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Focusing of Internal Tides by Near-inertial Waves

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Abstract

The refraction of internal waves by lower-frequency near-inertial waves has been predicted theoretically, but never observed before. Here, we report observations of semi-diurnal internal tides generated by the rough topography of the Central Indian Ridge, in the presence of a strong, lower-frequency near-inertial wave field generated by a tropical storm. The semi-diurnal internal tide energy is trapped within upward-propagating bands with a periodicity close to the inertial period. A ray-tracing model suggests that this trapping results from the internal tide refraction by the shear associated with near-inertial waves. This yields a strong increase of the internal tide energy and shear in space-time regions where the background flow focuses the rays, leading to the formation of caustics. This mechanism may increase vertical mixing generated by baroclinic tides in the vicinity of mid-ocean ridges in tropical regions.

1 Introduction

41 Internal tides are believed to be a key source of the vertical mixing necessary to sustain the
42 global overturning circulation (Wunsch and Ferrari, 2004). Internal tides develop as the result
43 of the interaction of barotropic tides with bottom ocean topography. Abyssal generation of
44 internal tides from mid-ocean ridges represents about 25% of the total power dissipated by
45 barotropic tides (Egbert and Ray, 2001). Mid-ocean ridges are characterized by their rough
46 small-scale topography, and internal tides generated there are associated with relatively small
47 horizontal and vertical wavelengths (St. Laurent and Garrett, 2002, St. Laurent and Nash,
48 2004). It is estimated that ~30% of internal tide energy generated at mid-ocean ridges is
49 dissipated locally near the bottom (Polzin et al., 1997) while the remaining ~70% propagates
50 away from the source and contributes to background mixing away from the oceanic bottom
51 and source region (St. Laurent and Garrett, 2002).

52 The dissipation mechanism for the ~70% of internal tide energy radiated away from the
53 bottom remains unclear. Recent studies have shown that the scattering of low vertical mode
54 internal tides by the sub-inertial mesoscale field (Rainville and Pinkel 2006, Chavanne et al.,
55 2010, Zaron and Egbert, 2014, Ponte and Klein, 2015) or topographic features (Kelly et al.,

56 2013), and the subsequent generation of higher vertical modes, is a possible dissipation
57 mechanism. In parallel, theoretical and idealized numerical studies suggest that short
58 horizontal and vertical wavelength internal waves with upward energy propagation, including
59 internal tides, can be strongly affected by the ubiquitous lower-frequency near-inertial waves
60 with upward phase propagation such as those generated by storms (Broutman and Young,
61 1986, Broutman, 1986, Broutman et al., 1997, Vanderhoff et al., 2008). The varying
62 background shear associated with lower-frequency (near-inertial) waves create waveguides
63 that can focus the higher frequency internal waves into regions where the high resulting shear
64 can lead to high energy dissipation (Broutman 1986). The time-dependence of the inertial
65 wave field is crucial in this mechanism, which is thus fundamentally different from the
66 classical critical layer mechanism (Olbers, 1981). This process could in particular be efficient
67 close to the surface, where strong near-inertial motions occur in response to changing winds.
68 The formation of caustics (i.e. space-time regions where a time varying background flow
69 focuses the internal tide energy) has to our knowledge not yet been observed from in-situ
70 data.

71 About 100 GW of M2 barotropic tidal energy is converted to baroclinic tides by the rough
72 small scale topography on the Central Indian Ocean Ridge (Egbert and Ray, 2001, Nycander,
73 2005, Melet et al., 2013), making this region a hotspot for turbulent vertical mixing in the
74 global ocean (Whalen et al., 2012). Tropical storms also form in this region, such as the
75 developing Dora cyclone observed during the 2007 Cirene cruise (Vialard et al., 2009). The
76 generation of downward propagating near-inertial waves by the Dora cyclone was observed
77 and described by Cuypers et al. (2013). The Cirene cruise area is thus a region where there are
78 both strong internal tides and wind-generated inertial waves, which could lead to the
79 formation of caustics. Internal tides generated in this region have also not yet been described
80 from in-situ measurements. The goal of this article is to characterize internal tides generated

81 at the Central Indian Ocean Ridge and their potential modulation by near-inertial internal
82 waves.

83

84 **2 Data and methods**

85 The Cirene cruise and observations are described in detail in Vialard et al. (2009). The
86 cruise was conducted in two legs, each involving a 10-12 days station near 8°S, 67°30'E
87 (January 14-26 and February 4-15). The Dora tropical Storm formed in the interval between
88 the two legs. In this study we use conductivity temperature depth (CTD) casts performed
89 roughly every 20 minutes down to 500 m and CTD and Lowered-Acoustic Doppler Current
90 Profiler (L-ADCP) profiles down to 1000 m roughly every 6 hours. Post-processed CTD data
91 have 1 m vertical resolution and post-processed L-ADCP data provide horizontal currents at 8
92 m vertical resolution.

93 In order to get a more comprehensive picture of semi-diurnal (hereafter, SD) internal
94 tide generation and propagation in the Cirene region, we used the Gerkema (2002) 2D linear
95 internal tide generation model. The model requires the prescription of the barotropic tidal
96 flux, the topographic section in the model plane and the buoyancy frequency profile $N(z)$.
97 Barotropic fluxes were specified from the TPXO.7.1 global tidal model (Egbert and Erofeeva,
98 2002). We considered the two main SD tidal components (M2, S2), which contribute to
99 98.5% of the semi diurnal variance. The model plane was chosen along the SD ellipses (135°
100 from parallels), which are very eccentric, almost exactly aligned along the strongest
101 topographic gradient with weaker topographic fluctuations in the perpendicular direction (Fig.
102 1a). This configuration makes the 2D assumption reasonable. The model domain spans 800
103 km centered on the Cirene station, with 400 m horizontal resolution, and 60 degrees of
104 freedom in the vertical. The topographic profile is a linear interpolation from the 1 to 12 km
105 resolution Smith and Sandwell (1997) bathymetry (V18.1). The stratification profile $N(z)$ is

106 obtained from the cruise-average of CTD profiles down to 1000 m, completed with
 107 climatological values derived from the World Ocean Atlas 2013 (Locarnini et al., 2013,
 108 Zweng et al., 2013) down to the bottom.

109 The interaction between the SD internal tide and Near-inertial internal Waves (NIWs
 110 hereafter) is investigated using an Eikonal equation as in Broutman (1986). Eikonal equations
 111 describe the evolution of a test internal wave, in our case the internal tide, in a slowly varying
 112 medium in time and/or space. The wave is described in terms of density fluctuations
 113 $\rho' = a \exp(i\theta)$ where $\mathbf{k} = (k, 0, m) = \nabla_{\mathbf{x}} \theta$ is the wavenumber and $\omega_i = \partial_t \theta$ the intrinsic
 114 frequency. The background medium is defined from the velocity \mathbf{U} and stratification N , \mathbf{k} and
 115 ω_i are related through the internal waves dispersion relationship:

$$116 \quad \omega_i^2 = \frac{N^2 k^2 + f^2 m^2}{k^2 + m^2} \quad (1)$$

117 Since the medium is moving with the velocity \mathbf{U} , there is a Doppler shift between ω_i and the
 118 ω_0 frequency observed from a fixed point: $\omega_0 = \mathbf{U} \cdot \mathbf{k} + \omega_i$. We also assume that there are no
 119 background and wave energy density variations along the horizontal. The Eikonal equations
 120 governing the ray energy propagation read:

$$121 \quad \mathbf{c}_g = \frac{d\mathbf{x}}{dt} = \nabla_{\mathbf{k}} \omega_0 \quad (2)$$

$$122 \quad \frac{dm}{dt} = -\frac{\partial \omega_0}{\partial z} = -\frac{k^2 N N_z}{(k^2 + m^2) \omega_i} - k U_z \quad (3)$$

$$123 \quad \frac{\partial A}{\partial t} + \nabla_{\mathbf{x}} [\mathbf{c}_g A] = 0 \quad (4)$$

124

125 Equations (2) and (3) yield the ray trajectory and the evolution of the vertical
 126 wavenumber along the ray. The third equation expresses the conservation of the wave action
 127 $A = E/\omega_i$ with the energy density E , related to density fluctuations amplitude a through

128 $E = 2PE \left[1 + \left(\frac{fm}{Nk} \right)^2 \right]$ and $PE = \frac{1}{4} \frac{g^2 a^2}{\rho_0 N^2}$ the Potential Energy density. The set of equations is

129 integrated numerically using a 4th order Runge-Kutta scheme with partial steps. We integrate
130 the volume element of a ray tube $V = 1/A$ rather than the wave action equation, as in Hayes
131 (1970). This quantity vanishes at caustics (ray convergence points) rather than displaying a
132 singular behavior. The initial conditions for the internal tide and the NIW are based on the
133 observations (supplementary material). We will discuss the validity of this ray-tracing
134 approach in section 4.

135

136 **3 Results**

137 **3.1 Near-inertial waves**

138 As previously described in Cuypers et al. (2013), the Dora tropical storm passed close to
139 the Cirene area between the two legs and generated a strong NIW response during the second
140 leg. This is reflected in the Kinetic Energy (KE) spectrum of 0-500 m averaged L-ADCP
141 velocities, which displays a clear near-inertial (1.2f +/-0.15f) peak during the second leg, but
142 not during the first (Fig.1b). The PE spectrum also displays a near-inertial peak during the
143 second leg, with much less energy than on the KE spectrum due to the weak PE/KE ratio of
144 NIWs.

145 NIWs are characterized by upward phase (and downward energy) propagation. Their
146 velocity fluctuations can hence be isolated by extracting upward phase propagating signals as
147 in Cuypers et al. (2013) (Fig. 2ab). This confirms that the NIW signal is much stronger during
148 the second leg, and that the downward energy propagation can be tracked using the WKB
149 method (dashed lines in Fig. 2b, see Cuypers et al. 2013 for details).

150 **3.2 Strong semi-diurnal tide signal**

151 Cirene data also indicate a clear tidal signal during both legs. The L-ADCP sampling
152 frequency ($\sim 4 \text{ day}^{-1}$) is not sufficient to resolve the SD internal tide peak. The frequent CTD
153 measurements ($\sim 3 \text{ hour}^{-1}$) however allow estimating the Potential Energy (PE) spectrum over
154 a wider frequency range. The PE spectrum displays a diurnal peak and the dominant peak is
155 associated with the SD internal tide for both legs. The frequency resolution is however not
156 sufficient to distinguish the close M2 and S2 periods.

157 The 2D internal tide model (ITide model in the following) allows to better
158 understanding the origin of the relatively strong SD tidal signal at the Cirene site. It displays a
159 complex array of narrow SD internal tide beams generated by the rough topography (Fig. 3a).
160 The Cirene station ($x=0 \text{ km}$) is exactly co-located with an energy near-surface local
161 maximum, corresponding to the convergence and surface reflection of two internal tidal
162 beams generated from two neighboring ridges at $x=65 \text{ km}$ and $x=-60 \text{ km}$. The vertically-
163 integrated energy flux (Fig. 3b) indeed displays two strong gradients near these positions,
164 implying strong internal tide generation there. The red lines on Fig. 3a indicate the rays
165 emanating from those ridges, computed using the WKB theory in the simple case of a
166 stationary beam with no background velocity, as in the ITide model simulation. In this case
167 the ray slope is simply defined by $dz/dx = \pm \left[(\omega^2 - f^2) / (N(z)^2 - \omega^2) \right]^{1/2}$. The ray trajectory
168 closely matches the energy beam of the ITide model, even in the pycnocline where the WKB
169 hypothesis (slow variation of stratification) in principle breaks down. This suggests that the
170 eikonal equations (2-4) can be used to study internal semi diurnal tide packets trajectories.

171 The ITide model can be validated from Fig. 3d and e, which compare the time-averaged
172 SD PE in the model and Cirene data for both legs. The model is linear with uniform
173 background, and thus generates very focused beams associated with strong PE spatial
174 gradients. Beams are likely less focused in reality because of dissipation and internal

175 reflections (Gerkema and Van Haren, 2012) and we hence compare the model average within
176 5 km (roughly half the beam width) of the Cirene station to Cirene observations. Observed
177 and simulated average PE profiles match reasonably well for both legs. There is a reasonable
178 match between the ITide model and the data which give some confidence that the internal
179 tides are indeed generated from the nearby ridges. We will show in the next section how the
180 NIW and time mean background currents during leg2 strongly alter the internal tide
181 propagation.

182 **3.3 Focusing of internal tides by near-inertial internal waves**

183 In order to characterize SD tide spatial scales, the density field was also separated into
184 ρ'_{up} with upward and ρ'_{dn} with downward phase propagation. The SD internal tide phase as a
185 function of depth $\varphi(z)$ was obtained from demodulation of ρ'_{dn} , over the first half of leg2
186 (Feb 4-Feb 10), before the tide could significantly interact with the NIW packet. The nearly
187 linear shape of $\varphi(z)$ between 250 m and 500 m (not shown) implies an initial vertical
188 wavenumber $m_0 = 8.06 \times 10^{-3}$ rad m⁻¹ with a $[8.03 \times 10^{-3} - 8.08 \times 10^{-3}]$ 95% confidence
189 interval. Applying the dispersion relationship (1) with the average N value in this depth range
190 leads to a horizontal wavelength of ~ 23 km. The same procedure applied to the first leg (when
191 there was hardly any NIW signal) yields similar numbers: $m_0 = 9.84 \times 10^{-3}$ $[9.82 \times 10^{-3} -$
192 $9.86 \times 10^{-3}]$ rad m⁻¹ and ~ 18 km. The internal tide horizontal wavelength is a bit larger in the
193 ITide model, where the horizontal wavenumber spectrum of buoyancy fluctuations (not
194 shown) displays a broad peak at ~ 30 km. This may result from the strong presence of small-
195 scale abyssal hills in the Central Indian Ridge which can enrich the high wavenumber modes
196 (Melet et al., 2013) but which are smoothed in the Smith and Sandwell (1997) bathymetry
197 used in the ITide model.

198 We have band pass filtered ρ'_{dn} using an elliptical filter (Park, 1987) in the $[0.8\omega_{SD} -$

199 1.2 ω_{SD}] frequency band in order to isolate the SD tide signal. The PE for an internal tide with
200 upward energy propagation (downward phase propagation) is $PE_{dn} = \frac{1}{2} \frac{g^2}{\rho_0 N^2} \langle \rho'_{dn}{}^2 \rangle_{T_{SD}}$ and
201 is represented in Fig. 2d and 2e for leg1 and leg2 respectively. During the first leg, the SD PE
202 roughly follows the spring / neap cycle of the barotropic tide (displayed in Fig. 2g), with most
203 SD tide signal between the 19th and 25th of January.

204 There is also a clear modulation of the SD tide by the spring/neap cycle during the
205 second leg, with much less energy after the 10th of February (Fig. 2e and red curve on 2h).
206 The most striking feature is however a very clear modulation of the SD energy with a period
207 close to the NIW period down to 350 m. This modulation propagates upward (Fig. 2e) and the
208 similarity with the NIW current variations (Fig. 2b) is striking.

209 The observed SD energy maxima vertical propagation velocity (roughly $1 \times 10^{-3} \text{ m s}^{-1}$
210 between 150 m and 300 m) is not consistent with the group velocity of the internal tide. This
211 group velocity should indeed be about one order of magnitude larger, considering the
212 horizontal wavelength of 20 km and mean value of N for this depth range. There is a clear
213 scale separation between the NIW and SD internal tide, not only in frequency (SD \sim 7f), but
214 also in horizontal wavenumber. The NIW is characterized by a \sim 250-300 km horizontal
215 wavelength (Cuyppers et al., 2013), about one order of magnitude larger than the internal tide.
216 In the vertical, this scale separation is not as clear, with a ratio of the internal tide to the NIW
217 vertical wavelength of order one (see supplementary material). Previous studies have however
218 used the WKB hypothesis successfully in such a marginal cases (e.g. Rainville and Pinkel,
219 (2006), Chavanne, (2010) or Sheen et al (2015)). Sartelet (2003) has more specifically shown
220 a very good match between a high resolution numerical model and the eikonal equation in a
221 very similar context to the one we study in the present paper. She concludes that “ray theory
222 performs remarkably well even when the scale separation between the background wave and

223 the gravity wave breaks down completely both in the vertical and time”. We have therefore
224 used equations (2) to (4) confidently to model the propagation of an internal tide ray
225 emanating from the eastern ridge in Fig. 3 and propagating westward. The rays emanating
226 from the western ridge do not cross the Cirene station when background NIW and time-mean
227 currents are considered and are therefore not discussed in the following. We only consider the
228 mean observed background velocity for the first leg (Fig. 2g). For the second leg, we consider
229 the mean observed background velocity (Fig. 2h) or alternatively both the mean current and
230 the NIW velocity fluctuations (Fig. 2f). The initial conditions and NIW velocity fluctuations
231 for the second leg (Fig. 2c) were constructed to mimic observations (compare Fig. 2b and 2c):
232 details are given as supplementary material.

233 For the first leg simulation, the ray trajectories are invariant, the slope of the rays
234 decreases within the pycnocline while the tidal PE increases (Fig. 2g). The overall observed
235 increase in amplitude of the SD tide associated with the spring-neap cycle toward the end of
236 the first leg (Fig. 2d) is reproduced with a lag of ~ 1 to 2 days relative to barotropic forcing,
237 consistent with the propagation time of SD tides from the generation sites to the Cirene
238 station. Observations however display shorter time scale fluctuations, possibly associated with
239 the weak NIW activity during the first leg, not considered in our calculations. These
240 fluctuations could also result from interactions with remotely generated internal tides (Kelly
241 and Nash, 2010).

242 The PE distribution obtained when only considering the second leg average current in
243 the Eikonal model is similar to that of the first leg, just showing a spring-neap cycle (Fig. 2h),
244 and very different from the observed modulation of SD tides at near-inertial frequencies (Fig.
245 2e). Including the refraction by NIWs gives a picture in much better qualitative agreement
246 with observations, with upward propagating bands of SD PE. The second leg ray trajectories
247 are shown on Fig. 2c, with the prescribed NIW field as a background. SD tide packets have an

248 initial group velocity c_{gz} , which exceeds the NIW phase speed C_{NIW} when they penetrate the
249 NIW packet, as revealed by the slope of the rays relative to the slope of the NIW phase.
250 Broutman et al. (1997) classified the different interactions between an inertial wave and a
251 shorter wave as a function of the $r = c_{gz} / C_{NIW}$ ratio before the interaction. For $r \gg 1$ (as in
252 our case, for which $r \cong 10$), a strong convergence of wave action is expected, leading to a
253 strong increase in the amplitude of the short waves. The internal tide rays form a caustic when
254 they approach NIW shear maxima. Positive NIW shear indeed induces a strong refraction,
255 characterized by a strong vertical wavenumber modulus increase (Eq. 3), and a group velocity
256 (e.g. ray slope) that becomes lower than C_{NIW} . The region of strong refraction propagates
257 upward along with the region of strongest NIW shear. The time variations of the ray slopes
258 lead to the formation of caustics, defined as the envelope of space-time regions where the rays
259 converge. After the caustic, the rays interact with the NIW field a second time in the upper
260 part of the water column (above 150 m depth). The negative NIW shear induces a vertical
261 group velocity increase and generates a second caustic. Following Broutman (1986), we can
262 define caustics as points where $c_{gz} = C_{NIW}$, i.e. with a zero internal tide group velocity in a
263 frame of reference moving at the NIW phase speed. In Fig. 2c these caustics are plotted as
264 continuous magenta curves, choosing the arbitrary numerical criterion that c_{gz} and C_{NIW}
265 differ by less than 1%.

266 The refraction and focusing of the internal tide by the NIW generates an alternation of
267 shadowed regions with no tidal energy and higher energy bands (Fig. 2f), with the NIW
268 period. The strongest PE increase is observed when rays approach caustics. As discussed by
269 Broutman (1986) and Vanderhoff (2008), the assumption of a slowly-varying medium breaks
270 down near caustics, yielding unrealistically high wave amplitude. Broutman (1986) suggests a
271 corrected maximum amplitude at the caustic, assuming that the wave envelope is given by an

272 Airy-function. Fig. 2i displays the PE computed using Broutman’s correction along a selected
 273 ray (whose trajectory in (z,t) and (x,z) plane is represented as a thick black dashed curve in
 274 Fig. 2f and 3f respectively). The ray experiences strong refraction and PE increase at the two
 275 caustics. We also show the internal tide Richardson number
 276 $Ri = N^2 / u_z^2 = k^{-2} (2 / \rho_0)^{-1/2} \omega_i (1 - f^2 / \omega_i^2) A^{-1}$, whose values below $1/4$ characterize
 277 potential shear instability. The strong refraction of the first ray yields a more than an order of
 278 magnitude decrease of the Ri values, relative to when refraction by NIWs is not considered.
 279 While Ri (~ 1) remains above the critical value, this strong localized increase of the shear will
 280 combine with other background shear variations, potentially leading to “bursts” of intense
 281 vertical mixing. This underlines the potential consequences of this ray-focusing mechanism
 282 on vertical mixing in the upper ocean.

283

284 **4 Discussion**

285 The Cirene data is characterized by strong NIW activity during the second leg, after the
 286 passage of tropical storm Dora, with scales of ~ 3 days and ~ 300 km (Cuypers et al. 2013).
 287 The data also reveals a strong semi-diurnal internal tide activity of ~ 20 km horizontal scale,
 288 generated by the interaction of barotropic tides with two nearby ridges. Here, we report the
 289 first observations of SD internal tide amplitude modulation by the larger-scale, lower-
 290 frequency NIW fluctuations. Ray tracing using the Eikonal equation suggests that the NIW
 291 velocity field can strongly focus the internal tide energy along caustics and induce large space
 292 and time fluctuations of the internal tide amplitude and wavenumber at NIW spatio-temporal
 293 scales, as first theoretically described by Broutman et al. (1997) and Vanderhoff et al. (2008).
 294 Most of the internal tide rays are strongly refracted by the NIW shear, giving rise to two
 295 caustic points in the upper 150 m. The strong decrease of the Richardson number at the

296 caustic points is potentially conducive to “bursts” of vertical mixing there.

297 The ray tracing approach in this paper neglects potentially important effects such as ray
298 scattering by stratification or currents horizontal heterogeneities, the three-dimensional
299 character of the SD energy radiation associated with real bathymetry, and possible nonlinear
300 effects such as wave breaking or energy leaks to the diurnal frequency by PSI (Sun and
301 Pinkel, 2013) which can lead to the degradation of the I-Tide beam in the pycnocline (Gayen
302 and Sakar, 2013). It is hence not surprising that the agreement between Figs. 2e and f is only
303 qualitative. Yet the ray tracing captures the observed SD energy trapping along bands that
304 display vertical propagation at a speed close to the NIW phase speed, while those bands
305 disappear when the refraction mechanism is neglected (Fig. 2h). We also examined alternative
306 mechanisms for semi-diurnal tide modulation at near-inertial frequencies in the
307 supplementary information. NIW-induced stratification variations are too weak to explain the
308 observed variations. Considering beams of finite width instead infinite width as presented
309 here yields advection of the rays that can induce SD tide amplitude variation at a given point,
310 due to the varying distance between the beam centers and observational point. Those
311 variations however form an oscillation of the PE maximum around 130 m depth, unlike
312 observations which show NIW-modulation in the form of energy bands observed over several
313 hundreds of meters. We hence believe that the mechanism proposed by Broutman et al.
314 (1986) is the most plausible explanation for the SD modulation at near-inertial frequencies
315 that we observe. Fully non-linear simulations and turbulent dissipation measurements during
316 future field studies will however be needed to ascertain the validity of this hypothesis and to
317 estimate the internal tide dissipation rate associated with this process. It is worth noting that
318 this process is fundamentally different from the direct generation of harmonics at the sum of
319 tidal and inertial frequencies as observed by Davies and Xing (2003), which would not
320 produce a modulation of the internal tide energy at the near inertial period.

321 The process described above can enhance vertical mixing in the ocean by concentrating
322 the tidal energy and/or decreasing the vertical wavelength such that the flow can become
323 super critical. This process can only affect internal tides when there is a sufficient scale
324 separation between the internal tide and NIWs. The tropical region, where the semi-diurnal
325 frequency is much higher than f , is favorable for this scale-separation. This process also
326 requires a strong high-frequency wind variability that can generate NIWs, as for example in
327 regions where deep atmospheric convection and low-pressure systems frequently form. The
328 dissipation of internal tides by this process will mostly occur in the upper 150m, where most
329 NIW energy is concentrated (Furuichi et al. 2008). Small-scale internal tides are mostly
330 generated near mid-ocean ridges such as the Indian and mid-Atlantic ridges and the East
331 Pacific rise (St. Laurent and Garrett 2002). The intersection of these generation sites with the
332 tropical band represents a large fraction of the global internal tide generation (Egbert and Ray,
333 2001, Nycander, 2005, Melet et al., 2013). There is hence a potential for this focusing process
334 to be relevant on a global scale, confirmation of which will rely on further observational and
335 modeling work.

336

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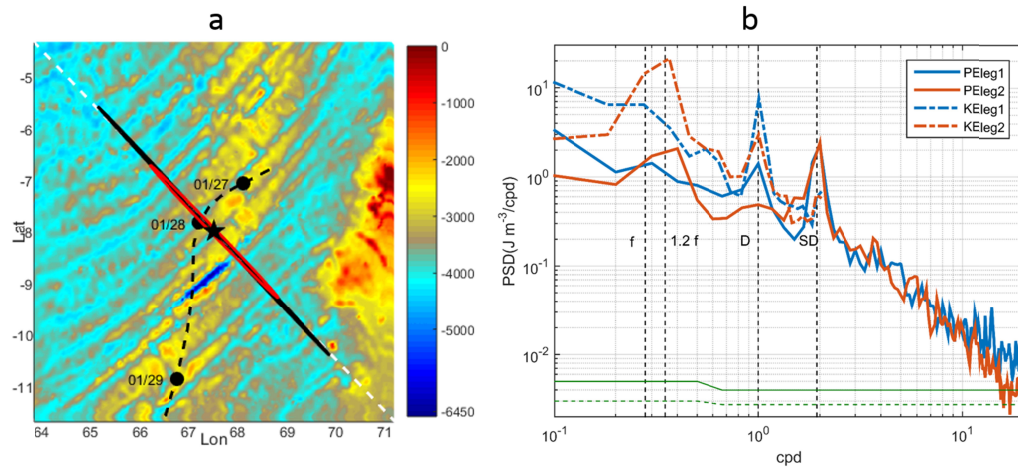
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343 (Jerome.Vialard@ird.fr). This is PMEL contribution no. 4434.

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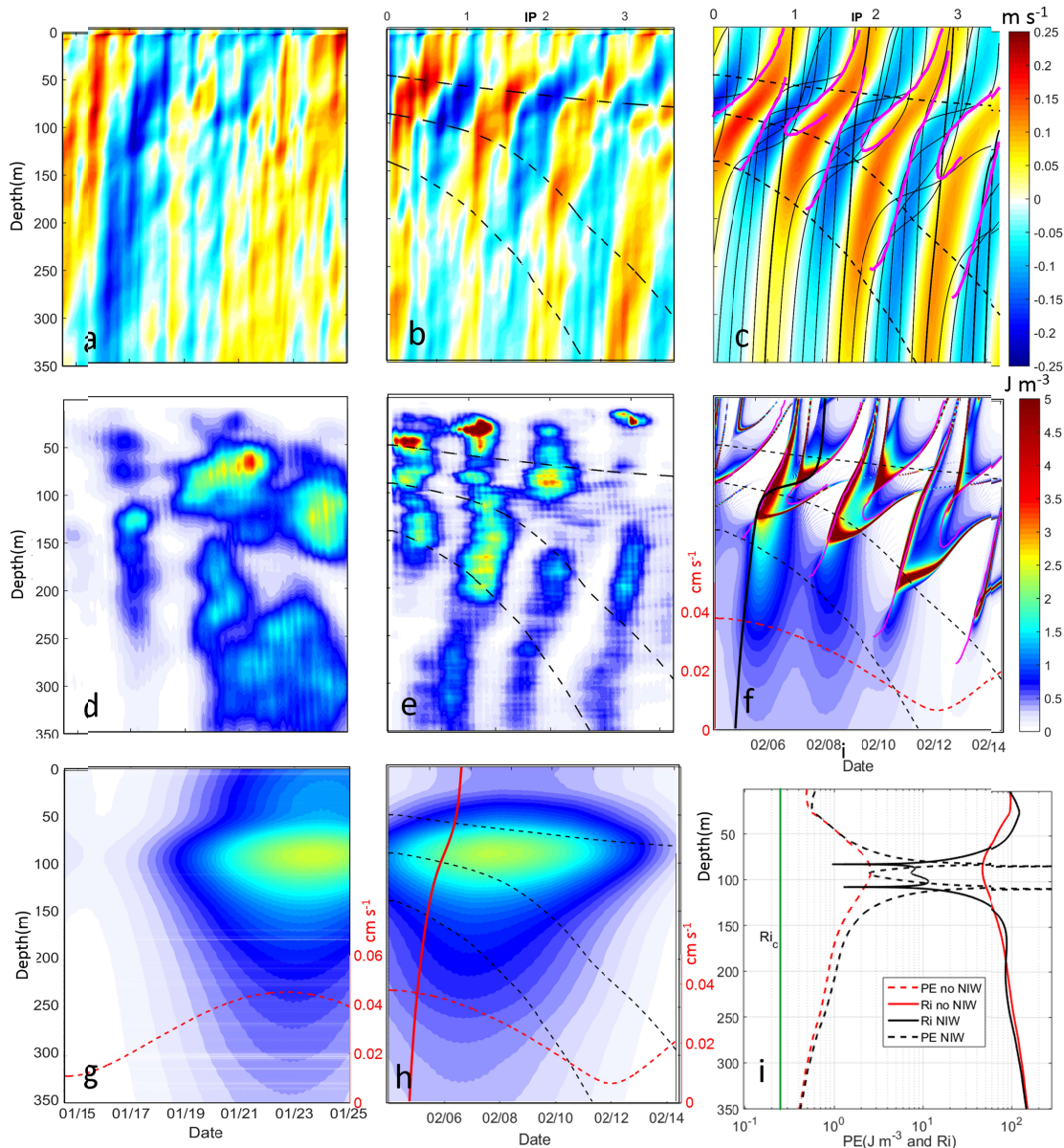
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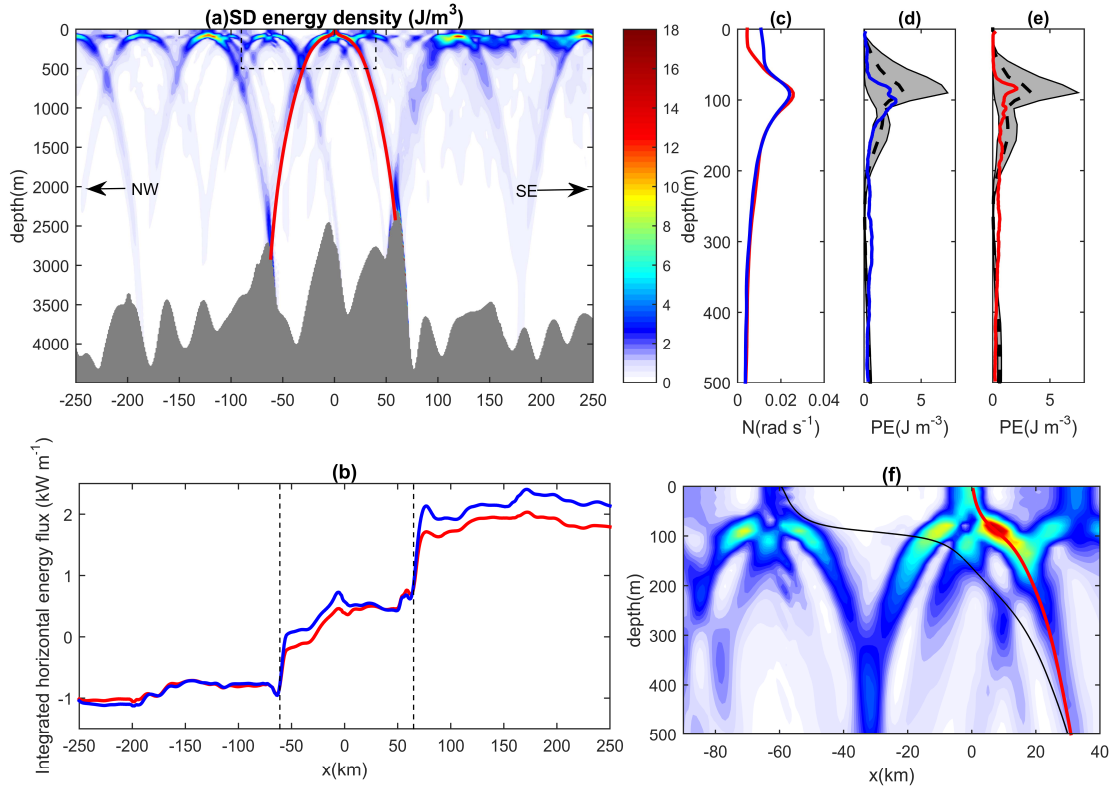
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Figure 1 (a) Bathymetry in the Cirene region. The black star indicates the Cirene measurement position (8°S 67.5°E). The black dashed line is the trajectory of cyclone Dora from the IBTraCs database (Levinson et al. 2010). The black and red ellipses respectively represent the M2 and S2 barotropic tidal ellipses, the white dashed line represents the two-dimensional numerical model plane; (b) 0 to 500 m depth averaged spectrum from Cirene observations during leg1 (in blue) and leg2 (in red): Kinetic Energy (KE) as dashed lines, Potential Energy (PE) as solid lines. The vertical black dashed lines indicate the inertial (f), near-inertial ($1.2f$), tidal Diurnal (D) and Semi-Diurnal (SD) frequencies. The distance between the green dashed line and the green plain line (bottom line) represents the upper (lower) bound of the 95% confidence interval.



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Figure 2 (a) leg1 and (b) leg2 L-ADCP velocities with upward phase propagation projected on the semi-diurnal tidal ellipse axis. (c) Prescribed idealized NIW velocity field for Eikonal equation integrations for leg2. (d) leg1 and (e) leg2 observations of the internal wave Potential Energy PE_{dn} (see text for details). (f), leg2 and (g) leg1 PE along ray trajectories in the Eikonal model, (h) same as (f) when refraction by the NIW shear is not considered. (i) PE (dashed line) and Richardson number Ri (plain line) along the plain red (see panel h) and plain black (see panels f) ray trajectories. The black lines in (c) indicate some internal tide ray trajectories. The black dashed lines in (b), (c), (e), (f), (h) represent the NIW packet envelope. The magenta plain lines in (c) and (f) represent the estimated caustic curves. For panels (b), (c), (e), (f), (h) the upper axis is in near-inertial periods (IP). The red dashed lines at the bottom of panels (f), (g) and (h) represent the SD barotropic tidal envelope of the current amplitude the corresponding scale is on the right (left) of (g),(h) ((f)), .



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 440 Figure 3 (a) Semi-diurnal tide energy density vertical section along the white dashed line in figure 1a in the
 441 numerical model. The red lines indicate internal tide ray trajectory in the absence of background velocity. (b)
 442 depth-averaged horizontal energy flux in the numerical model, leg1 in blue, leg2 in red. (c) Mean Brunt Väisälä
 443 frequency during leg1 in blue, leg2 in red. (d) leg1 and (e) leg2 mean observed PE density, in red and blue for
 444 observations, as a black dashed line for the model average within 3 km of observations with the range of
 445 variation in this region indicated as a grey shading. (f) zoom corresponding to the dashed frame on panel (a),
 446 with the trajectory of an internal wave ray in the absence of background velocity in red dashed line and when the
 447 idealized near-inertial wave velocity field and background current are included in black dashed line. The
 448 corresponding trajectories in the (z,t) plane are represented in Fig. 2f and h.