355 intercomparison among the different forecast cases. Like the conditional moments, the distribution
 356 of sampled D is skewed. Dotted lines indicate the 10% confidence level, meaning that 90% of the
 357 sample D values lie above these curves. Also shown are values of D using parameters estimated
 358 from the theoretical moments derived in Appendix A. For all three initial conditions x_o , D peaks
 359 at lags of about $\tau = 0.6$ but remains large for longer lead times. Not surprisingly, D is largest for
 360 forecasts from $x_o = 5$, since the multiplicative noise is relatively largest in that case, but D generally
 361 cannot be ignored even in the case of the much less extreme initial condition $x_o = 1$; the median
 362 (50% confidence level) values of D are close to the theoretically estimated values.

363 The skew of the forecast ensemble also affects the predicted risk ratio R of extreme positive
 364 and negative deviations from the ensemble-mean forecast. For simplicity, we define these

365 extremes as ± 2 standard deviations from the ensemble-mean forecast. That is, for $\tilde{x} =$
366 $(x - x_{mean})/\sigma_f$, where x_{mean} and σ_f are the mean and standard deviation of the conditional *pdf*,
367 respectively, we define $R = p(\tilde{x} = +2)/p(\tilde{x} = -2)$. Fig. 7 shows R as a function of forecast lead
368 time for the three initial conditions x_o using Eq. 2 and parameters estimated from the control
369 ensemble with 50000 members (black circles; see Fig. 5). Note that in the Gaussian (Ornstein-
370 Uhlenbeck) case, $R = 1$ at all lead times. In the SGS case, R is clearly not equal to 1, and for small
371 S , is related approximately to S as

$$372 \quad R \approx \exp \left[S \frac{\tilde{x} (\tilde{x}^2 - 3)}{3} \right] = \exp \left[\frac{2}{3} S \right]. \quad (4)$$

373 We tested this simple approximation to R using our 100- and 200-member ensembles. **Fig. 7**
374 shows that the approximation appears to be valid for conditionals skews smaller than about 0.5.
375 As with the conditional moments and D , the distributions of R estimated from our ensemble sets
376 are skewed. In all cases, the values of R estimated from the *pdfs* shown in Fig. 5 are well within
377 the range of ensemble values.

378

379 5. Discussion and Concluding Remarks

380 We have argued that forecast ensembles in the climate system must be generally skewed, in part
381 because the chaotic system dynamics responsible for the ensemble spread are generally state
382 dependent. We showed that forecast ensembles are skewed even in the simplest scalar linear
383 dynamical system with state-dependent noise, and even when the initial and the long-lag
384 (“climatological”) ensembles are not skewed. The ensemble skew S is related to two other
385 important aspects of ensemble asymmetry: the difference D between the expected mean and most
386 likely forecast, and the risk ratio R of extreme positive and negative deviations from the ensemble-
387 mean forecast. In our simple system, which has relevance in weather and climate dynamics (e.g.,

388 Sardeshmukh et al. 2015), these relationships can be made explicit. This motivates us to propose
389 S as a generally useful metric of ensemble forecast information, in addition to the ensemble mean
390 and spread.

391 State-dependent (multiplicative) and state-independent (additive) noises can both increase
392 forecast spread. Unlike additive noise, however, multiplicative noise can make the spread even
393 larger than the climatological standard deviation and generate skewed and heavy tailed forecast
394 distributions. These effects can cause predictions of extreme anomaly risks to strongly differ from
395 similar forecasts based on Gaussian statistics.

396 In our illustrative examples, the multiplicative noise effects on the forecast *pdfs* peaked at
397 forecast lead times $\tau \sim 1$, and were relatively large when the initial anomaly was large ($x_o = 5$).
398 The basic reason for this is not hard to understand. When the initial anomaly is large, the
399 multiplicative noise is also relatively large, and can push the system to even larger values than
400 additive noise. The effect becomes smaller after $\tau \sim 1$ in all forecast cases as the ensemble drifts
401 towards climatology and x becomes smaller.

402 Analysis of a multivariate CAM-noise system (Sardeshmukh and Penland 2022,
403 manuscript in progress) suggests that one may reasonably assume results from this simple system
404 to be generally relevant even in multi-dimensional systems with state-dependent noise. Further,
405 Thompson et al. (2017) have shown that homogenization of a system driven by linear CAM noise
406 rigorously converges to one driven by an α -stable Lévy process, whose self-similar properties of
407 the *pdf* are well-known. If the CAM noise process were to be far from self-similar itself, one
408 would not expect it to be such a good approximation of a Lévy process as it is (Penland and
409 Sardeshmukh 2012; Gottwald 2021).

410 The fact that the largest multiplicative noise effects occur for forecasts from extreme initial
411 conditions highlights the importance of properly accounting for them in predictions of extreme
412 events. We should note that in a multi-component system such extreme initial conditions need not
413 be associated with extremely large-amplitude initial anomaly vectors as in our scalar 1-component
414 system, but merely have a large projection on the optimal initial vectors for anomaly growth over
415 the global forecast domain or a desired sub-domain. Such optimal vectors, usually identified with
416 the dominant singular vectors (SVs) of the system's perturbation evolution operator, have found
417 extensive application in initial ensemble design and predictions (e.g., Buizza and Palmer 1995).
418 Multiplicative noise in the evolution from an initial SV perturbation to an extreme large-amplitude
419 perturbation would then generate large skew in the forecast *pdfs* even from small initial amplitudes
420 in this vector scenario, unlike in the scalar 1-component scenario.

421 Finally, because multiplicative noise effects on forecast *pdfs* depend on the initial condition
422 as well as forecast lead time, one cannot account for them using state-independent *a-posteriori*
423 probabilistic bias corrections. For example, it may be tempting to correct for the undesirable
424 symmetric U shapes of rank histograms in many ensemble forecasting systems through state-
425 independent adjustments of the ensemble spread. However, in reality such a symmetry may also
426 arise from a conflation of spuriously skewed conditional *pdfs* in forecasts from positive and
427 negative initial conditions (see Fig 2c), as also noted by Hamill (2001). One may think of
428 addressing such issues by stratifying the *a-posteriori* corrections with respect to initial conditions
429 and forecast lead times. This is possible in principle, but impractical in high dimensional systems.
430 A more practical and physically grounded approach is to attempt to represent such state-dependent
431 noise effects, however crudely, in the forecast model itself through stochastic parameterizations of
432 the state-dependent chaotic model tendencies. As mentioned in the introduction, such approaches

433 have already proven beneficial in operational ensemble forecasting. The analysis of this paper
434 further highlights the need to continue improving such parameterizations to continue improving
435 the prediction and dissemination of extreme anomaly risks.

436

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438 was partially funded through NOAA's Weather Prediction Office.

439

440

Appendix A

441

442

443 Given Eq. (1), the Stratonovich form of the Fokker-Planck equation for the conditional probability

444 that the system is found between x and $x+dx$ at time t when the initial condition at $t=0$ is x_o is

445

$$446 \quad \frac{\partial p(x,t|x_o)}{\partial t} = -\frac{\partial}{\partial x} \left[\left(L + \frac{1}{2}E^2 \right) x p(x,t|x_o) \right] + \frac{1}{2} \frac{\partial^2}{\partial x^2} \{ [(Ex + g)^2 + b^2] p(x,t|x_o) \}. \quad (\text{A1})$$

447

448 This equation is also obeyed by the stationary probability density function (*pdf*) $p(x)$ for x with the

449 time derivative on the left-hand side set to zero. Integrating the equation for the stationary *pdf*

450 once over x and using the fact that $p(x)$ vanishes as x gets infinitely large yields

451

$$452 \quad p(x) = \frac{1}{N} [(Ex + g)^2 + b^2]^{-(\nu+1)} \exp \left[q \arctan \left(\frac{Ex+g}{b} \right) \right] \quad (\text{A2a})$$

453

454 where $\nu = -[(L/E^2) + 1/2]$, $q = 2g\nu/b$, and N is a normalization constant

455

$$456 \quad N = \frac{2\pi}{E} (2b)^{-(2\nu+1)} \frac{\Gamma(2\nu+1)}{\Gamma(\nu+1-iq/2)\Gamma(\nu+1+iq/2)}. \quad (\text{A2b})$$

457

458 Note that in Eq. (A2) the deterministic parameter L occurs only in terms of ν . It is therefore possible

459 to factor it completely out of the equation. In Sardeshmukh et al. (2015), the parameter

460 $\lambda = -(L + E^2/2)$ was used to rescale E^2 , g , and b^2 , rendering a three-parameter version of (A2) that

461 is completely equivalent numerically. Note that λ is the rate that the autocorrelation of x decays
 462 and is therefore easily estimated from data.

463 Eq. (1) has been developed for an anomalous state variable so that the stationary mean $\langle x \rangle$
 464 is zero. As shown in Sardeshmukh et al (2015), the parameters E , g , and b can be estimated from
 465 the moments of $p(x)$ in terms of λ . Recalling the general definitions of variance σ^2 , skew S , and
 466 excess kurtosis K :

$$467 \quad \sigma^2 = \langle (x - \langle x \rangle)^2 \rangle \quad (\text{A3a})$$

$$468 \quad S = \langle (x - \langle x \rangle)^3 \rangle / \sigma^3 \quad (\text{A3b})$$

$$469 \quad K = \langle (x - \langle x \rangle)^4 \rangle / \sigma^4 - 3, \quad (\text{A3c})$$

470
 471 the following expressions for E , g , and b can be derived analytically and evaluated sequentially:

$$473 \quad E^2 = \left(\frac{2}{3}\right) \frac{\left[\frac{K - \frac{3}{2}S^2}{K - S^2 + 2}\right] \lambda}{\lambda} \quad (\text{A4a})$$

$$474 \quad g = S \sigma \left(\frac{1 - E^2/\lambda}{2E}\right) \lambda \quad (\text{A4b})$$

$$475 \quad b^2 = 2\sigma^2 \left[1 - \frac{\lambda E^2}{2} - \frac{(1 - \lambda E^2)^2}{8\lambda E^2} S^2\right] \lambda. \quad (\text{A4c})$$

476
 477 The system considered here has a stationary *pdf* that is symmetric, i.e., $g = 0$. We have
 478 shown numerically that even in this case the conditional *pdf* is skewed. However, it is not clear if
 479 the conditional probability $p(x, t | x_0)$ is strictly a *Stochastically Generated Skewed* (SGS)
 480 distribution of the form (A2). If so, the parameters of $p(x, t | x_0)$ could be estimated from expressions
 481 (A3) and (A4), with all of the moments in Eq. (A3), including the means and standard deviations

482 σ , replaced by the conditional moments. Recall that the conditional probability $\rho(x, t|x_o)$ obeys
 483 the same equation as the stationary probability $\rho(x)$, retaining the time derivative. For our system,
 484 that is

485

$$486 \frac{\partial p(x,t|x_o)}{\partial t} = -\frac{\partial}{\partial x} \left[\left(L + \frac{1}{2}E^2 \right) x p(x, t|x_o) \right] + \frac{1}{2} \frac{\partial^2}{\partial x^2} \{ [E^2 x^2 + b^2] p(x, t|x_o) \}. \quad (\text{A5})$$

487

488 The time derivative makes it more difficult to solve for $\rho(x, t |x_o)$. However, it is easy to derive
 489 solvable equations for the conditional moments via integration by parts. The equation for the n^{th}
 490 conditional moment at time $t = \tau$ given an initial condition x_o at $t = 0$ reads

491

$$492 \frac{\partial \langle x^n(\tau) | x_o \rangle}{\partial \tau} = n \left(L + \frac{1}{2}E^2 \right) \langle x^n(\tau) | x_o \rangle$$

$$493 + \frac{1}{2} n(n-1) [E^2 \langle x^n | x_o \rangle + b^2 \langle x^{n-2} | x_o \rangle] . \quad (\text{A6})$$

494

495 Solving Eq. (A6) for the first four conditional moments in terms of the stationary variance σ^2 , and
 496 using the fluctuation-dissipation relation to eliminate b , we find

497

$$498 \langle x(\tau) | x_o \rangle = \exp \left[\left(L + \frac{1}{2}E^2 \right) \tau \right] x_o \quad (\text{A7a})$$

$$499 \langle x^2(\tau) | x_o \rangle = \exp \left[2(L + E^2) \tau \right] x_o^2 + (1 - \exp \left[2(L + E^2) \tau \right]) \sigma^2 \quad (\text{A7b})$$

$$500 \langle x^3(\tau) | x_o \rangle = \exp \left[3(L + \frac{3}{2}E^2) \tau \right] x_o^3 +$$

$$501 \frac{3(L+E^2)}{(L+2E^2)} \sigma^2 (1 - \exp \left[2(L+2E^2) \tau \right]) \exp \left[\left(L + \frac{1}{2}E^2 \right) \tau \right] x_o \quad (\text{A7c})$$

$$502 \langle x^4(\tau) | x_o \rangle = \exp \left[4(L + 2E^2) \tau \right] x_o^4$$

503
$$- 6 \frac{(L+E^2)}{(L+3E^2)} \sigma^2 (\sigma^2 - x^2_o) \exp[2(L+E^2) \tau] (1 - \exp[2(L+3E^2) \tau]$$

504
$$+ 3 \sigma^4 \frac{(L+E^2)}{(L+2E^2)} (1 - \exp[4(L+2E^2) \tau]) \quad . \quad (A7d)$$

505

506 Using the conditional moments in Eqs. (A3) and (A4) to estimate the conditional parameters
507 $E^2(x, \tau | x_o)$, $g(x, \tau | x_o)$ and $b^2(x, \tau | x_o)$ in terms of λ , one can evaluate a conditional, unnormalized
508 version of Eq. (A2a). Translating x appropriately by $\langle x | x_o \rangle$ and normalizing numerically, we
509 compare the SGS estimation of $p(x, \tau | x_o)$ with the estimation of $p(x, \tau | x_o)$ from the raw histogram
510 (Fig. 5). We leave analytical verification of their equivalence to those with more patience and/or
511 better symbolic manipulators than we have. (Such hardy souls would have to contend with Gamma
512 functions of temporally-varying, complex argument in the normalization factor contributing to the
513 left side of Eq. (A5))

514

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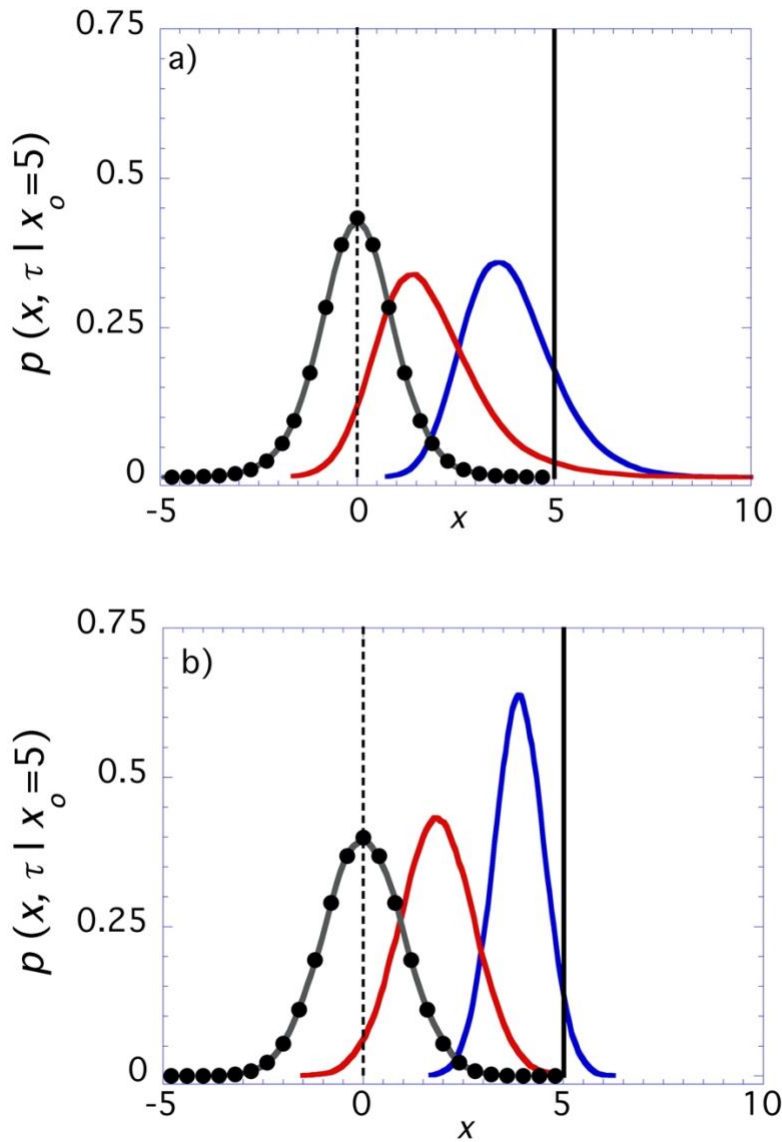
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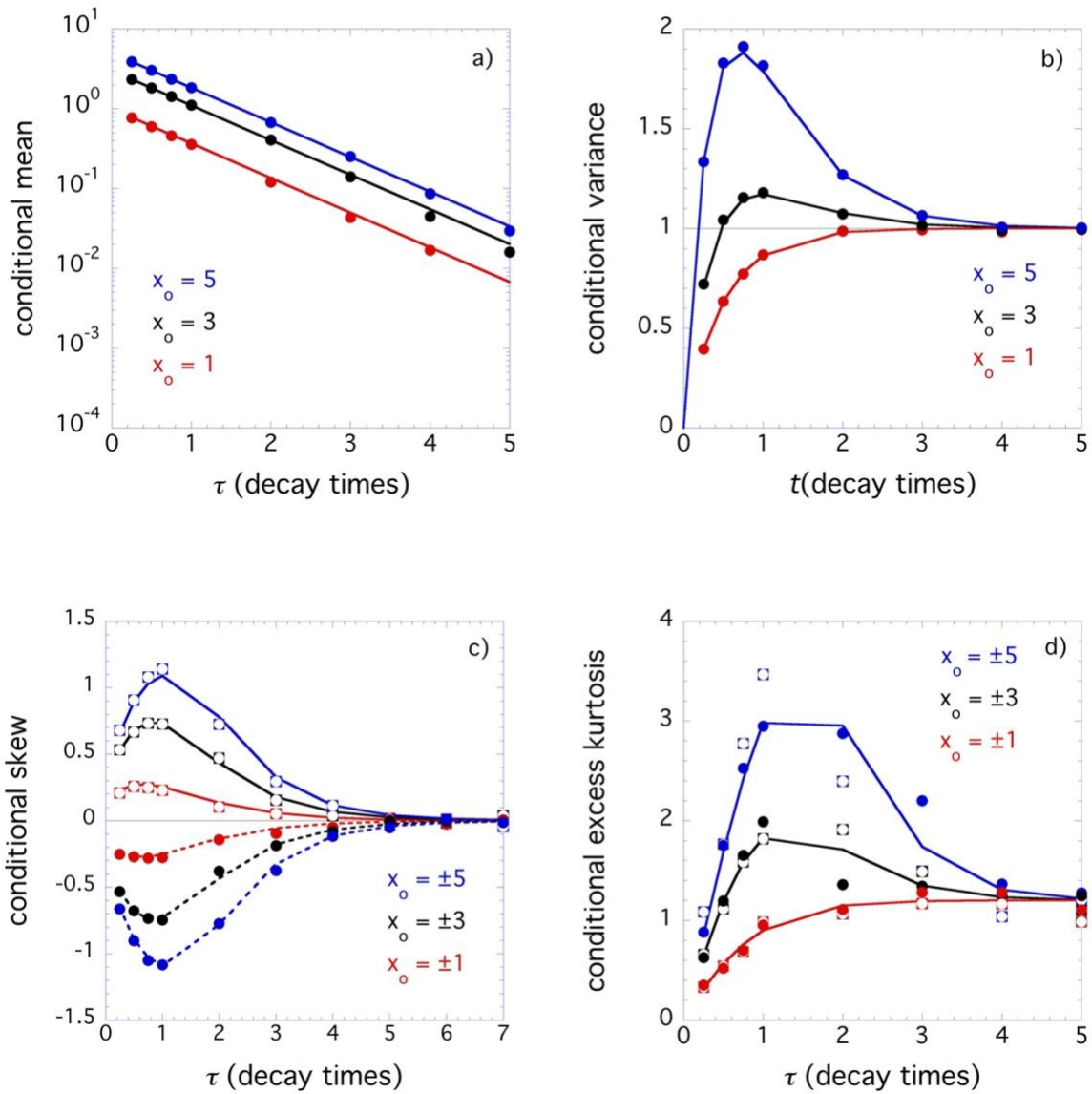
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653



654

655 **Fig. 1:** Snapshots of conditional (forecast) probability density functions (*pdfs*) of ensemble
 656 forecasts starting from an initial condition $x_o = 5$ (solid vertical line), at lead times of $\tau = 0.25$ (
 657 blue line), 1 (red line), and 7 (black line). Analytical stationary *pdf* is shown with black
 658 symbols, and its zero mean as a dotted vertical line. (a) Top: Stochastically Generated Skew
 659 process. (b) Bottom: Ornstein-Uhlenbeck Gaussian process. All distributions have been
 660 smoothed with a five-point smoother. Note that the conditional *pdfs* are skewed and heavy tailed
 661 in (a) but Gaussian in (b).

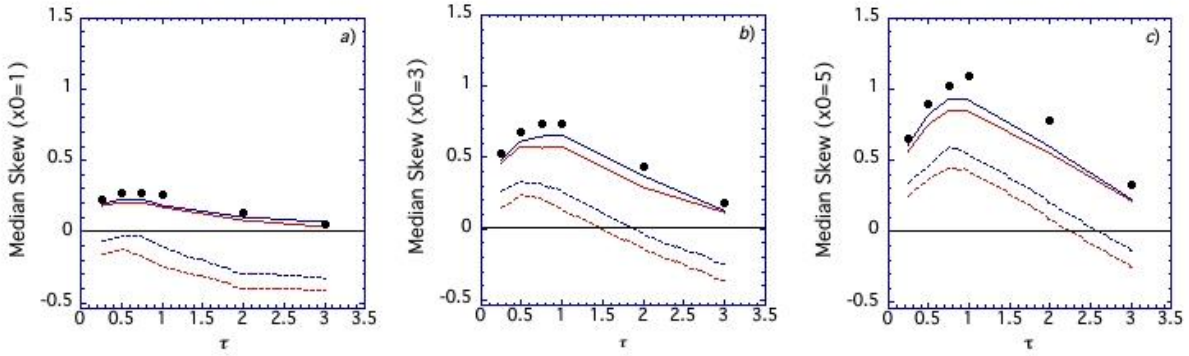


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664 **Fig. 2:** Conditional moments as a function of forecast lead time. Symbols: Using forecast values
 665 from the ensemble members. Lines: Analytical values. *a*) Conditional mean for $x_o = 1$ (red lines),
 666 3 (black lines), and 5 (blue lines). *b*) Same as *a*) but for conditional variance. *c*) Conditional skew
 667 for $x_o = 1, 3,$ and 5 (open symbols and solid lines) and for $x_o = -1, -3,$ and -5 (solid symbols and
 668 dotted lines). *d*) Same as *c*) but for conditional excess kurtosis.

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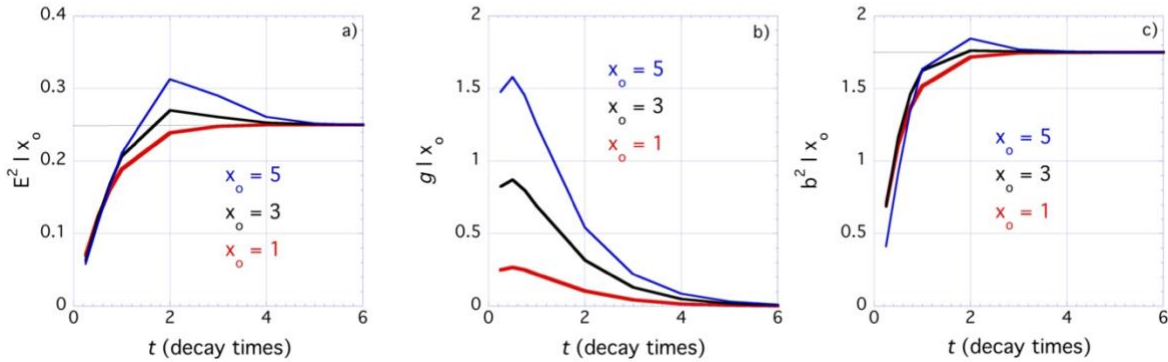


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672 **Fig. 3:** Sample conditional skew estimated from 500 100-member ensembles (red lines) and 250
 673 200-member ensembles (blue lines) for initial conditions $x_o = 1, 3,$ and 5 . Black dots: values
 674 estimated from expressions derived in Appendix A. Lines: Median skew of the ensembles. Dotted
 675 lines: 10% confidence levels based on number of 100-member (red) and 200-member (blue)
 676 ensembles.

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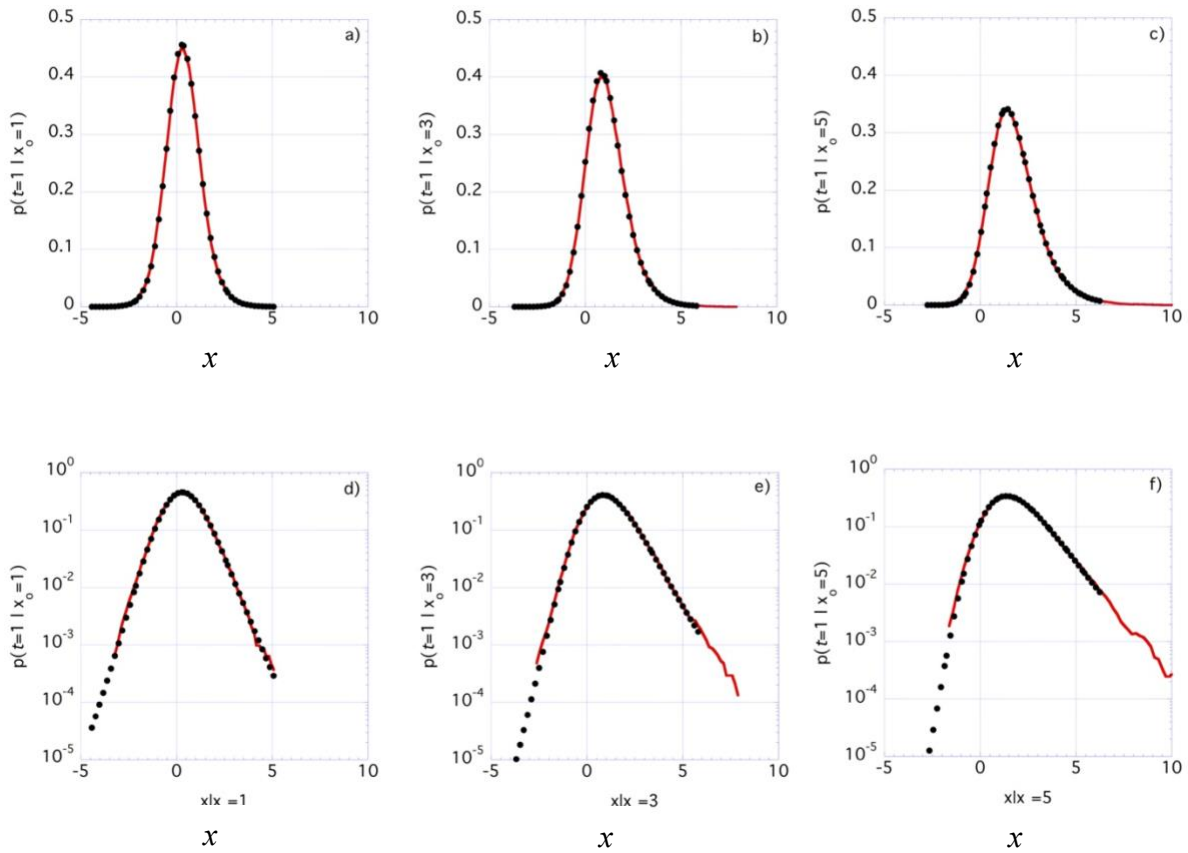


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680 **Fig. 4:** Conditional SGS parameters as a function of forecast lead time for initial conditions $x_o =$
 681 1 (red line), 3 (black line), and 5 (blue line). a) E^2 given x_o , b) g given x_o . c) b^2 given x_o .
 682 Climatological values are $E^2 = 0.25$, $g = 0.$, and $b^2 = 1.75$.

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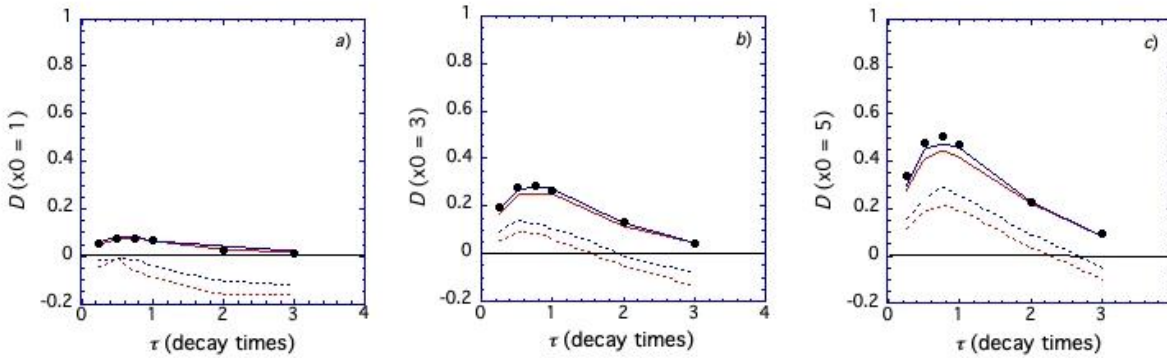
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687 **Fig. 5:** Conditional probability distribution functions (*pdfs*) at lead time $\tau = 1$ estimated using
 688 forecast values from the ensemble members (red lines) and derived analytically using the
 689 estimated conditional SGS parameters (black symbols). Top row (*a-c*): For initial conditions x_0
 690 $= 1, 3,$ and $5,$ on a linear scale. Bottom row (*d-f*): As in the top row, but on a logarithmic scale.

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696 **Fig. 6:** Distributional bias D , i.e. conditional pdf mean minus the pdf mode, as a function of lead
697 time for $a) x_o = 1$, $b) x_o = 3$, and $c) x_o = 5$. Solid lines: Eq. (3) median of 500 100-member ensembles
698 (solid red lines) and of 250 200-member ensembles (solid blue lines). Dashed lines: 10%
699 confidence levels. Symbols: Values estimated from Fig. 5 and Fig. 2a.

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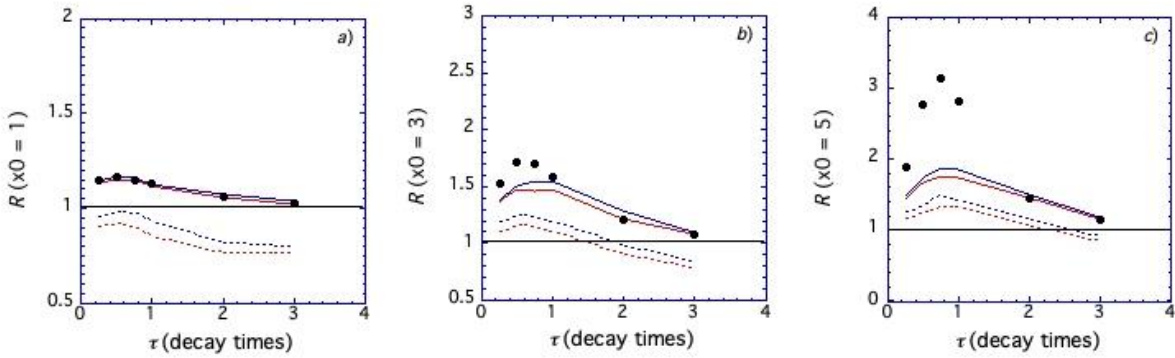
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709 **Fig. 7:** Conditional risk ratio R defined in Eq. (4) as a function of lead time for *a*) $x_o = 1$, *b*) $x_o =$

710 3, and *c*) $x_o = 5$. Black circles: Values estimated from *pdfs* shown in Fig. 5. Approximations of R

711 using only skew (Eq. 4) are shown for the 100-member (red lines) and 200-member (blue lines)

712 ensemble sets. Solid Lines: median approximate values of R . Dashed Lines: 10% confidence

713 levels. Accurately estimated quantities (symbols) are well within the range of ensemble values;

714 the 90% confidence levels are off the scale of the graphs.