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# **NOAA Technical Report NESS 81**

# Improved Algorithm for Calculation of UTM and Geodetic Coordinates

Washington, D.C. September 1980

U.S. DEPARTMENT OF COMMERCE National Oceanic and Atmospheric Administration National Environmental Satellite Service

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U.S. DEPARTMENT OF COMMERCE Philip M. Klutznick, Secretary SILVER SPRING CENTER

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#### IMPROVED ALGORITHM FOR CALCULATION OF UTM AND GEODETIC COORDINATES

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#### ABSTRACT

Expression of the equations for a UTM (Gauss-Krüger) projection in terms of Jacobian elliptic functions, rather than their series expansions, allows the projection to be used over wider zones than the standard 6 deg strips, and thus makes it applicable to satellite data from the NOAA A-G series. An efficient iterative solution method for either UTM or geodetic coordinates is developed using a complex-arithmetic version of Newton's method. The method can be used for longitudes up to 90 deg from the central meridian.

#### 1. Description of UTM projection

The UTM (Universal Transverse Mercator) projection is commonly used world-wide for the production of large-scale (i.e.  $\ge 1:1,000,000$ ) maps. The coordinate system is described in detail in many references (e.g. Richardus and Adler, 1972; Maling, 1973, U.S. Army, 1973). It is a Gauss-Krüger projection with the following standards and modifications:

- (1) The world is divided, with minor exceptions, into 60 zones with principal meridians every 6 deg of longitude. The meridian for zone 1 is 177 deg W.
- (2) Along the principal meridian the scale is set to 0.9996 instead of 1.0, in order to minimize the average error in the zone.
- (3) To avoid the use of negative numbers, the easting coordinate is incremented by  $5 \times 10^5$ . Similarly, in the southern hemisphere, the northing coordinate is incremented by  $10^7$ .

#### 2. UTM equations in closed form

Description of the UTM projection coordinates requires elliptic functions and integrals. Normally these have been expressed as series expansions, and the existent calculations methods were designed for hand computation with extensive use of tables. The equivalent computer programs (e.g. Quinones, 1970a, 1970b) typically have code which is cumbersome and whose accuracy decreases markedly outside the 6 deg longitude zone.

The more formal descriptions of the Gauss-Krüger equations which have appeared in the literature, have presented the results in terms of series expansions (e.g. Redfearn, 1948; Lee, 1962, 1976). Unfortunately these converge slowly and are often not sufficiently accurate. In this paper I adapt Lee's (1976) closed form expressions for these equations to a reasonably efficient computer solution which is accurate to machine precision. The method can be used for all longitudes up to and including 90 deg from the central meridian.

#### 2.1. Notation

a	semi-major axis of the ellipsoid
Ь	semi-minor axis of the ellipsoid
k	eccentricity of the ellipsoid $[k^2 = (a^2 - c^2)/a^2]$
φ	geodetic latitude
ψ	isometric latitude $[\psi = \arctan(\sin\phi) - k \operatorname{arctanh}(k \sin\phi)]$

 $\lambda$  longitude, from principal meridian of projection

х,у	northing and easting coordinates of Gauss-Krüger projection
u,v	northing and easting coordinates of intermediate projection
ζ	$\psi + i\lambda$
Ζ	x+iy
W	<i>u</i> + <i>iv</i>
m	parameter of elliptic functions $[m = k^2]$
sn,cn,dn	Jacobian elliptic functions

K,E	elliptic integrals of first and second kinds (complete when no argument is specified)
$\alpha, \beta, \gamma$	parameters used in calculation of complete elliptic integrals

#### 2.2. Gauss-Krüger equations

The general equation for a transverse mercator projection of a ellipsoid, expressed in terms of elliptic functions, is:

$$\zeta = \arctan(\operatorname{sn} w) - k \arctan(k \operatorname{sn} w) \tag{1}$$

where  $\operatorname{sn} w$  is the Jacobian elliptic function  $\operatorname{sn}(w|m)$  (see Abramowitz and Stegun, 1964, ch. 16; Byrd and Friedman, 1971). The Gauss-Krüger projection (one of a general class of transverse mercator projections) is defined in terms of w as:

$$\frac{z}{a} = (1-m) \int_0^w \mathrm{dn}^{-2} t \, dt$$

or

$$\frac{z}{a} = \mathcal{E}(w|m) - m \,\frac{\operatorname{sn} w \operatorname{cn} w}{\operatorname{dn} w} \tag{2}$$

where E(w|m) is an elliptic integral of the second kind.

#### 2.3. Some useful approximations and identities

From the wealth of information about elliptic functions and integrals, I have made use of the following:

#### 2.3.1. q-series approximations

Define the nome q as:

$$q = \exp\left[-\frac{\pi \mathbf{K}'}{\mathbf{K}}\right]$$

The quarter periods K, K' are complete elliptic integrals of the first kind with parameters m, 1-m respectively.

$$K = K(m) = \int_{0}^{\frac{\pi}{2}} (1 - m \sin^2 \theta)^{-\frac{1}{2}} d\theta$$
$$K' = K(1 - m)$$

From these, a rapidly converging expression for snw is:

$$\operatorname{sn} w = \frac{2\pi}{m^{\frac{1}{2}} K} \sum_{n=0}^{\infty} \frac{q^{n+\frac{1}{2}}}{1-q^{2n+1}} \sin\left[ (2n+1) \frac{\pi w}{2K} \right]$$

Similar approximations exist for cnw, dnw etc., as well as for an elliptic integral of the second kind:

$$\mathbf{E}(w|m) = \mathbf{Z}(w|m) + \frac{w\mathbf{E}}{\mathbf{K}}$$

E is the complete elliptic integral of the second kind, and Z(w|m) is Jacobi's zeta function:

$$E = E(m) = \int_{0}^{\frac{\pi}{2}} (1 - m \sin^2 \theta)^{\frac{1}{2}} d\theta$$
$$Z(w|m) = \frac{2\pi}{K} \sum_{n=1}^{\infty} \frac{q^n}{1 - q^{2n}} \sin \frac{n\pi w}{K}$$

#### 2.3.2. Gauss' arithmetic-geometric mean methods

Start with the triple  $(\alpha_0, \beta_0, \gamma_0)$ . Then

$$\alpha_n = \frac{\alpha_{n-1} + \beta_{n-1}}{2}$$
$$\beta_n = \sqrt{\alpha_{n-1}\beta_{n-1}}$$
$$\gamma_n = \alpha_n^2 - \beta_n^2$$

until  $\alpha_n = \beta_n$  to the desired accuracy.

To evaluate complete elliptic integrals, we start with (Bulirsch, 1965; Hart, et al., 1968):

$$\alpha_0 = 1 \quad \beta_0 = \sqrt{1 - m} \quad \gamma_0 = m$$

Then

$$K = K(m) = \frac{\pi}{2\alpha_n}$$
$$E = E(m) = K \left[ 1 - \sum_{j=0}^n 2^{j-1} \gamma_j \right]$$

The algorithm can be checked by Legendre's relation:

$$\mathbf{E}\mathbf{K}' + \mathbf{E}'\mathbf{K} - \mathbf{K}\mathbf{K}' = \frac{\pi}{2}$$

2.3.3. Identities

 $sn^2w + cn^2w = 1$  $m sn^2w + dn^2w = 1$ 

#### 2.3.4. Derivatives

 $\frac{d}{dw} \operatorname{sn} w = \operatorname{cn} w \operatorname{dn} w$  $\frac{d}{dw} \operatorname{cn} w = -\operatorname{sn} w \operatorname{dn} w$  $\frac{d}{dw} \operatorname{dn} w = -m \operatorname{sn} w \operatorname{cn} w$  $\frac{d}{dw} \operatorname{E}(w|m) = \operatorname{dn}^2 w$ 

#### 3. Solution of UTM equations

We define the *forward transform* as solving for (x,y) given  $(\phi,\lambda)$  and the *inverse transform* as the reverse process. The principal difficulty in the forward transform is solving equation (1) for w, given  $\zeta$ . Likewise, the principal difficulty in the inverse transform is solving equation (2) for w, given z. Once w is known, neither the solution of (1) for  $\zeta$  nor (2) for z presents any formidable difficulty. The q-series approximations in 2.3.1 typically converge to  $10^{-18}$  accuracy in 10 or 11 iterations.

The numerical computation for the elliptic integrals, as described in section 2.3.2, need be performed only once for any ellipsoid.

There are two reasonable choices for solving (1) or (2) when w is unknown.

#### 3.1. Infinite series method

The traditional method is to express the appropriate equation in its Taylor series expansion, and then to invert the power series (Van Orstrand, 1910). For example, Lee (1962) combines (1) and (2) to express  $z = f(\zeta)$  as:

$$\frac{z}{a} = \zeta - \frac{1}{3! (1-m)} \zeta^3 + \frac{(5-m)}{5! (1-m)^3} \zeta^5 - \frac{(61+26m+m^2)}{7! (1-m)^5} \zeta^7 + \dots$$
(3)

The difficulties with this approach are that:

- (1) No general equations for the coefficients in the series expansions of the Jacobian elliptic functions are known, so one must work them out from the Taylor series. The algebra can get difficult.
- (2) Similarly, the formulas for inversion of a series become more and more cumbersome as the order increases.
- (3) The series converges slowly; for example the 7th order term in (3) still has a magnitude exeeding 5 km at 45 deg N.
- (4) It is therefore difficult to use the series method where our accuracy specifications are arbitrary, or where we might wish to use the Gauss-Krüger projection over a zone wider than 6 deg (Snyder, 1979). Therefore the traditional methods of calculating UTM coordinates are inappropriate for many kinds of satellite data, including those from the NOAA A-G series.

#### 3.2. Iterative method

For these reasons I resort instead to an iterative scheme. This has the disadvantage of increasing the computation time, but gives us the capability to solve either the forward or inverse transform to the limit of our computing machine's precision. The method is accurate for all longitudes up to 90 deg away from the central meridian of the projection, although the projection itself is generally not suitable at more than 10 or 15 deg away. The choice of numerical methods is influenced by two factors:

- (1) The equations are well-behaved and have analytic derivatives. Furthermore, the identities between the Jacobian elliptic functions allow the first derivatives to be easily evaluated as a by-product of the evaluations of the functions.
- (2) For all cases we can select a very good initial guess by letting k = 0 and solving the equation for a spherical rather than ellipsoidal earth.

Under these circumstances Newton's method will give us very rapid convergence.

#### 3.2.1. The forward transform

Rewrite (1) as:

 $f(w) = \operatorname{arctanh}(\operatorname{sn} w) - k \operatorname{arctanh}(k \operatorname{sn} w) - \zeta = 0$ 

Then, using sections 2.3.3 and 2.3.4:

$$f'(w) = \frac{1-m}{\operatorname{cn} w \operatorname{dn} w}$$

By Newton's method (see Lanczos, 1957):

$$w_{n+1} = w_n - \frac{f(w_n)}{f'(w_n)}$$

The starting guess  $w_0$  is found by letting k = 0. Then m = 0 and  $\operatorname{sn} w = \operatorname{sin} w$ , and:

 $w_0 = \arcsin(\tanh \zeta)$ 

#### 3.2.2. The inverse transform

Rewrite (2) as:

$$f(w) = \mathbf{E}(w|m) - m \, \frac{\operatorname{sn} w \operatorname{cn} w}{\operatorname{dn} w} - \frac{z}{a} = 0$$

Then

$$f'(w) = \mathrm{dn}^2 w - m \left[ \mathrm{cn}^2 w - (1-m) \, \frac{\mathrm{sn}^2 w}{\mathrm{dn}^2 w} \right]$$

When k = 0, solution of (2) leads to:

$$\frac{z}{a} = \frac{w}{2} + \frac{\sin(2w)}{4}$$

To a first order approximation sin(2w) = 2w, so our initial guess  $w_0$  is:

$$w_0 = \frac{z}{a}$$

#### 4. Subprograms

There are only two subprograms with which the user need interface. They are written in the C programming language (Kernighan and Ritchie, 1978) but would be easily translated into other languages, as the C code is relatively algorithmic. The routines are utmf and utminv, to calculate the forward and inverse transforms respectively. They are listed in the Appendix, along with the necessary supporting routines.

Arguments for utmf are:

utmf (lc, zone, lat, lon, north, east) int lc, \*zone; double lat, lon, \*north, \*east;

The input variables are:

lc ellipsoid code (see Table 1)

zone UTM reference zone (if zone = 0, the routine will calculate the appropriate zone) lat, lon geodetic coordinates in radians (positive in northern, eastern hemispheres)

Output variables are:

north northing

east easting

For utminv, the arguments are in a different order:

utminv (hflag, lc, zone, north, east, lat, lon) int hflag, lc, zone; double north, east, \*lat, \*lon; In this case zone must be non-zero, as it is part of the UTM coordinate specification; north and east are input variables and lat and lon are output variables. The variable hflag, when negative, specifies that the northing coordinate has been incremented by  $10^7$ .

code	ellipsoid	major axis (m)	minor axis (m)	eccentricity
0	sphere of Int'l volume	6370949	6370949	.000000
1	International	6378388	6356912	.081992
2	Clarke 1866	6378206	6356584	.082272
3	Clarke 1880	6378249	6356515	.082483
4	Everest	6377276	6356075	.081473
5	Bessel	6377397	6356079	.081697
6	Modified Everest	6377304	6356103	.081473
7	Australian National	6378160	6356775	.081820
8	Airy	6377563	6356257	.081673
9	Modified Airy	6377342	6356036	.081673
10	Walbeck	6376896	6355835	.081207
11	Southeast Asia	6378155	6356773	.081813
12	Krasovskiy	6378245	6356863	.081813
13	GTDS Model	6378140	6356755	.081820

# Table 1 Ellipsoid Options for UTM Routines

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#### APPENDIX

Here are listed the source codes for the routines in the UTM programs. The codes are in alphabetical order, but the structure of the routines is as follows:

- 1. The major subprograms utmf: calculates UTM coordinates from geodetic utminv: calculates geodetic coordinates from UTM
- Miscellaneous routines related to the UTM projection utmorig: longitude of principal meridian utmzone: zone associated with a given longitude
- 3. Miscellaneous routines related to earth coordinates
  - geodet: geodetic latitude ( $\phi$ ) from isometric ( $\psi$ )
  - isomet: isometric latitude ( $\psi$ ) from geodetic ( $\phi$ )
  - sphcon: axes (a, c) and eccentricity (k) for a given ellipsoid
- 4. Some transcendental and complex functions
  - complex: elementary operations (place ahead of main)
    - atanh: inverse of hyperbolic tangent (real)
    - casin: complex arc sine
    - catan: complex arc tangent
    - catanh: complex arc tanh
    - cpowr: complex number to real power
    - csin: complex sine
    - ctanh: complex hyperbolic tangent
- 5. Elliptic integrals and elliptic functions
  - e comp: complete elliptic integral of 1st or 2nd kind (K or E)
  - e2ci: complex elliptic integral of 2nd kind (with real parameter)
  - isn: complex Jacobian elliptic function sn
  - agm1: Gauss' arithmetic-geometric mean from initial triple
  - jzeta: Jacobi's zeta function
  - jseries: general routine for q-series approximations
  - giper: quarter periods K, K' for Jacobian elliptic functions and other q-series approximations
  - Gauss-Krüger and transverse Mercator coordinates
    - gkfor: Gauss-Krüger coordinates z from isometric coordinates  $\zeta$
    - gkinv:  $\zeta$  from z
    - gkf: Gauss-Krüger coordinates z from intermediate coordinates w
    - gk: Gauss-Krüger coordinates from w, snw
    - tmw: intermediate coordinates w from isometric coordinates  $\zeta$
    - tmwi: intermediate coordinates w from Gauss-Krüger coordinates z
    - tmerc: isometric coordinates  $\zeta$  from snw
- 7. Miscellaneous routines

6

abend: terminates pro	ogram with	error	message
-----------------------	------------	-------	---------

- cnewton: complex version of Newton's method
- getpi: returns double precision value of  $\pi$
- macheps: returns machine epsilon
- zerobr: finds a zero of an arbitrary real function (from Brent, 1973)

All of the routines are coded in the C programming language (Kernighan and Ritchie, 1978). This language is close enough to Algol to make translation into other languages relatively simple. The following conventions and notations may be somewhat peculiar:

- 1. C makes explicit the distinction between the value of a variable and its address. Only a variable whose *address* is passed to a subprogram may be changed by the subprogram. These have an asterisk (\*), meaning "pointer to", prepended to the variable name in the subprogram. Variables passed by value may be changed in a subprogram, but these changes will not be passed back to the calling program. The prefix & to a variable in a calling program means "address of".
- 2. The notation i op= j; is the same as

i = i op j;

- 3. The notation
  - i++;
  - applies to integers only and means

i = i + 1;

The ++ after the variable indicates that the incrementation takes place after the use of the variable in an operation.

#### 1. abend

```
#define BELL 07
 abend (msg, dump)
 char *msg;
 int dump;
 /*
 * function: abend
   purpose : terminate a program if an irrecoverable
          error occurs, and print explanatory message
          if (dump = = 1) cause a core dump also
   usage : abend ("message to be printed", dump);
 */
      putchar (BELL);
      printf ("\n\tABEND: %s\n", msg);
      if (dump = = 1) {
           printf ("\t(core dumped)\n");
           fclose (stdout):
           abort (); /* dump core for debugging */
     }
     else
           exit ();
2. agm1
/*
     returns arithmetic-geometric mean of a, b
     also calculates sum 2^{(k-1)} c sub k
*/
double
          agm1 (a, b, c, c2sum)
double
          a, b, c, *c2sum;
     double
                fabs(), macheps(), sqrt();
                sum, h, tk, eps;
     double
     char
                str[80];
     if (a^*b < = 0.) {
          sprintf (str, "agm1: a %e b %e", a, b);
          abend (str, 0);
     sum = 5.e - 1 * c;
    tk = 1;
```

eps = macheps();while (fabs  $(a-b) > a^*eps$ ) { h = a;a = (a + b) / 2;b = sqrt (h \* b);c = (a-b) \* (a+b);sum + = tk \* c;tk \* = 2;\*c2sum = sum; return (a); } 3. atanh 11 inverse hyperbolic tangent \*/ double atanh (x) double X: double log(), hold; char str[80]; /\* check range \*/ hold = x \* x;if (hold < 0. || hold > = 1.) { sprintf (str, "atanh: argument = %e", x); abend (str, 1); return  $(5.e-1 * \log ((1. + x) / (1. - x)));$ 4. casin /\* complex arcsin \*/ complex casin (z) complex z; extern int errno:

complex result;

double alpha, beta, xp, xm, y2, l, r; double sqrt(), asin(), log(); char str[80];  $xp = z.re^*z.re + 2.*z.re + 1.;$  $xm = z.re^*z.re - 2.*z.re + 1.;$  $y_2 = z.im * z.im;$ l = sqrt (xp + y2);if (errno = = EDOM)abend ("casin: bad sqrt for I", 1); r = sqrt (xm + y2);if (errno = EDOM)abend ("casin: bad sqrt for r", 1); alpha = 5.e - 1 \* (1 + r);beta = 5.e - 1 \* (1 - r);

/\* correct for slight rounding error \*/
if (fabs(beta) > 1. && fabs(beta) < 1.01)
beta = l > r ? 1 : -1;

result.re = asin (beta); if (errno == EDOM) { sprintf (str, "casin: bad asin() arg %.18e", beta); abend (str, 0);

return (result);

#### 5. catan

#### /\*

```
complex arctan
*/
complex catan (z)
complex z;
{
complex result;
```

doublex, y, xsqysq, twoy;doublelog(), atan2();

```
extern int errno;
```

```
x = z.re;
y = z.im;
if (x == 0. && y == 1.)
abend ("catan: z^2 = -1", 0);
```

```
return (result);
```

#### 6. catanh

```
/*
    complex arctanh
    uses relation arctanh z = -i arctan iz
*/
complex catanh (z)
complex z;
{
    complex i, hold;
    complex cmult(), catan();
    i.re = 0;
    i.im = 1;
    hold = cmult (i, z);
    i.im = -1;
    return (cmult (i, catan (hold)));
}
```

#### 7. cnewton

/\*

\*/

```
complex version of Newton's method
```

solves for complex zero of f(z) given initial guess z

```
the function f (z, ofzp) returns f(z)
and 1/f'(z) in ofzp
```

```
complex fz, ofzp, h, zold, cr, cmult(), csub();
double macheps(), cabs();
double eps, me;
int maxit;
char str[80];
```

```
me = 2. * macheps();
maxit = 50;
zold = z;
fz = (*f) (z, &ofzp);
```

while (maxit - -) {

/\* check if z is a root \*/
if (cabs (fz) <= tol) return (z);</pre>

```
/* calculate new approximation */
h = cmult (fz, ofzp);
z = csub (z, h);
```

```
/* see how close new approx. is */
eps = tol + me * cabs(z);
cr = csub (z, zold);
if (5.e-1 * cabs(cr) < eps) return (z);</pre>
```

```
zold = z;
fz = (*f) (z, &ofzp);
```

/\* if we get to here, we didn't converge \*/

```
sprintf (str,
"cnewton, no convergence: eps %e .5 cr %e |f(z)| %e",
eps, 5.e-1*cabs(cr), cabs(fz));
abend (str, 0);
```

#### 8. complex

}

```
typedef struct {
              re, im;
      double
 }
complex;
complex cadd (z1, z2)
complex z1, z2;
     complex result;
     result.re = z1.re + z2.re;
     result.im = z1.im + z2.im;
     return (result);
complex csub(z1, z2)
complex z1, z2;
     complex result;
     result.re = z1.re - z2.re;
     result.im = z1.im - z2.im;
     return (result);
}
complex cmult (z1, z2)
complex z1, z2;
{
     complex result;
     result.re = z1.re * z2.re - z1.im * z2.im;
  • result.im = z1.re * z2.im + z1.im * z2.re;
     return (result);
)
complex cdiv (z1, z2)
complex z1, z2;
     double denom;
     complex result;
     denom = z2.re * z2.re + z2.im * z2.im;
     result.re = (z1.re * z2.re + z1.im * z2.im) / denom;
     result.im = (z_2.re * z_1.im - z_1.re * z_2.im) / denom;
    return (result);
double carg (z)
complex z;
ł
    double atan2();
    return (atan2 (z.im, z.re));
}
complex conjg (z)
complex z;
{
    complex result;
```

```
result.re = z.re;
result.im = -z.im;
```

```
return (result);
```

```
9. cpowr
```

```
/*
    complex number to a real power
*/
complex cpowr (z, p)
complex z;
double p;
{
    extern int errno;
    complex result;
    double pow(), carg(), cabs(), r, nt, cos(), sin(), rn;
    r = cabs (z);
```

```
result.re = rn * cos (nt);
result.im = rn * sin (nt);
return (result);
```

```
10. csin
```

}

```
/*
    complex sin
*/
complex csin (z)
complex z;
{
    complex result;
    double x, y, sin(), cos(), sinh(), cosh();
    x = z.re;
```

y = z.im; result.re = sin(x) \* cosh(y); result.im = cos(x) \* sinh(y);

return (result);

#### 11. ctanh

}

```
/*

complex tanh

*/

complex ctanh (z)

complex z;

{

complex result;

double sin(), cos(), sinh(), cosh();

double denom, twox, twoy;
```

```
twox = 2. * z.re;
```

```
twoy = 2. * z.im;
denom = cosh (twox) + cos (twoy);
result.re = sinh (twox) / denom;
result.im = sin (twoy) / denom;
```

return (result);

#### 12. e2ci

}

```
/*
     complex elliptic integral of 2nd kind
*/
complex e2ci (m, w)
double
          m;
complex w;
     complex jzeta(), cadd(), z;
                k, kp, ek, e_comp(), q;
     double
     static double
                     mlast, e;
     giper (m, &k, &kp, &g);
     z = jzeta (k, q, w);
     if (m != mlast || m == 0.) {
           e = e \operatorname{comp}(m, 2);
           mlast = m;
     ek = e / k;
     w.re * = ek;
     w.im * = ek;
     return (cadd (z, w));
13. e_comp
1*
     complete elliptic integral
     parameter m
     kind 1 or 2
*/
double
           e_comp (m, kind)
double
           m;
int
           kind;
     /* uses a.g.m. method*/
     double sqrt(), agm1();
     double
               b, c, pi;
     static double
                     mlast, a, sum;
     getpi (&pi);
     if (kind != 1 \&\& kind != 2)
           abend ("e comp: kind must be 1 or 2", 0);
     if (m = = 1.) {
           if (kind = = 1)
                abend ("e_comp: 1st kind, m = 1", 0);
           return (1.);
```

```
if (m = = 0.) return (pi/2);
     if (m != mlast) {
          mlast = m;
          c = sqrt(m);
          a = 1;
          b = sqrt ((1.+c) * (1.-c));
          c = m;
          a = agm1 (a, b, c, \&sum);
     return (kind = = 1 ? pi/(2.*a): (pi/(2.*a)) * (1.-sum));
14. geodet
/*
     finds latitude as function of isometric latitude
     and eccentricity
*/
double
           psi_trn, k_tran;
double
          geodet (psi, k)
double
          psi, k;
                zerobr(), isomd(), tanh(), asin();
     double
     double
                phi, pi2, tol;
     char str[80];
     extern double psi trn, k tran;
     extern int
                      errno;
     /* first check the special case k = 0 */
     if (k = = 0) {
           phi = asin (tanh (psi));
           if (errno = EDOM) {
                sprintf (str, "geodet: asin arg > 1, psi %f", psi);
                abend (str, 1);
           return (phi);
     psi trn = psi;
     k tran = k;
     getpi (&pi2);
     pi2 / = 2;
     tol = 0;
     if (psi = = 0.) return (0.);
     else if (psi > 0.)
           phi = zerobr (0., pi2, tol, isomd);
     else
           phi = zerobr (-pi2, 0., tol, isomd);
     return (phi);
           isomd (phi)
double
double
           phi;
```

```
extern double psi trn, k tran;
```

double isomet();

return (psi\_trn - isomet (phi, k\_tran));

#### 15. getpi

}

/\* sets double precision value of pi

```
Usage: getpi (&pi)
```

\*x:

```
*/
getpi(x)
```

double

```
{
```

```
static int first;
static double holdpi;
double atan();
```

```
if (first != 1) {
holdpi = 4.0 * atan(1.0);
first = 1;
```

\*x = holdpi;

```
16. gk
```

}

#### /\*

}

```
calculates G-K coordinates from w, sn w
*/
complex gk (m, w, snw)
double m;
complex w, snw;
```

complex cdw, cadd(), cmult(), csub(), e2ci(), cpowr(), t1, t2, one, num, denom, ms;

```
t1 = e2ci (m, w);
```

```
one.re = 1;
one.im = 0;
num = cmult (csub (one, snw), cadd (one, snw));
ms.re = m * snw.re;
ms.im = m * snw.im;
denom = cmult (csub (one, ms), cadd (one, ms));
cdw = cpowr (cdiv (num, denom), 5.e-1);
```

```
t2 = cmult (snw, cdw);
t2.re *= m;
t2.im *= m;
```

return (csub (t1, t2));

#### } 17. gkf

```
/*
    calculates G-K coordinates from intermediate coordinates
*/
complex gkf (a, m, w)
double a, m;
complex w;
{
    complex gk(), jsn(), snw, z;
}
```

```
z = gk (m, w, snw);
       z.re * = a;
       z.im * = a;
       return (z);
 }
 18. gkfor
 /*
       finds Gauss-Krüger coordinates given psi, lon
 */
 gkfor (k, a, psi, lon, x, y)
 double k, a, psi, lon, *x, *y;
 /*
      k
            eccentricity
      a semi-major axis
      psi isometric latitude
      lon longitude (from principal meridian)
      x, y northing, easting
 */
      complex zeta, z, w, tmw(), gkf();
      zeta.re = psi;
      zeta.im = lon;
      /* find w */
      w = tmw (k, zeta);
      /* now calculate z */
      z = gkf(a, k^*k, w);
      *x = z.re;
      *y = z.im;
19. gkinv
/*
     finds psi, lon given Gauss-Krüger coordinates
*/
gkinv (k, a, x, y, psi, lon)
double k, a, x, y, *psi, *lon;
/*
     see gkfor() for explanation of variables
*/
{
     complex zeta, z, w, snw;
     complex jsn(), tmwi(), tmerc();
     z.re = x / a;
     z.im = y / a;
     /* find w */
     w = tmwi (k, z);
     /* now calculate zeta */
    snw = jsn (k^*k, w);
    zeta = tmerc (k, snw);
```

```
*psi = zeta.re:
```

snw = jsn (m, w);

```
*lon = zeta.im;
```

#### 20. isomet

finds isometric latitude as a function of latitude and eccentricity

#### \*/

/\*

```
double
           isomet (lat, eccen)
double
           lat, eccen;
     double
                 x, atanh(), sin(), pi2, fabs();
     char
                 str[80];
```

```
/* check range of arguments */
if (eccen < 0. \parallel eccen > = 1.) {
      sprintf (str, "isomet: eccentricity = %e",
           eccen):
     abend (str, 1);
}
```

```
getpi (&pi2);
pi2 / = 2;
if (fabs (lat) > pi2) {
      sprintf (str, "isomet: latitude %.16f", lat);
      abend (str, 1);
}
```

x = sin (lat);return (atanh (x) - eccen \* atanh (eccen \* x));

#### 21. jseries

#### /\*

}

expands an arbitrary complex series (of the type found in Jacobian elliptic functions) until convergence \*/ complex jseries (start, q, v, qf, vf) int start; double q, (\*qf)(); complex v, (\*vf)(); /\* start = 0 or 1= real parameter for real qf()q v = complex parameter for complex vf() \*/ { int n; double h, aold, anew; complex z, x, cadd();n = start;anew = z.re = z.im = 0;aold = 999;while (aold != anew) { h = (\*qf) (n, q);x = (\*vf) (n, v);n++; x.re \* = h;x.im \* = h;

```
z = cadd (z, x);
            anew = z.re * z.re + z.im * z.im;
      return (z);
 22. jsn
 /*
      complex Jacobian elliptic function sn u
 */
 complex jsn (m, u)
 double
           m;
 complex
           u;
      double
                 pi, k, kp, q, fact, qsnf(), sqrt();
      complex
                v, sn;
      complex vsnf(), jseries(), csin(), ctanh();
      if (m = = 0.)
           return (csin (u));
      if (m = 1.0)
           return (ctanh (u));
      qjper (m, &k, &kp, &q);
      getpi (&pi);
      fact = pi / (2.*k);
      v.re = fact * u.re;
      v.im = fact * u.im;
      sn = jseries (0, q, v, qsnf, vsnf);
      fact = 2. * pi / (sqrt(m) * k);
      sn.re *= fact;
     sn.im *= fact;
     return (sn);
/*
     coefficients for q-series expansion of sn
*/
double
           qsnf (n, q)
int n;
double
           q;
     double
                np, dp, pow();
     np = n;
     np += 5.e-1;
     dp = 2^*n + 1;
     return (pow(q,np) / (1. - pow(q,dp)));
}
/*
     argument for q-series expansion of sn
*/
complex vsnf (n, v)
int n;
complex v;
{
     complex
               csin(), arg;
     double
               d;
```

aold = anew;

```
d = 2 * n + 1;
arg.re = v.re * d;
arg.im = v.im * d;
```

return (csin (arg));

#### 23. jzeta

/\* Jacobi's zeta function \*/ complex jzeta (k, q, w) double k, q; complex w; double pi, exp(), qfzeta();

complex v, h, vfzeta(), jseries();

if (k = = 0.)h.re = h.im = 0;

#### else {

```
getpi (&pi);
v.re = w.re * pi /k;
v.im = w.im * pi / k;
h = jseries (1, q, v, qfzeta, vfzeta);
h.re *= 2 * pi / k;
h.im *= 2 * pi / k;
```

```
}
```

return (h);

## }

/\*

coefficients for sin series for jzeta() \*/ double qfzeta (n, q) int n; double q; { double qp, pow();

qp = pow (q, (double) n);return  $(qp / (1. - qp^*qp));$ 

#### /\*

}

argument for sin-series for jzeta() \*/ complex vfzeta (n, v) int n; complex v; { complex csin(); v.re \* = n;

```
v.im *= n;
```

```
return (csin (v));
}
```

```
24. macheps
```

/\*

\*/ {

}

/\*

```
double macheps()
      returns machine epsilon
      double s, one;
      static double x;
      static int first;
     if (first != 1) {
           first = 1;
           one = 1;
          x = 1;
          s = 2;
          while (s > one) {
               x *= 5.e-1;
               s = one + x;
          }
     }
     return (x);
25. gjper
```

calculates quarter periods K & K' and nome q for Jacobian elliptic functions \*/ qjper (m, k, kp, q) double m, \*k, \*kp, \*q;

static double mlast, klast, kplast, qlast;

sqrt(), e\_comp (), exp(); double double pi, ms;

getpi (&pi);

```
if (m = = 0.) {
     *k = pi / 2.;
     *q = 0;
     *kp = 1.e23;
}
```

## else {

}

```
if (mlast == m) {
     *k = klast;
     *kp = kplast;
     *q = qlast;
}
```

else { mlast = m; $k = klast = e_{comp} (m, 1);$ ms = sqrt(m); $kp = kplast = e_{comp} ((1-ms) * (1.+ms), 1);$ \*q = qlast = exp (-pi \* kplast / klast);}

} 26. sphcon See Table 1. /\* information about specific ellipsoids \*/ sphcon (1, major, minor, eccen) int l; double \*major, \*minor, \*eccen; /\* code for ellipsoid choices (1 -- input) 0 Sphere of equal volume International 1 2 Clarke 1866 Clarke 1880 3 Everest 4 5 Bessel Modified Everest 6 Australian National 7 8 Airy 9 Modified Airy 10 Walbeck Southeast Asia 11 12 Krasovskiy GTDS Model 13 (output) major semi-major axis minor semi-minor axis eccen[defined as sqrt  $((a^*a - b^*b) / (a^*a))]$ \*/ char str[80]; double eccensq, sqrt(); in some of the values that follow, the minor axis is computed from the flattening \*/ switch (1) { case 0: \*major = \*minor = 6.370949; break; case 1: \*major = 6.378388; \*minor = 6.3569119461; break; case 2: \*major = 6.3782064; \*minor = 6.3565838;break; case 3: \*major = 6.378249145; \*minor = \*major \* (1. - 1./2.93465e2);break;

case 4:

\*major = 6.3772763452; \*minor = 6.3560754133; break; case 5: \*major = 6.377397155; \*minor = 6.3560789628; break: case 6: \*major = 6.377304063; \*minor = \*major \* (1. - 1./3.008017e2);break; case 7: \*major = 6.37816; \*minor = \*major \* (1. - 1./2.9825e2);break: case 8: \*major = 6.377563396; \*minor = \*major \* (1. - 1./2.99324965e2);break; case 9: \*major = 6.37734189; \*minor = \*major \* (1. - 1./2.99324959e2);break; case 10: \*major = 6.376896; \*minor = 6.3558348467; break; case 11: \*major = 6.378155; \*minor = 6.3567733205; break; case 12: \*major = 6.378245; \*minor = 6.3568630188; break; case 13: \*major = 6.37814; \*minor = 6.356754786253143; break; default: sprintf (str, "sphcon: ellipsoid code %d unimplemented", l); abend (str, 0); eccensq = (\*major - \*minor) \* (\*major + \*minor) / (\*major \* \*major); if (eccensq < 0)sprintf (str, "sphcon - e sq %f, maj %f, min %f", eccensq, \*major, \*minor); abend (str, 1); \*major \*= 1.e6; \*minor \*= 1.e6;\*eccen = sqrt (eccensq);

27. tmerc

/\* returns transverse mercator coordinates as function of sn w \*/ complex tmerc (k, snw) double k: complex snw; {

complex ksnw, catanh(), csub(), t1, t2;

ksnw.re = k \*snw.re; ksnw.im = k \* snw.im;

t1 = catanh (snw);t2 = catanh (ksnw);t2.re \* = k;t2.im \* = k;

return (csub (t1, t2));

#### 28. tmw

/\*

}

iteratively finds w as function of k, zeta \*/ complex z\_tran; double k tran;

```
complex
         tmw (k, zeta)
double
         k;
complex zeta;
```

extern complex z tran; extern double k\_tran;

complex cnewton(), tmfd(), w, casin(), ctanh(); double tol, pi2; double sqrt(), macheps(); double fabs():

tol = sqrt (macheps());getpi (&pi2); pi2 / = 2;

/\* initial guess \*/ w = casin (ctanh (zeta));

/\* check for quadrant compatibility \*/ if (fabs (zeta.im) > pi2) abend ("tmw: long. diff > pi/2", 0);

 $k_{tran} = k;$  $z_{tran} = zeta;$ 

w = cnewton (w, tmfd, tol);

```
return (w);
```

```
}
/*
```

returns difference between transverse mercator intermediate coordinates and zeta

returns (in ofpw) 1/f' \*/ complex tmfd (w, ofpw) complex w, \*ofpw; extern complex z\_tran; extern double k\_tran; complex snw, fw, cnw, dnw, ms, num, denom, one: complex jsn(), cdiv(), cmult(), cadd(), cpowr(), tmerc(); double m; one.re = 1;one.im = 0; $m = k_{tran} * k_{tran};$ snw = jsn (m, w); $fw = csub (tmerc (k_tran, snw), z_tran);$ cnw = cpowr (cmult (csub (one, snw), cadd (one, snw)), 5.e-1); ms = snw;ms.re \*= m; ms.im \*= m;dnw = cpowr (cmult (csub (one, ms), cadd (one, ms)), 5.e-1); num = cmult (cnw, dnw);denom.re = 1. - m;denom.im = 0;\*ofpw = cdiv (num, denom); return (fw); 29. tmwi iteratively finds w as function of k, z complex z\_tran; double k\_tran;

complex tmwi (k, z) double k; complex z;

}

/\*

\*/

extern complex z\_tran; extern double k tran;

complex cnewton(), casin(), tmid(); complex W; double pi2, tol; double sqrt(), macheps();

tol = sqrt (macheps());getpi (&pi2); pi2 / = 2;

 $k_{tran} = k;$  $z_{tran} = z;$ 

/\* initial guess \*/

\*/ {

```
w = z;
```

w = cnewton (w, tmid, tol);

return (w);

/\*

\*/

retturns difference between transverse mercator intermediate coordinates and z

returns (in ofpw) reciprocal of first derivative complex tmid (w, ofpw)

complex w, \*ofpw;

extern complex z\_tran; extern double k\_tran;

complex snw, cn2w, dn2w, hold, ms, fw, one; complex jsn(), csub(), gk(), cmult(); double m, mp;

 $m = k_tran * k_tran;$ mp = (1.+k tran) \* (1.-k tran);snw = jsn (m, w);

fw = csub (gk (m, w, snw), z tran);

/\* calculate derivative \*/ snw = cmult (snw, snw);one.re = 1;one.im = 0;cn2w = csub (one, snw);ms = snw;ms.re \* = m;ms.im \* = m;dn2w = csub (one, ms);hold = cdiv (snw, dn2w);hold.re \* = mp;hold.im \*= mp; hold = csub (cn2w, hold);hold.re \*= m; hold.im \*= m; hold = csub (dn2w, hold); \*ofpw = cdiv (one, hold);

return (fw);

#### 30. utmf

/\*

computes the UTM coordinates (major zone) for a point whose lat-long coordinates are known \*/

#define N FALSE 1.e7 #define E FALSE 5.e5 #define SCALE F 9.996e - 1

utmf (1, lat, lon, zone, north, east) int 1, \*zone; lat, lon, \*north, \*east; double /\*

#### variables:

(input) ellipsoid code [see sphcon() for tabulation] 1 zone (if != 0) major zone lat latitude in radians (+ in north hem) lon longitude in radians (+ in east hem)

(output) zone major zone (if = = 0 on input) northnorthing east easting

double /\* eccentricity of ellipsoid \*/ eccen. /\* semi-major axis (m) major, \*/ /\* longitude origin of zone \*/ orig, \*/ minor, /\* minor axis, not used isomet(), utmorig();

int utmzone(), abs();

sphcon (l, &major, &minor, &eccen); if (\*zone = = 0) \*zone = utmzone (lon);orig = utmorig (\*zone);

gkfor (eccen, major, isomet (lat, eccen), lon-orig, north, east);

\*north \*= SCALE F; \*east \*= SCALE F;

if (lat < 0.) \*north += N FALSE; \*east + = E FALSE;

#### 31. utminv

computes geodetic coordinates for a point whose UTM coordinates are known

#### \*/

#define N\_FALSE 1.e7 #define E FALSE 5.e5 #define SCALE F 9.996e - 1

utminv (hflag, l, zone, north, east, lat, lon) int hflag, l, zone; double north, east, \*lat, \*lon; /\*

#### variables:

(input) hflag (if negative) in so. hem., add N FALSE to northing 1 ellipsoid code [see sphcon() for tabulation] zone (if != 0) major zone lat latitude in radians (+ in north hem) lon longitude in radians (+ in east hem)

(output) zone major zone (if = = 0 on input) northnorthing east easting

```
double
          eccen,
                           /* eccentricity of ellipsoid */
                      /* semi-major axis (m)
     major,
                                                   */
     orig,
                /* longitude origin of zone */
     minor,
                      /* minor axis, not used
                                                   */
                /* isometric latitude */
     psi,
                /* longitude difference
     dlon,
                                             */
     geodet(),
     utmorig();
```

int utmzone(), abs();

gkinv (eccen, major, north, east, &psi, &dlon);

\*lon = dlon + orig; \*lat = geodet (psi, eccen);

#### 32. utmorig

/\*

}

returns origin (in radians) of designated UTM zone \*/

```
double utmorig (zone)
int zone;
{
```

double d, pi;

getpi (&pi); d = zone \* 6 - 183; return (d \* pi / 1.8e2);

#### 33. utmzone

/\*

returned value is UTM zone corresponding to lon (radians)

#### \*/

int utmzone (lon) double lon; {

> int geod[3], z, i; double pi, fabs();

getpi (&pi); /\* check range \*/ if (fabs (lon) > pi) abend ("utmzone -- abs (lon) > pi", 0); rdms (lon, geod); z = geod[0] / 6;

/\* set i = 0 if on zone boundary \*/
i = geod[0]%6 + geod[1] + geod[2] == 0 ? 0:1;

if (lon < 0) z += 31 - i; else z += 30 + i; return (z);

#### 34. zerobr

```
/*
```

\*/

}

finds zero of a function by Brent's algorithm (from Brent, 1973)

double zerobr (a, b, t, f)double a, b, t, (\*f)();

> double c, d, e, fa, fb, fc, tol, m, p, q, r, s, eps; double macheps(); double fabs(), log(); int maxfun; char str[80];

eps = 2 \* macheps();

fb = fabs (b) >= fabs (a) ? fabs (b) : fabs (a); fo = fabs (b) >= fabs (a) ? fabs (b) : fabs (a); fo = 5.e-1 \* t + 2 \* eps \* fb; s = log (fabs (b - a) / tol) / log (2.);maxfun = s \* s + 1;

fa = (\*f)(a); fc = fb = (\*f)(b);if (fb == 0) return (b); if (fa == 0) return (a); if (fb\*fa > 0) abend ("zerobr(): root not spanned", 1);

while (maxfun - -) {

if  $((fb > 0 \&\& fc > 0) \parallel (fb <= 0 \&\& fc <= 0))$  { c = a; fc = fa; d = e = b - a;} if (fabs(fc) < fabs(fb)) { a = b; b = c; c = a; fa = fb; fb = fc; fc = fa;} tol = eps \* fabs(b) + t; m = (c - b) / 2;

if  $(fabs(m) < tol \parallel fb == 0)$  return (b);

```
/* see if bisection is forced */
      if (fabs(e) < tol \parallel fabs(fa) <= fabs(fb)) d = e = m;
      else {
           s = fb/fa;
           if (a = = c) \{ /* \text{ linear interpolation } */
               p = 2 * m * s;
                q = 1 - s;
           }
           else { /* inverse quadratic interpolation */
               q = fa/fc;
                r = fb/fc;
                 p = s * (2 * m * q * (q-r) - (b-a) * (r-1));

q -= 1;

q *= (r-1) * (s-1);
           }
            \begin{array}{ll} \text{if } (p>0) \ q=-q;\\ \text{else} \qquad p=-p; \end{array} 
         s = e;
           e = d;
           if (2^*p < 3^*m^*q - fabs(tol^*q) \&\& p < fabs(s^*q/2))
           d = p/q;
           else d = e = m;
      }
     a = b;
     fa = fb;
     .
     fb = (*f)(b);
abend ("zerobr(): did not converge",1);
```

}

#### (Continued from inside front cover)

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