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CALIBRATION OF THE ADVANCED MICROWAVE SOUNDING UNIT-A RADIOMETERS FOR NOAA-N AND NOAA-N'

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NOAA Technical Report NESDIS 106

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CALIBRATION OF THE ADVANCED MICROWAVE SOUNDING UNIT-A RADIOMETERS FOR NOAA-N AND NOAA-N'

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ABSTRACT

After successful launches of NOAA-K, -L, and -M, receptively, in 1998, 2000, and 2002, NOAA will launch its next two satellites, NOAA-N and -N' within next few years. Both will carry the Advanced Microwave Sounding Unit-A (AMSU-A) instruments. This report concerns the analysis and evaluation of the thermal-vacuum chamber calibration data from the two AMSU-A flight models for NOAA-N and -N'. These pre-launch calibration data were analyzed to evaluate the instrument performance, including calibration accuracy, nonlinearity, and temperature sensitivity. Great effort was taken to understand the instrument's radiometric performance as a function of instrument temperature. The calibration data provide a base for derivation of the calibration parameters input data sets (CPIDS) which will be incorporated into the NOAA operational calibration algorithm for producing the AMSU-A IB data sets.

The nonlinearity parameters, which will be used for correcting the nonlinear contributions from an imperfect square-law detector, were determined from this data analysis. The existence and magnitude of nonlinearity in each channel were established and simulated with a quadratic formula for modeling the nonlinear contributions developed in the analysis of the NOAA-KLM AMSU-A pre-launch calibration data. The model was characterized by a single parameter u, values of which were obtained for each channel via least-squares fit to the data. Quadratic corrections which would be expected from the on-orbit data after the launch of AMSU-A into space were simulated. In these simulations, the cosmic background radiance corresponding to a cold space temperature 2.73K was adopted as one of the two reference points of calibration. The largest simulated nonlinear correction is about 2 K. Experience learned from examining the NOAA-15, -16, and -17 AMSU-A on-orbit data provides a better understanding of the AMSU-A performance in space and helps process these pre-launch calibration data. The calibration information presented in this report will be useful for post launch on-orbit verification of the AMSU-A instrument performance.

1. INTRODUCTION

On 13 May 1998, the NOAA-K, which is designated NOAA-15 after the launch, was successfully launched into a circular, near-polar orbit with an altitude of 833-km above the Earth and an inclination angle of 98.7° to the Equator. NOAA-15 carries the first of a series of new microwave total-powerradiometers, the Advanced Microwave Sounding Units (AMSU-A and AMSU-B), which provide 20 channels for atmospheric temperature and moisture soundings. NOAA-L/16, which was launched on September 28, 2000, carries the second AMSU-A. The NOAA-M/17 AMSU-A was successfully launched on June 24, 2002. NOAA-17 is the first NOAA satellite that has been launched into an orbit to cross the Equator at 1000 and 2200 local time. NOAA-N and -N' will be launched in 2004 and 2008, respectively.

The AMSU-A is divided into two physically separate modules, each of which operates and interfaces with the spacecraft independently. The AMSU-A₁ module uses two independent antenna-radiometer systems (A 1-1 and Al-2) to provide 12 channels in the range of 50.3 to 57.3 GHz oxygen band for retrieving the atmospheric temperature profiles from the Earth's surface to about 50 kilometers (or ¹ mb), and another channel at 89 GHz. The AMSU-A2, which has 2 channels at 23.8 and 31.4 GHz, are used to identify precipitation and correct the effect of surface emissivity, atmospheric liquid water, and water vapor on temperature sounding. These window channels are also used to derive rain rate, sea ice concentration, and snow cover.

Table ¹ lists some of the NOAA-N AMSU-A main channel characteristics, including channel frequency, number of bands, 3-dB RF band width, radiometric temperature sensitivity (or NEAT), antenna beam efficiency, polarization, and field of view (FOV) angular beam width for each channel. More detailed information on the AMSU-A radiometers is reported elsewhere [1]. Each of the AMSU-A antenna systems requires to have a nominal FOV of 3.3° at the half-power points and to cover a cross-track scan of ±48°20' (to beam centers) from the nadir direction with 30 Earth FOVs per scan line. Beam positions ¹ and 30 are the outermost scan positions of the Earth views, while beam positions 15 and 16 (at $\pm 1.67^\circ$ from nadir) straddle the nadir. Views of cold space and a blackbody target at the end of scan provide onboard calibration once each scan (8 sec).

The U. K. Meteorological Office provides the AMSU-B for humidity sounding. It has two channels at 89 and 150 GHz, respectively, and three channels around the 183 GHz water vapor line, detailed description of the AMSU-B radiometric performance has been given elsewhere [2].

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All of the AMSU-A flight models were tested and calibrated in a thermal-vacuum (TV) chamber by the contractor. These pre-launch TV calibration data were evaluated and analyzed to derive the calibration parameters input data set (CPIDS) which is used in the NOAA operational calibration algorithm to produce the AMSU-A IB data sets. Particularly, there is a small nonlinearity (of the order of ¹ K or less), which cannot be evaluated by the two-point calibration but is determined from the pre-launch calibration data.

In this study, the TV test data from the NOAA-N and -N' AMSU-A instruments are analyzed and evaluated. The same procedure developed for NOAA-K AMSU-A calibration [1] is closely followed. Experience gained from examining the NOAA-15 and NOAA-16 AMSU-A on-orbit data provides a better understanding of the AMSU-A performance in space and helps improve the analysis of the pre-launch calibration data. Some samples of AMSU-A on-orbit data are presented to demonstrate the dynamic variation of the calibration coefficients and calibration counts as a function of instrument temperature.

In the following sections, the results from the data analysis are presented. Instrument performances evaluated in this analysis include calibration accuracy, nonlinearity, and radiometric temperature sensitivity (or NEAT, the noise-equivalent temperature). Section ¹ gives an introduction and section 2 presents a brief description of the TV chamber test data. The calibration algorithm is described in section 3. The NOAA-N results are presented in section 4, while Appendix B has the NOAA-N' results. Conclusion and discussion are in section 5. Tables of CPLDS for NOAA-N and -N' are given in Appendices A and B, respectively. A brief description of the NOAA Level 1B data is given in Appendix C, which also contains some samples of AMSU-A on-orbit calibration coefficients and radiometric blackbody and space counts.

2. DESCRIPTION OF CALIBRATION DATA

Aerojet (now Northrop Grumman), the AMSU-A contractor, took the calibration data in a TV test chamber using the full scan mode. In this mode, AMSU-A scans through 30 Earth FOVs, the cold target, and the internal warm blackbody target once every 8 seconds. It takes one sample at each Earth FOV and two samples at the cold and warm targets, respectively. Since the scene calibration target was fixed at FOV 6 (31 40' from nadir), only radiometric counts from FOV 6 can be used for calibration. Each antenna system looks at its own individual cold, warm, and scene targets. A series of Platinum Resistance Thermometers (PRTs) were used to monitor the temperatures of these

Table 2. Number of PRTs in each calibration target and channels provided by individual antenna systems. The last row gives the number of scans collected in the calibration for each system.

Note: For Al-1 & Al-2: N=400, 400, 550, 725, 725, and 900 when

the scene target temperature at T_s =84, 130, 180, 230, 280, and 330K, respectively.

calibration targets. The numbers of PRTs used to measure the physical temperatures of the scene, cold, and warm calibration targets in each antenna system are given in Table 2, which also lists the channels provided by each AMSU-A antenna system. Channels 9-14 have both primary and secondary phase locked loop oscillators (called PLLO #1 and PLLO #2, respectively) built-in. The PLLO #2 will be used for backup if PLLO #1 fails in operation. Invar high-Q cavity stabilized local oscillators [3] are used in other channels.

The physical temperatures of scene and cold targets measured by individual PRTs were provided in Kelvin (K) on Aerojet's data packages. However, the data from the PRTs monitoring the warm blackbody targets are given in counts, which are proportional to the blackbody temperatures. One should note that the scene and cold targets used in the TV chamber will not be carried into space. The PRT counts from the warm blackbody targets must be converted to PRT temperatures. The normal approach of deriving the PRT temperatures from counts is a two-step process: (1) the resistance of each PRT in ohms is computed by acount-to-resistance look-up table provided by the manufacturer. In this study, we used a polynomial representation of the count-to-resistance

relationship provided by Aerojet; and (2) the individual PRT temperature in degrees Celsius is obtained from an analytic PRT equation [4], which is described in Appendix A. However, the two steps can be compressed to a single step with negligible errors. This single step process, which has been used in the NOAA-KLM calibrations, computes the PRT temperatures directly from the PRT counts, using a cubic polynomial

$$
T_{Wk} = \sum_{j=0}^{3} f_{kj} C_k^j
$$
 (1)

where T_{wk} and C_k represent the temperature and count of the PRT k. The polynomial coefficients, f_{kj} , are derived for each PRT in this study. Equation (1) also applies to 47 other housekeeping temperature sensors, such as the mixers, the IF amplifiers and the local oscillators. These f_{ki} coefficients for all PRTs and housekeeping sensors in NOAA-N AMSU-A1 S/N 109 and AMSU-A2 S/N 105 are listed in Appendix A (Tables A-l and A-2, respectively), while the corresponding ones for the NOAA-N' AMSU-A are listed in Appendix B.

The mean internal blackbody temperature, T_w , is calculated from the individual PRT temperatures,

$$
T_{W} = \frac{\sum_{k=1}^{m} W_{k} T_{Wk}}{\sum_{k=1}^{m} W_{k}} + \Delta T_{W}
$$
 (2)

where m represents the number of PRTs for each antenna system (as listed in Table 2) and W_k is a weight assigned to each PRT k. The quantity ΔT_w is a warm load correction factor, which is derived for each channel from the TV test data at three instrument temperatures (low, nominal, and high). The procedure for determining the ΔT_W values has been described elsewhere [5]. For each AMSU-A antenna system, a special set of calibration data was taken to determine the ΔT_w values, which are given in Appendix A (Table A-3). The W_k value, which equals 1 (0) if the PRT k is determined good (bad) before launch. Similarly, the mean temperatures of cold targets are defined in the same way as in Equation (2), except without the term ΔT_w .

The TV calibration data were taken at three instrument temperatures and the scene target was cycled through six temperatures 84, 130, 180, 230, 280, and 330K, respectively, at each instrument temperature. At each of the scene target temperatures, calibration data were acquired for enough number of scans to assure an effective temperature sensitivity less than 0.03K. The number of scans

required depends upon the expected NEAT of channel 14 (which has the largest NEAT) and is therefore a function of the scene target temperature. Actual numbers of scans taken at individual temperatures are given in Table 2. The uncertainties in knowledge of brightness temperatures and measurement errors were obtained by Aerojet [6, 7]. These are given Table 3.

Source of Error		AMSU-A2:Ch. 1 and 2	AMSU-A1: Ch. 3 - 15		
	Bias(K)	Random (k)	Bias(K)	Random (k)	
Warm Target	±0.122 -0.050		-0.050	±0.122	
Cold Target	0.024	± 0.105	0.024	±0.091	
Scene Target	0.002	±0.101	0.002	± 0.090	

Table 3. Uncertainty in brightness temperatures and measurement errors.

Antenna beam widths at all channels were also measured and the values for NOAA-N are listed in Table 1. Antenna beam efficiency at each channel frequency is greater than 95% and Table ¹ lists the calculated values.

3. CALIBRATION ALGORITHM

In this study, calculations of radiometric measurements and variables related to the calibration process are all performed in radiance with dimension of $mW/(m^2-sr-cm^{-1})$, but the final results are presented in temperatures. All instruments flown on NOAA satellites produce measurements in radiance. Conversion between brightness temperature and radiance was performed using the full Planck function, instead of the Rayleigh-Jeans approximation. This also eliminates any possible conversion inaccuracy that may occur, particularly in the region of the space cosmic back-ground temperature -2.73K, where the Rayleigh-Jeans approximation breaks down.

For each scan, the blackbody radiometric counts C_w are the averages of two samples of the internal blackbody. Similarly, the space radiometric counts C_c are the average of two samples of the space target. To reduce noise in the calibrations, the C_x (where X=W or C) for each scan line were convoluted over several neighboring scan lines according to a weight function [1]

$$
\overline{C}_X = \frac{\sum_{i=-n}^{n} W_i C_X(t_i)}{\sum_{i=-n}^{n} W_i}
$$
 (3)

where t_i (when $i \ne 0$) represents the time of the scan lines just before or after the current scan line and t_0 is the time of the current scan line. One can write $t_i = t_0 + i\Delta t$, where $\Delta t = 8$ seconds for AMSU-A. The 2n+l values are equally distributed about the scan line to be calibrated. Following the NOAA-KLM operational preprocessor software, the value of n=3 is chosen for all AMSU-A antenna systems. A set of triangular weights of $1, 2, 3, 4, 3, 2$ and 1 is chosen for the weight factor W_i that appears in Equation (3) for the seven scans with $i = -3, -2, -1, 0, 1, 2,$ and 3.

The calibration algorithm [1], which takes into account the nonlinear contributions due to an imperfect square-law detector, converts scene counts to radiance R_s as follows,

$$
R_{S} = R_{W} + \left(R_{W} - R_{C}\right) \left(\frac{C_{S} - \overline{C}_{W}}{\overline{C}_{W} - \overline{C}_{C}}\right) + Q \tag{4}
$$

where R_w and R_c are the radiance computed from the PRT blackbody temperature T_w and the PRT cold target temperature T_c , respectively, using the Planck function. The C_s is the radiometric count from the Earth scene target. The averaged blackbody and space counts, $\overline{C_W}$ and $\overline{C_C}$, are defined by Equation (3). The quantity Q, which represents the nonlinear contribution, is given by $[1]$

$$
Q = u\left(R_w - R_c\right)^2 \frac{\left(C_s - \overline{C}_w\right)\left(C_s - \overline{C}_c\right)}{\left(\overline{C}_w - \overline{C}_c\right)^2}
$$
(5)

where u is a free parameter, values of which are determined at three instrument temperatures (low, nominal, and high). After launch of the instruments, the u values at the actual on-orbit instrument temperatures will be interpolated from these three values. For channels 9-14 (AMSU-A1-1), two sets of the u parameters are provided; one set is for the primary PLLO#l and the other one for the redundant PLLO#2.

The quantity R_s in Equation (4) represents the radiometric scene radiance of individual channels. For users of NOAA Level 1B data, a simplified formula, which converts C_s directly into R_s , is

presented in Appendix C. It should be noted that the ratios in Equations (4) and (5) will eliminate the effect of any linear variation in the radiometric counts on R_s . The channel gain, G, is defined as

$$
G = \frac{\overline{C}_W - \overline{C}_C}{R_W - R_C} \tag{6}
$$

The quantity G varies with instrument temperature, which is defined as the RF Shelf temperature for each AMSU-A antenna system. For a fixed instrument temperature, G is approximately constant. The first two terms in Equation (4) constitute a linear two-point calibration equation, if the quadratic term is negligible. Let R_{SL} denote these two terms,

$$
R_{SL} = R_W + \left(R_W - R_C\right) \left(\frac{C_S - \overline{C}_W}{\overline{C}_W - \overline{C}_C}\right) \tag{7}
$$

The linear scene radiance R_{SL} are calculated from the TV chamber test data. The results are given in the following section.

4. RESULTS

4.1 *Calibration Accuracy*

The calibration accuracy, ΔR , is defined as the difference between the scene PRT radiance R_{spr} and the radiometric scene radiance. It is calculated from the equation

$$
\Delta R = \frac{1}{N} \left[\sum_{i=1}^{N} \left(R_{sprt} - R_{SL} \right)_i \right]
$$
 (8)

where N is the number of scans at a specific scene temperatures (N= 120 for AMSU-A2 channels but ranging from 400 to 900 for AMSU-A1 channels).

Equation (7) would be a good representation of the microwave radiometric scene radiance for a perfect square-law detector. Any deviation of the quantity R_{SL} from the measured scene PRT radiance R_{sprt} indicates the presence of either nonlinearity in the radiometer system or some other potential source of contamination in the calibration data.

Figure 1 shows the calculated calibration accuracies in temperature, ΔT (corresponding to ΔR), versus the scene PRT temperature for channels ¹ through 15. The AMSU-A specification requires $\Delta T = 2.0$ K for channels 1, 2, and 15, and $\Delta T = 1.5$ K for all other channels. The results in Figure 1 are better than the specification at all channels.

4.2 *Nonlinearity*

The ΔT patterns (Figure 1) show clearly the nonlinearity patterns which can be represented by the quadratic formula Q defined in Equation (5). The nonlinearity of a channel is normally defined as the residuals from a least-squares fit of its scene PRT radiance as a linear function of the radiometric radiance $R_{\rm SL}$ (Equation 7). The nonlinearity is defined as the differences (or residuals) between the R_{spr} and the best-fit results from a linear equation in the form LinFit = a + b R_{SL} (where a and b are the intercept and slope). These residuals, which are defined as the measured nonlinearity $Q (= R_{sort})$ - LinFit), are shown in Figure 2. The largest (absolute) Q value on each curve is defined as measured nonlinearity that can be compared to those as defined in the AMSU-A specification.

The results in Figure 2 show that the maximum (absolute) Q values are about 0.7K. The AMSU-A specification requires $Q = 0.5K$ for channels 1, 2, and 15; and $Q = 0.375K$ for other channels. Some results in Figure 2 do not meet the specifications. Most of the curves in Figure 2 have two roots, representing solutions of a quadratic equation, which can be written in the form

$$
Q = u\big(R_S - R_1\big)\big(R_S - R_2\big) \tag{9}
$$

where R_1 and R_2 represent the two roots. One can obtain a similar equation from Equation (5) by replacing the counts by its individual radiance, since the radiometric counts are proportional to radiance in a first-order approximation. The resultant equation represents a different straight line intersecting the same curve at R_w and R_c . Once the parameter u is determined, it can be used with any pair of roots to calculate Q. We applied Equation (9) to fit the quadratic curves in Figure 2 to obtain the u values with the two roots extracted from each plot. Table 4 gives the best-fit u values at three instrument temperatures for individual channels.

Figure 1. NOAA-N AMSU-A: Calibration accuracies at three instrument temperatures.

Figure 2. NOAA-N AMSU-A: Residuals after least-squares fit of the scene PRT radiance as a function of R_{SI} as discussed in the text.

Table 4. NOAA-N nonlinearity parameters u in dimension of (m²-sr-cm⁻¹)/mW. **Table 4. NOAA-N nonlinearity parameters u in dimension of (m2-sr-cm 'VmW.** AMSU-A1-2 S/N 109 channels: **AMSU-A2 S/N 105 channels: AMSU-A1-2 S/N 109 channels:** AMSU-A2 S/N 105 channels:

AMSU-A1-1 S/N 109 channels: PLLO#1

AMSU-A1-1 S/N 109 channels: PLLO#2

4.3 Simulated Quadratic Corrections to On-Orbit Data

Once the u values are determined, Equation (9) can be used to simulate the quadratic contributions which are expected from on-orbit data. This can be accomplished by replacing the roots R_1 and R_2 with the radiance $R_{2,73}$ (corresponding to cold space temperature 2.73K) and R_w , respectively. The simulated results are shown in Figure 3 for all channels at three instrument temperatures which are listed at the top of the plots. One should note that the instrument temperatures are different for the AMSU-A antenna systems. The simulated quadratic contributions displayed in Figure 3 are larger than those shown in Figure 2. This is expected because the separation of the two reference calibration points is increased in the cases of simulations. Noticeable quadratic contributions appear in all channels, particularly the AMSU-A1-1 channels. The largest ones are approximately 2 K at several channels. It is important to note that the effect of instrument temperature on the quadratic contributions is nonlinear in general.

Similarly, Figure 4 shows the calculated results which are associated with the redundant PLLO#2 built into channels 9-14. The left column of Figure 4 displays the calibration accuracies, ΔT , which are similar to Figure ¹ and the middle column (corresponding to Figure 2) shows the residuals ofthe least-squares fit. The right column shows the simulated quadratic corrections, Q, which would occur in on-orbit data.

4.4 Temperature Sensitivity

The temperature sensitivity (or NEAT) specification for AMSU-A channels are listed in Table 1. It is defined as the minimum change of a scene brightness temperature that can be detected. In practice, it is calculated as the standard deviation of the radiometer output (in K), when an antenna system is viewing a scene target at 300K. The calculated NEAT values are shown in Figure 5 together with the AMSU-A specifications. These calculated NEAT values correspond to a scene temperature of 305K, because they are the average values of two measurements taken at the scene target temperatures of 280K and 330K, respectively (no calibration data were taken at 300K). All of the calculated NEAT values in Figure 5 are better than those given in the AMSU-A specification. Actually, all NEAT values measured at both 280 and 330K are better than the specification. The calculated NEAT values at other instrument temperatures are shown in Figure 6. In general, the NEAT value decreases as the instrument temperature becomes colder.

Figure 3. NOAA-N AMSU-A: Simulated nonlinearity as expected from on-orbit data.

NOAA-N: AMSU-A2 S/N 105 RF-Shelf Temperature (C): xx= -6.2, **=12.2, ++=30.7 AMSU-A1 S/N 109 RF-Shelf Temperature (C): xx= -2.3, **=18.2, ++=38.0

Figure 4. NOAA-N AMSU-A: Calibration results withe PLLO #2 in Channels 9-14, including calibration accuracy, measured nonlinearity, and simulated nonlinearity.

AMSU-A1 S/N 109, AMSU-A2 S/N 105 NEdT (K): xx = Measured, ** = Specification

Figure 5. NOAA-N AMSU-A: Comparison of the measured NEAT values with specification.

NOAA-N: AMSU-A2 S/N 105 RF-Shelf Temperature (C): xx= -6.2, **=12.2, ++=30.7 AMSU-A1 S/N 109 RF-Shelf Temperature (C): xx= -2.3, **=18.2, ++=38.0

Figure 6. NOAA-N AMSU-A: Measured NEAT values at three instrument temperatures.

4.5 Radiometric Counts at Zero Radiance

Figure 7 shows the plots of TV scene radiometric counts versus the scene PRT temperatures for all channels. The radiometric counts increase linearly with scene PRT temperatures ranging from 84 to 330K and good linear relationships exist between the radiometric counts and the scene PRT temperatures in the range 84 to 330K. One can extrapolate these linear relationships to 0 K (corresponding to zero radiance) and obtain the intercepts for these plots. The intercept for each data point can also be computed from Equation (7) by setting $R_{SL}= 0$ and solve for C_s , which is denoted by C_{Sint} as

$$
C_{\text{S int}} = \overline{C}_W - G R_W \tag{10}
$$

where G is the channel gain defined in Equation (6). The calculated C_{Sint} values at three instrument temperatures for each channel are shown in Figure 8. These are the mean values of all calculations performed with all available calibration data (from 120 to 900 scans as the scene temperature varies from 84 to 330K). At each instrument temperature, the variation in the calculated C_{Sint} values is relatively small and all C_{Sint} values are positive. Figure 8 shows that the instrument temperature has a big impact on the magnitude of C_{Sint}, which increases as the instrument temperature decreases.

4.6 Channel gains

Values of the channel gains as defined in Equation (6) are also calculated from the TV calibration data. The gain values are converted into dimension of count/K. Figure 9 shows the calculated channel gains at three instrument temperatures, which are listed at the top of Figure 9. One should note that the gains of Channel 15 are smaller than those of other channels by about a factor 2 and that the abnormal behavior of channel ¹ gain as a function of instrument temperature. Similarly, Figure 10 shows the calculated results of channel gain, radiometric counts, and intercept counts as a function of the scene PRT temperature when PLLO#2 was used.

5. CONCLUSION AND DISCUSSION

The TV chamber calibration data for the NOAA-N AMSU-A, including A1 S/N 109 and A2 S/N 105, and the NOAA-N' AMSU-A, A1 S/N 105 and A2 S/N 107, were analyzed to derive the CPIDS which will be used in the NOAA operational calibration algorithm to produce the AMSU-A IB data

Figure 7. NOAA-N AMSU-A: Radiometric counts versus scene PRT temperature at three instrument temperatures

Figure 8. NOAA-N AMSU-A: Intercept counts at three instrument temperatures.

NOAA-N: AMSU-A2 S/N 105 RF-Shelf Temperature (C): xx= -6.2, **=12.2, ++=30.7 AMSU-A1 S/N 109 RF-Shelf Temperature (C): xx=-2.3, **=18.2, ++=38.0

Figure 9. NOAA-N AMSU-A: Channel gains at three instrument temperatures

NOAA-N: AMSU-A2 S/N 105 RF-Shelf Temperature (C): xx= -6.2, **=12.2, ++=30.7 AMSU-A1 S/N 109 RF-Shelf Temperature (C): xx=-2.3, **=18.2, ++=38.0

Figure 10. NOAA-N AMSU-A: Calibration results with PLLO #2 in Channels 9-14, including gains, radiometric counts, and intercept counts at three instrument temperatures.

sets. The results show that both sets of instruments meet the AMSU-A specification in calibration accuracy and temperature sensitivity, but the measured nonlinearities at several channels do not meet the AMSU-A specification.

The nonlinearity exists at all channels. A quadratic formula with a single parameter u is used to simulate these nonlinear contributions. The u values at three instrument temperatures were obtained from the pre-launch calibration data. Using the best-fit u values, the quadratic corrections which would be expected from the on-orbit data were simulated. In the simulations, the cold space radiance corresponding to 2.73K was adopted as one of the two reference calibration points (the other one is the internal blackbody temperature). The largest simulated nonlinear correction is about 2 K as shown in Figure 3.

Experience gained from examining the NOAA-15 and NOAA-16 AMSU-A on-orbit data provides a better understanding of the AMSU-A performance in space and helps process the pre-launch calibration data. Some samples of on-orbit AMSU-A data are presented in Appendix C to demonstrate the dynamic variations of the calibration coefficients and calibration counts as a function of scan number over an orbit. The time series of the calibration coefficients shows why they must be calculated at each scan lines and that the pre-launch values, which do not have the dynamic variation induced by the on-orbit changes in instrument temperature, can not be employed to convert the radiometric counts into temperatures.

In general, the qualities of the calibration data are quite good. Particularly, the instrument temperatures were stabilized at pre-selected values with total variation less than $\pm 0.5K$ during each test cycle. This renders the measured instrument nonlinearities more reliable. The results presented in this study confirm the good quality of the AMSU-A instruments for both NOAA-N and -N'. The calibration information presented in this report will be immensely useful for post launch on-orbit verification of the AMSU-A instrument performance.

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APPENDIX A

NOAA-N CPIDS : COEFFICIENTS OF AMSU-A1 S/N 109 AND AMSU-A2 S/N 105

A.1 *Polynomial Coefficientsfor Converting PRT Counts into Temperatures*

The two-step process for deriving the PRT temperatures from PRT counts C_k is briefly described here. First the count C_k from PRT k is converted into resistance r_k (in ohms) by a polynomial

$$
r_k = \sum_{i=0}^{3} A_i C_k^i
$$
 (A-1)

where the coefficients A_i for individual PRTs and temperature sensors were provided by Aerojet. Once the resistance r_k is known, then one can calculate the PRT temperature t (in Celsius) from the Callendar-Van Dusen equation [6], which is given by

$$
\frac{r_t}{r_o} = 1 + \alpha \left[t - \delta \left(\frac{t}{100} - 1 \right) \frac{t}{100} - \beta \left(\frac{t}{100} - 1 \right) \left(\frac{t}{100} \right)^3 \right]
$$
 (A-2)

where: r_i = resistance (in ohms) at temperature t (\degree C) of the blackbody target

 r_0 = resistance at t = 0°C (supplied by the manufacturer via Aerojet)

 $\beta = 0$ for $t > 0$ °C, and 0.11 for $t < 0$ °

 α and δ are constants provided by the manufacturer via Aerojet.

Calculation shows that the error is negligible by setting $\beta = 0$ in Equation (A-2). In this study, it is assumed $\beta = 0$ in all cases, then Equation (A-2) is simplified into a quadratic equation in t. In such case, one can solve the quadratic equation for t in terms of r_t . Then the PRT temperature in degree Kelvin is obtained from $t + 273.15$. By this way, data sets of PRT temperatures versus counts for individual PRTs are computed. Then Equation (1) is applied to fit these data sets for obtaining the polynomial coefficients f_{ki} for individual PRTs and housekeeping sensors. These best-fit coefficients are listed in Tables A-l and A-2, respectively, for AMSU-A1 S/N 109 and AMSU-A2 S/N 105. Test calculations show that these polynomials are highly accurate in reproducing temperatures of PRTs and sensors with errors in order of 0.01 K.

Table A-l. NOAA-N AMSU-A1 S/N 109: Polynomial coefficients for converting PRT counts into temperatures.

PRT	Description	Fk0	Fk1 Fk ₂		Fk3
No.		(K)	(K/count)	(K/count^2)	(K/count^3)
$\mathbf{1}$	Al-1 MOTOR	264.1476	1.743266E-03	3.348404E-09	1.664095E-14
\overline{c}	AI-2 MOTOR	263.1678	1.731684E-03	3.541911E-09	1.427281E-14
3	AI-1 FEED HORN	263.4037	1.756592E-03	3.025439E-09	2.592502E-14
$\overline{4}$	AI-2 FEED HORN	263.7585	1.736896E-03	3.524135E-09	9.021489E-15
5	AI-1 RF MUX	263.2661	1.734418E-03	3.627309E-09	1.270486E-14
6	AI-2 RF MUX	263.6731	1.710861E-03	3.745664E-09	1.954603E-14
$\overline{7}$	CH. 3 DRO	263.2463	1.736519E-03	3.450645E-09	1.511849E-14
8	CH. 4 DRO	263.0471	1.729353E-03	3.629986E-09	1.108148E-14
$\mathsf g$	CH. 5 DRO	263.4903	1.732073E-03	3.660650E-09	1.059084E-14
10	CH. 6 DRO	263.4986	1.743907E-03	3.378800E-09	1.731537E-14
11	CH. 7 DRO	263.4426	1.739730E-03	3.231359E-09	1.741037E-14
12	CH. 8 DRO	263.1023	1.739080E-03	3.435317E-09	1.558292E-14
13	CH. 15 GDO	262.9301	1.733167E-03	3.619563E-09	1.102405E-14
14	CH. 9 thru 14 PLO #2	262.2288	1.782365E-03	3.927203E-09	1.074884E-14
15	CH. 9 thru 14 PLO #1	263.3337	1.730256E-03	3.892761E-09	6.229168E-15
16	Not Used	263.3337	1.730256E-03	3.892761E-09	6.229168E-15
17	CH. 3 MIXER/IF	262.5283	1.728151E-03	3.730241E-09	1.055708E-14
18	CH. 4 MINER/IF	262.6262	1.735880E-03	3.482120E-09	1.463006E-14
19	CH. 5 MINER/IF	263.0344	1.729980E-03	3.817571E-09	1.028369E-14
20	CH. 6 MIXER/IF	263.3519	1.737770E-03	3.370750E-09	1.684445E-14
21	CH. 7 MIXER/117	263.5179	1.740044E-03	3.604026E-09	8.980608E-15
22	CH. 8 MIXER/IF	262.9976	1.731195E-03	3.710931E-09	1.056478E-14
23	CH. 9 thru 14 MIXER/IF	262.8521	1.733479E-03	3.670787E-09	1.256491E-14
24	CH. 15 MIXER/IF	263.6473	1.747099E-03	3.570097E-09	5.402493E-15
25	CH. 11 thru 14 IF AMP	263.3202	1.739659E-03	3.556506E-09	1.392298E-14
26	CH. 9 IF AMP	263.3429	1.735609E-03	3.941727E-09	8.254678E-15
27	CH. 10 IF AMP	263.0992	1.736573E-03	3.728037E-09	1.006676E-14
28	CH. 11 IF AMP	263.4255	1.735551E-03	3.770763E-09	5.931839E-15
29	DC/DC CONVERTER	263.7365	1.730352E-03	4.034858E-09	5.881655E-15
30	CH. 13 IF AMP	263.3358	1.738157E-03	3.671254E-09	1.245797E-14
31	CH. 14 IF AMP	263.1836	1.740586E-03	3.337613E-09	1.651644E-14
32	CH. 12 IF AMP	263.4875	1.735195E-03	3.754301E-09	1.074719E-14
33	AI-1 RFSHELF	264.1607	1.745635E-03	3.690702E-09	1.142299E-14
34	A1-2 RF SHELF	263.4113	1.743746E-03	3.503678E-09	1.460422E-14
35	DETECTOR/PRE-AMP	263.7069	1.735678E-03	3.727127E-09	1.026177E-14
36	AI-1 WARM LOAD 1	254.4792	1.632697E-03	5.925882E-09	2.790891E-14
37	AI-1 WARM LOAD 2	254.9357	1.632574E-03	5.837758E-09	2.895840E-14
38	AI-1 WARM LOAD 3	254.6461	1.628482E-03	5.867548E-09	2.882307E-14
39	AI-1 WARM LOAD 4	254.7101	1.628115E-03	5.919407E-09	2.831376E-14
40	A1-1 WARM LD CENTER	254.6003	1.605916E-03	8.652163E-09	$-4.137281E-14$
41	A1-2 WARM LOAD 1	255.4882	1.617997E-03	4.542892E-09	6.722034E-14
42	A1-2 WARM LOAD 2	254.5689	1.632253E-03	5.841521E-09	2.896778E-14
43	A1-2 WARM LOAD 3	254.4614	1.630541E-03	5.753884E-09	2.950770E-14
44	AI-2 WARM LOAD 4	254.7519	1.634664E-03	5.835678E-09	2.836517E-14
45	A1-2 WARM LD CENTER	254.8793	1.635115E-03	5.835015E-09	2.919876E-14

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Table A-2. NOAA-N AMSU-A2 S/N 105: Polynomial coefficients for converting PRT counts into temperatures.

PRT	Description	Fk0	Fk1	Fk ₂	Fk3
No.		(K)	(K/count)	(K/count ^{^2})	(K/count^3)
	Scan Motor	263.5758	1.767645E-03	3.794875E-09	1.261675E-14
2	Feedhorn	263.1411	1.764382E-03	3.661103E-09	1.305692E-14
3	RF Diplexer	263.1864	1.752475E-03	3.844197E-09	1.227221E-14
4	Mixer/IF CH1	263.8281	1.754974E-03	3.802008E-09	1.077093E-14
5	Mixer/IF CH2	263.7458	1.756229E-03	3.727035E-09	1.308870E-14
6	CHIDRO	263.9707	1.758280E-03	3.718871E-09	1.248273E-14
7	CH ₂ DRO	263.4596	1.756287E-03	3.826887E-09	1.288159E-14
8	Compensat Motor	263.9833	1.757519E-03	3.769221E-09	1.130973E-14
9	Sub Reflector	262.8663	1.756615E-03	3.798051E-09	1.160802E-14
10	DODC Converter	263.8742	1.763420E-03	3.708442E-09	1.357872E-14
11	RF Shelf	263.9671	1.752322E-03	4.039884E-09	9.326652E-15
12	Det. Pre-Amp	263.4358	1.750480E-03	3.754149E-09	1.171233E-14
13	Warm Load Ctr	254.7616	1.653192E-03	5.988277E-09	3.101395E-14
14	Warm Load #1	254.5737	1.652519E-03	5.994369E-09	3.214458E-14
15	Warm Load #2	254.6311	1.658779E-03	6.059266E-09	3.010172E-14
16	Warm Load #3	254.8459	1.655427E-03	6.099508E-09	3.020449E-14
17	Warm Load #4	254.8488	1.652000E-03	6.055910E-09	3.032410E-14
18	Warm Load #5	254.7966	1.647523E-03	6.068692E-09	2.974657E-14
19	Warm Load #6	253.9048	1.657694E-03	5.929873E-09	3.021740E-14

A.2 *Warm Load Correction*

The in-flight warm load correction (WLC) was calculated according to a formula developed by Aerojet [6-9]. For each AMSU-A antenna system, a special set of calibration data were acquired by setting the temperature of its variable scene target equal to that of the internal blackbody (warm) target. The physical temperature T_w of the internal blackbody target was determined from the PRT counts as described in Section 2. The radiometric temperature T_{wrad} of the blackbody (in-flight warm load) was calculated (for each scan in the data set) by the formula

$$
T_{wrad} = T_{spr} + \left(T_{spr} - T_c\right) \left(\frac{C_w - C_s}{C_s - C_c}\right) \tag{A-3}
$$

where: $T_{\text{spr}} = \text{PRT}$ temperature of the variable scene target,

 T_c = PRT temperature of the cold target,

 C_w = the average of two radiometric counts from the warm target,

 C_c = the average of two radiometric counts from the cold target, and

 C_s = radiometric counts from the variable scene target.

One should note that temperatures, T_{spr} and T_c , from the scene and cold targets are used as the two reference calibration points in Equation (A-3) to calculate the radiometric temperature of the warm target. The in-flight warm load correction factor ΔT_w was computed from the formula,

$$
\Delta T_W = \frac{1}{N} \left[\sum_{i=1}^{N} \left(T_{wrad} - T_W \right)_i \right] \tag{A-4}
$$

where N represents the number of scans in the data set. The ΔT_w values at three instrument (RF Shelf) temperatures for each AMSU-A antenna system are listed in Table A-3. For AMSU-A1, the ΔT_w values for both PLLO#1 and PPLO#2 were calculated and listed. These ΔT_w values will be used in NOAA-N AMSU-A operational calibration algorithm.

A.3 *Nonlinearity parameters*

The nonlinearity is discussed in Section 4.2 and the values of the nonlinearity parameter u are listed in Table 4.

Table A-3. NOAA-N AMSU-A warm load corrections (K) at three instrument temperatures.

Instrument	Ch.1	Ch.2	Instrument	Ch.3	Ch.4	Ch.5	Ch.8	
Temp(C)			Temp(C)					
-6.24	0.169	0.025	-2.35	-0.293	-0.131	-0.126	-0.173	
12.18	0.125	0.046	18.30	0.016	-0.159	-0.092	-0.225	
30.66	0.066	0.019	38.26	-0.107	-0.222	-0.240	-0.265	

AMSU-A2 S/N 105 channels: AMSU-A1-2 S/N 109 channels:

AMSU-A1-1 S/N 109 channels: PLLO#l:

Instrument Ch.6 Ch.7 Ch.9 Ch.10 Ch.11 Ch.12 Ch.13 Ch.14 Ch.15					
Temp(C)					
-2.33		$\begin{array}{cccccc} 0.482 & 0.447 & 0.445 & 0.441 & 0.449 & 0.462 & 0.467 & 0.510 & 0.434 \end{array}$			
18.26		0.227 0.243 0.232 0.241 0.219 0.249 0.226 0.206 0.215			
38.18		0.296 0.266 0.268 0.281 0.272 0.258 0.195 0.295 0.213			

AMSU-A1-1 S/N 109 channels: PLLO#2:

A.4 *Correction to In-orbit Cold Space Calibration*

For on-orbit cold space calibration, there is an uncertainty due to antenna side lobe interference with the Earth limb and spacecraft. The contribution from this uncertainty should be added to the cold space cosmic background temperature of 2.73K. Therefore, the "effective " cold space temperature T_{EC} can be represented by

$$
T_{EC} = 2.73 + \Delta T_C \tag{A-5}
$$

where ΔT_c represents the contribution due to antenna side lobe interference with the Earth limb and spacecraft. Aerojet made estimates of ΔT_C for individual channels and its values for NOAA-N AMSU-A [6, 7] are listed in Table A-4. These ΔT_c values will be used initially and the final optimal values for ΔT_c will be determined from the antenna pattern corrections [10].

Channel	ΔT_C (K)	$\Delta C_{\rm W}$ (count)
$\mathbf{1}$	0.46	15
$\overline{2}$	0.58	19
3	1.46	12
4	1.46	11
5	1.65	15
6	1.22	8
7	1.25	8
8	1.68	11
9	1.22	10
10	1.22	10
11	1.22	14
12	1.22	18
13	1.22	25
14	1.22	42
15	1.18	3

Table A-4. Values of NOAA-N AMSU-A ΔT_c and limits of blackbody count variations.

A.5 *Limit ofBlackbody Counts Variation*

For each scan, the blackbody count C_w is the average of two samples. If the two samples of the blackbody differ by more than a pre-set limit of blackbody count variation ΔC_w , the data in the scan will not be used. The ΔC_W values for individual channels are listed in Table A-4. These ΔC_W values, which equal approximately 3σ (where σ is the standard deviation of the internal blackbody counts), are calculated from the TV calibration data.

A.6 *Pre-launch Determined Weight Factors wkforthe Internal Blackbody PRTs*

The weight factors w_k (see Equation 2) assigned to individual PRTs in the internal blackbody targets are listed in Table A-5. All NOAA-N AMSU-A PRTs are good.

Table A-5. Pre-launch determined weight factors w_k assigned to NOAA-N AMSU-A PRTs in blackbody targets.

Antenna System	W_1	W_2	W_3	$\rm W_4$	W_5	W_6	W_7
AMSU-A2							
AMSU-A1-1							
$AMSU-A1-2$							

A.7 *Conversion Coefficients of Analog Data*

AMSU-A instrument has an analog telemetry bus to monitor key temperatures and voltages through the spacecraft. The resolutions of the analog telemetry received on the ground is 20 mV for one part in 256. To convert the analog data into physical quantities y, one must multiply the analog values by 0. 02V (20 mV) to obtain the measured output, x, in volts and then uses the conversion equation,

$$
y = B + M x \tag{6}
$$

where the values of B and M are given in Tables A-6 and A-7 for NOAA-N AMSU-A.

UIIS	Description	у	M	B
1	Scanner Motor Temperature	\mathbf{C}	68.027	$\mathbf{0}$
\overline{c}	Comp Motor Temperature	\mathbf{C}	68.027	$\overline{0}$
3	R.F. Shelf Temperature	\mathcal{C}	68.027	$\mathbf{0}$
$\overline{4}$	Warm Load Temperature	\mathbf{C}	68.027	Ω
5	Comp. Motor Current (average)	mA	46.6	$\mathbf{0}$
6	Ant Drive Motor Current (ave)	mA	46.6	Ω
7	Signal Processing +15Vdc	V	4.315	$\mathbf{0}$
8	Antenna Drive +15Vdc	V	4.315	Ω
9	Signal Processing -1 5Vdc	V	2.504	-22.562
10	Antenna Drive - 15Vdc	\mathbf{V}	2.504	-22.562
11	Mixer/IF Amplifier +10 Vdc	$\overline{\mathbf{V}}$	2.889	$\mathbf{0}$
12	Signal Processing +5Vdc	V	1.667	Ω
13	Antenna Drive +5Vdc	V	1.667	Ω
14	Local Oscillator +lOVdc. (ch.1)	V	2.861	Ω
15	Local Oscillator +IOVdc (ch.2)	V	2.861	$\mathbf{0}$

Table A-6. NOAA-N AMSU-A2 S/N 105: Analog data conversion coefficients.

Table A-7. NOAA-N AMSU-A1 S/N 109: Analog data conversion coefficients.

Spacecraft Connector $P6 - Pin#$	Description	y	\mathbf{m}	$\mathbf b$
3	Al-1 Scanner Motor Temperature	$\mathbf C$	68.027	$\bf{0}$
22	A1-2 Scanner Motor Temperature	\mathbb{C}	68.027	Ω
$\overline{2}$	A1-1 RF Shelf Temperature	$\mathbf C$	68.027	Ω
21	A1-2 RF Shelf Temperature	$\mathbf C$	68.027	\mathbf{O}
$\overline{4}$	Al-1 Warm Load Temperature	$\mathbf C$	68.027	\bf{o}
23	A1-2 Warm Load Temperature	$\mathbf C$	68.027	\mathbf{O}
8	Al-1 Antenna Drive Motor Current	mA	23.3	Ω
27	A1-2 Antenna Drive Motor Current	mA	23.3	\circ
$\mathbf{1}$	Signal Processing (+15VDC)	\vee	4.315	Ω
9	Antenna Drive (+15 VDC)	\vee	4.315	Ω
29	Signal Processing (-15 VDC)	$\mathbf v$	2.504	-22.562
28	Antenna Drive (-15 VDC)	\mathbf{v}	2.504	-22.562
34	Receiver Amplifiers (+8 VDC)	\mathbf{V}	2.5	Ω
12	Signal Processing (+5 VDC)	V	1.667	Ω
10	Antenna Drive (+5 VDC)	\vee	1.667	Ω
17	Receiver Mixer/IF (+10 VDC)	V	2.889	\mathbf{O}
16	Phase Lock Loop Ch 9/14 (+15 VDC)	V	4.315	Ω
33	Phase Lock Loop Ch 9/14 (-15VDC)	\mathbf{v}	2.504	-22.562
13	Ch 3 L.O. Voltage (50.3 GHz)	V	2.861	\mathbf{O}
30	Ch 4 L.O. Voltage (52.8 GHz)	V	2.861	\mathbf{O}
14	Ch 5 L.O. Voltage (53.596 GHz)	\mathbf{V}	2.861	\mathbf{o}
31	Ch 6 L.O. Voltage (54.4 GHz)	\mathbf{V}	2.861	\mathbf{O}
15	Ch 7 L.O. Voltage (54.94 GHz)	\mathbf{v}	2.861	\mathbf{O}
32	Ch 8 L.O. Voltage (55.5 GHz)	\mathbf{v}	2.861	\mathbf{O}
25	PLLO Primary Lock Detect (PLO #1)			\mathbf{o}
6	PLLO Redundant Lock Detect (PLO #2)			\mathbf{O}
18	Ch 15 L.O. Voltage (89.0 GHz)		4.315	$\mathbf 0$

 $mA =$ milliamps

APPENDIX B

NOAA-N' CPIDS : COEFFICIENTS OF AMSU-A1 S/N 105 AND AMSU-A2 S/N 107

In this appendix, the NOAA-N' AMSU-A CPIDS are presented in a similar fashion as those for NOAA-N AMSU-A given in the main text and Appendix A. Tables of coefficients and plots of results from analysis of the TV calibration data follow the same order as that of NOAA-N calibration steps as closely as possible. The NOAA-N' analog conversion coefficients are identical to those of NOAA-N (Tables A-6 and A-7).

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Table B-2. NOAA-N' nonlinearity parameters u in dimension of (m²-sr-cm⁻¹)/mW. **Table B-2. NOAA-N' nonlinearity parameters u in dimension of (m2-sr-cm ')/mW.**

AMSU-A1-1 S/N 105 channels: PLLO#1

AMSU-A1-1 S/N 105 channels: PLLO#2

Table B-3. NOAA-N' AMSU-A warm load corrections (K) at three instrument temperatures.

Instrument Temp.(C)	Ch.1	Ch.2	Instrument Temp.(C)	Ch.3	Ch.4	Ch.5	Ch.8
-7.51	0.092	0.084	-2.35	0.095	0.027	0.034	0.031
11.54	-0.053	-0.041	18.30	0.242	0.000	-0.014	0.014
30.28	-0.161	-0.123	38.26	-0.022	-0.090	-0.093	-0.095

AMSU-A2 S/N 107 channels: AMSU-A1-2 S/N 105 channels:

AMSU-A1-1 S/N 105 channels: PLLO#l:

Instrument Ch.6 Ch.7 Ch.9 Ch.10 Ch.11 Ch.12 Ch.13 Ch.14 Ch.15 Temp(C)					
-2.15		0.324 0.385 0.312 0.242 0.348 0.302 0.291 0.288 0.308			
17.70		$\begin{bmatrix} 0.311 & 0.380 & 0.321 & 0.228 & 0.326 & 0.315 & 0.319 & 0.329 & 0.304 \end{bmatrix}$			
38.48					$\vert 0.112 \vert 0.092 \vert 0.078 \vert -0.091 \vert 0.084 \vert 0.080 \vert 0.077 \vert 0.087 \vert 0.060$

AMSU-A1-1 S/N 105 channels: PLLO#2:

Table B-4. NOAA-N' AMSU-A1 S/N 105: Polynomial coefficients for converting PRT counts into temperatures.

Table B-5. NOAA-N' AMSU-A2 S/N 107: Polynomial coefficients for converting PRT counts into temperatures.

Figure B-l. NOAA-N' AMSU-A : Calibration accuracies at three instrument temperatures.

Figure B-2. NOAA-N' AMSU-A : Residuals after least-squares fit of the scene PRT radiance as a function of R_{SL} as discussed in the text.

Figure B-3. NOAA-N' AMSU-A: Simulated nonlinear corrections which are expected from on-orbit data.

NOAA-N': AMSU-A2 S/N 107 RF-Shelf Temperature (C): xx= -7.5, **=11.5, ++=30.3 AMSU-A1 S/N 105 RF-Shelf Temperature (C): xx= -2.0, **=17.9, ++=38.5

Figure B-4. NOAA-N' AMSU-A : Calibration results withe PLLO #2 in Channels 9-14, (a) calibration accuracy, (b) measured nonlinearity, and (c) simulated nonlinearity.

AMSU-A1 S/N 105, AMSU-A2 S/N 107 NEdT (K): $xx =$ Measured, $** =$ Specification

Figure B-5. NOAA-N' AMSU-A: Comparison of the measured NEAT values with specification.

Figure B-6. NOAA-N' AMSU-A: Measured NEAT values at three instrument temperatures.

Figure B-7. NOAA-N' AMSU-A: Radiometric counts versus scene PRT temperature at three instrument temperatures.

Figure B-8. NOAA-N' AMSU-A: Intercept counts at three instrument temperatures.

NOAA-N': AMSU-A2 S/N 107 RF-Shelf Temperature (C): xx= -7.5, **=11.5, ++=30.3 AMSU-A1 S/N 105 RF-Shelf Temperature (C): xx= -2.0, **=17.9, ++=38.5

Figure B-9. NOAA-N' AMSU-A: Channel gains at three instrument temperatures.

NOAA-N': AMSU-A2 S/N 107 RF-Shelf Temperature (C): xx= -7.5, **=11.5, ++=30.3 AMSU-A1 S/N 105 RF-Shelf Temperature (C): xx= -2.0, **=17.9, ++=38.5

Figure B-10. NOAA-N' AMSU-A: Calibration results with PLLO #2, including gains, radiometric counts, and intercept counts at three instrument temperatures.

APPENDIX C

NOAA LEVEL 1B **DATA AND SAMPLES OF AMSU-A ON-ORBIT DATA**

The NOAA Polar Orbiter Level IB data are raw data that have been quality controlled and assembled into discrete data sets, to which Earth location and calibration information are appended but not applied. For simplification of application, one can rewrite Equation (4) as,

$$
R_S = a_0 + a_1 C_S + a_2 C_S^2 \tag{C-1}
$$

where the calibration coefficients a_i (where $i = 0, 1,$ and 2) can be expressed in terms of R_w , G, $\overline{C_w}$ and $\overline{C_c}$. This can be accomplished by rewriting the right-hand side of Equation (4) in powers of C_s and equates the a_i 's to the coefficients of same powers of C_s . The results are,

$$
a_0 = R_w - \frac{\overline{C}_w}{G} + u \frac{\overline{C}_w \overline{C}_c}{G^2}
$$
 (C-2)

$$
a_1 = \frac{1}{G} - u \frac{\overline{C}_C + \overline{C}_W}{G^2}
$$
 (C-3)

and

$$
a_2 = u \frac{1}{G^2} \tag{C-4}
$$

These calibration coefficients will be calculated at each scan line for all channels and appended to the IB data. With these coefficients, one can simply apply Equation (C-l) to obtain the scene radiance R_s . Users, who prefer brightness temperature instead of radiance, can make the simple conversion,

$$
T_S = B^{-1}(R_S) \tag{C-5}
$$

where $B^{-1}(R_s)$ is the inverse of the Planck function for radiance R_s . The T_s is the converted brightness temperature.

The coefficients defined in Equations (C-2) to (C-4) are functions of instrument temperature. In general, they are not constant and should be re-calculated at each scan. To demonstrate the dynamic features of the calibration coefficients, the a_0 , a_1 , and a_2 values calculated from one orbit of the NOAA-16 AMSU-A IB data are shown in Figure C-l. These time series of calibration coefficients show the changes over the period. Therefore, they must be calculated at each scan line and their values obtained from pre-launch data [6-9] can not be employed to convert the radiometric counts into temperatures.

Figure C-2 shows the instrument (RF Shelf) temperature variation during the same orbital period. It shows that RF Shelf temperature changed only about 2 K during the orbital period, but it is large enough to induce signification variation in the a_0 , a_1 , and a_2 coefficients. The RF Mux temperature, which is used as a backup instrument temperature, is also shown in Figure C-2. Similarly, the effect ofthe instrument temperature on the cold space and blackbody radiometric counts is also significant. These are shown in Figures C-3 and C-4, where the space and blackbody radiometric counts versus the scan line number are displayed. The variations in the space and warm radiometric counts are mainly due to changes in the instrument temperatures. Analysis of these NOAA-16 AMSU-A onorbit data provides us with a better understanding of the AMSU-A instrument performance in space and helps process the pre-launch calibration data.

Figure C-1. NOAA-16 AMSU-A on-orbit data : Calibration coefficients, a_0 , a_1 , and a_2 values at Channel 6 over one orbit.

Figure C-2. NOAA-16 AMSU-A on-orbit data: RF-Shelf and RF Mux temperatures over one orbital period.

Figure C-3. NOAA-16 AMSU-A on-orbit data: Radiometric space counts over one orbital period.

Figure C-4. NOAA-16 AMSU-A on-orbit data: Radiometric blackbody counts over one orbital period.

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