

1 Fog Formation during Gravity Currents Interacting with 2 Coastal Topography

3 **Stef L. Bardoel¹ • Daniel V. Horna Muñoz^{1,2} • Andrey A. Grachev^{1,3*} • Raghavendra
4 Krishnamurthy^{1,4} • Leonardo P. Chamorro⁵ • Harindra J.S. Fernando¹**

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7 **Abstract** An interesting mixing-fog event was identified during the C-FOG field campaign,
8 where a cold front arriving from the north-east collided with the Downs peninsula in
9 Ferryland, Newfoundland, to produce misty/foggy conditions. A comprehensive set of
10 field observations suggests that this collision caused turbulent mixing of nearly saturated
11 ambient air with an almost saturated cold front, creating conditions for mixing fog. To
12 delve into physical processes underlying this phenomenon, laboratory experiments were
13 performed on the interaction of lock-exchange induced gravity currents with a (rectangular)
14 obstacle. Instantaneous velocity and density fields were obtained using particle image
15 velocimetry and planar laser-induced fluorescence. The observations suggest that the
16 obstacle starts affecting the approaching gravity-current propagation at an upstream
17 distance of $2H$ and, upon collision, the mixing is taking place over a horizontal spatial
18 extent of $0.83H$, where H is the depth of the ambient fluid layer. The time for larger-scale
19 turbulent stirring to permeate to the smallest scales of turbulence and activate the

✉ Stef L. Bardoel
sbardoel@nd.edu

¹ Department of Civil and Environmental Engineering and Earth Sciences, University of Notre Dame, Notre Dame, IN, USA

² Department of Science and Centro de Investigación y Tecnología del Agua (CITA), Universidad de Ingeniería y Tecnología (UTEC), Barranco, Lima, Peru.

³ NOAA Physical Sciences Laboratory / Cooperative Institute for Research in Environmental Sciences, University of Colorado, Boulder, CO, USA

⁴ Pacific Northwest National Laboratory, Richland, WA, USA

⁵ Department of Mechanical Science and Engineering, University of Illinois, Urbana, IL, USA

* Permanent Address: Boundary Layer Research Team/Atmospheric Dynamics & Analytics Branch, DEVCOM Army Research Laboratory, WSMR, NM, USA

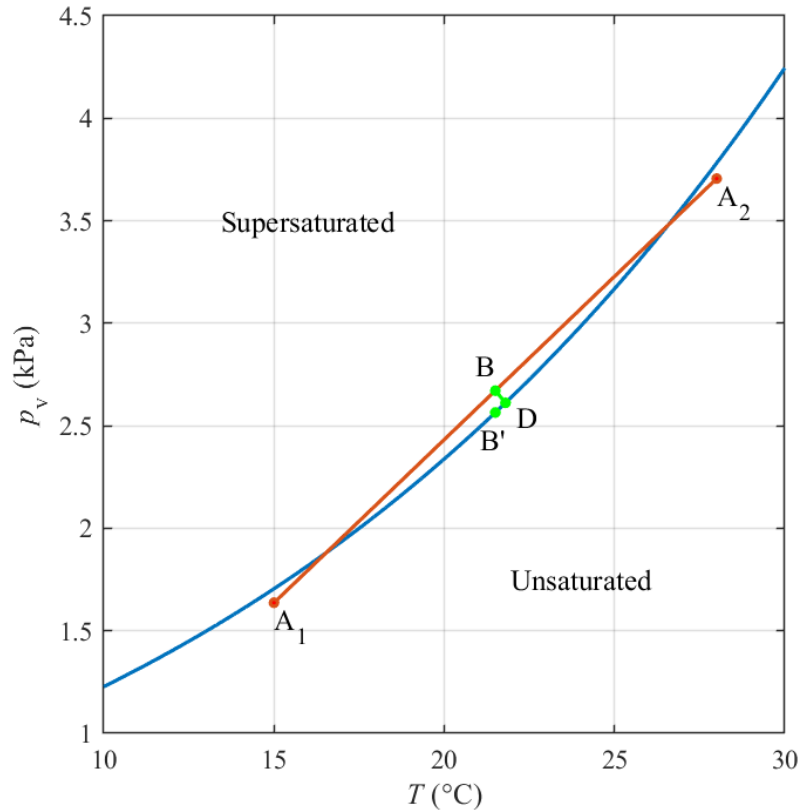
20 condensation nuclei was estimated as $3t^*$, where $t^* = \sqrt{H/g'}$ is the intrinsic time scale of
21 the gravity current and g' the reduced gravity. Extrapolation of laboratory results to field
22 conditions showed a good agreement with observations.

23 **Keywords** C-FOG case study · Gravity currents · Mixing fog · Topography · Turbulent
24 mixing

25 **1 Introduction**

26 Fog typically forms when air close to the earth's surface becomes slightly supersaturated
27 and produces a layer of very small suspended water droplets (or clouds) in contact with the
28 surface. The AMS Glossary (2020) defines fog in terms of visibility, a condition that
29 reduces visibility below 1 km (0.62 miles). The extent of visibility reduction is considered
30 as the 'intensity' of fog. The usual fog formation mechanism is the deposition of water
31 vapour on hygroscopic aerosol nuclei (condensation nuclei CN) under favourable
32 conditions, such as reduction of temperature that 'activates' CN leading to droplet growth
33 (Gultepe et al. 2007). Dry atmospheric aerosols with a typical size of $\sim 0.1 \mu\text{m}$ are small
34 enough to scatter wavelengths of visible light preferentially, producing colours and give
35 opalescent appearance to the atmosphere (known as haze, visibility 2-5 km). On the other
36 hand, activated droplets are too large to yield differential scattering, lead to visibility
37 impairment, and give a white appearance. These include mist ($\sim 1 \mu\text{m}$, visibility 1-2 km)
38 and fog ($\sim 1\text{--}30 \mu\text{m}$, visibility < 1 km). There are many types of fog, three main types
39 being radiation fog, advection fog, and mixing fog (Fernando et al. 2020). *Radiation fog*
40 appears when net outgoing radiation cools the ground surface so as to drop the temperature
41 of overlying air below the dew point. *Advection fog* (or *movement fog*, a term used when
42 fog is localized) forms when relatively warm air overrunning a colder surface is cooled to
43 saturation by air-surface exchange processes (warm fog) or cooler air moving over a
44 warmer surface is saturated by evaporation (cold fog). Advection fog is not a feature of a
45 particular synoptic type of wind speed regime. The mixing of two near-saturated air masses
46 of different temperatures may produce supersaturation and, therefore, *mixing fog* (Taylor
47 1917), which is the focus of this paper.

48



49

50 **Fig. 1** A diagram of water vapour (partial) pressure as a function of temperature, illustrating the mixing of
 51 moist air masses A_1 and A_2 to produce fog at B

52 The mechanism of mixing-fog formation is best explained using the curvature of the water
 53 vapour pressure p_v (ordinate) and temperature T (abscissa) diagram (Rhode 1962), as
 54 illustrated in Fig. 1. The blue line represents the saturation vapour pressure p_s curve for
 55 water vapour, which is described by the Clausius-Clapeyron equation for a perfect gas
 56 $d(\ln p_s)/dT = L(T)/RT^2$, where $L(T)$ is the latent heat of evaporation, which is a
 57 function of T , and R the universal gas constant. Air parcels located to the left of this curve
 58 are supersaturated and to the right are unsaturated. The mixing line of two near-saturated
 59 air masses of different temperatures (A_1 and A_2) is shown in red, and in favourable
 60 situations of mixing the two air masses lead to B, which is supersaturated and
 61 preconditioned for fog (for a derivation of mixing line for a special case, see Schumann
 62 1996). The liquid water content (LWC) of B can be estimated by considering the difference
 63 of water vapour pressure of point B and the saturation water vapour pressure B' at the same
 64 temperature. However, due to the release of latent heat, the temperature of air-mass state
 65 B increases during condensation and may reach D instead. Therefore, the overall mixing

66 process is nonlinear and the *LWC* is determined by the difference in water vapour pressure
67 of points B and D. Fog is expected when the mixing curve traverse to the left of the
68 saturation curve, and the appearance is more of cirrus nature. Persistence of fog requires
69 that the state of the final mixture is to the left of the saturation curve, otherwise fog is short
70 lived (Paoli and Shariff 2016).

71 It is also noted that the depiction in Fig. 1 does not consider the droplet activation by CN,
72 which may occur under unsaturated or supersaturated conditions (Gultepe et al. 2007). An
73 interesting aside is the case where A_1 and A_2 are further apart, as in the case of a humid jet
74 engine exhaust in colder air in the upper atmosphere. In such cases, the mixing line first
75 crosses the saturation curve of ice, and then into the supersaturation to produce contrails
76 when the so-called Schmidt-Appleman criterion is satisfied (Paoli and Shariff 2016). The
77 thread of the argument leading to this criterion is analogous to mixing-fog formation,
78 except that the airmasses involved are in a different p_v - T regime and thus physical
79 processes at play can be different.

80 An example is mixing fog that forms during the meeting of warm or cold masses at a front.
81 Byers (1959) defined three fog categories related to fronts, namely pre-frontal, post-frontal,
82 and frontal-passage. Pre-frontal fog occurs before a warm front, whence warm rainfall
83 evaporates into colder air close to the ground and increases humidity toward saturation.
84 Post-frontal fog shortly follows the passing of a cold front, also due to evaporation of rain.
85 Fog behind a cold front, however, is not as widespread because precipitation bands of cold
86 fronts have smaller spatial scales (Gultepe et al. 2007). The third and of interest to this
87 paper is frontal-passage fog, which occurs during the mixing of nearly saturated cold and
88 warm air masses (e.g. Fig. 1). Petterssen (1941) noted that it is impossible to form dense
89 fog by mixing alone, because of the relatively small amount of condensation during mixing.
90 According to Roach (1994), however, dense mixing fog is possible when the temperature
91 difference between the warm and cold air masses is $\sim 10^\circ\text{C}$. The formation of dense fog is
92 also abetted by radiative cooling. George (1940) noted that enhanced mixing of air masses
93 at rough terrain increases the likelihood of fog. Thus, in order to successfully predict
94 coastal mixing fog, it is crucial to understand the interaction of fronts with coastal
95 topography.

96 Cold fronts belong to the class of flows driven by density differences or gravity currents.
97 Numerical models of gravity-current/obstacle interactions have been developed for both
98 2D and 3D cases (Gonzalez-Juez and Meiburg 2009; Tokyay et al. 2012; Nasr-Azadani
99 and Meiburg 2014; Tokyay and Constantinescu 2015; Jung and Yoon 2016; Nasr-Azadani
100 et al. 2018), mostly focussing on confined and unconfined turbidity currents, with model
101 validations conducted against laboratory data. Theoretical developments have been made
102 using shallow water models (Rottman et al. 1985; Lane-Serff et al. 1995), but these models
103 are incapable of rapid transient flow adjustments and mixing at the obstacle. The influence
104 of background rotation on gravity currents with and without sloping surfaces has been
105 studied numerically and using linear and non-linear theories (Hunt et al. 2005) and
106 laboratory experiments (Mahalov et al. 2000). The influence of background turbulence and
107 bottom (surface) friction on gravity current frontal mixing has been modelled by Noh and
108 Fernando (1991, 1993). Nevertheless, experimental laboratory studies on gravity currents
109 interacting with topography have been sparse. Lane-Serff et al. (1995) studied the
110 interaction of a gravity current with a triangular obstacle, where the flow was found to split
111 at the obstacle to a reflected hydraulic jump and an overflow at the obstacle. More recently,
112 Wilson et al. (2018, 2019) performed experimental studies on the interaction of lock-
113 exchange generated gravity currents with a rectangular obstacle. Measurements of
114 instantaneous velocity fields using ultrasonic Doppler velocity profilers showed that
115 velocity profiles in the unobstructed currents have a log-Gaussian shape. In contrast, those
116 of obstructed flow close to the obstacle are more Gaussian. Downstream of the obstacle,
117 the maximum velocity and turbulence intensity were 48% and 28%, respectively, less than
118 unobstructed experiments; i.e., the obstacle decreased the downstream current's maximum
119 velocity and turbulence, possibly by vertical spreading of turbulence activity and due to
120 enhanced dissipation. The height and variance of the maximum speed and turbulence were
121 larger at the obstacle, suggesting its possible role in enhanced mixing and entrainment of
122 ambient fluid into the gravity current.

123 The present study was motivated by field observations of the “Toward Improving Coastal
124 Fog Prediction” (C-FOG) project (Fernando et al. 2021), which included twelve land-based
125 and three ship-based Intensive Operational Periods (IOPs). An interesting, conceivably a
126 mixing-fog, event occurred during IOP7 (0000 UTC 16 September to 1430 UTC 17

127 September 2018), where a cold front from the north-east collided on a peninsula (Downs
128 promontory in the town of Ferryland) protruding into the Atlantic Ocean. It is hypothesized
129 that this collision caused warm near-saturated ambient air to mix with colder near-saturated
130 air at the cold front, thus producing mist and fog. In this paper, the observations from IOP7
131 will be presented in Section 2, together with a description of the observational area and
132 field measurements. Given sparse previous observations on turbulence generated by a
133 gravity current impinging on a topography, a laboratory experiment was conducted to study
134 the impingement of a lock-exchange released gravity current on a rectangular obstacle. The
135 experiment is described in Section 3, followed by results in Section 4 and a discussion in
136 Section 5. The conclusions are given in Section 6.

137 **2 Field Observations**

138 The C-FOG field program involved measurements in eastern Canada at three main coastal
139 land sites on the Avalon Peninsula, Newfoundland (NL) and one in Nova Scotia (NS).
140 Additionally, measurements were conducted aboard an instrumented research vessel R/V,
141 Hugh R. Sharp, that tracked on the coastal areas of NL and NS. The fog-climatology based
142 site selection and the details of the C-FOG field program, including IOPs, are given in
143 Fernando et al. (2021). The information given below is relevant to IOP7 only.

144 **2.1 Instrumentation**

145 During IOP7, the Ferryland site registered relatively short fog events, while other land sites
146 or the R/V did not record fog. The Ferryland site consisted of three auxiliary sites, the
147 Downs site (mentioned earlier, 32 m above sea level a.s.l.), Battery site (3 m a.s.l.), Beach
148 House site and Judges Hill site (129 m a.s.l. on a nearby hill). Downs and Battery were the
149 two most densely instrumented. Figure 2 is a visual overview of the Ferryland sites.

150 Figure 3 shows the instrumentation at the Downs and Battery sites used for this study. The
151 Downs site had a 16.2 m tower with four levels (2, 5, 10 and 15 m above ground level
152 a.g.l.) of instrumentation, with each level having a fast response three-axis ultrasonic
153 anemometer (Model 81000, R.M. Young) for velocity components and virtual temperature
154 at 20 Hz and four slow response HC2S3 (Campbell Scientific) temperature and relative

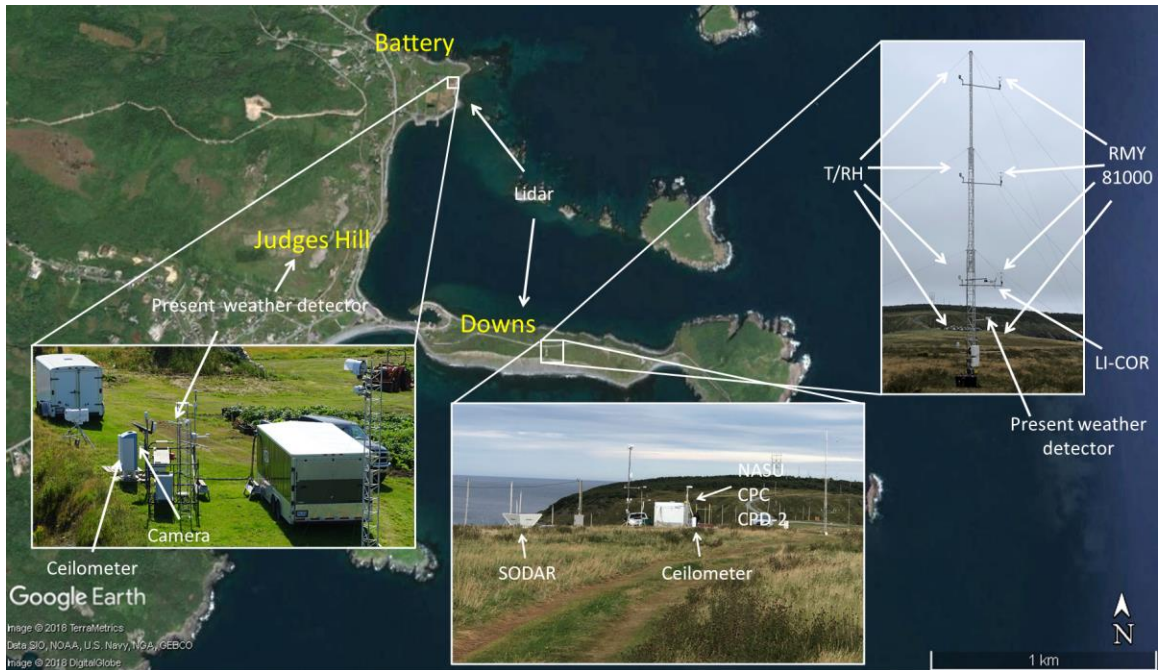


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156 **Fig. 2** Overview of Ferryland C-FOG sites

157 humidity (T/RH) measurements at 1 Hz. The tower also had a LI-COR Inc. fast response
 158 (20 Hz) open-path infrared gas analyzer (LI-7500A) for water vapour fluxes mounted at 5
 159 m. The visibility was measured by a Present Weather Detector (PWD22, Vaisala Inc.), with
 160 data recorded at every 15 s. Other details of the turbulence measurements, data processing,
 161 and data-quality criteria for the measurements at Downs during the C-FOG field campaign
 162 can be found in Grachev et al. (2018, 2021). The site also had a Scintec MFAS
 163 SODAR/RASS wind and temperature profiling system as well the Naval Postgraduate
 164 School (NPS) Aerosol Sampling Unit (NASU) microphysics trailer housed with a modified
 165 CDP-2 (Droplet Measurement Technologies) for cloud/fog droplet size spectrum (2-50
 166 μm) measurements. Two Halo Photonics Streamline XR Doppler lidars, at the Downs and
 167 Battery sites, were conducting co-planar range-height indicator scans to measure winds and
 168 aerosol backscatter. The dual doppler configuration setup and general lidar info is
 169 described by Newsom and Krishnamurthy (2020) and Vassallo et al. (2021).

170 The instrumentation at the Battery site included a CL31 ceilometer (Vaisala Inc.) that
 171 measured aerosol backscatter for cloud base height (CBH) or fog extent observations.
 172 There was also a Present Weather Detector (PWD52, Vaisala Inc.) for visibility
 173 measurements and a camera that took a photo of the Downs once every five minutes. Judges
 174 Hill had a PWD22 (Vaisala Inc.) for visibility measurements. The Beach House site had



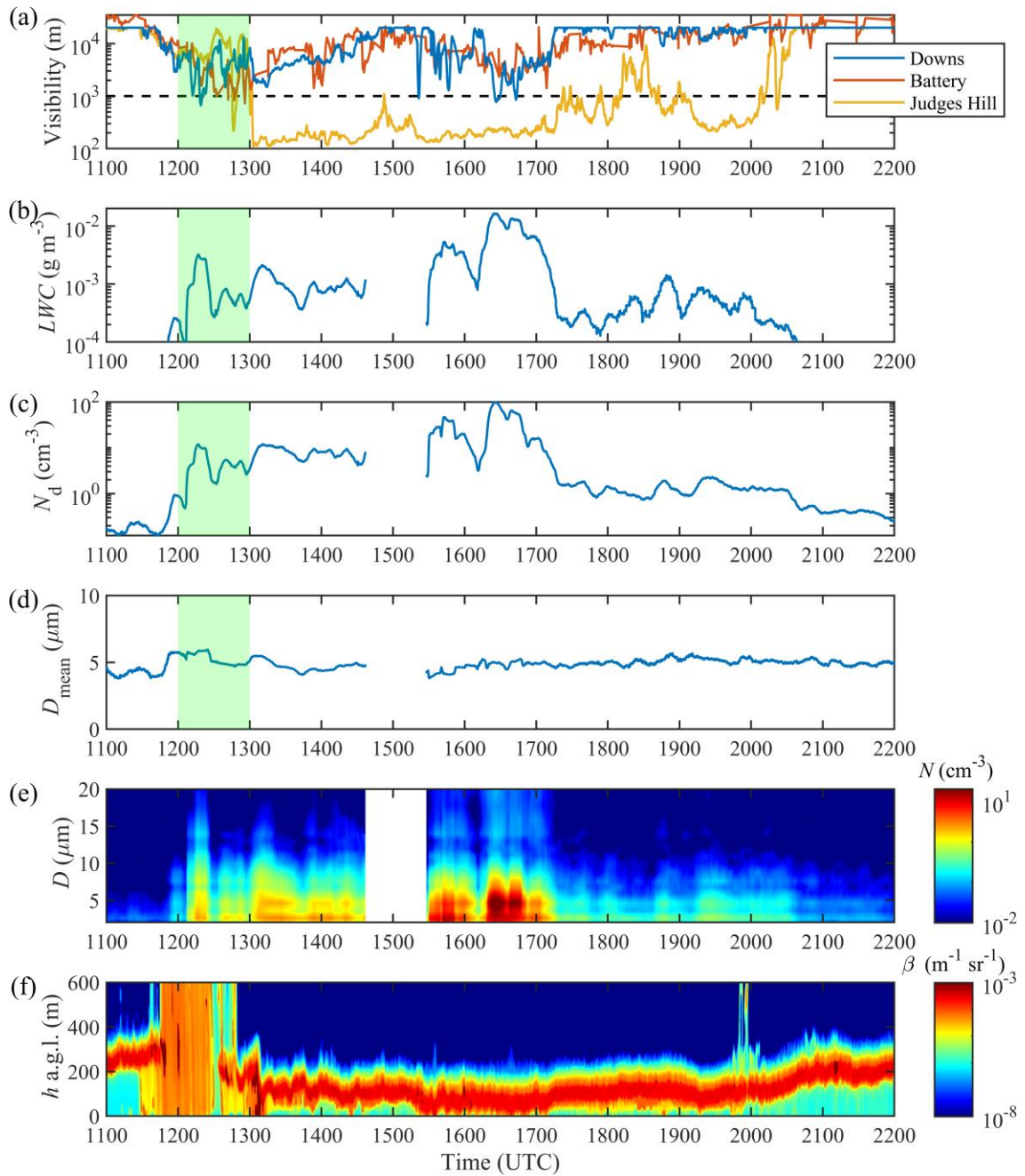
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176 **Fig. 3** Instrument map of the Ferryland sites: the Downs, Battery, Judges Hill. Details of the instrumentation
 177 are given in the text

178 only a scanning Doppler Lidar.

179 **2.2 Fog observations during IOP7**

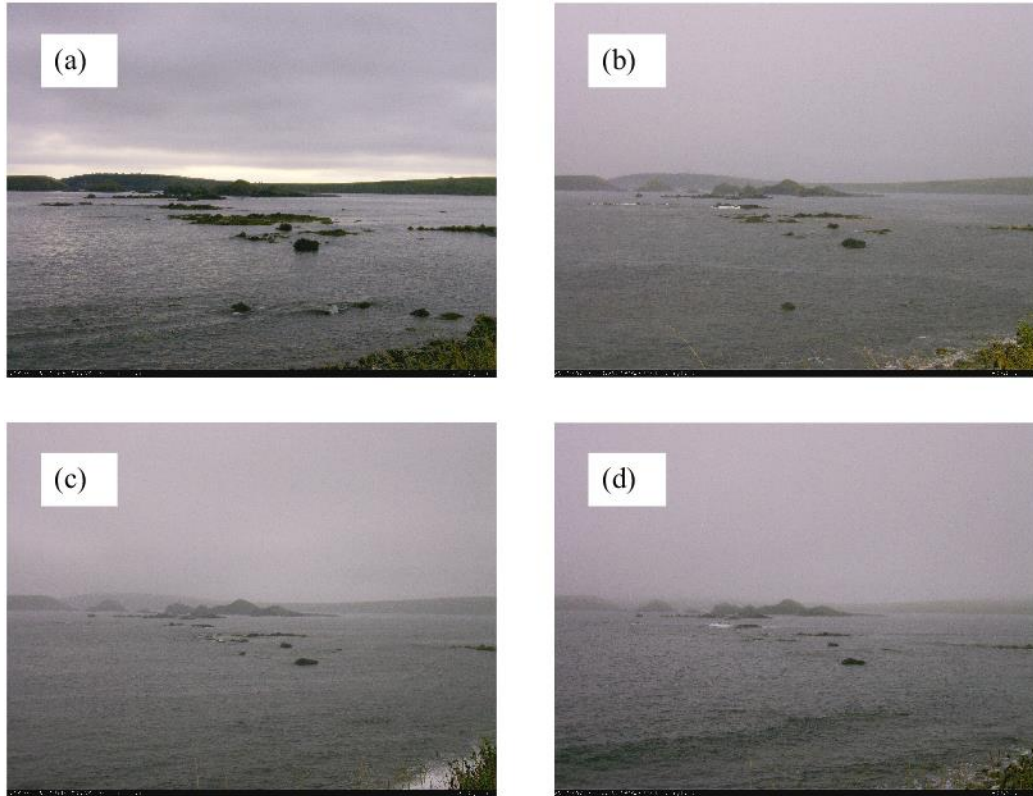
180 At 0000 UTC 16 September (local time = UTC – 2.5 h) the winds were westerly and ~ 10
 181 m s⁻¹. The wind speed steadily decreased to ~ 2.5 m s⁻¹ until 0900 UTC and remained
 182 constant until 1130 UTC (to be shown later, Fig. 7a). The visibility of all Ferryland sites
 183 observed by PWDs was similar until this time (Fig. 4a) but then started to drop at all
 184 stations. At ~ 1200 UTC, fog/mist appeared at the low-lying coastal stations (Downs and
 185 Battery) with visibility teetering around and dropping below 1 km. The *LWC* and the
 186 droplet number concentration N_d at Downs first spiked at the time of first fog/mist
 187 occurrence and then dropped somewhat until 1300 UTC where visibility dropped again
 188 and *LWC* and N_d increase (Fig. 4b,c), with the appearance of fog/mist again. The mean
 189 droplet diameter also increased from 4 to 6 μm at ~ 1200 UTC (Fig. 4d). This agrees well
 190 with a study conducted on the Californian west-coast (Goodman 1977), where coastal fog
 191 was observed to have a mean droplet diameter ranging from 4.5 to 10 μm . Inspection of
 192 ceilometer backscatter (Fig. 4f) and the observations described below show that there is an
 193 intense mixing event starting at ~ 1145 UTC but without a clear cloud ceiling even after



194

195 **Fig. 4** Time series of (a) visibility, (b) liquid water content, (c) droplet number density, (d) mean droplet
 196 diameter, (e) droplet number density spectrum, and (f) ceilometer backscatter profile during IOP7 (16
 197 September 2018). The visibility was measured with a PWD, the backscatter profile with a ceilometer, and
 198 the other variables are measured with a CDP-2

199 the appearance of fog at 1200 UTC and until about ~ 1230 UTC (ceilometer at the Downs
 200 malfunctioned, and hence data from the Battery site is shown). In Fig. 4f, there appears to
 201 be a stratus cloud layer from 200-300 m prior to the event, and there is no gradual cloud

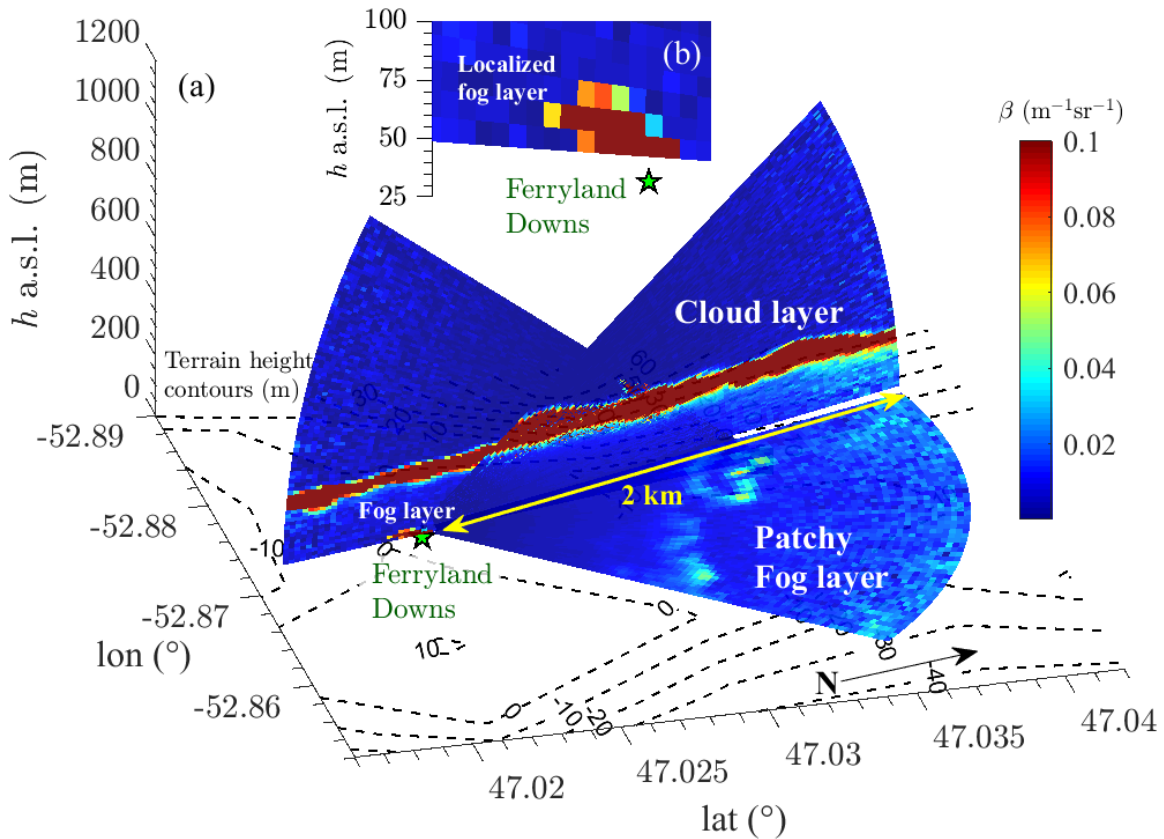


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203 **Fig. 5** Visibility observations at Downs during IOP7 (16 September 2018); photos from a camera at the
 204 Battery site at time (a) 1100 UTC, (b) 1220 UTC, (c) 1310 UTC, (d) 1640 UTC

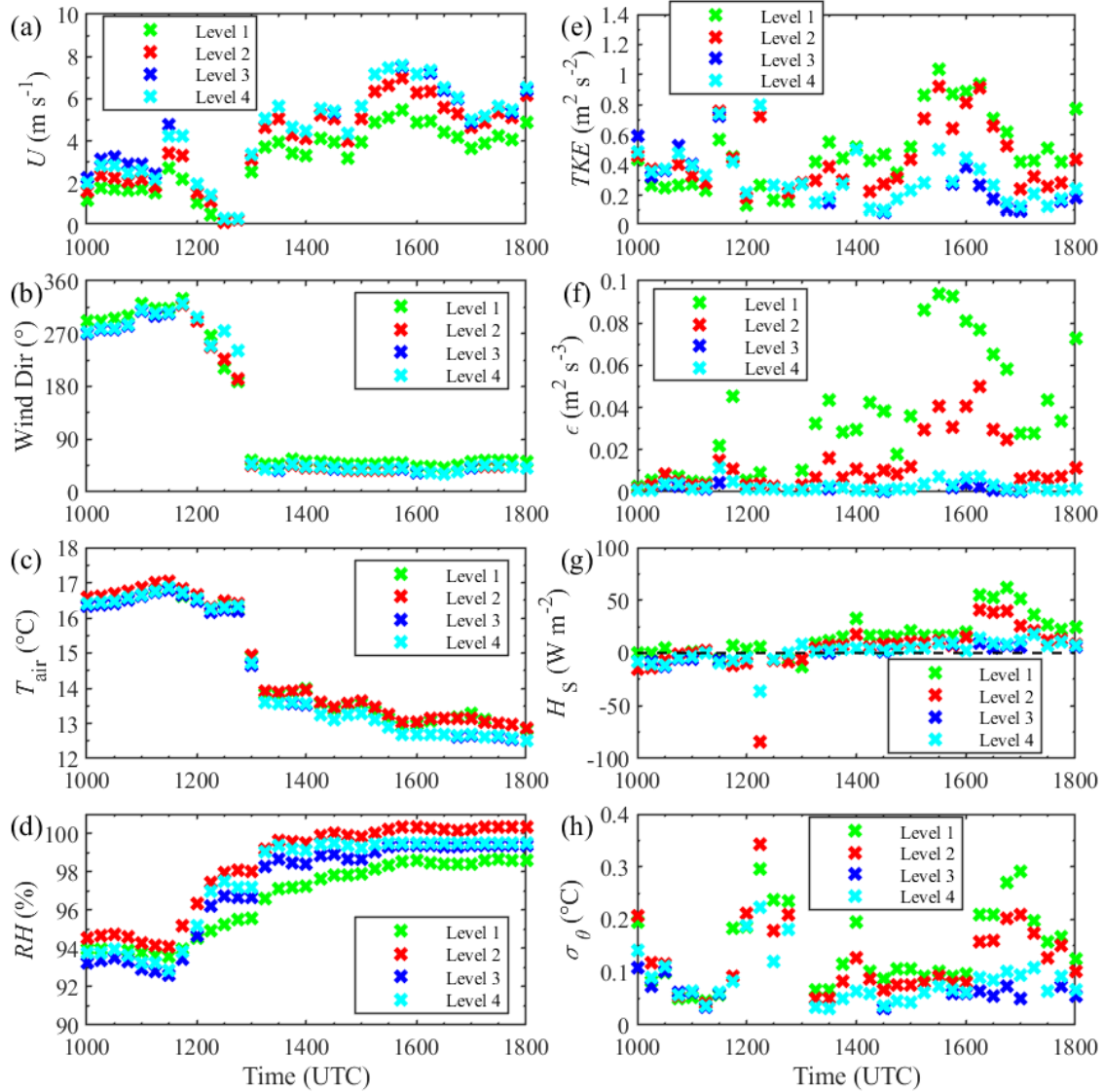
205 lowering leading to the event; instead, an abrupt mixing event appeared, following which
 206 the cloud layer re-established, whereupon another mixing event followed at 1300 UTC.
 207 Visual observations by a camera directed at Downs also confirmed a fog layer with a
 208 diffused top (Fig 5b,c). Range-corrected backscatter profiles from Doppler lidars give a
 209 view of the fog event. A scan taken at 1215 UTC shows the presence of a fog layer that is
 210 ~ 45 m thick above Downs (Fig. 6b).

211 The cloud base height was approximately 270 m a.s.l. or 240 m a.g.l. (Fig. 6a). The vertical
 212 mixing evident spanning the entire air column above Downs Topography evident in Fig.
 213 4f is not clear here, but a fog layer over Downs and the remnant cloud layer are very clear.
 214 Since Doppler lidars were scanning at an angle away from Downs, the Doppler lidar beams
 215 were not attenuated as observed by the ceilometer measurements (vertically pointing) in
 216 Fig 4f. Overall, the results suggest that the fog event observed in IOP7 is not due to stratus
 217 lowering. Tower data at Downs indeed confirmed the occurrence of an intense turbulent
 218 mixing event at 1145 UTC with (i) a gradual change of winds from westerly to north-



219
 220 **Fig. 6** Range-corrected backscatter from Doppler lidars located at Downs and Battery sites at 1215 UTC
 221 during IOP7 (16 September 2018). (a) shows range-corrected backscatter profiles for the entire range of the
 222 lidars, and (b) inset shows a zoomed section of backscatter measurements above Downs. The fog layer above
 223 Downs is about 45 m thick. The base of the cloud layer is ~ 270 m a.s.l. The dashed contours represent the
 224 terrain height in m. (adapted from Fernando et al. 2020)

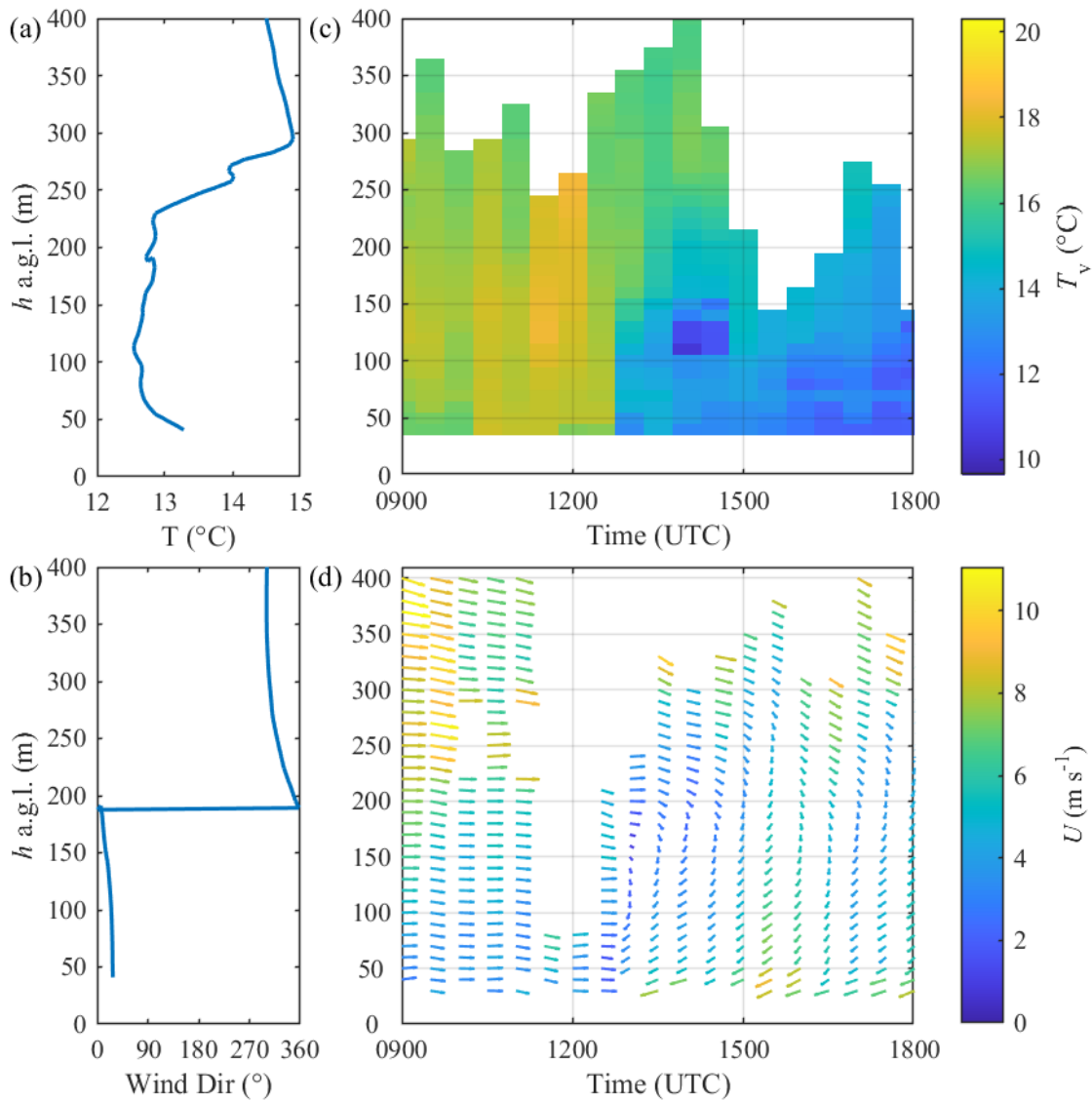
225 westerly with a temporary gust and then stagnation during 1215-1300 UTC (Fig. 7a,b); (ii)
 226 a gradual drop of near-surface T (Fig. 7c); (iii) an increase of RH (Fig. 7d); and (iv) increase
 227 of turbulent kinetic energy (TKE), its dissipation rate (ϵ) as well as the rms temperature
 228 fluctuations (σ_{θ}) (Fig. 7e,f,h). No significant changes of surface heat flux occurred except
 229 at the two lowest levels indicating momentarily large negative fluxes, perhaps due to
 230 undercutting of a cold front as a paint stripper and raising colder air past the sensors (Fig.
 231 7g). All these observations are consistent with the possibility of a local turbulent mixing
 232 event produced by mixing of two air masses of near saturation (existing air mass $RH \sim$
 233 94%, after mixing rising to ~ 98 -100%) that produced fog. The temperature difference
 234 between the airmasses was $\sim 4^{\circ}\text{C}$. The absence of fog at the Judges Hill site during the
 235 event confirms that this is a local event confined to lower altitudes (and unrelated to



236

237 **Fig. 7** Time series of (a) wind speed, (b) wind direction, (c) air temperature, (d) relative humidity, (e) *TKE*,
 238 (f) *TKE* dissipation rate, (g) sensible heat flux and (h) standard deviation of sonic temperature. The data was
 239 collected from a 16.2 m flux tower at Downs with four levels during IOP7 (16 September 2018). The wind
 240 speed, wind direction, and turbulence statistics are calculated from 15-min averaged 20 Hz ultrasonic
 241 anemometer measurements. Temperature and relative humidity were measured using 1 Hz T/RH sensors.
 242 Four levels are 2 (level 1), 5 (level 2), 10 (level 3) and 15 (level 4) m a.g.l. (also see Fernando et al. 2020 and
 243 Grachev et al. 2021)

244 possible stratus lowering) at least during the stagnation event (1215-1300 UTC). We
 245 hypothesize that a north-easterly cold front arriving approximately along the coast and
 246 impinging on the headland protruding to the ocean is a possible cause for this event. The
 247 appearance of fog ~ 15 minutes after the initiation of the event at 1145 UTC supports the



248

249 **Fig. 8** Vertical profiles during IOP7 (16 September 2018) of (a) temperature and (b) wind direction as
 250 measured by a radiosonde. The profile was taken at 1445 UTC at the Downs site. SODAR/RASS profiles of
 251 (c) virtual temperature and (d) wind vectors from the Downs site

252 idea that this is a mixing fog event, since the event is triggered during the arrival of frontal
 253 intrusion at the Downs topography, wherefore turbulent mixing is initiated. Thereafter
 254 some time is needed for frontal-scale mixing to permeate to microphysical levels to spawn
 255 fog at sub-Kolmogorov scales. This will be addressed using a laboratory experiment later.

256 A somewhat broader perspective of the above event could be obtained using Sodar/RASS
 257 measurements at Downs. This system measures wind components and temperature above

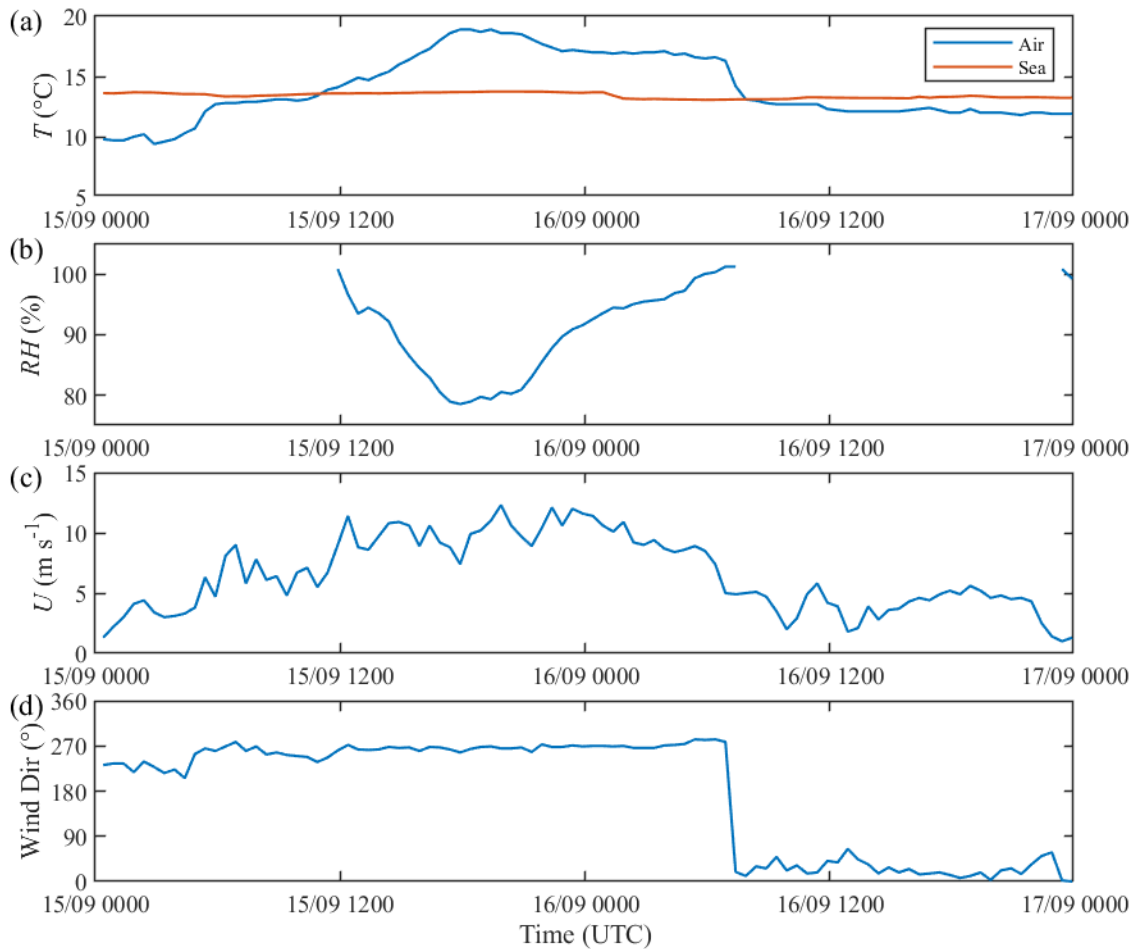
258 ~ 30-40 m, and 30 min averaged wind vectors and temperature are shown in Fig. 8c,d. The
259 Sodar/RASS registered a change of winds to north-easterly after 1300 UTC, indicating that
260 the changes of wind velocity at lower levels at 1130 UTC in Fig. 7a,b is due to the arrival
261 of the shallow gravity current nose at ~ 1145 UTC, followed by a trailing current of greater
262 thickness (for a review, see Simpson 1999). The front impinges on the Downs topography,
263 stagnates momentarily during the flow adjustment, accelerates due convergence over
264 topography, and causes mixing at low levels (Fig. 7). The turbulent front is then advected
265 past the topography, establishing a quasi-steady trailing current, which appears to occur at
266 ~ 1300 UTC. This transition is evident from wind velocity and T records in Figs. 7a,b and
267 8, which show a flow establishment at ~ 1300 UTC. From Sodar/RASS data (Fig. 8c,d),
268 it appears that the thickness of the trailing gravity current, estimated using the height where
269 T drops to background values, is about ~ 200-250 m, which is in agreement with the
270 radiosonde launch made at 1445 UTC shown in Fig. 8a,b.

271 To further investigate whether intense mixing at Downs at 1145 UTC is due to the arrival
272 and impingement of a gravity current nose with the topography, it is possible to compare
273 measurements with the theoretical trailing flow velocity u_g of gravity currents (Simpson
274 1999)

$$275 \quad u_g = \sqrt{gh \frac{\Delta T}{T_0}} \quad (1)$$

276 where g is the gravitational acceleration, h the height of the gravity current, ΔT the
277 temperature difference and T_0 the reference temperature. Figure 7c indicates $\Delta T \sim 4$ K,
278 $T_0 = 288$ K, and using $h \sim 250$ m, it is possible to calculate $u_g = 5.8$ m s⁻¹. This
279 approximately agrees well with the observed gravity current trailing velocity of ~ 6 m s⁻¹
280 (registered after 1145 UTC). As evident from Noh and Fernando (1991, 1993), the estimate
281 (1) is only approximate, as other factors such as background turbulence and surface friction
282 play a role in determining frontal propagation.

283 To reassert the gravity-current attributes of flow preceding the fog event, key
284 meteorological parameters from a coastal buoy located north of St. Johns are shown in Fig.
285 9. Note the sudden drop of the temperature starting at 0630 UTC September 16 with a sharp

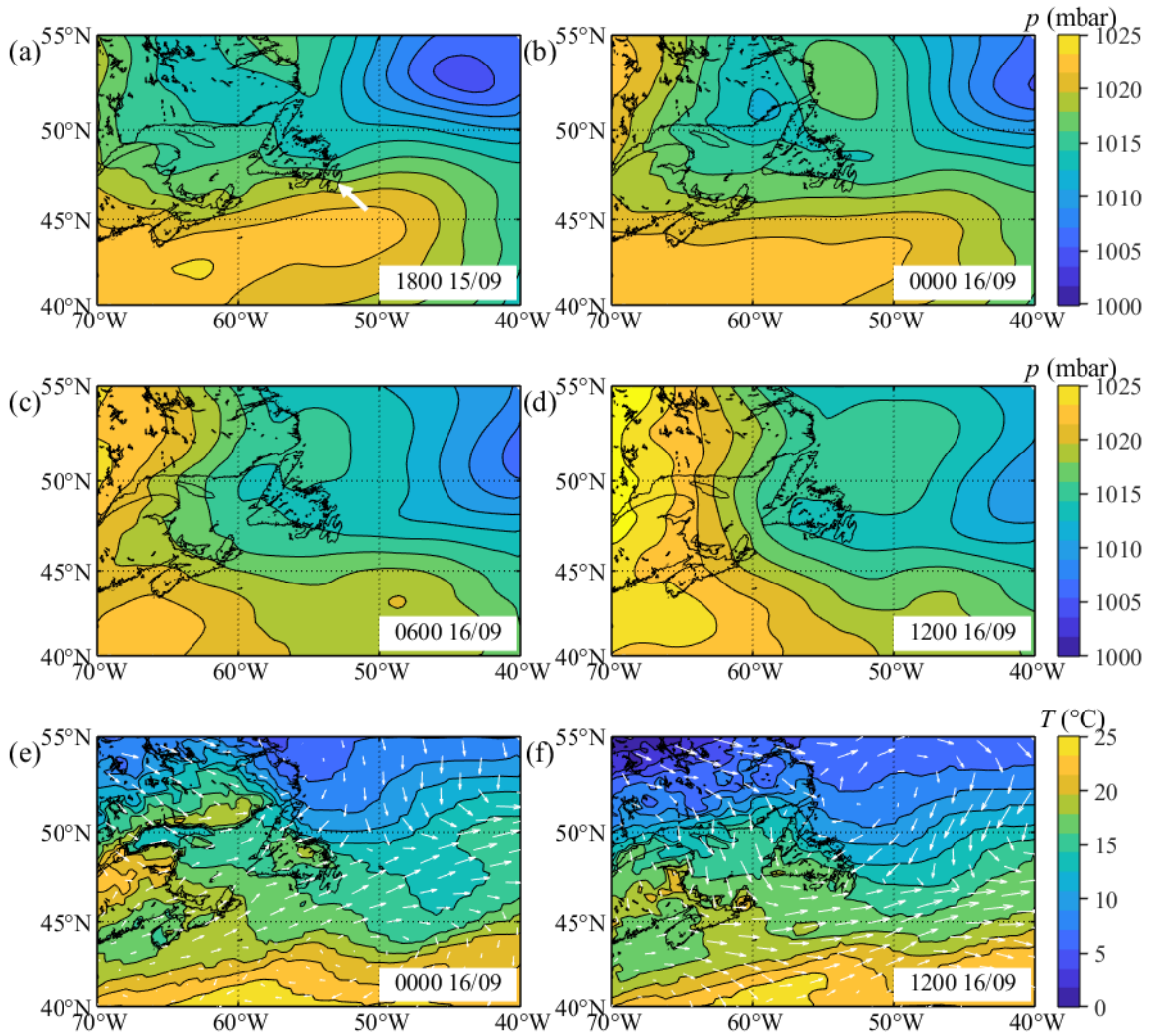


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287 **Fig. 9** Time series of (a) air and sea surface temperature, (b) relative humidity, (c) wind speed, and (d) wind
 288 direction recorded by a buoy located 6 km east of St. John's. The data was curated by the Smart Atlantic
 289 Alliance through the Memorial University of Newfoundland

290 change of wind direction from westerly to north-easterly and a drop of velocity to ~ 5
 291 m s^{-1} , which can be identified with a south-westward propagating colder gravity current. If
 292 this current were to continue south-westward with a frontal flow velocity $\sim 5 \text{ m s}^{-1}$, this
 293 gravity current is expected to arrive in Ferryland located $\sim 60 \text{ km}$ south of St. Johns within
 294 about 4.5 hours or around the observed time. The front is nearly saturated, and its arrival in
 295 Ferryland raises the relative humidity therein close to saturation following the mixing fog
 296 event.

297 The synoptic conditions surrounding the event are of interest, given our claim that a
 298 gravity-driven rather than a pressure-gradient driven flow is dominant after 1145 UTC.
 299 Figure 10 shows synoptic maps of sea level pressure, 2-m air temperature, and 10-m wind



300

301 **Fig. 10** NARR (a-d) sea level pressure plots and (e-f) 2 m air temperature and 10 m wind vectors for
 302 various dates. The time stamps are given in the plots. The white arrow in (a) points to Downs

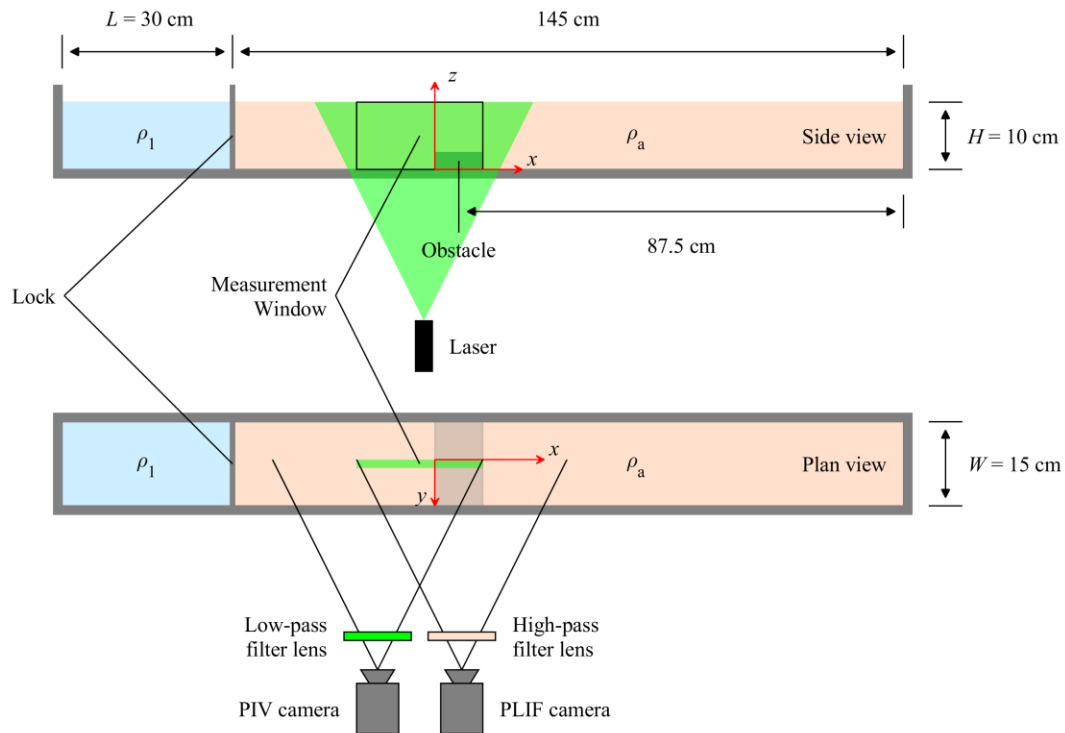
303 vectors during IOP7, indicating an eastward moving Low about 5° north of Ferryland (at ~
 304 53° N) on September 15, consistent with the easterly flow observed before switching to
 305 northerly at 1130 UTC on 16 September. The low has cleared the general area by 1200
 306 UTC 16 September, paving way to the north-easterly gravity current. Figure 10e,f shows
 307 that cold air in the wake of the Low is moving south-westward as a frontal system,
 308 undercutting warmer air, and its gravity driven nature is supported by (i) the broad
 309 agreement of flow velocity with the theoretical formula described above, and (ii) the
 310 southward propagation of cold air mass in the 2-m temperature maps of Fig. 10b that
 311 indicates ~ 16°C air over Ferryland at 0000 UTC is replaced by ~ 12°C air at around 1200
 312 UTC.

313 After 1300 UTC, fog at Downs and Battery subsided, and the visibility gradually increased
314 to ~10 km, but a strong fog presence was noted at the Judges Hill site where the visibility
315 dropped to ~ 300 m and remained low until 2030 UTC, except for some sporadic (> 1 km)
316 increases of visibility (Fig. 4a). This observation is consistent with the ceilometer
317 observations of Fig. 4f that shows the development a stratus cloud layer shrouding the
318 Judges Hill instrumentation (~129 m a.s.l.). This stratus-based fog development process is
319 different from that at low levels, which is due to mixing induced by a gravity current
320 impinging on the (Downs) topography. Prior to 1145 UTC September 16, the cloud base
321 was too high (~200-250 m) for cloud-shrouding to affect visibility at the Judges Hill.

322 At Downs, from 1300 UTC until about 1400 UTC, there was an appreciable increase of
323 the *TKE*, modest increase of σ_θ , and elevated level of *TKE* dissipation rate ε , pointing to
324 turbulence in the trailing sheared gravity current. As mentioned, the drop of *T*, increase of
325 *RH*, and a change of wind direction to a quasi-steady state indicate that at this time, the
326 front has cleared the topography. In the trailing flow, the cloud layer is low, with its base
327 at ~ 100 m (and the estimated cloud top at ~200-250 m), but the clouds stay clear from the
328 ground. Between 1400 UTC and 1600 UTC, there is an increase of turbulence activity,
329 possibly because of an increase of local wind speed. Intermittent stratus lowering occurred
330 between 1530 and 1700 UTC, briefly enveloping Battery and Downs sites, as indicated in
331 Fig. 4a,f, and observed from the camera (Fig. 5d). During this stratus lowering period, in
332 general there were increased levels of *TKE*, ε , and σ_θ , pointing to enhanced turbulence
333 responding to an increase of wind speed. At ~ 1700 UTC, the cloud layer started rising,
334 even gradually clearing fog at Judges Hill. Complete clearing of fog at Judges Hill occurred
335 after 2000 UTC, whence the base of the stratus cloud was about ~150 m. In all, local small-
336 scale activities dominated Ferryland during and after fog/mist formation processes.

337 **3 Laboratory Experiments**

338 The laboratory experiments were directed at understanding the initial stages of the
339 interaction of a gravity-current induced front with an obstacle to shed light on observations
340 made during IOP7.



341

342 **Fig. 11** The experimental setup for producing lock-exchange gravity currents. The gravity current was created
 343 by lifting the lock. PIV and PLIF techniques were used for flow diagnostics

344 **3.1 Experimental (lock-exchange) configuration**

345 The experiments were carried out in a 175 cm long, 15 cm wide, and 30 cm high Plexiglas
 346 tank. A gate (lock) was placed 30 cm from the left side of the tank that separated the dense
 347 fluid (density ρ_1) from the lighter fluid (ρ_a). The gravity current was generated by lifting
 348 the lock instantaneously. Figure 11 shows a schematic of the experimental configuration.
 349 A 1 W 520 nm continuous-wave diode laser was used to generate a laser sheet that
 350 illuminated the centre vertical section of the tank. The velocity and density fields were
 351 measured simultaneously using a time-resolved PIV/PLIF system. Two separate cameras
 352 were used for the particle image velocimetry (PIV) and planar laser-induced fluorescence
 353 (PLIF) measurements. The cameras were synchronized with a hardware trigger. Low- and
 354 high-pass filters (in wavelength) were mounted on the PIV and PLIF cameras to filter out
 355 unwanted signals.

356 Separate experiments were conducted with (obstructed) and without (unobstructed)
357 topography. For obstructed runs, a 10 cm long and 2 cm high rectangular Plexiglas block
358 that spans the tank width was placed at the centre of the tank. Ten independent replications
359 were conducted with identical experimental parameters for each unobstructed and
360 obstructed case to aid ensemble averaging. The fluid layer depth was $H = 10.0$ cm, which
361 was selected to mimic the length scale ratio as in the field; for this case, the obstacle height
362 to H ratio is 0.2, and in the field it is about 0.15. The dense fluid was prepared with a salt
363 solution ($\rho_1 = 1002.3$ kg m⁻³), and an aqueous ethanol solution with $\rho_a = 993.4$ kg m⁻³
364 was used as the lighter fluid. The densities were chosen such that the refractive indices
365 between the two fluids are matched to prevent laser-beam distortions. The density and
366 refractive indices were measured with a Mettler Toledo Densito 30 PX densitometer and a
367 Leica handheld analogue refractometer. Hannoun et al. (1988), Strang and Fernando (2001),
368 and Xu and Chen (2012) provide further details on the refractive-index matching technique.

369 **3.2 Velocity measurements**

370 Two-dimensional instantaneous velocity fields were obtained using high-frame-rate PIV.
371 Both the dense and lighter fluids were seeded with hollow glass spheres with a median
372 diameter of 10 μm . The 1 W 520 nm continuous-wave laser illuminated the particles. An
373 IDS UI-3360CP-M USB 3.0 camera, equipped with a 2048 x 1088 pixels CMOS sensor
374 and a 50 mm f/2.0 lens, was used at a frame rate of 40 Hz. A low-pass filter mounted on
375 the lens filtered out the fluorescence of the PLIF dye from the laser light. The particle
376 frames were processed using the MATLAB PIVlab package (Thielicke 2014; Thielicke
377 and Stamhuis 2014; Thielicke and Stamhuis 2019), which is based on the iterative
378 multigrid image deformation method (Scarano 2001). The sampling frequency and
379 interrogation window size were chosen such that they met the one-quarter rule (Adrian
380 1991). Consequently, the window size of the final selection was 16 x 16 pixels with a 50%
381 overlap.

382 **3.3 Density measurements**

383 Time-resolved PLIF measurements provided density fields within gravity currents.
384 Rhodamine 6G (R6G) was chosen as the fluorescent dye because of its high quantum

385 efficiency, high resistance to photo-bleaching, and because its absorption peak of 525 nm
386 (Crimaldi 2008) is close to the laser wavelength (520 nm). The dye was added only to the
387 lighter fluid at an initial concentration of $100 \mu\text{g L}^{-1}$. The fluorescence intensity was
388 recorded by an IDS uEye UI-1220-C USB 2.0 camera with a 752 x 480 pixels CMOS
389 sensor. A high-pass filter with a cut-off wavelength of 550 nm was mounted on the PLIF
390 camera lens to filter out the laser light. The majority of the fluorescence was retained
391 because the fluorescence spectrum of R6G has a peak at 555 nm (Penzkofer and Leupacher
392 1987). The R6G concentrations were calculated from the grey values using the calibration
393 technique described by Xu and Chen (2012). Local density values were derived from the
394 measured R6G concentration using a calibration curve.

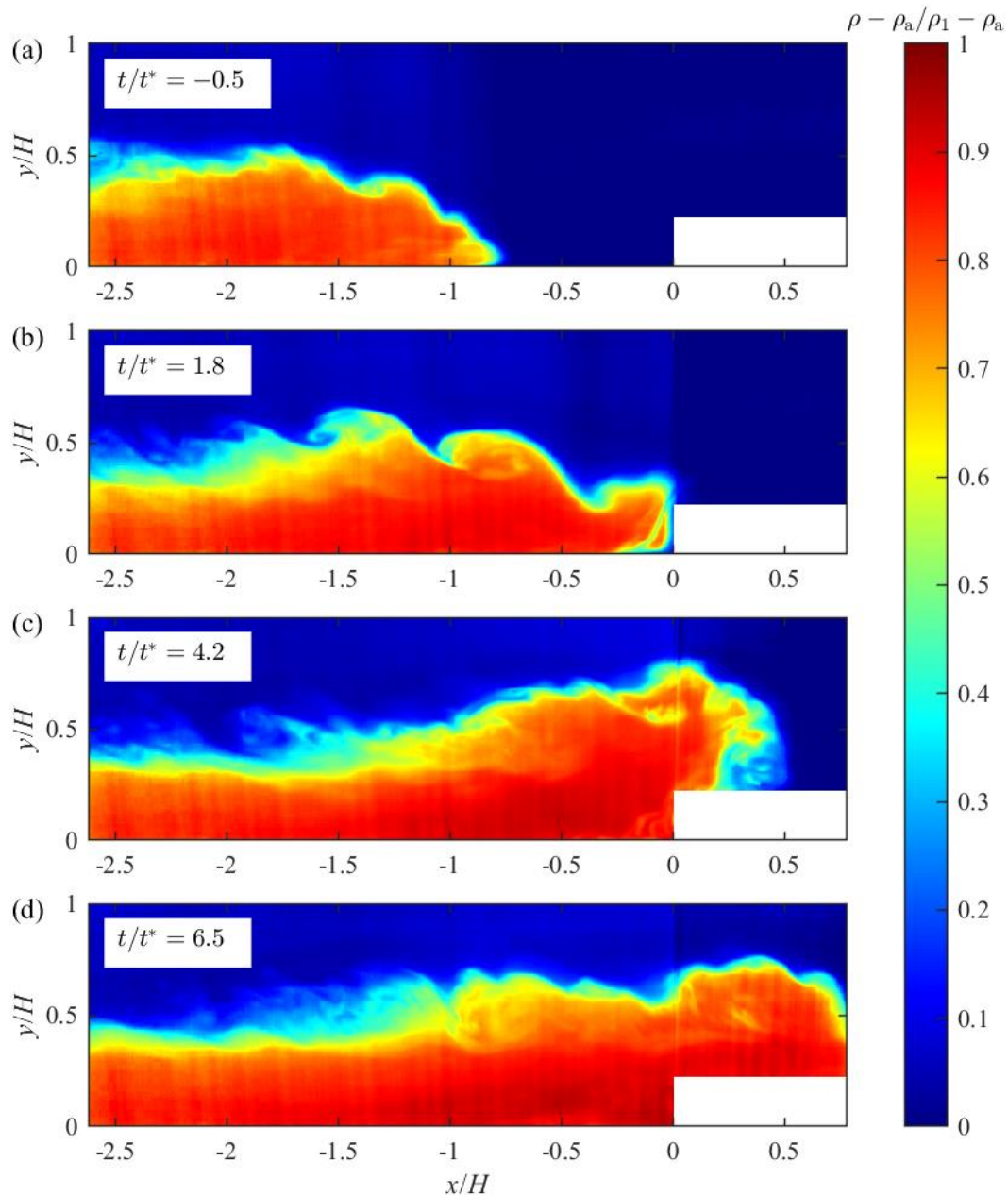
395 **3.4 Phase-aligned ensemble averaging technique**

396 Turbulence statistics of obstructed and unobstructed gravity currents were obtained by
397 ensemble averaging because of the nonstationarity and spatial inhomogeneity of gravity
398 currents. Similar to Zhong et al. (2018, 2020), we applied the phase-aligned ensemble
399 averaging technique (PAET), which aligns the time and space coordinates of each
400 realization to minimize jitter. It iteratively maximizes the cross-correlation of individual
401 realizations with the ensemble-averaged field by shifting the realizations along space and
402 time axes. In this study, the alignment was only necessary for time since the horizontal
403 variations are fixed by the position of the obstacle (or centre of the tank), and vertical
404 variations are limited by the tank bottom and the water surface. More details of PAET as
405 applied to gravity currents are given in Zhong et al. (2020).

406 **4 Observations**

407 **4.1 Observations of obstructed gravity currents**

408 The gravity current, formed after lifting the lock, propagated towards the obstacle and
409 impinged on it. An example of the instantaneous normalized density field of an obstructed
410 run is shown in Fig. 12. Based on the horizontal frontal velocity u_F the flow could be
411 divided into four stages. In the first stage, the gravity current propagated independently of
412 the obstacle at a constant speed. In stage two, as the gravity current approached the obstacle,

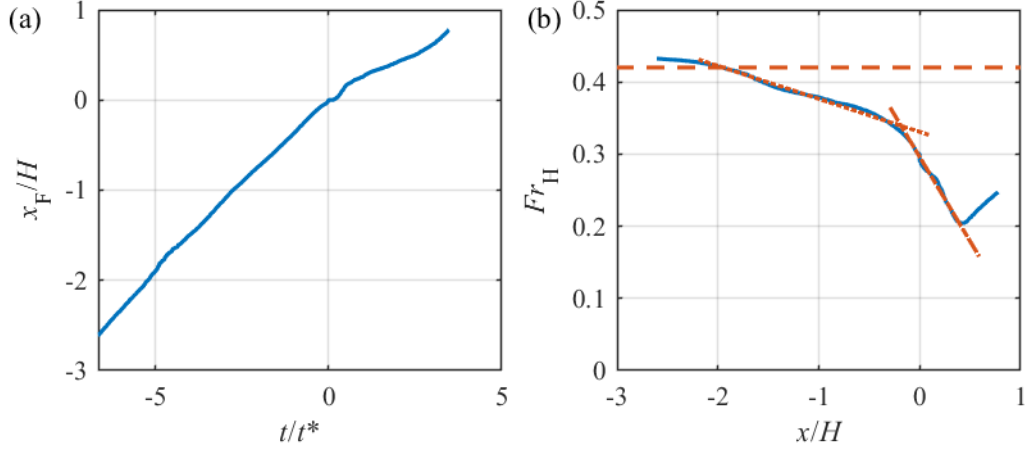


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Fig. 12 The instantaneous normalized density field for an obstructed run. The time origin is set to the instance where the gravity current makes contact with the obstacle

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the horizontal front velocity gradually decreased with time. The collision of the gravity current with the obstacle is the third stage, which is characterized by a low horizontal frontal velocity and near stagnation. A part of the current deflected vertically similar to a jet, reaching a maximum height of about $0.8H$. A vortex was formed at the leading edge of the obstacle that provided enhanced mixing. The current was also partially reflected as a hydraulic jump moving upstream. At the start of the fourth stage, the gravity current collapsed and continued to propagate over the obstacle.



423

424 **Fig. 13** Propagation of the gravity current toward the obstacle, with (a) normalized frontal position as a
 425 function of time and (b) normalized frontal velocity as a function of normalized frontal position (shown by
 426 lines, the different phases). The dashed lines indicate various phases of propagation described in the text. For
 427 this experiment, $\rho_1 = 1002.3 \text{ kg m}^{-3}$, $\rho_a = 993.4 \text{ kg m}^{-3}$, and $H = 10 \text{ cm}$

428 According to the hydraulic (inviscid) analysis of Benjamin (1968), an unobstructed gravity
 429 current in its energy conserving mode has a height of $h_g = 0.5H$ and a Froude number of
 430 $Fr_H = \frac{u_F}{\sqrt{g'H}} = 0.5$, where u_F is the frontal velocity, $g' = g(\rho_1 - \rho_a)/\rho_0$ the reduced
 431 gravity, and $\rho_0 = (\rho_1 + \rho_a)/2$ is a reference density. Previous laboratory and numerical
 432 experiments have shown that in practice Fr_H has values between 0.36 and 0.45 for gravity
 433 currents produced by lock-exchange (Zhong et al. 2018). In the present unobstructed
 434 experiments, it was found that $Fr_H \approx 0.42$, which agrees well with the observation that
 435 gravity currents, after their initial development, typically have a constant velocity for about
 436 5-10 lockbox lengths L (Meiburg and Kneller 2010). This is called the slumping phase. In
 437 our case, the obstacle is located $3L$ from the lock, and hence observations are in the
 438 slumping phase. Figure 13 shows ensemble-averaged results for a run, where the
 439 normalized frontal position x_F relative to the obstacle leading edge ($x = 0$) is shown as a
 440 function of time. Here x_F was measured using PLIF by locating the largest x -value of the
 441 $\rho^* = (\rho - \rho_a)/(\rho_1 - \rho_a) = 0.1$ contour.

442 The normalization variables were selected based on the intrinsic velocity, length, and time
 443 scales of unobstructed runs: $\sqrt{g'H}$, H and $t^* = \sqrt{H/g'}$, respectively. Note that near the
 444 obstacle the horizontal frontal speed is significantly reduced (Fig. 13a,b) as the front is

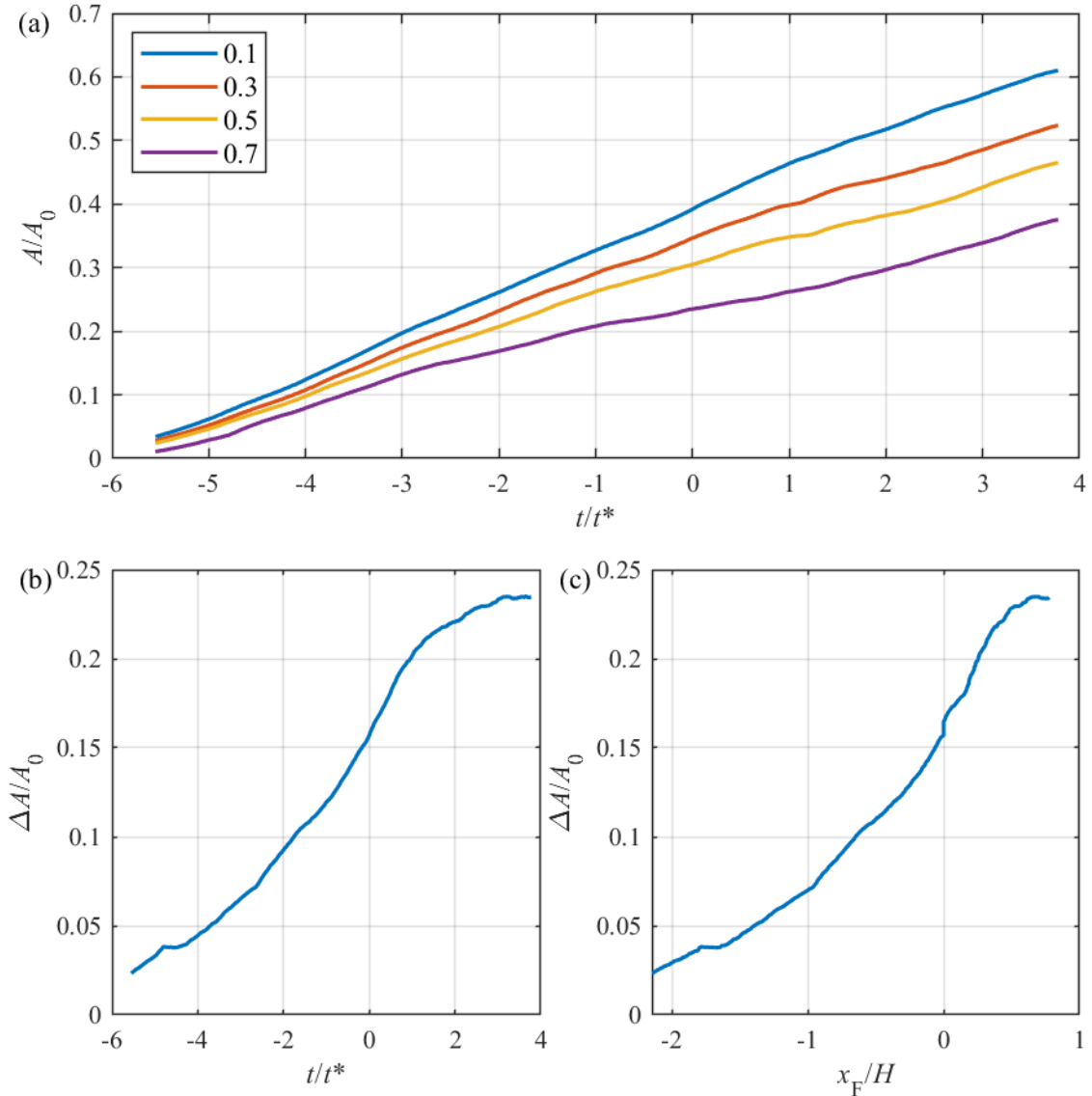
445 deflected upward and start overrunning the obstacle (Fig. 12c,d).

446 The Froude number Fr_H can be interpreted as a measure of the ratio of the kinetic energy
447 of fluid parcels at the gravity current front ($\sim u_F^2$) to the potential energy acquired by their
448 rise over the height of the current ($\sim g'H$) (c.f., Lane-Serff et al. 1995). Fr_H is an important
449 quantity in specifying an unobstructed gravity current, but when it interacts with an
450 obstacle an additional parameter $1/G = u_F / \sqrt{g'd} = Fr_H (\sqrt{H/d})$ becomes important (d
451 being the obstacle height). In general, the interaction between a gravity current and a 2D
452 obstacle is a multi-parameter problem, determined by upstream Fr_H , d/H and other
453 geometric ratios; for discussions, see Baines (1998) and Vassallo et al. (2021). The
454 variation of Fr_H as a function of the front position x_F/H is shown in Fig. 13b, where in the
455 unobstructed phase the Froude number maintains an approximately constant value of
456 $Fr_H = 0.42$ until the gravity current is at a distance $2H$ away from the obstacle (i.e., the
457 first stage). As mentioned, this is valid for all runs conducted. The flow evolution beyond
458 the first stage is expected to be determined by additional parameters such as d/H .

459 As evident from Fig. 13b, during the second stage, the Froude number based on local
460 u_F slowly decreases with time to about 0.35, and this upstream influence is due to stronger
461 return flows established in the proximity of the obstacle. When the gravity current collides
462 with the obstacle in stage three, Fr_H decreases significantly to a minimum value of 0.2 or
463 $G \approx 0.45$. The decrease is associated with the vertical deflection of the gravity current
464 during the collision. The Froude number slowly increases again as the gravity current
465 continues to propagate over the obstacle, after adjusting to the new propagation state. Given
466 the interest of this paper is mixing during the collision stage, detailed studies were not
467 conducted on how local Fr_H varies with d/H and other geometric parameters or the
468 properties of the layer overflowing or reflected back from the obstacle.

469 **4.2 Mixing**

470 To quantify enhanced mixing at the obstacle between the gravity current front and
471 background fluid, the variability of mixed fluid volume was measured. Given the high
472 space-time variability and steep concentration gradients involved, direct measurements of



473

474 **Fig. 14** Time series of (a) four different isopleths; (b) the difference of $\rho^* = 0.1$ and $\rho^* = 0.7$ isopleths as a
 475 function of normalized time; (c) the difference of the $\rho^* = 0.1$ and $\rho^* = 0.7$ isopleths as a function of
 476 normalized frontal position x_F/H

477 molecular-scale mixing rates (or dissipation rates of a scalar) are untenable with the
 478 resolution of the current (and other available) techniques, but the amount of mixed fluid
 479 present can be quantified in an integral sense by measuring the time evolution of the density
 480 field over the domain of the gravity current. In our case, because of the limited field of
 481 view, it was not possible to probe the entire gravity current, but a sound quantitative
 482 estimate of mixing could be obtained by considering the time evolution of mixed fluid
 483 within the field of view. Figure 14a shows the normalized volumes of ‘mixed’ fluid per

484 unit tank width, where A_0 is the volume per tank width of the measurement window
 485 excluding the obstacle and A the volume per tank width of fluid parcels with a density at
 486 or above a given density threshold. The density thresholds in Fig. 14a are $\rho^* = (\rho -$
 487 $\rho_a)/(\rho_1 - \rho_a) = 0.7, 0.5, 0.3,$ and 0.1 . Curves for a given ρ^* threshold will be referred to
 488 as an isopleth. Note that the fractional area A/A_0 of the isopleths increases with time because
 489 more gravity current fluid is arriving in the field of view. If no further mixing takes place
 490 in the field of view, then the fraction A/A_0 ought to increase linearly with time t (i.e.,
 491 dA/dt is constant) in Stage 1, given the constant speed of the gravity current feeding the
 492 mixing area. When mixing takes place, dA/dt is also dependent on the rate of mixing,
 493 whence fluid with lesser densities are generated due to mixing.

494 As pointed out by Hacker et al. (1996), mixing in a flow can be quantified by considering
 495 the evolution of isopleths. That is, the volume of fluid for a given normalized density range
 496 can be reckoned by subtracting the A/A_0 values corresponding to two isopleths. This is
 497 illustrated in Fig. 14b,c for temporal and spatial characteristics. For example, the volume
 498 of mixed fluid in the range of $0.1 \leq \rho^* < 0.7$ can be found by taking the difference of A/A_0
 499 values of 0.1 and 0.7 isopleths, with the divergence of two isopleths being an indicator of
 500 how the density distribution changed due to mixing by reducing the A/A_0 ratio
 501 corresponding to the $\rho^* = 0.7$ and increasing that of the $\rho^* = 0.1$ isopleth. In order to
 502 quantify the mixing, we will focus on the divergence of the $\rho^* = 0.1$ and 0.7 isopleths,
 503 $\Delta A/A_0$ by taking the difference of A/A_0 at a given instant. In Fig. 14 until $t/t^* \approx -1.5$
 504 the rate of increase of $\Delta A/A_0$ is approximately constant. Between $t/t^* \approx -1.5$ and $t/t^* \approx$
 505 1.5 , the divergence of the isopleths increases significantly faster with time (Figs. 14b,c),
 506 which coincides with the arrival of the front at the obstacle (Stage 3) wherein mixing
 507 enhances due to the collision of the gravity current with the topography at $t/t^* = 0$. After
 508 $t/t^* \approx 1.5$ the divergence decreases again, and thus there is less mixing taking place.

509 The period $\Delta t/t^*$ where enhanced mixing takes place can be evaluated by finding the
 510 period beyond which the slope of $\Delta A/A_0$ drops to half of its maximum. When applied to
 511 Fig. 14, its normalized value could be evaluated as $\Delta t/t^* \approx 3$. This period corresponds to
 512 frontal positions over the range $-0.5 < x/H < 0.33$, which surrounds the leading edge of
 513 the obstacle. These temporal and spatial scales of enhanced mixing, when applied to IOP7,

514 yield a duration of ~ 2.8 min over a length of ~ 400 m for $H = 2h_g = 450$ m, $g' = 0.15$
515 m s^{-2} . This indicates that scalar (water vapour and temperature) mixing is permeated to
516 molecular scales quite rapidly after large-scale turbulence is generated by impingement of
517 gravity current on the topography, thus facilitating the activation of condensation nuclei
518 within the smallest scales of turbulence (i.e., Kolmogorov scale) and appearance of fog
519 over a short time scale. According to Fernando and Hunt (1996), the time scale for
520 homogenization when two dissimilar concentrations are brought together in a turbulent fluid
521 is ℓ/σ , where ℓ is the integral length scale and σ the *rms* velocity of turbulence. For the
522 case of Downs topography, this can be estimated using $\ell = 30$ m and $\sigma = \sqrt{TKE} \sim 0.2$
523 m s^{-1} as 2.5 min, which is in broad agreement with the laboratory-based estimates presented
524 above. Note that the gravity current front approaches Downs topography at ~ 1130 UTC
525 (Fig. 7a-d), mixing becomes intense at ~ 1200 UTC (Fig 7e-h) and we expect the fog to
526 appear soon thereafter.

527 **5 Discussion**

528 The laboratory experiment described above was intended to provide physical insights and
529 relevant scales pertinent to gravity-current/topography interactions. In the traditional
530 (reductionist) approach to modelling, the study of complex natural flows is made tractable
531 by employing simplified yet realistic geometries in designing model configurations. In our
532 case, the long narrow stretch of Downs peninsula with a ‘tadpole’ headland (Fig. 2) was
533 approximated by a two-dimensional topography with a rectangular cross section, the
534 approach flow being normal to the long axis, notwithstanding in reality the approach flow
535 is about 45° to the long axis. The laboratory working fluid was water, which has been a
536 customary working fluid in previous laboratory simulations of atmospheric flows. In
537 particular, Chen et al. (1996, 1999) showed that the equations for the conservation of
538 momentum, mass, and buoyancy in a water-tank model are similar to those in the
539 atmosphere: the counterparts of the potential temperature and Exner function in the
540 atmosphere correspond to the specific volume and pressure, respectively, of the
541 experimental fluid (also see Berman et al. 1995). The displacements, velocity, and
542 temperature fields play similar roles in the atmosphere and in the model. Therefore, with
543 suitable matching of dimensionless parameters, it is possible to mimic natural flow

544 processes using water tank experiments.

545 For the laboratory case of Fig. 11, the governing variables are g' , H , the width w_0 and the
546 height d of the obstacle, and the kinematic viscosity ν of the working fluid. The ratios
547 d/w_0 and d/H were maintained at 0.2, close to that of Downs topography 0.15. With some
548 manipulation, the frontal velocity of the gravity current prior to the influence of the obstacle
549 and away from the lock can be written as $u_F/\sqrt{g'H} = \Psi(Re)$, where $Re = u_F H/\nu$ can be
550 construed as the Reynolds number. Independence on the Reynolds number (or the
551 Reynolds number similarity; Barenblatt 1996) applies at high Re , whence $u_F = C\sqrt{g'H}$,
552 where C is a constant. Previous and current experiments show a constant $C = 0.36-0.45$,
553 implying Reynolds number similarity. The value of Re for the current experiments is about
554 9300 whereas for IOP7 $Re \sim 10^9$. Breidenthal (1981) proposed that Reynolds-number
555 independent mixing is achieved when $Re > 3000$, an aspect that has been further
556 discussed by Princevac et al. (2005) and Zhong et al. (2018). Given both laboratory and
557 field Re are greater than the critical value above, we may assume that our laboratory
558 experiments can provide useful information on mixing in natural gravity currents. It is also
559 noted that the highest achievable Reynolds numbers in lock-exchange laboratory tanks are
560 of the order ($10^3 - 10^5$).

561 Worthy of mentioning are the shallow water models and high-resolution numerical
562 simulations on gravity/turbidity currents propagating past topographies, for which
563 substantial literature exists. Layered (1.5, 2, and 2.5) shallow water models have been
564 developed and refined since Rottman et al. (1985) to describe the approach, overflow and
565 reflected phases (Lane-Serff et al. 1995), but these models cannot capture the initial impact
566 and vertical flow deflection let alone mixing down to dissipation scales. Direct numerical
567 simulation (DNS) and Large-Eddy Simulation (LES) of lock-exchange (e.g., Gonzalez-
568 Juez and Meiburg 2009; Gonzalez-Juez et al. 2009; Nasr-Azadani et al. (2014, 2018) using
569 DNS; and Tokay and Constantinescu 2015; Bhaganagar 2017; Zhou and
570 Venayagamoorthy 2017; Wu and Ouyang 2020 using LES) and steady constant flux
571 (Tokay and Constantinescu (2015) using LES) gravity and turbidity currents interacting
572 with obstacles have been reported, with identification of fine-scale details of flow evolution
573 around the obstacle. These studies, however, did not identify the time scale of scalar mixing

574 (homogenization at small scales) that is interest to our study, which was prompted by the
575 field observations.**6 Conclusions**

576 An interesting short-lived mixing-fog event occurred during the IOP7 of the C-FOG project,
577 where a cold front arriving from the north-east collided on the Downs peninsula to produce
578 foggy/misty conditions. The collision caused turbulent mixing of nearly saturated ambient
579 air with the almost saturated cold front, creating fog. The conditions were such that fog
580 lasted for a short period of time, and the final mixture of airmasses quickly reached the
581 unsaturated region of the vapour pressure/temperature diagram.

582 Earlier in the day, the conditions were clear at the three Ferryland sites until 1200 UTC,
583 whence visibilities fluctuated and dropped < 1 km at the low-lying stations. At the same
584 time, the liquid water content, droplet number concentration, mean diameter, and total
585 aerosol count increased. The tower data showed a drop of temperature, an increase of
586 relative humidity, and an increase of *TKE* and its dissipation rate, and *rms* temperature
587 fluctuations. These observations suggest the possibility of a local turbulent mixing event
588 triggered by topographically induced mixing of two nearly saturated air masses with
589 different temperatures that led to fog. Measurements of a Sodar/RASS system, a
590 meteorological flux tower, a meteorological buoy located ~ 60 km upstream and synoptic
591 weather maps gave perspectives of the cold front. They pointed to a shallow gravity current
592 nose with a trailing current of about 200-250 m high arriving prior to the fog event. This
593 was confirmed by a radiosonde launch that followed. The observed velocity of the trailing
594 current agreed with the theoretical velocity.

595 Motivated by these observations, laboratory experiments were performed to study the
596 interaction of gravity currents with a rectangular obstacle to confirm enhanced mixing at
597 the topography as well as quantify the timescale for molecular-scale mixing due to the
598 collision (which is representative of the time for fog appearance following condensation
599 nuclei activation). Gravity currents were generated by lifting a (proverbial) lock that
600 separated a dense saline solution from a lighter ethanol solution that served as the ambient
601 fluid. The refractive index matching of the fluids made it possible to use optical techniques
602 for flow diagnostics. Instantaneous velocity and density fields were measured

603 simultaneously using PIV/PLIF techniques. Molecular-level mixing was quantified by
604 considering the divergence of a set of concentration isopleths. Given the logistical
605 constraints, the ratios of the obstacle height to the fluid layer depth and the aspect ratio of
606 the obstacle were kept fixed in the experiments reported in this paper, approximately
607 matched with the field conditions. Future studies ought to study the effect of these
608 parameters.

609 The propagation of the obstructed gravity current and its interaction with the obstacle
610 during experiments could be divided into four stages, namely, propagation independent of
611 the obstacle, approach, collision and continued propagation. Most of the mixing occurs in
612 the third stage due to turbulent eddies (vertical structures) generated during the collision.
613 An analysis based on isopleths confirmed increased (molecular-scale) mixing during the
614 collision, with enhanced mixing starting at $t/t^* = -1.5$ and reaching a maximum at
615 $t/t^* = 1.5$, and decreasing rapidly thereafter. The period of enhanced mixing was $\Delta t/t^* \approx$
616 3 and the extent around the obstacle that showed increased mixing was $-0.5 < x_F/H <$
617 0.33. When applied to the IOP7 case, the time scale and length scale of enhanced mixing
618 were ~ 2.8 min and ~ 400 m, respectively. The smaller space-time scales involved in the
619 natural fog situation studied in this paper points to the difficulty of simulating mixing fog
620 using mesoscale numerical weather prediction (NWP) models, since critical underlying
621 processes of fog formation belong to the sub-grid category. The success of fog predictions
622 in NWP models, therefore, is expected to be sensitively dependent on the efficacy of sub-
623 grid parameterizations.

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