

Statistical Evaluation of Satellite Ocean Color Data Retrievals

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ABSTRACT

We develop a statistical approach to evaluate the performance of the ocean color data processing system for satellite-derived ocean color data products based on temporal stability of retrievals. We use the Multi-Sensor Level-1 to Level-2 (MSL12) ocean color data processing system to obtain the normalized water-leaving reflectance $\rho_{wN}(\lambda)$ spectra from the Visible Infrared Imaging Radiometer Suite (VIIRS) measurements. The deviations of $\rho_{wN}(\lambda)$ spectra from temporally and spatially averaged values are investigated, and the statistics with respect to various retrieval parameters are collected, including the solar-sensor geometry (solar-zenith, sensor-zenith, and relative azimuth angles), and various ancillary data (surface wind speed, surface atmospheric pressure, water vapor amount, and ozone concentration). The performance of MSL12 is also evaluated with respect to other intermediate retrieval parameters. The study shows that MSL12 produces statistically consistent VIIRS ocean color retrievals in the global open ocean, with respect to retrieval geometry parameters, as well as the ancillary inputs.

Keywords: Ocean color remote sensing, atmospheric correction, data processing, and VIIRS.

1. Introduction

27 Ocean color remote sensing from space is a rapidly growing field with several new polar
28 orbiting instruments launched and recently deployed for routine global ocean property
29 measurements and monitoring, including the Visible Infrared Imaging Radiometer Suite (VIIRS)
30 on the Suomi National Polar-orbiting Partnership (SNPP) and NOAA-20 satellites, the Ocean
31 and Land Colour Instrument (OLCI) on the Sentinel-3A and Sentinel-3B satellites, and the
32 Second-Generation Global Imager (SGLI) on the Global Change Observation Mission-Climate
33 (GCOM-C) satellite. In particular, VIIRS-SNPP, launched on October 28, 2011, is the first
34 VIIRS in the series that are all planned to provide decades long consistent global atmosphere,
35 land, cryosphere, and ocean Environmental Data Records (EDR or Level-2 data) (*Goldberg et*
36 *al.*, 2013). To ensure accurate routine production of global ocean optical, biological, and
37 biogeochemical properties, a considerable effort has been made towards sensor on-orbit
38 instrument calibration (*Cao et al.*, 2013; *Sun and Wang*, 2015; 2016), on-orbit vicarious
39 calibration using the in situ optics data from the Marine Optical Buoy (MOBY) in the waters off
40 Hawaii (*Clark et al.*, 1997; *Wang et al.*, 2016), development and improvement of ocean color
41 retrieval algorithms (*Jiang and Wang*, 2014; *Wang and Jiang*, 2018; *Wang et al.*, 2012; *Wang*
42 *and Son*, 2016), and VIIRS ocean color product data evaluation and validation (*Barnes et al.*,
43 2019; *Hlaing et al.*, 2013; *Wang et al.*, 2013).

44 Deriving global ocean optical, biological, and biogeochemical property data from space-
45 based measurements is a very complicated and challenging task, which requires close attention to
46 instrument calibration (including vicarious calibration), algorithm development, and the
47 validation of the results (*McClain*, 2009). The spaceborne radiometers directly measure the top
48 of the atmosphere (TOA) radiances, which in addition to the water color signal also include
49 radiance contributions from various light scattering processes in the atmosphere, and reflection
50 and refraction from the water surface. Extracting the ocean color spectra from the TOA radiance
51 spectra is known as atmospheric correction (*Gordon and Wang*, 1994a; *IOCCG*, 2010; *Wang*,
52 2007), and it involves subtraction of these different radiance contributions, which are much
53 larger than the ocean color signal (*Gordon and Wang*, 1994a; *IOCCG*, 2010; *Wang*, 2007).

54 Furthermore, these terms depend on multiple parameters, such as solar-sensor geometry, but also
55 physical or meteorological conditions, such as surface wind speed, atmospheric pressure,
56 atmospheric ozone concentration, and column water vapor amount (*Ramachandran and Wang,*
57 2011). Due to the overall complexity and the multiple parameters involved in the atmospheric
58 correction process, comparisons with in situ ocean color measurements are integral and
59 important to maintain good quality of the ocean color satellite data retrievals (*Werdell and*
60 *Bailey, 2005*).

61 In situ measurements, however, also have significant limitations. First, the number of in situ
62 measurements is several orders of magnitude less than the number of satellite remote sensing
63 retrievals. Second, in situ measurements are usually confined to a few stations in the regions of
64 interest, while cruises with more variety of sampling stations are not as frequent. This means that
65 most geographical ocean regions and times are never represented in the available in situ data.
66 Furthermore, in situ data also involve measurement uncertainties (*Werdell and Bailey, 2005;*
67 *Zibordi et al., 2009; Zibordi et al., 2015*), sometimes even with large uncertainties in data
68 quality.

69 One possibility to address these limitations is to perform an inter-sensor comparison (*Barnes*
70 *and Hu, 2016; Wang et al., 2002; Zibordi et al., 2006*). However, it is not easy to attribute the
71 discrepancies to a particular sensor, especially since the older and more thoroughly studied
72 sensors often have also suffered more degradation due to longer time in service. In addition,
73 differences in sensor calibration may be difficult if not impossible to disentangle from the
74 differences in performance of the retrieval algorithms.

75 The main idea of this work is to use ocean color data from the same sensor over large
76 swaths of the global open ocean, where water conditions are relatively uniform, and temporal
77 changes are more gradual, to evaluate the statistical consistency of the ocean color retrievals.
78 Yet, even within the relatively uniform conditions of the open ocean, ocean optical, biological,
79 and biogeochemical properties can have a large spatial variation over longer length scales, as
80 well as a significant seasonal variation. Therefore, it is the deviation from the spatial and

81 temporal average, or anomaly, that can provide the information about the biases and consistency
82 of ocean color data retrievals for the particular space-time domain. It is assumed that, for perfect
83 retrievals, little to no deviations from such an average are expected.

84 The relevant spatial and temporal length scales of the average need to be long enough to
85 provide enough data for a meaningful and consistent average (given that often retrievals are not
86 possible or are masked out due to clouds (*King et al.*, 2013), high sun glint contamination (*Wang
87 and Bailey*, 2001), straylight effect (*Jiang and Wang*, 2013), high solar- and sensor-zenith angles
88 (*Mikelsons and Wang*, 2019), or other conditions. The length scale must be smaller than the
89 relevant scale of physical phenomena in water (such as ocean currents, mesoscale eddies, algal
90 blooms, etc.). The time scale for calculating the average also needs to be shorter than the
91 corresponding time scales of physical processes. At the very least, it should be able to resolve the
92 seasonal variation, which is the most important temporal cycle resolvable by daily satellite
93 observations.

94 In this study, it is assumed that the ocean physical processes with comparatively short-time
95 scales, such as diurnal changes, are relatively small in the open ocean over the hourly time scale.
96 The orbital period of polar orbiting Earth observing satellites is about 100 minutes, and the
97 variation of the local time of satellite overpass from day to day is roughly of the same time scale.
98 Thus, we assume that the diurnal changes are comparatively small within this time scale. Indeed,
99 matchups with in situ data measured within hours of satellite overpass are often used to check
100 data quality (*Wang et al.*, 2009b; *Werdell and Bailey*, 2005).

101 It should be pointed out that this method has limitations — it can only measure satellite data
102 statistical consistency. It cannot, for example, detect any long-term trends in ocean color
103 measurements, nor any changes or irregularities of the radiance spectral shape (*Wei et al.*, 2016).

104 **2. Methodology**

105 *2.1. The MSL12 ocean color data processing system*

106 The Multi-Sensor Level-1 to Level-2 (MSL12) is the current routine ocean color data
107 product retrieval system used at NOAA. In particular, VIIRS-SNPP and VIIRS-NOAA-20 global
108 ocean color products have been routinely produced using the MSL12 ocean color data processing
109 system since their successful launches in October 2011 and November 2017, respectively.
110 MSL12 is an enterprise ocean color data processing system, and the software is based on the
111 NASA Sea-viewing Wide Field-of-view Sensor (SeaWiFS) Data Analysis System (SeaDAS)
112 version 4.6 with several updates and modifications. MSL12 is designed to provide consistent
113 ocean color retrievals by employing the same retrieval algorithms for multiple satellite sensors
114 (Wang, *et al.*, 2002). Specifically, in this study, we use the near-infrared (NIR)-based
115 atmospheric correction in MSL12 (Gordon and Wang, 1994a; Wang, *et al.*, 2013) with the
116 improved NIR ocean reflectance correction algorithm (Jiang and Wang, 2014), which combines
117 the three other NIR water reflectance correction algorithms (Bailey *et al.*, 2010; Ruddick *et al.*,
118 2000; Wang, *et al.*, 2012). However, it should be noted that for global open ocean waters the NIR
119 ocean reflectance contribution is generally negligible, thus the NIR reflectance correction in the
120 ocean color data processing is not important.

121 The main VIIRS ocean color data products derived by MSL12 are the normalized water-
122 leaving radiance spectra, $nL_w(\lambda)$, which are equivalent to the water-leaving radiances measured
123 on the ocean surface assuming no atmosphere and the Sun at the zenith (Gordon, 2005; Morel
124 and Gentili, 1996; Wang, 2006). The $nL_w(\lambda)$ spectra are subject to destriping to remove striping
125 artifacts (Mikelsons *et al.*, 2014). The destriped $nL_w(\lambda)$ spectra are then used to further derive
126 ocean biological and biochemical property data, such as chlorophyll-a (Chl-a) concentration (Hu
127 *et al.*, 2012; O'Reilly *et al.*, 1998; Wang and Son, 2016), water diffuse attenuation coefficient at
128 the wavelength of 490 nm $K_d(490)$ (Lee *et al.*, 2005; Wang *et al.*, 2009a), and for the domain of
129 the photosynthetically available radiation (PAR) $K_d(\text{PAR})$ (Son and Wang, 2015). Other data
130 products, such as aerosol optical depth (AOD) (Wang *et al.*, 2005) and the aerosol Angstrom
131 coefficient (Angstrom, 1929), characterize aerosol properties, and are also used in atmospheric
132 correction (IOCCG, 2010; Wang, *et al.*, 2005). In this study, we convert the $nL_w(\lambda)$ spectra to the

133 normalized water-leaving reflectance spectra, defined as $\rho_{wN}(\lambda) = \pi nL_w(\lambda)/F_0(\lambda)$, where $F_0(\lambda)$ is
 134 the extraterrestrial solar irradiance (*Thuillier et al.*, 2003). This way, the ocean color spectra can
 135 be more equally compared spectrally (*Gordon and Wang*, 1994a).

136 On input, MSL12 requires the TOA radiances measured by the sensor, i.e., VIIRS Sensor
 137 Data Records (SDR, or Level-1B data), as well as ancillary data characterizing the atmospheric
 138 and ocean surface conditions to aid the retrievals (*Ramachandran and Wang*, 2011). The sensor-
 139 measured TOA radiances need to be properly calibrated (*Cao, et al.*, 2013; *Sun and Wang*, 2015;
 140 2016). In addition, complete geolocation information is required (latitude and longitude for each
 141 retrieval sample), as well as information about retrieval geometry, which includes the solar- and
 142 sensor-zenith angles, as well as the relative azimuth angle.

143 The main ancillary data are the sea level atmospheric pressure, surface wind speed, the
 144 amount of ozone and water vapor in the atmosphere, integrated over the altitude (*Ramachandran*
 145 *and Wang*, 2011). Using more accurate ancillary data, such as those produced by the NOAA
 146 Global Forecasting System (GFS) has been shown to increase the accuracy of the ocean color
 147 products (*Ramachandran and Wang*, 2011).

148 2.2. Evaluation approach and criteria

149 The VIIRS-SNPP SDR (or Level-1B data) from the year 2016 were processed to produce
 150 global ocean color product data using the MSL12 ocean color data processing system. We work
 151 with $\rho_{wN}(\lambda)$ spectra, but the entire procedure can also be applied to the normalized water-leaving
 152 radiance $nL_w(\lambda)$ spectra (or remote sensing reflectance $R_{rs}(\lambda)$ spectra, defined as $R_{rs}(\lambda) =$
 153 $nL_w(\lambda)/F_0(\lambda)$), since these only differ by constant conversion factors. Data obtained with other
 154 retrieval algorithms, or even from different sensors can also be analyzed using this method.

155 The first step is to establish a baseline for temporal and spatial averaged $\rho_{wN}(\lambda)$ data. For
 156 each day of retrievals, the global Level-2 $\rho_{wN}(\lambda)$ data are binned into a global Level-3 daily
 157 average (*Campbell et al.*, 1995). We use a particular kind of grid that covers the entire globe with
 158 bins of nearly equal area. Within this grid, the bins are arranged into rows, such that an integer
 159 number of rows fit within the 180° of latitude. For example, the 9 km bins used in this study

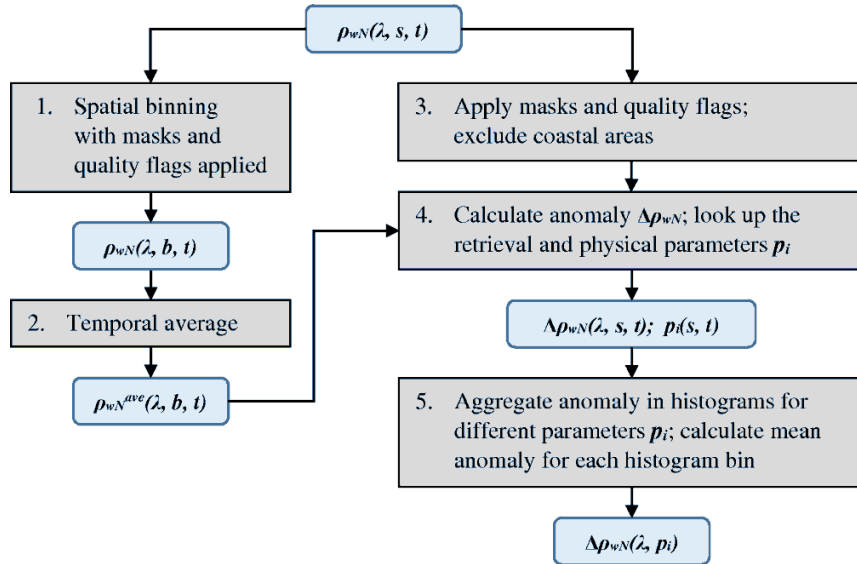
160 cover 5 minutes in latitude. Thus, 2160 rows of the 9 km bins span the entire 180° of latitude.
 161 The longitudinal size of bins is determined as the closest match to the extent in the latitude (thus
 162 ensuring nearly square shaped bins) that also fits an integer number of bins within the row
 163 spanning the entire 360° of longitude, for each row of bins. Thus, with 9 km bins, there are 4320
 164 bins per one row on the equator. This number decreases towards the poles as a cosine function of
 165 the latitude.

166 In the binning process, we use several masks and data quality flags, e.g., land, cloud (*Wang*
 167 *and Shi*, 2006), straylight (*Jiang and Wang*, 2013), high sun glint contamination (*Wang and*
 168 *Bailey*, 2001), etc., to discard poor quality retrievals due to various conditions in order to obtain
 169 the most accurate binned data product for all ocean color data, e.g., $\rho_{wN}(\lambda)$ [or $nL_w(\lambda)$], Chl-a,
 170 $K_d(490)$, etc.

171 In the following step, a weighted temporal moving average is calculated for each bin using
 172 the daily binned data for the time average period $T_a = 17$ days from 8 days before and 8 days
 173 after the particular day. We weight the time average with a cosine function of time difference Δt ,
 174 measured in days, i.e., $w(\Delta t) = \cos[\pi\Delta t/(T_a + 1)]$. This choice of weighting function emphasizes
 175 the importance of the data closer to a given time, however, other weighting functions can also be
 176 used, with qualitatively similar results. The choice of time period T_a is motivated by the SNPP
 177 revisit cycle of 16 days.

178 Once the spatially binned and time averaged baseline data are obtained, the anomaly can be
 179 calculated. But, before that, we filter the $\rho_{wN}(\lambda)$ data to discard the coastal areas from the
 180 statistical analysis due to their much shorter temporal and spatial scales of variability. To have a
 181 simple criterion, we choose to only use the ocean color retrievals from locations with water depth
 182 exceeding 1 km (i.e., global deep-water regions). Furthermore, we again exclude data with high
 183 sun glint conditions, and other conditions severely affecting the retrieval quality. However, we
 184 include data with high solar- and sensor-zenith angles to estimate the effect of such conditions on
 185 the derived ocean color products.

186 For each of the filtered retrieval data points, the corresponding spatial-temporal average
 187 calculated in the previous step is subtracted, yielding the anomaly $\Delta\rho_{wN}(\lambda, s, t)$ (s and t denote
 188 space and time dependence, respectively). We also look up the corresponding values of various
 189 geometrical, ancillary, and physical parameters $p_i(s, t)$ for each retrieval. The anomaly, $\Delta\rho_{wN}(\lambda, s,$
 190 $t)$ is then recorded in histograms to analyze its dependence on parameters p_i . For each histogram,
 191 we calculate the mean anomaly $\Delta\rho_{wN}(\lambda, p_i)$ over all retrievals that fall in the same bin of the
 192 parameter p_i . The total number of data points in each histogram bin of the dependent parameter p_i
 193 is also recorded. The major steps of this statistical analysis are schematically depicted in Fig. 1.



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Figure 1. Diagram summarizing the steps of the statistical analysis. Indices s and t denote space and time dependence of the Level-2 data, respectively. The index b denotes spatial bin. The parameters p_i denote all retrieval and physical parameters included in this study. The quality flags in the step 3 may differ from those in the step 1.

200 For good quality ocean color retrievals, the mean anomaly $\Delta\rho_{wN}(\lambda, p_i)$ [further denoted as
 201 $\Delta\rho_{wN}(\lambda)$] is expected to be close to zero across the entire range of any dependent parameter p_i . A
 202 significant departure from zero in anomaly (compared to the retrieval accuracy requirements)
 203 indicates either an underestimate or an overestimate of $\rho_{wN}(\lambda)$ spectra under certain conditions
 204 characterized by the corresponding dependent variable p_i . We note that the accuracy requirement
 205 of satellite-derived $\rho_{wN}(\lambda)$ spectra for the global open ocean is that $\Delta\rho_{wN}(\lambda)$ in the blue (443 nm

206 band) $< \sim 0.001$ (or 5%) (Gordon and Wang, 1994a; IOCCG, 2010; Wang, 2007). In addition to
207 the mean anomaly, we also collect statistics and monitor the variance of the anomaly. Lower
208 variance usually implies more precise retrievals for a given parameter range.

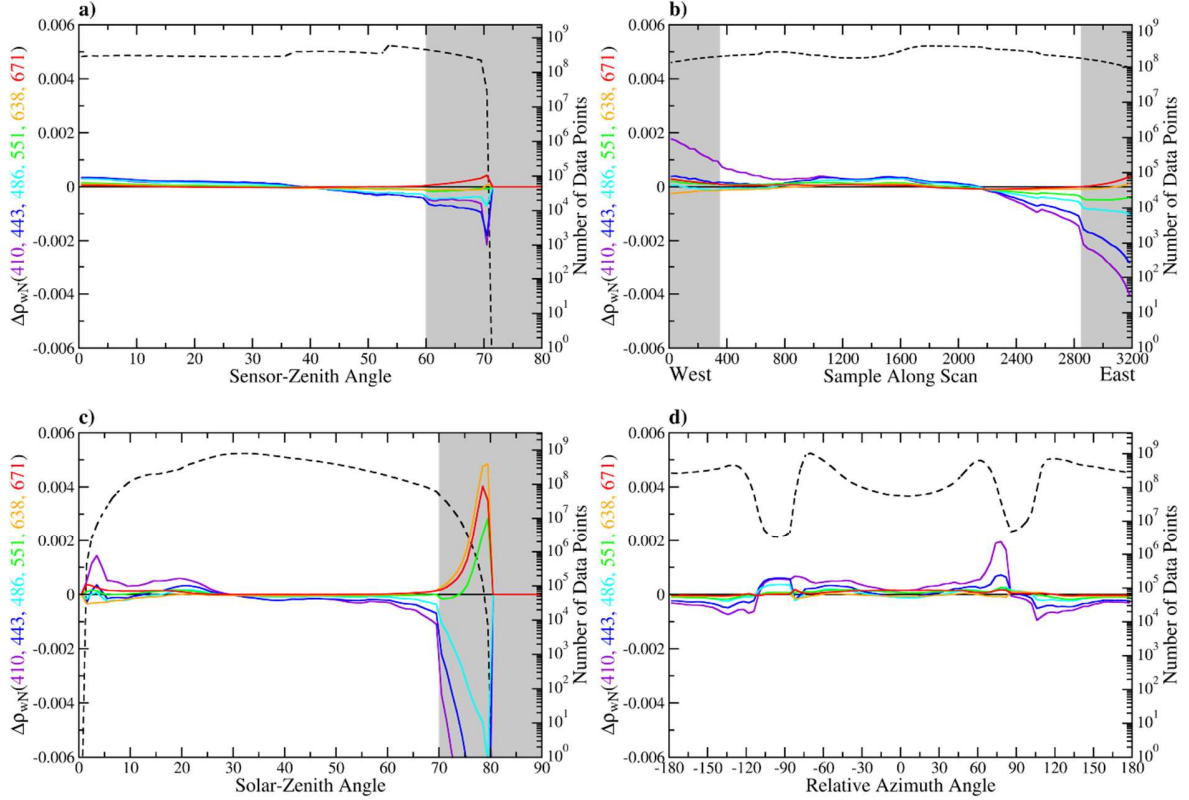
209 We note that results for the reflectance anomaly $\Delta\rho_{wN}(\lambda)$ are not sensitive to the exact choice
210 of T_a , or the type of time averaging used, nor to the spatial bin size. However, the variance of the
211 $\Delta\rho_{wN}(\lambda)$ increases with averaging over longer time period ($T_a=33$), and larger spatial bin size
212 (18km), because natural fluctuations and changes of $\rho_{wN}(\lambda)$ over longer temporal and spatial
213 scales are included into the average.

214 **3. Results**

215 The relevant parameters to ocean color retrievals can roughly be divided into three groups:
216 the parameters describing the solar-sensor geometry with respect to the observable water surface,
217 the ancillary parameters, and the intermediate parameters describing the retrieval conditions.

218 *3.1. Solar-sensor geometry effect*

219 We first evaluate the consistency of retrievals with respect to the solar-sensor geometry
220 angles. Figure 2a shows the dependence of $\Delta\rho_{wN}(\lambda)$ on the sensor-zenith angle. Unsurprisingly,
221 $\Delta\rho_{wN}(\lambda)$ is smaller for lower values of this angle, but it becomes significantly negative for
222 sensor-zenith angles over 60° , which means that $\rho_{wN}(\lambda)$ spectra are underestimated for these
223 conditions. From results in Fig. 2a, it appears that the currently used threshold value for sensor-
224 zenith angle of 60° in MSL12 is a reasonable choice to separate good retrievals from poor or
225 questionable ones.



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Figure 2. Dependence of $\Delta\rho_{wN}(\lambda)$ on (a) the sensor-zenith angle, (b) the sample along the scan, (c) the solar-zenith angle, and (d) the relative azimuth angle. The solid colored lines show the anomaly $\Delta\rho_{wN}(\lambda)$. The black dashed lines, along with the right ordinates, indicate the number of retrievals. The gray shaded areas in panels (a) and (b) indicate where the large sensor-zenith angle flag ($> 60^\circ$) is triggered by MSL12 to warn about questionable data quality of retrievals. Similarly, the gray shaded areas in panel (c) indicate the large solar-zenith angle flag ($> 70^\circ$) in MSL12.

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Table 1. Mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for three ranges of sensor-zenith angle (θ).

λ (nm)	MAD[$\Delta\rho_{wN}(\lambda)$] $\times 10^4$		
	$\theta \leq 20^\circ$	$20^\circ < \theta \leq 60^\circ$	$\theta > 60^\circ$
410	2.1	2.1	5.2
443	2.7	2.0	7.6
486	2.2	1.4	4.2
551	1.0	0.5	1.4
638	0.6	0.6	1.1
671	0.3	0.2	1.6

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To quantify the deviations of $\Delta\rho_{wN}(\lambda)$ from zero, we show the results for the mean absolute

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deviation of $\Delta\rho_{wN}(\lambda)$ over the three ranges of sensor-zenith angle (less than 20° , 20° to 60° , and

237 larger than 60°) in Table 1. Note that since $\Delta\rho_{wN}(\lambda)$ is an anomaly, its mean value over the entire
 238 range of the histogram is zero.

239 A closely related parameter to sensor-zenith angle is the position or sample along the scan.
 240 The VIIRS-SNPP swath extends up to about 3040 km across the direction of the flight path and
 241 is symmetric about the nadir. This corresponds to the sensor-zenith angle ranging from 0° at the
 242 center of the scan up to 70° at the swath edge. The VIIRS-SNPP medium spatial resolution bands
 243 (M-bands) used for ocean color retrievals have 3200 samples per scan. Figure 1b shows the
 244 dependence of $\Delta\rho_{wN}(\lambda)$ from the sample along the scan. Here, we count the number the samples
 245 in the direction from west to east, with the middle sample roughly corresponding to the nadir.
 246 Again, for most locations along the swath, the mean anomaly is minimal. However, near the
 247 swath edges, $\Delta\rho_{wN}(\lambda)$ is significantly increased. Predictably, the results are very similar to the
 248 sensor-zenith angle dependence. However, the sample along the scan dependence reveals more
 249 information, since it also captures asymmetry between the two sides of a scan, which would get
 250 averaged out on the sensor-zenith angle dependence (Fig. 2a). For this reason, we choose the
 251 sample along the scan in favor of sensor-zenith angle in the further discussion. Table 2
 252 summarizes the results for the mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for the five regions along the
 253 swath, defined by the five ranges of sample numbers: [0–399], [399–1199], [1200–1999], [2000–
 254 2799], and [2800–3199].

255 **Table 2.** Mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for the five regions along the swath.

λ (nm)	MAD[$\Delta\rho_{wN}(\lambda)$] $\times 10^4$				
	0–399	400–1199	1200–1999	2000–2799	2800–3199
410	13	3.7	2.1	4.7	26
443	2.5	1.8	2.6	3.7	19
486	0.7	1.0	2.1	2.3	8.3
551	1.3	1.2	0.9	1.6	4.5
638	1.7	0.6	0.6	0.9	0.6
671	1.7	0.6	0.4	0.7	1.1

256 Figure 2c shows the dependence with respect to the solar-zenith angle. Similar to the sensor-
 257 zenith angle dependence, we see significant deviations pointing to poor retrievals for values of
 258 the solar-zenith angle above 70° , which is also used as a threshold value in MSL12 for flagging
 259 the poor quality data due to this condition. In addition, we see some deviations from the average
 260 for very low values of solar-zenith angle. This corresponds to the relatively smaller number of
 261 retrievals in the tropics (only 1.3% of all retrievals have solar-zenith angle less than 10°),
 262 acquired almost exclusively on the west side of the scan (since VIIRS-SNPP is flying in the polar
 263 sun synchronous orbit with afternoon equatorial overpass). Thus, this deviation is also related to
 264 the sample across the scan dependence in Fig. 2b, for an approximate range of samples in the
 265 range 0–400. Table 3 summarizes the mean absolute deviations of $\Delta\rho_{wN}(\lambda)$ for the three ranges of
 266 solar-zenith angle: less than 20° , 20° to 70° , and larger than 70° .

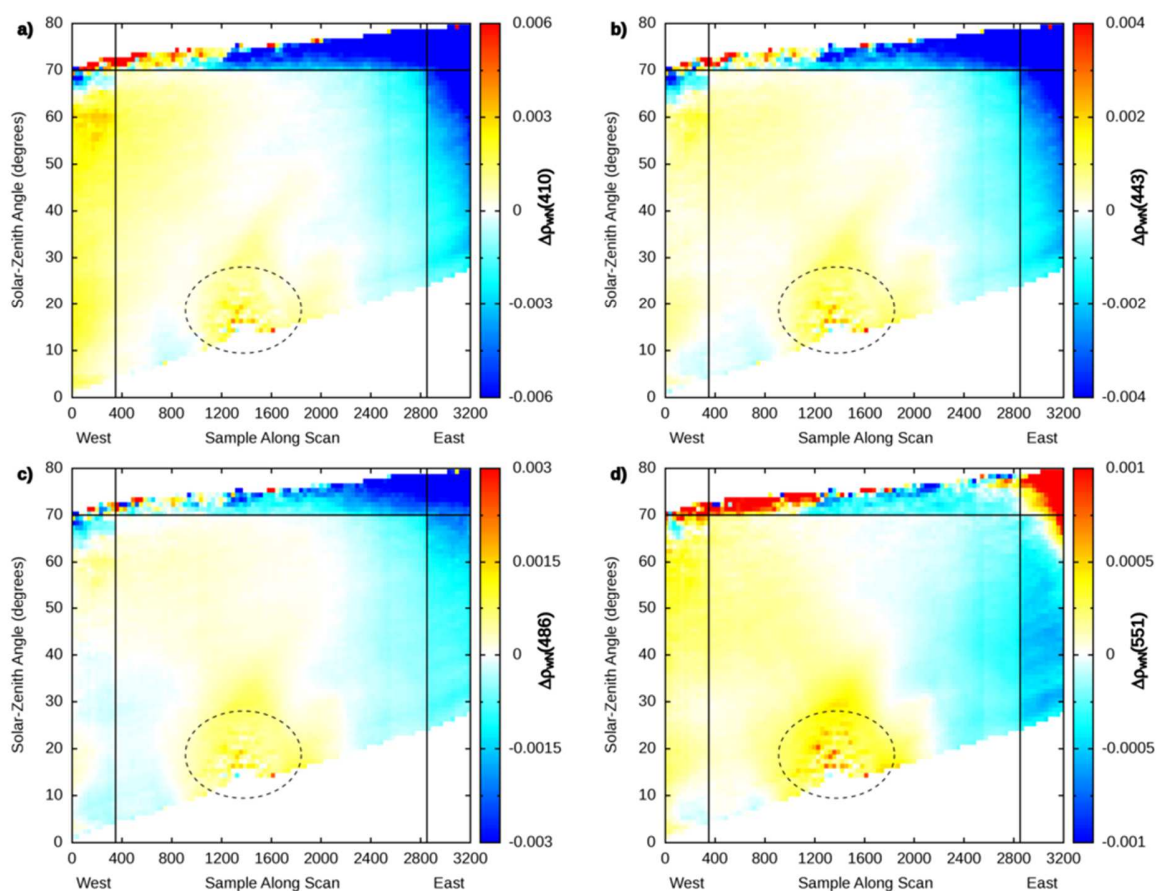
267 **Table 3.** Mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for the three ranges of solar-zenith angle (θ_0).

λ (nm)	MAD[$\Delta\rho_{wN}(\lambda)$] $\times 10^4$		
	$\theta_0 \leq 20^\circ$	$20^\circ < \theta_0 \leq 70^\circ$	$\theta_0 > 70^\circ$
410	4.7	2.0	5.2
443	1.4	1.7	3.2
486	1.0	0.8	1.6
551	1.0	0.5	0.2
638	0.9	0.4	0.6
671	1.3	0.1	0.4

268 The dependence of $\rho_{wN}(\lambda)$ anomaly on the relative azimuth angle is shown in Fig. 2d. We
 269 define it as an angle ranging from -180° to $+180^\circ$ between the azimuthal directions of the solar
 270 reflection and the sensor direction. Thus, a small (near zero) relative azimuth angle indicates a
 271 possibility of solar Fresnel reflection from water surface (sun glint) reaching the sensor (if solar-
 272 and sensor-zenith angles are similar). The number of data points shows two dips near $\pm 90^\circ$. This
 273 is because near the equator, the relative azimuth angle can only fall within two narrow ranges of
 274 values corresponding to the samples on the east and the west side of the scan. For higher
 275 latitudes, these two ranges of the relative azimuth angle can broaden and also shift with a

276 seasonal modulation. However, the two ranges of relative azimuth angle near $\pm 90^\circ$ do not have
277 many retrievals. Unsurprisingly, with fewer data points for statistics, $\Delta\rho_{wN}(\lambda)$ also shows
278 somewhat more noisy behavior near these two dips.

279 We also investigate how different combinations of geometric retrieval parameters affect the
280 ocean color retrievals. In particular, we focus on the combined effect of different values of the
281 solar-zenith angle and the sample along the scan. In Fig. 3a, we show the mean anomaly for the
282 short-blue band $\rho_{wN}(410)$ with respect to both solar-zenith angle and the sample along the scan.
283 Note that low values of solar-zenith angle are only possible for the westernmost part of the
284 swath. The oval in dashed line indicates an approximate parameter region with high sun glint,
285 and thus sharply decreased the number of retrievals and increased noise in the data. Significant
286 deviations from average are encountered for the large values of solar-zenith angle, and near both
287 edges of the swath. It appears that high solar-zenith angles have a more pronounced impact on
288 retrievals in the eastern part of the swath. Figures 3b, 3c, and 3d show a very similar pattern for
289 $\rho_{wN}(443)$, $\rho_{wN}(486)$, and the green band $\rho_{wN}(551)$, respectively, although the amplitude of the
290 anomaly $\Delta\rho_{wN}(\lambda)$ is decreasing with increasing wavelength λ . Indeed, the similarity of patterns
291 seen in Fig. 3 indicates that the $\rho_{wN}(\lambda)$ anomalies are quite correlated across the spectrum
292 (IOCCG, 2010; Wang and Gordon, 2018).



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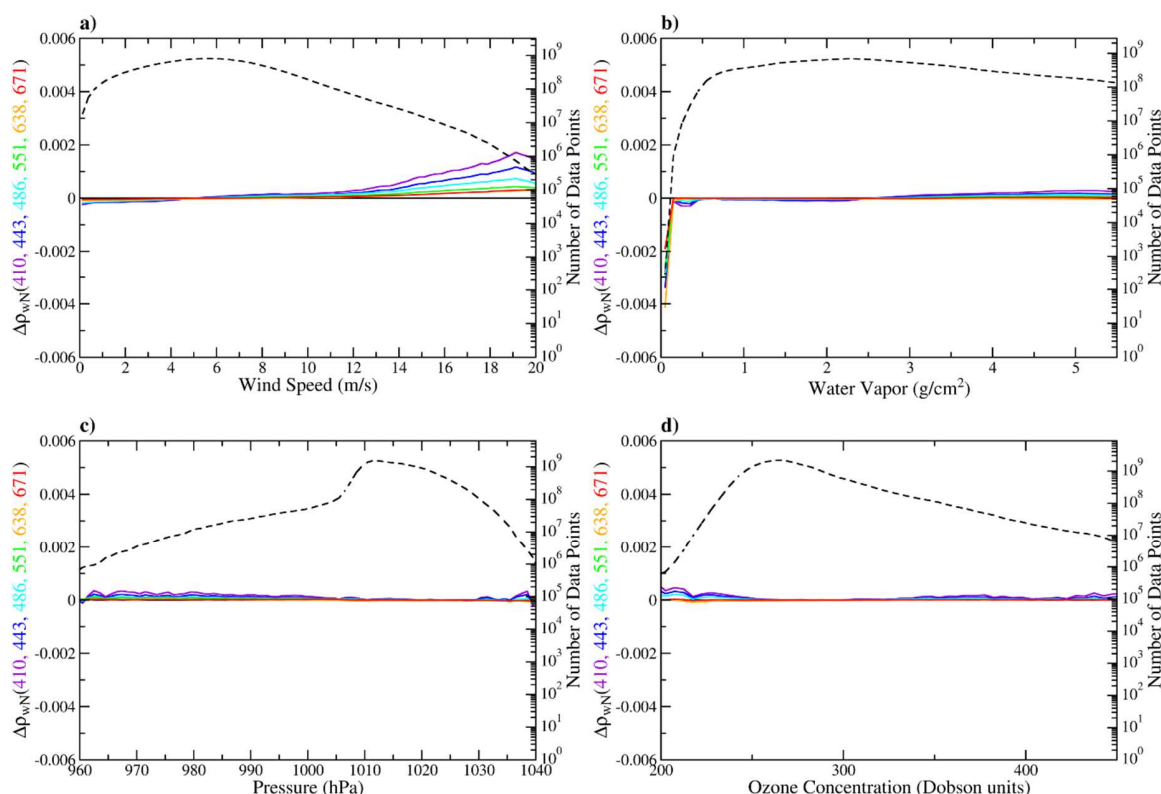
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Figure 3. Dependence of $\Delta\rho_{wN}(\lambda)$ on the solar-zenith angle and the sample along the scan for (a) $\Delta\rho_{wN}(410)$, (b) $\Delta\rho_{wN}(443)$, (c) $\Delta\rho_{wN}(486)$, and (d) $\Delta\rho_{wN}(551)$. The dashed oval line indicates the parameter region with high sun glint conditions, where fewer retrievals result in more noisy data. The vertical lines indicate the boundary when sensor-zenith angle exceeds 60° , and the horizontal line indicates the onset of high solar-zenith angle at 70° .

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Figure 4. Dependence of $\Delta\rho_{wN}(\lambda)$ on (a) the wind speed, (b) the amount of water vapor in the atmosphere, (c) the sea level atmospheric pressure, and (d) the ozone concentration. The solid colored lines show $\Delta\rho_{wN}(\lambda)$. The black dashed lines, along with the right ordinates, indicate the number of retrievals.

306 3.2. Effects of ancillary data inputs

307 Next, we examine the statistical consistency with respect to some physical parameters
 308 affecting the ocean color retrievals. The results in the previous section confirm that good quality
 309 ocean color retrievals are obtained for solar-zenith angles and sensor-zenith angles of $\leq 70^\circ$ and
 310 $\leq 60^\circ$, respectively. Therefore, in the further analysis, we only use the ocean color retrievals with
 311 solar- and sensor-zenith angles meeting these requirements.

312 Figure 4a shows the dependence of the deviation or $\rho_{wN}(\lambda)$ anomaly $\Delta\rho_{wN}(\lambda)$ with respect to
 313 the surface wind speed. For wind speed up to about 14 m/s, the deviation from average value is
 314 negligible. However, it increases for higher wind speeds, which occur for a smaller number of
 315 retrievals. In the MSL12 ocean color data processing, there are three radiance components that
 316 require the input of wind speed for surface roughness characteristics, i.e., ocean whitecap

317 radiance (*Gordon and Wang, 1994b*), sun glint radiance (*Wang and Bailey, 2001*), and Rayleigh
 318 scattering radiance (*Gordon and Wang, 1992; Wang, 2002; 2016*). However, examining the
 319 scenes with whitecaps in detail, we have concluded that MSL12 $\rho_{wN}(\lambda)$ spectra data are
 320 unchanged, or not impacted by the whitecaps radiance contribution with high wind speed. In
 321 fact, the TOA whitecap radiance contribution is generally not significant (*Moore et al., 2000*).
 322 Other possible errors are from sun glint contamination correction (*Wang and Bailey, 2001*) that
 323 uses the Cox and Munk (*Cox and Munk, 1954*) model with the input of wind speed for sun glint
 324 radiance estimation, as well as Rayleigh scattering radiance computation that also uses the Cox
 325 and Munk model (*Cox and Munk, 1954*) for surface roughness (wind speed) (*Gordon and Wang,*
 326 *1992; Wang, 2002; 2016*). Thus, it is possible that there may be some uncertainties related to the
 327 Cox and Munk (1954) model for high wind speed (*Zhang and Wang, 2010*). In addition, we note
 328 that high winds have a slight correlation with sparse clouds, which may increase variability in the
 329 satellite-retrieved $\rho_{wN}(\lambda)$ values.

330 Figure 4b shows the dependence on the amount of the water vapor in the atmospheric
 331 column. No obvious deviation is seen here, except for rarely encountered values of very low
 332 humidity, where insufficient data yields noisy results. Figure 4c similarly indicates no anomalous
 333 behavior with respect to the sea level atmospheric pressure, which is used as one of the inputs for
 334 the Rayleigh radiance computation (*Wang, 2005; 2016*). Likewise, Figure 4d shows no
 335 significant deviation with respect to the amount of the ozone in the atmosphere. Therefore, the
 336 MSL12 ocean color data processing system performed perfectly for VIIRS-derived $\rho_{wN}(\lambda)$
 337 spectra with respect to the ancillary inputs of water vapor, ocean surface atmospheric pressure,
 338 and ozone amount, as well as to the wind speed up to about 14 m/s.

339 To better quantify the anomaly $\Delta\rho_{wN}(\lambda)$, we summarize the mean absolute deviation of
 340 $\Delta\rho_{wN}(\lambda)$ for wind speeds less than and exceeding 14 m/s in Table 4. We recall that the accuracy
 341 requirement for $\rho_{wN}(\lambda)$ at the blue 443 nm band is within ~ 0.001 (or 5%). From Table 4, MSL12
 342 retrieved VIIRS-SNPP $\rho_{wN}(\lambda)$ spectra meet this goal even for high wind speeds.

343 **Table 4.** Mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for low and high wind speeds.

λ (nm)	MAD[$\Delta\rho_{wN}(\lambda)$] $\times 10^4$	
	wind speed ≤ 14 m/s	wind speed > 14 m/s
410	1.1	8.6
443	0.8	5.9
486	0.6	4.2
551	0.3	2.5
638	0.3	1.5
671	0.2	1.6

344 3.3. Effects of intermediate retrieval parameters

345 The glint coefficient, computed from the Cox and Munk model (Cox and Munk, 1954; Wang
346 and Bailey, 2001), is a more precise indicator of the degree to which the sun glint affects the
347 retrievals. In addition to the combination of the solar- and sensor-zenith angles, and the relative
348 azimuth angle, the glint coefficient computation also includes the wind speed, which affects the
349 surface roughness, and thus also impacts sun reflection from the water surface (Cox and Munk,
350 1954; Wang and Bailey, 2001). The dependence on the glint coefficient in Fig. 5a shows some
351 increase in $\Delta\rho_{wN}(\lambda)$ for stronger glint conditions.

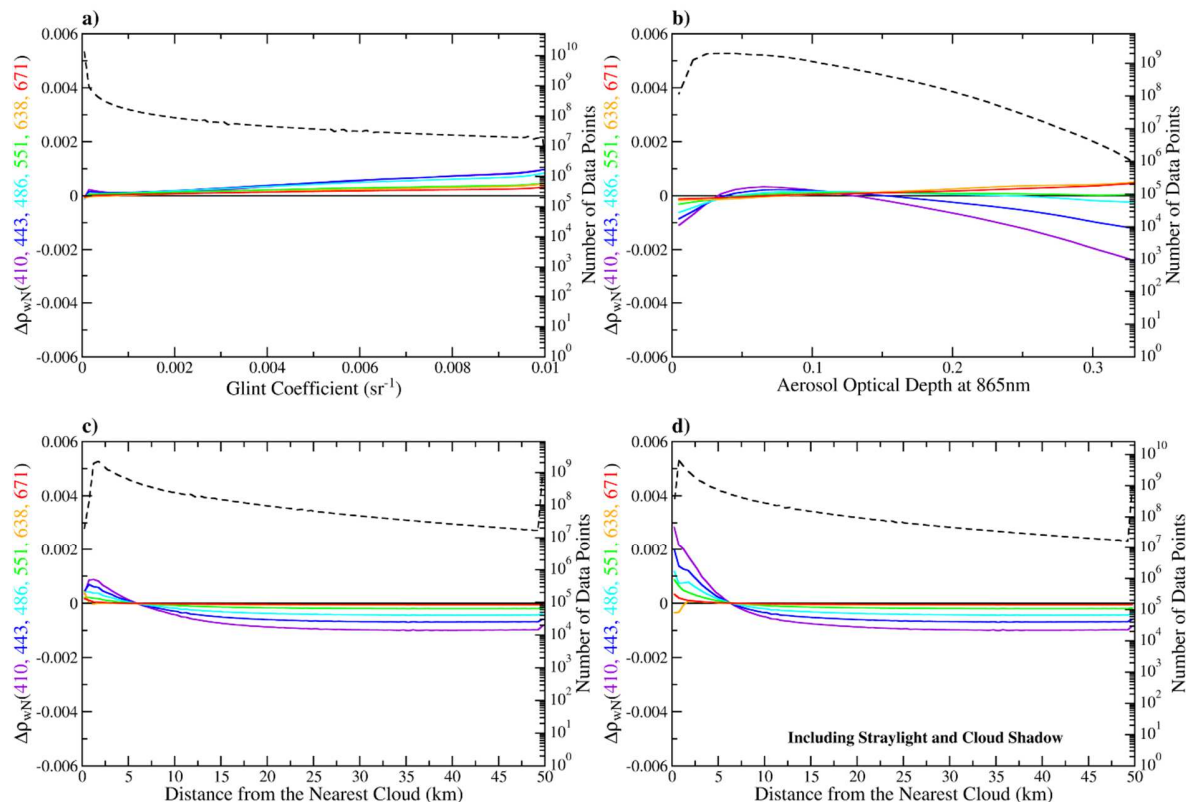
352 **Table 5.** Mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for different ranges of AOD.

λ (nm)	MAD[$\Delta\rho_{wN}(\lambda)$] $\times 10^4$			
	AOD < 0.1	$0.1 < \text{AOD} \leq 0.2$	$0.2 < \text{AOD} \leq 0.3$	AOD > 0.3
410	2.9	1.7	9.4	22
443	2.1	1.1	4.0	11
486	1.5	1.3	0.3	2.6
551	0.9	1.1	0.8	1.0
638	0.9	1.1	3.1	5.2
671	0.5	0.9	2.2	4.7

353 The AOD is a parameter quantifying the aerosols in the atmosphere above the retrieval
354 location. Aerosol optical property data are by-products from atmospheric correction (Gordon and
355 Wang, 1994a; Wang, et al., 2005). The AOD measures the extent of light scattering and
356 attenuation by aerosols. Higher AOD usually corresponds to more difficult retrieval conditions
357 (Gordon and Wang, 1994a; IOCCG, 2010; Wang, 2007). In particular, cases with heavy aerosols

358 such as smoke and dust are masked out. In these conditions, for AOD higher than 0.2, $\rho_{wN}(\lambda)$
 359 spectra are slightly overestimated for the red bands, and they are slightly underestimated for the
 360 blue bands (Fig. 5b). However, for extremely low AOD cases (e.g., < 0.03) $\rho_{wN}(\lambda)$ spectra at the
 361 short wavelengths (410 and 443 nm) are also slightly underestimated (Fig. 5b). This is likely due
 362 to the correlations with other impact factors, e.g., cases with low aerosols in the Southern Ocean
 363 where retrievals are also associated with large solar-zenith angles (see Fig. 2c). Table 5
 364 summarizes the mean absolute deviation of $\rho_{wN}(\lambda)$ for these four ranges of AOD: less than 0.1,
 365 0.1–0.2, 0.2–0.3, and larger than 0.3. We note that for $\text{AOD} \leq 0.3$, the mean anomalies of VIIRS-
 366 derived $\rho_{wN}(\lambda)$ spectra are mostly within ~ 0.0005 , except for the short blue band $\rho_{wN}(410)$ with
 367 the value of ~ 0.001 , which is significantly better than the accuracy requirement of ~ 0.001 (or
 368 5%) for the blue band reflectance $\rho_{wN}(443)$.

369 The next panel, Fig. 5c, shows that reflectance anomaly $\Delta\rho_{wN}(\lambda)$ spectra are slightly
 370 elevated for retrievals within ~ 5 – 10 km from clouds. The remnants of the straylight effect (*Jiang*
 371 *and Wang*, 2013) are the most likely cause for this. Table 6 shows the increases in $\Delta\rho_{wN}(\lambda)$ for
 372 the three ranges (0–5 km, 5–15 km, and 15–30 km) of distance from the nearest cloud, using the
 373 average values for retrievals further than 30 km from clouds [$\Delta\rho_{wN}^*(\lambda)$] as baseline values. To
 374 confirm the effect of the proximity to the clouds, we have also repeated the same study by
 375 including the data affected by straylight and cloud shadow conditions. These results are shown in
 376 Fig. 5d, and indeed exhibit a more pronounced increase in $\Delta\rho_{wN}(\lambda)$ spectra for retrievals near
 377 clouds, as compared to Fig. 5c, where data affected by straylight and cloud shadow are masked
 378 out.



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Figure 5. Dependence of $\Delta\rho_{wN}(\lambda)$ on (a) sun glint coefficient, (b) aerosol optical depth, (c) distance from the nearest cloud, and (d) distance from the nearest cloud, including the retrievals affected by straylight and cloud shadow conditions. The solid colored lines show $\Delta\rho_{wN}(\lambda)$. The black dashed lines, along with the right ordinates, indicate the number of retrievals.

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Table 6. Mean absolute deviation of $\Delta\rho_{wN}(\lambda)$ for the three ranges of distance from the nearest cloud (DNC).

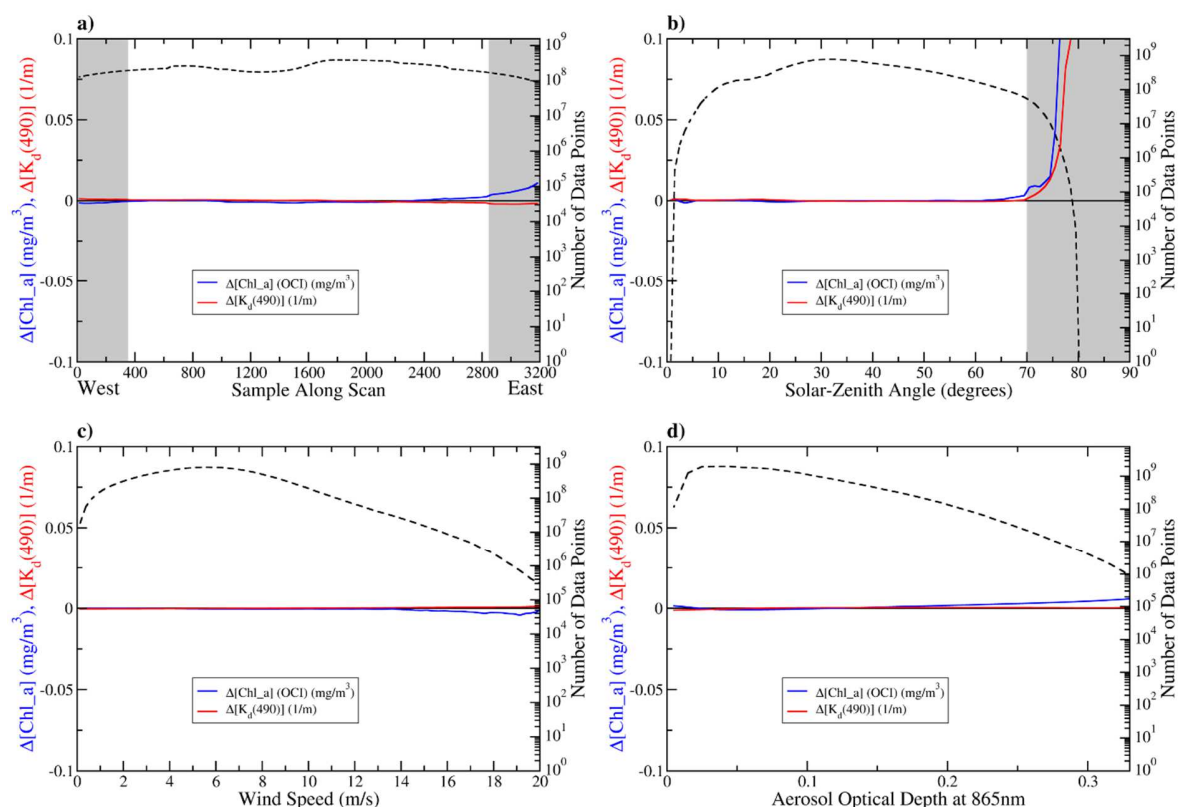
λ (nm)	$MAD[\Delta\rho_{wN}(\lambda) - \Delta\rho_{wN}^*(\lambda)] \times 10^4$		
	$DNC \leq 5$ km	$5 \text{ km} < DNC \leq 15$ km	$15 \text{ km} < DNC \leq 30$ km
410	15	6.3	0.5
443	11	4.4	0.4
486	6.8	2.8	0.3
551	3.1	1.3	0.2
638	0.3	0.3	0.0
671	0.4	0.4	0.1

387 *3.4. Impact on derived biological and biogeochemical products*

388 While normalized water-leaving reflectance $\rho_{wN}(\lambda)$ spectra are the central ocean color
389 measurements, the impact of $\Delta\rho_{wN}(\lambda)$ on the $\rho_{wN}(\lambda)$ -derived ocean biological and biogeochemical
390 products, such as Chlorophyll-a concentration (*Hu, et al., 2012; O'Reilly, et al., 1998; Wang and*
391 *Son, 2016*), and water diffuse attenuation coefficient at the wavelength of 490 nm $K_d(490)$
392 (*Wang, et al., 2009a*), are also important to evaluate. Consequently, we want to estimate how
393 much these derived quantities are affected by any systematic biases in $\rho_{wN}(\lambda)$ with respect to
394 various retrieval parameters.

395 However, both Chl-a and $K_d(490)$ range over several orders of magnitude, and their
396 frequency distributions are very skewed. Consequently, the mean anomaly is not a good measure
397 for data consistencies, as it is heavily impacted by few infrequent localized events (such as algae
398 blooms). In order to overcome this difficulty, we have instead opted to evaluate the median
399 anomaly dependence on various retrieval parameters for Chl-a and $K_d(490)$.

400 Figure 6a shows the median anomaly for Chl-a and $K_d(490)$ versus the sample along the
401 scan (also related to sensor-zenith angle), with little to no impact in the derived quantities for all
402 values of this angle. This suggests that the calculation of Chl-a and $K_d(490)$ may be extended for
403 larger values of sensor-zenith angle. However, in Fig. 6b, the dependence on solar-zenith angle
404 shows significant positive bias for values larger than 70° , which is similar to the results for the
405 reflectances in Fig. 2c. Thus, the systematic biases in reflectances for large solar-zenith angles
406 translate into biases for the derived Chl-a and $K_d(490)$. The last two panels, Figs. 6c and 6d,
407 show the dependence of the median anomaly in Chl-a and $K_d(490)$ with respect to the wind speed
408 and the AOD. Comparing these with the results in Figs. 4a and 5b, respectively, we note that the
409 derived quantities are overall less sensitive to the biases in the reflectance spectra. Results for
410 other dependent variables (not shown) similarly show little to no systematic bias in derived Chl-a
411 and $K_d(490)$. In fact, across the all the parameter ranges in Figs. 2, 4 and 5 (with the exception of
412 solar-zenith angles exceeding 70°), the Chl-a anomaly is less than 0.01mg/m^3 , and $K_d(490)$
413 anomaly is less than 0.005m^{-1} .



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Figure 6. Dependence of Chl-a (OCI algorithm) and $K_d(490)$ on (a) the sample along the scan, (b) the solar-zenith angle, (c) the wind speed, and (d) the aerosol optical depth. The gray shaded areas in panel (a) indicate where the large sensor-zenith angle flag ($> 60^\circ$) is set in MSL12. Similarly, the gray shaded areas in panel (b) indicate the large solar-zenith angle flag ($> 70^\circ$) in MSL12.

420 4. Discussion

421 In the past, several studies for the regional ocean color data consistency have looked at the
422 variability of ocean color spectra with respect to different retrieval parameters (e.g., Barnes and
423 Hu, 2016). Instead, we have analyzed the deviation from the normal (average or median), or the
424 anomaly, for the global deep waters which display high natural variability of the ocean color
425 spectra across the different regions. To illustrate the importance of using the $\rho_{wN}(\lambda)$ spectra
426 anomaly, or the deviation from average, to assess the global data statistical consistency, we have
427 also repeated parts of this study using the $\rho_{wN}(\lambda)$ spectra directly, i.e., without subtracting the
428 average values. Results (not shown) confirmed our expectations, i.e. satellite-derived $\rho_{wN}(\lambda)$
429 spectra are a function of environmental variables, as well as various solar-sensor geometry

430 inputs. Thus, the variability of $\rho_{wN}(\lambda)$ is not a good measure of global data statistical consistency,
431 and is only appropriate in regional studies where the natural variability of the ocean color spectra
432 is rather small.

433 We have also analyzed the reflectance anomaly $\Delta\rho_{wN}(\lambda)$ with retrievals restricted to
434 oligotrophic waters (areas where multi-year Chl-a average is less than 0.1 mg/m³), which yield
435 very similar results. Since the oligotrophic waters do not extend to latitudes 40° from the
436 equator, the values of the solar-zenith angle are also generally smaller, and the retrieval
437 conditions are overall better. This is reflected in lower anomaly across the swath dependence.
438 The dependences on the physical ancillary parameters are also very close to those from the
439 global deep water discussed in the previous sections. Furthermore, we have also repeated the
440 analysis by restricting the retrievals to those with QA score (Wei *et al.*, 2016) higher than 0.6.
441 These results are even closer to those described in the previous sections, with only slightly
442 smaller systematic biases for the retrievals with high solar zenith angle and high wind speed.

443 It is noted that while the method in this study identifies anomalies in ocean color data
444 retrievals, it does not identify the exact sources of data inconsistencies. For example, this
445 analysis cannot distinguish between inconsistencies introduced by suboptimal sensor calibration,
446 and various parts of the atmospheric correction and retrieval algorithms. Any correlation of the
447 deviations in anomaly with particular retrieval parameters should be seen only as a hint for what
448 part of the retrieval process might need a further examination. Furthermore, although we have
449 attempted to investigate the dependencies with respect to the most relevant retrieval parameters,
450 these results are not exhaustive in the sense that there may be other significant parameters or
451 combinations of parameters that provide yet more information on the quality and consistency of
452 the retrievals.

453 **5. Conclusions**

454 We have analyzed recently reprocessed VIIRS-SNPP ocean color data for the entire 2016
455 year for statistical consistency over the global deep ocean. The results show very small to

456 negligible deviations from average values for most retrieval parameters. We note somewhat
457 increased $\rho_{wN}(\lambda)$ spectra in all bands for higher wind speed (> 14 m/s), and also in areas close to
458 clouds, likely due to the effect of straylight. We also find that $\rho_{wN}(\lambda)$ for the blue bands are
459 underestimated in the areas with heavy aerosol presence in atmosphere, while the red band is
460 overestimated in those conditions.

461 We have also demonstrated how the analysis presented in this work can be used to identify
462 and distinguish the regions of parameters (such as solar- and sensor-zenith angles) with a
463 systematic bias in the retrieved data, and how different versions of retrieval algorithms can be
464 evaluated based on retrieval consistency. Although this study only covers one year of data, we
465 have not observed any significant changes in any of our results for different years in previous
466 studies. Also, further analysis shows that varying the time and length scale of the average values
467 used to calculate the anomaly does not significantly impact the results presented in this study.
468 While the list of retrieval parameters considered and investigated in this study is not exhaustive,
469 it provides a comprehensive test for data statistical consistency, and means to identify the
470 systematic biases. This is particularly helpful in the design of better and more precise satellite
471 ocean color retrieval algorithms, as well as identification of what areas to look into where these
472 algorithms can be improved.

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