1	Wind, waves, and fronts: Frictional effects in a generalized Ekman model
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ABSTRACT

Ocean currents in the surface boundary layer are sensitive to a variety of pa-8 rameters not included in classic Ekman theory, including the vertical structure 9 of eddy viscosity, finite boundary layer depth, baroclinic pressure gradients, 10 and surface waves. These parameters can modify the horizontal and verti-11 cal flow in the near-surface ocean, making them of first-order significance 12 to a wide range of phenomenon of broad practical and scientific import. In 13 this work, an approximate Green's function solution is found for a model of 14 the frictional ocean surface boundary layer, termed the generalized Ekman 15 (or Turbulent Thermal Wind) balance. The solution admits consideration of 16 general, more physically realistic, forms of parameters than previously possi-17 ble, offering improved physical insight into the underlying dynamics. Closed 18 form solutions are given for the wind-driven flow in the presence of Coriolis-19 Stokes shear, a result of the surface wave field, and thermal wind shear, arising 20 from a baroclinic pressure gradient, revealing the common underlying phys-21 ical mechanisms through which they modify currents in the ocean boundary 22 layer. These dynamics are further illustrated by a case study of an idealized 23 two-dimensional front. The solutions, and estimates of the global distribution 24 of the relative influence of surface waves and baroclinic pressure gradients 25 on near-surface ocean currents, emphasize the broad importance of consider-26 ing ocean sources of shear and physically realistic parameters in the Ekman 27 problem. 28

29 1. Introduction

Diagnosing velocities in the ocean boundary layer is key to many issues of broad practical and scientific importance, from larval dispersion, to search and rescue, to the general ocean circulation. Today much of our understanding of boundary layer currents remains rooted in classic Ekman theory, which holds that, with some knowledge of the turbulent eddy viscosity, the ageostrophic ocean response is completely determined by the surface wind stress (Ekman 1905). However, despite the tremendous explanatory power of Ekman theory, basic observational confirmation of the structure of flow in the boundary layer has been challenging.

In response to discrepancies between the theory and observations, a large literature has devel-37 oped, focused on modifications to the classic Ekman theory. Broadly speaking the proposed mod-38 ifications to Ekman theory can be divided into local one-dimensional mechanisms, such as time-39 variability (Price et al. 1986; Schudlich and Price 1998; Price and Sundermeyer 1999; McWilliams 40 et al. 2009), vertical structure in eddy viscosity (Madsen 1977; Miles 1994; Grisogono 1995), or 41 finite boundary layer depth (Welander 1957; Stommel 1960; Lewis and Belcher 2004; Elipot and 42 Gille 2009), and mechanisms that involve non-local effects such as horizontal buoyancy gradi-43 ents (McPhaden 1981; Cronin and Kessler 2009), surface waves (Huang 1979; Jenkins 1986; Xu 44 and Bowen 1994; Lewis and Belcher 2004; Polton et al. 2005), and non-linearity (Stern 1965; 45 Niiler 1969; Thomas and Rhines 2002). Many of these proposed modifications have closed the 46 gap between theory and observations, however, generally analytic solutions are only available for 47 specific forms of parameters, limiting the possibility for inter-comparison of the various proposed 48 mechanisms, and critically, their application to realistic ocean fields. 49

Here we utilize a simple model of the viscous boundary layer, termed the generalized Ekman
 model (Cronin and Kessler 2009), or the Turbulent Thermal Wind balance (Gula et al. 2014). This

model contains many of the modifications to basic Ekman theory that have been proposed individually, and has already proven successful in explaining observed horizontal currents (Cronin and Kessler 2009), as well as modeled boundary layer vertical velocities (Gula et al., McWilliams et al., 2015). Previously, solutions to this model with physically realistic parameters required numerical methods, with analytic solutions available only for greatly simplified forms of the parameters (Bonjean and Lagerloef 2002; Cronin and Kessler 2009; McWilliams et al. 2015), limiting insight into the underlying dynamics.

In this manuscript we significantly extend these earlier results by providing an approximate so-59 lution to the generalized Ekman (Turbulent Thermal Wind) model that can accommodate a wide-60 range of physically realistic parameters, providing a unifying framework for many of the individu-61 ally proposed modifications to classic Ekman theory (section 2). Using this solution, two limiting 62 cases, corresponding to a surface wave field and horizontal buoyancy gradient, are explored to 63 further illuminate the underlying dynamics (section 3). The approximate solutions to these lim-64 iting cases reveal how these two ocean dynamic processes modify the Ekman solution in similar 65 ways, drawing a previously unnoted connection between these processes, and their accompanying 66 literatures. 67

In section 4 the solution is applied to an idealized front, illustrating how thermal wind shear in the presence of viscosity can alter both the Ekman layer flow as well as drive overturning circulations in the boundary layer (Garrett and Loder 1981; Thompson 2000; McWilliams et al. 2015). Estimates of the global distribution of wave and baroclinic pressure gradient effects on frictional boundary layer flow (section 5), and scaling analysis, suggest that these ocean dynamical processes can be expected to be of first order importance in determining near-surface currents for much of the world's oceans.

2. Theory

We consider steady, Boussinesq, flow in hydrostatic balance, where the complex horizontal velocity is denoted by $u \equiv u + iv$, and $\nabla \equiv \frac{\partial}{\partial x} + i\frac{\partial}{\partial y}$. Horizontal mixing is ignored, and vertical mixing is parameterized by a turbulent eddy viscosity, A_v , which is considered to be a specified parameter, allowed to vary vertically subject to moderate constraints imposed by the approximation technique utilized, as discussed below. The horizontal and vertical momentum equations are thus given by,

$$if\boldsymbol{u} = -\frac{1}{\rho_0}\nabla P + \frac{\partial}{\partial z} \left(A_v \frac{\partial \boldsymbol{u}}{\partial z} \right), \tag{1}$$

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$$0 = -\frac{1}{\rho_0} \frac{\partial P}{\partial z} + b.$$
⁽²⁾

Where the Rossby number, $\varepsilon = U/fL$, is assumed small, and therefore the non-linear advection 83 terms are excluded. Equation (2) expresses the hydrostatic balance, where $b = -g\rho/\rho_0$ is the 84 buoyancy, also considered to be a known quantity, allowed to vary in the horizontal and vertical. 85 Equation (1), a balance between the Coriolis acceleration, the pressure gradient force, and the 86 turbulent diffusive flux divergence provides the basic starting point for Ekman theory. Deriving 87 Ekman's 1905 result begins with a decomposition of the total velocity into a geostrophic velocity 88 in balance with the pressure gradient force ($u_g = i(\rho_0 f)^{-1} \nabla P$), and solving for the ageostrophic 89 velocity $(u_a = u - u_g)$ in a boundary layer with characteristic thickness $h_{Ek} = \sqrt{2A_v/f}$, the 90 Ekman depth, where it is assumed that A_v is vertically uniform and $\nabla b = 0$ (see for example 91 Gill 1982, section 9.6). Equation (1) is a second order linear ordinary differential equation for 92 velocity and so requires two boundary conditions on u, given for the classic Ekman problem by 93 $\rho A_v \partial u_a / \partial z = \tau_w$ at the surface, where τ_w is the surface wind stress, and $u_a \to 0$ as $z \to -\infty$. 94

Here we take a more general approach that does not require separating into geostrophic and ageostrophic components, by first vertically differentiating (1), and multiplying by $\rho_0 A_v(z)$ to form an equation for the stress, $\tau = \rho_0 A_v(z) \partial u / \partial z$, which we refer to as the generalized Ekman model (following Cronin and Kessler 2009),

$$A_{\nu}(z)\frac{\partial^2 \tau}{\partial z^2} - if\tau = \rho_0 A_{\nu}(z)\nabla b, \qquad (3)$$

100

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$$\boldsymbol{\tau}(0) = \boldsymbol{\tau}_{\boldsymbol{w}},\tag{4}$$

$$\boldsymbol{\tau}(-h) = 0. \tag{5}$$

The relationship of this model to various alternate Ekman layer formulations is discussed in detail by Cronin and Kessler (2009), however we further note that this is the same model termed the Turbulent Thermal Wind balance by Gula et al. (2014, and McWilliams et al. 2015) in their investigation of submesoscale cold filament dynamics.

The surface boundary condition, (4), is unchanged from the classic Ekman problem, however 105 the bottom boundary condition, (5), is posed as a no-stress condition, applied at a finite depth 106 z = -h, rather than the no-slip condition utilized in the classic Ekman problem. This formulation 107 of the problem maintains the classic Ekman transport, even in the presence of geostrophic shear 108 at the base of the layer (Cronin and Kessler 2009), and is applicable at low latitudes or in depth 109 limited seas (Stommel 1960; Bonjean and Lagerloef 2002). The solution technique utilized below 110 is a global method, as opposed to a local boundary layer expansion, and thus sufficiently far from 111 the boundary layer the solution will approach the inviscid limit. This gives a measure of flexibility 112 in the choice of an appropriate h, however on the basis of physical arguments, developed further 113 below, h should be chosen to be deeper than significant sources of geostrophic stress (defined in 114 section 3b), so as to avoid the creation of a spurious interior 'Ekman' layer. When $h \gg h_{Ek}$, as 115

is the case for most of the extra-tropics, the near-surface solution is insensitive to the particular bottom boundary condition, and we further note that for the solutions given below letting $h \rightarrow \infty$ results in simplified forms of the solutions that are equivalent to applying the bottom boundary condition $\tau \rightarrow 0$ as $z \rightarrow -\infty$. However, if a no-slip boundary condition is desired, the derivation follows directly from that given in Appendix A.

To solve this linear inhomogeneous ordinary differential equation with non-constant coefficients, we first approximate a solution to the homogenous formulation of equation (3) using the Wentzel-Kramer-Brillouin (WKB) method (Bender and Orszag 1978; Grisogono 1995), and then solve for the inhomogeneous solution using variation of parameters (Hidaka 1955; Berger and Grisogono 1998). A detailed derivation of the full solution is presented in Appendix A, however, briefly, the WKB method assumes the solution can be represented as:

$$\tau \propto e^{(S_0 + S_1 \delta + S_2 \delta^2 + \dots)\frac{1}{\delta}}.$$
(6)

Here, we use the physical optics approximation, and solve to first order (S_1) . The distinguished 127 limit for the small parameter δ is found to be $\delta \sim Ek^{1/2}$, where $Ek = A_v/fH^2$, the Ekman number. 128 The classic non-dimensionalization of (1), for constant A_{ν} , identifies H as the depth scale of the 129 interior flow, which for values typical of a stratified mid-latitude ocean ($A_v \sim 10^{-2} \,\mathrm{m^2 s^{-1}}$, $f \sim$ 130 10^{-4} s⁻¹, $H \sim 100$ m) gives $Ek \sim O(10^{-2})$. However, retaining vertical structure in A_v introduces 131 an additional vertical length scale, h_{Av} , into the problem. We thus have six physically relevant 132 terms $(A_v, f, u_g, u_a, H, h_{Av})$, with two physical dimensions (time and length). Application of the 133 Buckingham pi theorem (Buckingham 1914) then gives 4 non-dimensional parameters, 2 of which 134 take the form of an Ekman number, $\pi_1 = A_{\nu 0}/fH^2$ and $\pi_2 = A_{\nu 0}/fh_{A\nu}^2$. The latter of these is likely 135 to be a stricter constraint on the validity of the WKB expansion (Appendix B). 136

Formally, the use of the WKB approximation requires that the properties of the medium vary 137 more slowly than the solution (Bender and Orszag 1978), a condition which may be violated in 138 some geophysical flows. Further analysis of this requirement is given in Appendix A, however, 139 as discussed in Appendix B, we find good agreement between numerical and approximate solu-140 tions for a range of A_v profiles, and values of Ek, suggesting the utility of this solution (see also 141 Grisogono 1995). For simplicity we also require that $A_{\nu} > 0$ throughout the layer, so as to avoid 142 the additional complexity of singularities in the equation. This constraint, and the WKB condition 143 (A12), does not allow the direct application of the solution to cases where $A_v \rightarrow 0$ as $z \rightarrow 0$, as 144 for instance occurs in the K-Profile Parameterization (KPP) (Large et al. 1994). If necessary, this 145 restriction can be removed by patching the WKB solution to an inner solution, valid in a thin layer 146 as $A_v \rightarrow 0$ (as in Parmhed et al. 2005), discussed further in Appendix B. 147

Once the WKB solution to the homogenous problem is identified, the inhomogeneous solution can be found using variation of parameters, and stated in terms of a Green's function. The full solution (as derived in Appendix A) is thus given by,

$$\boldsymbol{\tau}(\boldsymbol{z}) = \boldsymbol{\tau}_{\boldsymbol{w}} \left(\frac{A_{\boldsymbol{v}}(\boldsymbol{z})}{A_{\boldsymbol{v}}(\boldsymbol{0})}\right)^{\frac{1}{4}} \frac{\sinh\left[\boldsymbol{\theta}(\boldsymbol{z})\right]}{\sinh\left[\boldsymbol{\theta}(\boldsymbol{0})\right]} + \int_{-h}^{0} G(\boldsymbol{z},\boldsymbol{s})\left[\boldsymbol{\rho}_{0}\nabla\boldsymbol{b}\right] \mathrm{d}\boldsymbol{s}.$$
(7)

¹⁵¹ Where,

$$\boldsymbol{\theta}(z) = \sqrt{if} \int_{-h}^{z} A_{\nu}(Z)^{-\frac{1}{2}} \,\mathrm{d}Z,\tag{8}$$

and, G(z,s) is the symmetric Green's function,

$$G(z,s) = \begin{cases} \frac{\sinh[\theta(z)]\sinh[\theta(s)-\theta(0)]A_{\nu}(s)^{\frac{1}{4}}A_{\nu}(z)^{\frac{1}{4}}}{\sinh[\theta(0)]\sqrt{if}} & \text{if } s > z\\ \frac{\sinh[\theta(s)]\sinh[\theta(z)-\theta(0)]A_{\nu}(s)^{\frac{1}{4}}A_{\nu}(z)^{\frac{1}{4}}}{\sinh[\theta(0)]\sqrt{if}} & \text{if } s < z. \end{cases}$$
(9)

¹⁵³ This general solution is a primary result of this manuscript.

Velocity shear follows directly from the definition of stress. However, equation (3) is a thirdorder linear ordinary differential equation in velocity, and hence to go from shear to velocity requires an additional boundary condition. Here, to determine velocity we use the solution for
 stress directly in the momentum equation (1),

$$if\boldsymbol{u} = -\frac{1}{\rho}\nabla P + \frac{1}{\rho}\frac{\partial\boldsymbol{\tau}}{\partial z},\tag{10}$$

where τ is now known through (7). This approach ensures that the vertically integrated ageostrophic velocity satisfies the classic Ekman transport relation.

160 a. Wind-Driven Component

The wind-driven component of the stress is given by the first term on the RHS of (7), which can be compared to the exact solution for the case where A_v is vertically constant (Stommel 1960; Bonjean and Lagerloef 2002),

$$\tau(z) = \tau_w \frac{\sinh\left[\sqrt{\frac{if}{A_v}}(z+h)\right]}{\sinh\left[\sqrt{\frac{if}{A_v}}h\right]}.$$
(11)

The parallels between the WKB approximation, (7), and the solution of the constant- A_{ν} problem (11) are apparent, with the leading order modification appearing in the argument of the hyperbolic functions, $\theta(z)$, given by (8). This term can be understood as introducing a stretched vertical coordinate system, defined by the integral in (8) (Lupini et al. 1975). Accordingly, $h_{Ek}(z) = \sqrt{2A_{\nu}(z)/f}$ defines a vertically localized Ekman depth, analogous to the local wavenumber found in WKB solutions to the wave equation (see for example Gill 1982, section 8.12).

The amplitude of the stress is also modified by vertical variations in viscosity, which appears as a ratio to the 1/4 power. For a slowly vertically decaying A_v , the effect of this term will be apparent only as $z \to -h$, where the ratio of $A_v(z)/A_v(0) \ll 1$. Figure 1 compares example vertical profiles of stress and velocity for the case of A_v decaying exponentially with depth, and for constant A_v , to illustrate the modification of the vertical structure arising from retaining a depth-dependent A_v . This depth dependent amplitude term allows the stress amplitude, and hence the ageostrophic ¹⁷⁶ velocity, to decay over a different vertical depth scale than the rotation of the stress vector, a ¹⁷⁷ feature which is commonly noted in observations but cannot be accommodated in classic Ekman ¹⁷⁸ theory (Price et al. 1986; Wijffels et al. 1994; Chereskin 1995; Price and Sundermeyer 1999). It is ¹⁷⁹ apparent that if A_{ν} is vertically constant in (7), the standard solution (11), as originally identified by ¹⁸⁰ Stommel (1960) in an investigation of the dynamics of the equatorial undercurrent, is immediately ¹⁸¹ recovered.

182 b. Inhomogenous Forcing

The second term on the RHS of (7) is a Green's function integral, which can accommodate 183 arbitrary vertical structure in both ocean sources of stress, which appear as inhomogenous forcing 184 terms in (3), as well as in the profile of $A_{\nu}(z)$, subject only to the constraints imposed by the 185 WKB method. The Green's function kernel takes the form of paired Ekman layers above and 186 below interior sources of stress (figure 2), which demonstrates how ocean sources of shear in the 187 presence of viscosity drive an ageostrophic frictional response felt throughout the entire boundary 188 layer (Hidaka 1955; Csanady 1982). Far from the boundaries the profile of the Green's function 189 is symmetric above and below interior shear, however approaching the boundaries of the domain 190 the shape of the Green's function becomes increasingly asymmetric, and the integral contribution 191 serves to satisfy the boundary conditions (4) and (5). 192

¹⁹³ The ageostrophic flow associated with several simplified forcings are shown in schematic form ¹⁹⁴ in figure 3. In the absence of wind-stress, with vertically uniform A_v and ∇b , Ekman layers are ¹⁹⁵ generated, both at the surface as well as at the base of the boundary layer, to satisfy the boundary ¹⁹⁶ conditions (Bonjean and Lagerloef 2002). A more physically realistic case is given in figure 3b, ¹⁹⁷ where a vertically decaying buoyancy gradient gives rise to both a surface Ekman layer, as well as ¹⁹⁸ a diffuse interior ageostrophic flow. The strength of the interior portion of the flow, for the situation

shown in figure 3b, scales as h_{Ek}/h relative to the surface ageostrophic flow, and is therefore often 199 assumed small and neglected. However, we note that this flow is necessary to balance the transport 200 in the surface Ekman layer so as to maintain the classic Ekman transport relation. Further, this flow 201 need not always be small, as illustrated in section 3a discussing surface wave effects, and hence 202 should be retained. The final panel shows the case of a decaying A_{ν} profile, with constant ∇b . 203 The resulting ageostrophic velocities are similar to those in figure 3b, however the associated 204 buoyancy fluxes will differ between the two cases, emphasizing how horizontal fluxes will be a 205 complex function of the spatial structure of both the background fields and A_{ν} , discussed further 206 in section 5b. 207

The total frictional ageostrophic response thus consists of a directly wind-forced component, as 208 well as an integral over Ekman-like responses to interior shear. Therefore, in order to understand 209 the oceanic response to wind-forcing it is also necessary to understand the ageostrophic frictional 210 response to ocean dynamical processes (Cronin and Kessler 2009). Recent observational work 211 has emphasized the importance of removing estimates of the geostrophic shear in order to isolate 212 the ageostrophic flow (Chereskin and Roemmich 1991; Polton et al. 2013; Roach et al. 2015), 213 however the analysis developed here suggests that to fully isolate the wind-driven component of 214 this flow it is also necessary to account for ageostrophic flow driven by the geostrophic shear 215 (section 3b). Further, although we have so far limited the discussion to shear which arises from 216 baroclinic pressure gradients, we note that any other forcing terms in the momentum equations will 217 act in a similar manner, and the case of Stokes shear from surface waves is discussed in section 3a. 218 In order to further illustrate the underlying dynamical mechanisms we now consider two limiting 219 cases representing important sources of shear in the ocean surface boundary layer which admit 220 further simplification of the full solution. 221

222 **3. Limiting Cases**

a. Stokes Shear: $h_s \ll h_{Ek}$

Surface waves modify the oceanic boundary layer in a variety of important ways (Xu and Bowen 224 1994; McWilliams et al. 1997; McWilliams and Restrepo 1999; Sullivan and McWilliams 2010; 225 Belcher et al. 2012; McWilliams et al. 2012). Here we focus on one particular aspect, termed the 226 Coriolis-Stokes force, which appears as an additional term in the Eulerian momentum equation that 227 arises from rotation acting on the Stokes drift, leading to a tilting of wave-orbitals in the along-228 crest direction (Polton et al. 2005). The Coriolis-Stokes force has been shown to significantly 229 modify flow in both the very near-surface layer, as well as throughout the entire Ekman layer 230 (Huang 1979; Jenkins 1986; Lewis and Belcher 2004; Polton et al. 2005; Aiki and Greatbatch 231 2012; McWilliams et al. 2014). 232

Equation (1) can be re-written to include the Coriolis-Stokes force as,

$$if(\boldsymbol{u}+\boldsymbol{u}_{\boldsymbol{s}}) = -\frac{1}{\rho_0} \nabla P + \frac{\partial}{\partial z} \left(A_v \frac{\partial \boldsymbol{u}}{\partial z} \right), \qquad (12)$$

with u_s the Stokes velocity, given by, $u_s(z) = \int_k 2\sigma k \chi(k) e^{2|k|z} dk$, and where σ is the wave fre-234 quency, k the wavenumber vector, and $\chi(k)$ the directional wave spectrum (Huang 1971). We make 235 the common simplifying assumption that u_s can be treated as a monochromatic wave such that 236 $u_s = U_0 e^{\frac{2}{h_s}} \hat{s}(t)$, where $h_s = (2|k|)^{-1}$, and $\hat{s}(t)$ is a unit vector in the direction of the waves, which 237 is not necessarily aligned with the local surface wind stress. The wavenumber, k, and amplitude, 238 U_0 are assumed to be known or parameterized. It is important to note that for a time-varying wave 239 field, the Coriolis-Stokes force initially accelerates an 'anti-Stokes' flow (McWilliams and Fox-240 Kemper 2013), with transients that decay as 1/ft (Lewis and Belcher 2004). In the steady-state 241 problem, including the Coriolis-Stokes force results in an additional forcing term on the RHS of 242 (3), perpendicular to the wave direction, given by $i f \rho A_{\nu} \partial u_s / \partial z$. This appears in the full solution 243

²⁴⁴ (7) within the Green's function integral, replacing the bracketed term with, $[\rho \nabla b + i f \rho \partial u_s / \partial z]$. ²⁴⁵ Hence, both horizontal buoyancy gradients and Stokes shear modify the standard Ekman solution ²⁴⁶ in mathematically identical ways.

In order to provide an asymptotic approximation to (7), we can take advantage of the scale separation between the typical depth scale of the surface waves, h_s , which is of order several meters, and h_{Ek} which is of order tens of meters, such that $h_s \ll h_{Ek}$. For simplicity in deriving the given form of (13), it is also assumed that $\tau_{CS}(-h) \sim 0$, and $h_{Av} \gg h_{Ek}$ where h_{Av} is the depth scale over which A_v varies, however neither of these assumptions are critical. After repeated integration by parts of (7) an asymptotic approximation is given by,

$$\boldsymbol{\tau}(z) \sim \left[\boldsymbol{\tau}_w - \boldsymbol{\tau}_{CS}(0)\right] \left(\frac{A_v(z)}{A_v(0)}\right)^{\frac{1}{4}} \frac{\sinh\left[\boldsymbol{\theta}(z)\right]}{\sinh\left[\boldsymbol{\theta}(0)\right]} + \boldsymbol{\tau}_{CS}(z), \qquad \frac{h_s^2}{h_{Ek}^2} \to 0.$$
(13)

²⁵³ The surface wave field therefore introduces a Coriolis-Stokes stress,

$$\boldsymbol{\tau}_{CS}(z) = -\boldsymbol{\rho}A_{\nu}\frac{\partial \mathbf{u}_{s}}{\partial z}(1+i\frac{1}{2}\frac{h_{Ek}^{2}(z)}{h_{s}^{2}})^{-1},$$
(14)

which is rotated $(90 + \Lambda)^{\circ}$ to the left of the wave direction (Northern Hemisphere) where $\Lambda \sim$ 254 $tan^{-1}(h_s^2/h_{Ek}^2)$ (Figure 4). This stress modifies the ageostrophic frictional response in two ways. 255 First, the Coriolis-Stokes stress can balance a portion of the applied surface wind stress, leading 256 to a total Ekman layer response which can be considered as forced by an effective stress, given 257 by the first bracketed term on the RHS of (13), rather than by the wind stress alone (Polton et al. 258 2005; McWilliams et al. 2014). Second, the Coriolis-Stokes stress directly affects a layer of depth 259 scale h_s , through the last term on the RHS of (13). The vertical divergence of this term, in (12), 260 drives near-surface ageostrophic velocities that tend to rotate the surface flow into the down-wave 261 direction (Fig. 4b,c). Together these two modifications introduce a boundary layer transport of 262 $-U_0h_s$, canceling the Lagrangian Stokes transport (see Polton et al. 2005, for a detailed discussion 263 of the frictional Coriolis-Stokes transport). 264

The results of this section confirm the analysis of Polton et al. (2005), and extend them to an arbitrary vertical structure of A_{ν} , subject to the aforementioned constraints. As discussed by Polton et al. (2005, their section 2c), (13) and (14) imply that in the limit $h_s^2/h_{Ek}^2 \rightarrow 0$, the wave modification to the Eulerian currents can be modeled solely through a modification to the surface boundary condition. The proceeding analysis confirms this result is fully independent of the particular form of vertical mixing, and consequently may be of general use in guiding observational or modeling studies where the Stokes layer is not directly resolved.

²⁷² b. Thermal wind shear: $h_{Ek} \ll h_{\rho}, h_{Av}$

A similar simplification of Eq. (7) can be found for the case of a horizontal buoyancy gradi-273 ent driving a thermal wind shear in the near-surface layer. We assume that the Ekman depth is 274 shallow relative to the depth scales over which the horizontal density gradient and A_v vary, ie. 275 $h_{Ek} \ll h_{\rho}, h_{A\nu}$. An example of the scales associated with a mesoscale frontal system can be found 276 from observations of the Azores front (Rudnick 1996), where, using parameters from Nagai et al. 277 (2006), $h_{Ek} \sim 15 \text{ m}$, $h_{Av} \sim 40 \text{ m}$, based on the depth of the transition layer below the mixed layer, 278 and $h_{\rho} \sim 100$ m, based on the depth of the thermocline and the observed geostrophic frontal veloc-279 ity. This limiting case is marginally valid for these parameter values, and thus can be considered 280 as requiring a fairly idealized frontal configuration (cf. Thomas and Lee 2005), included largely 281 for the insight it offers into the basic dynamics of (7), and for comparison with (13). 282

For simplicity it is also assumed that $A_{\nu}\nabla b \rightarrow 0$ at z = -h. If this assumption is not made the solution requires an additional bottom Ekman layer at z = -h in order to satisfy the bottom boundary condition (5), as shown schematically in Fig. 3a. Repeated integration by parts of equation (7) leads to an asymptotic approximation given by,

$$\boldsymbol{\tau}(z) \sim \left[\boldsymbol{\tau}_{w} - \boldsymbol{\tau}_{geo}(0)\right] \left(\frac{A_{v}(z)}{A_{v}(0)}\right)^{\frac{1}{4}} \frac{\sinh\left[\boldsymbol{\theta}(z)\right]}{\sinh\left[\boldsymbol{\theta}(0)\right]} + \boldsymbol{\tau}_{geo}(z), \qquad \frac{h_{Ek}^{2}}{h_{\rho,A_{v}}^{2}} \to 0.$$
(15)

²⁸⁷ Where,

$$\boldsymbol{\tau}_{geo}(z) = \rho A_{\nu} \frac{\partial \boldsymbol{u}_g}{\partial z} \left(1 - i h_{Ek}^2(z) \left[\frac{A_{\nu}'}{A_{\nu}} \frac{\nabla b'}{\nabla b} + \frac{1}{2} \frac{\nabla b''}{\nabla b} + \frac{3}{8} \frac{A_{\nu}''}{A_{\nu}} + \frac{3}{32} \left(\frac{A_{\nu}'}{A_{\nu}} \right)^2 \right] \right), \tag{16}$$

²⁸⁸ defines the geostrophic stress, with primes denoting vertical differentiation.

Closely paralleling the solution for the Coriolis-Stokes stress, (13), the modification of the sur-289 face boundary layer stress by horizontal buoyancy gradients also consists of two components. The 290 first is a modification to the Ekman layer, whereby the Ekman response is forced only by that 291 portion of the wind stress that is out of balance with the geostrophic stress, which again can be 292 considered as defining an effective surface stress, given by the first bracketed term on the RHS 293 of equation (15) (Thompson 2000; Nagai et al. 2006; Cronin and Kessler 2009). Thus, even in 294 the case of $\tau_w = 0$, thermal wind shear will drive an ageostrophic flow within the Ekman layer, 295 with implications for frontal spin-down (Garrett and Loder 1981; Csanady 1982; Thompson 2000; 296 Thomas and Rhines 2002), filament frontogenesis (Gula et al. 2014; McWilliams et al. 2015), and 297 near-surface fluxes (Thomas and Ferrari 2008), discussed in section 5b. It is worth noting that 298 advection of the horizontal buoyancy gradient by the ageostrophic frictional flow can modify the 299 buoyancy gradient and thereby feedback into the Ekman solution, which is discussed in further 300 detail in Thompson (2000), and McWilliams et al. (2015). 301

The second term on the RHS of equation (15) represents the turbulent stress that arises directly from a thermal wind shear in the presence of a viscosity, often termed the geostrophic stress, given by (16). The divergence of this term drives a weak flow throughout the entire layer with velocities that scale as h_{Ek}/h relative to the ageostrophic velocity in the Ekman layer, but with a vertically integrated transport that exactly cancels the transport in the Ekman layer driven by the surface geostrophic stress. The definition of geostrophic stress given here, (16), differs from that given by previous investigators, who, considering only vertically uniform A_v and ∇b , suggest $\tau_{geo}(z) =$ $\rho A_v \frac{\partial u_g}{\partial z}$. Including vertical structure in these parameters gives rise to four additional terms in the definition of geostrophic stress, bracketed in (16), which enter the asymptotic approximation at order $h_{Ek}^2/h_{\rho,A_v}^2$.

These additional terms are imaginary, and thus have the effect of rotating the geostrophic 312 stress vector slightly from the geostrophic shear vector. This is illustrated in figure 5, where 313 the geostrophic stress vector is rotated by an angle, λ , which scales as $\lambda \sim tan^{-1}(h_{Ek}^2/h_{\rho,Av}^2)$, 314 or equivalently, $\lambda \sim tan^{-1}(2Ek)$ (figure 6). Transport in the Ekman layer, U_{Ek}^T , is opposed by 315 geostrophic stress driven transport over the full boundary layer depth, U_{BL}^T . Surface velocity is 316 given by $u_{surf} = u_{Ek}(0) + u_{BL}(0)$, a combination of the Ekman ageostrophic velocity forced by 317 the effective surface stress (u_{Ek}) , and an interior ageostrophic velocity forced by the divergence of 318 the geostrophic stress (u_{BL}) . The direction of the near-surface frictional flow relative to the buoy-319 ancy gradient is consequently a function of both the angle of the geostrophic stress, determined 320 by vertical structure in A_v and ∇b , as well as the ratio $u_{BL}/u_{Ek} \sim h_{Ek}/h$. As a corollary to this, 321 a latitudinal dependence in λ appears implicitly through the Ekman depth, as $h_{Ek}^2/h_{\rho,A\nu}^2 \to \infty$ as 322 $f \rightarrow 0$, with the geostrophic stress vector becoming increasingly parallel to the buoyancy gradient 323 at low latitudes. 324

325 4. Frictional Secondary Circulation

The cross front circulation which arises from frictional effects, shown schematically in figure 3, acts to spin-down ocean fronts, and sharpen cold filaments, due to buoyancy fluxes associated with the ageostrophic velocities necessary to match the surface boundary condition, (4), in the presence

of a geostrophic shear (Garrett and Loder 1981; Thompson 2000; McWilliams et al. 2015). Fur-329 ther, in the case that $\nabla^2 b \neq 0$, convergences (divergences) of this cross-front ageostrophic circu-330 lations will drive negative (positive) vertical velocities in the boundary layer (Garrett and Loder 331 1981; Thompson 2000). These effects have been examined primarily in the context of subme-332 soscale dynamics, where the Rossby number, ε , is not small, and hence are generally diagnosed 333 within the context of non-linear models (eg. Nagai et al. 2006). However, recent comparisons with 334 modeled submesoscale eddies and filaments have suggested that vertical velocities in the bound-335 ary layer can be accurately diagnosed using this simple linear theory even at high ε (Ponte et al. 336 2013; Gula et al. 2014; McWilliams et al. 2015). At larger spatial scales, similar effects are also 337 suggested in an ocean global climate model (Cronin and Tozuka 2015). Therefore, friction acting 338 on the baroclinic component of the flow may be important to boundary layer dynamics across a 339 range of spatial scales. 340

³⁴¹ Vertical velocity for the generalized Ekman model is given by the standard relationship,

$$w(x, y, z) = -\hat{k} \cdot \nabla \times \frac{\boldsymbol{\tau}(x, y, z)}{\rho f} + L(x, y),$$
(17)

where L(x, y) is a constant of vertical integration chosen to fulfill a rigid lid boundary condition. To illustrate how the various components of the full solution enter the calculated vertical velocity we can utilize the simplified definition of stress given by (13), and (15), in (17), which for a two dimensional configuration, invariant in the y-direction, reduces to,

$$w(x,z) = -\frac{\partial}{\partial x} \left[\operatorname{Im} \left\{ \frac{\tau_{Eff}(x)}{\rho f} \left(\frac{A_{\nu}(x,z)}{A_{\nu}(x,0)} \right)^{\frac{1}{4}} \frac{\sinh\left[\theta(x,z)\right]}{\sinh\left[\theta(x,0)\right]} \right\} \right] - \frac{\partial}{\partial x} \left[\operatorname{Im} \left\{ \frac{\tau_{Int}(x,z)}{\rho f} \right\} \right] + L(x).$$
(18)

The first term on the RHS represents upwelling occurring within the Ekman layer, which is now forced by an effective stress, $\tau_{Eff} = \tau_w - \tau_{CS}(0) - \tau_{geo}(0)$. The second term on the RHS gives the boundary layer vertical velocity arising solely from the gradient of the interior forcing, ³⁴⁹ $\tau_{Int} = \tau_{CS} + \tau_{geo}$. When (18) is evaluated at the base of the layer, z = -h, it reduces to the ³⁵⁰ classic Ekman upwelling driven solely by the curl of the wind-stress, however, within the layer, ³⁵¹ both horizontal gradients in the forcing and horizontal gradients of the vertical structure can drive ³⁵² vertical velocities. For the Ekman layer, this can be envisioned as Ekman transport occurring along ³⁵³ contours of constant h_{Ek} , which, for a spatially varying A_v , have a vertical component.

Below the surface Ekman layer, where $\tau_{CS}(x,z) \approx 0$, and in the limit of $h_{Ek}^2/h_{\rho,A_v}^2 \to 0$, (18) 354 reduces to the scaling given by Garrett and Loder (1981), $w \sim g f^{-2} \rho^{-1} \partial (A_v \partial \rho / \partial x) / \partial x$. Thus, 355 vertical velocity in the boundary layer interior, outside the Ekman layer, is driven by gradients in 356 thermal wind shear and A_{ν} , with order $h_{Ek}^2/h_{\rho,A_{\nu}}^2$ modifications due to the gradient of the bracketed 357 terms in (16), reflecting the role of the vertical structure of A_v and ∇b in setting the direction of 358 the geostrophic stress vector (Section 3b, figure 6). Approaching the surface, vertical velocity 359 decays exponentially over an Ekman layer of depth scale h_{Ek} , with additional near-surface vertical 360 velocities in a thin layer of depth scale h_s driven by the horizontal divergence of the Coriolis-Stokes 361 stress. 362

To illustrate the secondary circulation that arises from the balance (1), we examine an idealized front in the x-z plane (200 km width, h=500 m), based on an approximation of the Frontal Air-Sea Interaction Experiment (FASINEX) data (Pollard and Regier 1992), similar to Thompson (2000). We set $\tau_w = 0$ and $\tau_{CS} = 0$ as the solution is linear and these effects are simply additive. The buoyancy in this model is given by,

$$b(x,z) = \frac{1}{2}b_f \tanh\left[\frac{\hat{z} - \alpha(\hat{x} - x_0)^3 - z_0}{d_0}\right] + \frac{1}{2}b_b \tanh\left[\frac{\hat{z} - z_0}{d_1}\right],$$
(19)

with values of parameters given in Table 1 and the hat notation indicating non-dimensionalized coordinates ranging from 0 to 1. Values of A_v are based on the approximation used in McWilliams et al. (2015), designed to be broadly consistent with KPP (Large et al. 1994). This is used simply to illustrate several general features of the solution that arise from horizontal and vertical structure in mixing across a frontal region, rather than provide an absolutely accurate diagnostic, and the qualitative discussion that follows is not sensitive to the detailed particulars of our choice of A_{ν} .

1

$$A_{\nu}(x,z) = A_{\nu 0}G(\zeta)\frac{\hat{h}(\hat{x})}{h_0} + A_{\nu b}, \quad \zeta = -\frac{\hat{z}}{\hat{h}(x)}, \tag{20}$$

374

$$G(\zeta) = \frac{27}{4} (1 + \zeta_0^2) (\zeta_0 + \zeta) (1 - \zeta)^2, \quad \zeta \le 1,$$
(21)

375

$$G(\zeta) = 0, \quad \zeta > 1. \tag{22}$$

³⁷⁶ *G* has a maximum value of 1 in the boundary layer, ζ_0 is a small parameter introduced to avoid a ³⁷⁷ singularity at z = 0, and \hat{h} is the surface boundary layer depth, taken here as,

$$\hat{h}(x) = h_0 + \delta_h \left(\tanh\left[\frac{\alpha(\hat{x} - x_0)^3}{d_0}\right] - \frac{1}{2} \right).$$
 (23)

All parameter values for equations (19-23) are given in Table 1. Figure 7 shows the structure of the idealized front, and the eddy viscosity, along with the associated along front geostrophic flow, implying $\varepsilon \sim 0.05$. We further assume w = 0 at z = 0 (rigid lid), and define an ageostrophic crossfront streamfunction such that $(u_{ag}, w) = (\psi_z, -\psi_x)$. Using the meridional momentum equation, Im[Eq. (1)], gives

$$\Psi = \frac{1}{\rho f} \tau^{y}(x, z). \tag{24}$$

The secondary overturning circulation arising from the geostrophic stress is found numerically, and shown in figure 7. This is a thermally direct circulation, with a counter-clockwise sense of rotation, that tends to tilt the front and restratify the near-surface (Thompson 2000). Downwelling velocities on the dense side of the front are stronger than the upwelling on the buoyant side of the front, consistent with previous findings (Samelson 1993; Thompson 2000). Streamlines are closed, indicating zero vertically integrated horizontal transport, as required to maintain the classic Ekman transport. This is a general result that does not depend on the frontal configuration. Note however that although the vertically integrated horizontal transport is zero, the associated fluxes need not be zero, as discussed further in section 5b. Further, the vertical buoyancy flux associated with the secondary overturning circulation can be non-zero (McWilliams et al. 2015), and hence may play a role in the general circulation through vorticity stretching of the interior.

To illustrate the importance of spatial variability in A_{y} , we decompose the total vertical velocity 394 field (figure 8a), w_{total} , into vertical velocities due to the gradient in the forcing ($\nabla^2 b$), which we 395 designate $w_{forcing}$, and the remainder which is a function only of the spatial structure in A_v , which 396 we designate w_{Av} , as discussed in relation to (18). For the particular frontal configuration examined 397 here w_{Av} is ~ 25% of w_{total} . However, locally near the base of the turbulent boundary layer (Fig 8, 398 dashed line) w_{Av} can be the dominant term, and hence may be of particular importance for vertical 399 fluxes into the near-surface layer. For a geostrophic stress, the ratio of vertical velocities is given 400 by, 401

$$\frac{w_{Av}}{w_{forcing}} \sim \frac{L_{forcing}}{L_{Av}},\tag{25}$$

where *L* indicates the relevant horizontal length scales. Observations suggest that horizontal length scales over which vertical mixing varies are comparable to frontal features (Dewey and Moum 1990; Nagai et al. 2006), and hence these effects may be first order in determining the vertical velocity in the boundary layer. A similar scaling holds within the Ekman layer, where for a surface wind stress aligned orthogonal to a horizontal gradient in A_{ν} , the ratio of vertical velocities at $z = -h_{Ek}$ is,

$$\frac{w_{Av}}{w_{forcing}} \sim \frac{L_{forcing}}{4L_{Av}}.$$
(26)

5. Discussion and Further Implications

The solutions presented here build upon prior work by allowing vertical variation in A_{ν} , as well 409 as realistic structure in ocean fields, such as ∇b and the Coriolis-Stokes force. Examination of 410 the solutions (7, 13, 15), suggests many ways in which including more physical realism in the 411 problem parameters can modify the expected ageostrophic flow, however to further motivate the 412 importance of this added complexity, we first consider scaling arguments relating the importance 413 of geostrophic stress and Coriolis-Stokes stress to wind stress. The global distribution of these 414 fields are then estimated using a combination of model output and reanalysis data. Finally, we 415 comment briefly on the importance of these modifications to determining horizontal fluxes in the 416 boundary layer. 417

418 a. Scaling and Geographic Distribution

The boundary conditions utilized here ensure that the classic Ekman transport relation is main-419 tained, even in the presence of ocean sources of stress. However, as demonstrated in Section 2, 420 ocean sources of stress can greatly modify the vertical structure of currents, and hence are fun-421 damental to understanding boundary layer dynamics. Determining the magnitude of both the 422 geostrophic stress and the Coriolis-Stokes stress depends critically on the value of A_{ν} , which 423 complicates their determination from observations. However, at low frequencies, variability in 424 near-surface A_v may be controlled by variability in the surface wind stress (Wenegrat et al. 2014). 425 Taking $h_{Ek} \sim u^*/f$ (Caldwell et al. 1972), where $u^* = \sqrt{\tau_w/\rho}$, gives $A_v \sim u^{*2}/f$, and hence the 426 ratio of the geostrophic stress to the surface wind stress can be scaled as, 427

$$\gamma_{GEO} = \frac{\tau_{geo}}{\tau_w} \sim \frac{\nabla b}{f^2}.$$
(27)

The direct proportionality of γ_{GEO} to ∇b , and independence from τ_w , highlights how the 428 geostrophic stress can be expected to be a ubiquitous forcing of ageostrophic flow at sharp frontal 429 features, and consequently may be fundamental for understanding horizontal heat flux at buoy-430 ancy fronts. The f^{-2} dependence indicates a rapid increase at low latitudes. Further, utilizing 431 the stratified Ekman depth scaling, $h_{Ek} \sim u^* / \sqrt{Nf}$ (Pollard et al. 1973), in (27) gives $\gamma_{GEO} \sim Bu$, 432 where Bu = NH/fL is the Burger number, defined such that $b \sim N^2H$. Determination of Bu is thus 433 dependent on the geometry of the particular front being considered, however, for many oceanic 434 flows, observations suggest $Bu \sim O(1)$, implying $\gamma_{GEO} \sim O(1)$ (Nagai et al. 2006; Boccaletti et al. 435 2007). 436

⁴³⁷ The Coriolis-Stokes stress can be scaled relative to the surface wind stress as,

$$\gamma_{CS} = \frac{\tau_{CS}}{\tau_w} \sim U_0 h_s \frac{\rho f}{\tau_w} \sim L a^{-2} \frac{h_s}{h_{Ek}}.$$
(28)

Therefore, γ_{CS} is proportional to the Stokes transport divided by the wind-driven Ekman transport (McWilliams and Restrepo 1999; Polton et al. 2005). Alternatively, this can be rewritten using the turbulent Langmuir number, $La = (u^*/U_0)^{1/2}$, which scales the ratio of wind forced production of turbulent kinetic energy (TKE) to the wave forced production of TKE (McWilliams et al. 1997; Grant and Belcher 2009), with typical values of 0.2-0.5 (Smith 1992; Belcher et al. 2012). This suggests that $\gamma_{CS} \sim O(1)$ for h_s/h_{Ek} of 0.04 - 0.25.

To form estimates of the global distributions of γ_{GEO} and γ_{CS} a combination of reanalysis data and model output is utilized. The total Stokes transport is found from the WaveWatch III (WWIII) model, reported every 6 hours on a 0.5° grid (Rascle et al. 2008; Rascle and Ardhuin 2013). For consistency with the WWIII model forcing, we utilize NCEP Climate Forecast System Reanalysis (CFSR) wind stress, temperature and salinity at 5 m, and horizontal currents at 5 and 15 m depth (Saha et al. 2006). To estimate the geostrophic stress we calculate buoyancy gradients from ⁴⁵⁰ monthly 5 m temperature and salinity (0.5° resolution), and then infer approximate monthly values ⁴⁵¹ of A_v using the surface boundary condition, $\rho A_v = \tau_w (\partial u/\partial z)^{-1}$, with the near-surface shear mag-⁴⁵² nitude approximated using CFSR velocities at 5 and 15 m, $\partial u/\partial z \sim |u(-5) - u(-15)|/10$. Alter-⁴⁵³ nate parameterizations of A_v were tested, including wind-stress only parameterizations (Wenegrat ⁴⁵⁴ et al. 2014) and bulk Richardson number closures (Pollard et al. 1973), and found to give sim-⁴⁵⁵ ilar results (not shown here). Monthly values of the Stokes transport and τ_{geo} over the period ⁴⁵⁶ 2001-2011 are then used to form climatologies of γ_{GEO} and γ_{CS} .

Figures 9 and 10 show the global seasonal climatology of γ_{GEO} and γ_{CS} , respectively. The dependence on latitude through the Coriolis frequency is apparent in both quantities, with γ_{GEO} peaking at low latitudes, and γ_{CS} dominating at higher latitudes (figure 11). Regional variability is also evident, with γ_{GEO} enhanced in boundary currents, along the equatorward edges of the subtropical gyres, and through much of the Indian ocean and eastern subtropical Pacific. These parameterized results can be compared to estimates derived from model output and alternate parameterizations of A_{ν} , which indicate similar spatial patterns (Chu 2015; Cronin and Tozuka 2015).

In the zonal average, and temporal average, γ_{CS} becomes larger than γ_{GEO} poleward of 15° 464 (figures 10 and 11), following a spatial pattern that in large part reflects the variability in Stokes 465 transport (McWilliams and Restrepo 1999). This latitudinal pattern may also reflect the effect 466 of the coarse resolution products utilized here on estimating γ_{GEO} , as the first baroclinic Rossby 467 radius at 15° is ~ 100km (Chelton et al. 1998), which is close to the resolved meridional Nyquist 468 wavelength, and hence ∇b may be underestimated at higher latitudes. Wide swaths of the world 469 oceans have $\gamma_{CS} \sim 0.25$, emphasizing how important these effects may be for Ekman layer currents. 470 Intensification of γ_{CS} in the southern ocean is also evident. In the Northern Hemisphere there is 471 a general enhancement of γ_{CS} in the eastern side of the ocean basins, with seasonal variability in 472

⁴⁷³ both extent and magnitude, resulting from enhanced Stokes transport associated with increased ⁴⁷⁴ wintertime wind-forcing.

The relative influences of the geostrophic stress and the Coriolis-Stokes stress can be consid-475 ered using the joint probability density function (PDF) of the monthly estimates of γ_{GEO} and γ_{CS} , 476 evaluated between $5^{\circ} - 73.5^{\circ}$ from 2001-2011 (figure 12). Consistent with the spatial maps, the 477 PDF has a broad peak at $\gamma_{CS} \sim 0.1 - 0.25$ with negligible γ_{GEO} . However, the distribution of γ_{GEO} 478 is long-tailed, reflecting its spatial and temporal inhomogeneity, evident in comparing an example 479 month (figure 13) and the climatological maps (figure 9). Considering the total relative change in 480 the effective surface stress arising from both the geostrophic stress and the Coriolis-Stokes stress, 481 $\gamma_T = \gamma_{GEO} + \gamma_{CS}$, 36% of all points have $\gamma_T > 0.25$. Together the estimates presented above, while 482 only a rough approximation, suggest that surface waves will be of O(1) importance for much 483 of the extra-tropics, while baroclinic pressure gradients will dominate at low latitudes, in frontal 484 systems, and potentially over shorter timescales, and smaller spatial scales, than resolved here, 485 specifically at the submesoscale, where geostrophic stress effects have been demonstrated to sig-486 nificantly modify the ageostrophic flow (Ponte et al. 2013; Gula et al. 2014; McWilliams et al. 487 2015). 488

489 b. Horizontal Fluxes

The proceeding analysis, and theory, highlights how ocean sources of stress can be expected to modify the frictional response within the near-surface layer, effecting the magnitude, direction, and vertical profile of the ageostrophic flow. These modifications to the ageostrophic velocity can often be approximated using the concept of an effective stress, τ_{Eff} (section 3), leading to a modified Ekman velocity scale of, $u_{Ek} \sim \tau_{Eff}/(\rho f h_{Ek})$. This has wide-ranging implications for horizontal advective fluxes, where for example, the geostrophic stress will always enhance heat flux down the buoyancy gradient relative to the classic Ekman solution, as well as for other dynamically important quantities such as the wind-work on the total ageostrophic flow, $\tau_w \cdot u_{Ek}$, which will be reduced for winds aligned with the surface frontal jet (down-front winds), and enhanced for winds aligned against the frontal jet (up-front winds).

⁵⁰⁰ A brief example, which highlights the role of vertical structure in A_{ν} , is given by considering ⁵⁰¹ the differential horizontal buoyancy flux across the surface Ekman layer, which can change the ⁵⁰² stratification, and hence the potential vorticity (PV), of the near surface layer (Thomas and Ferrari ⁵⁰³ 2008). A scaling for the frictional flux of the vertical component of PV due to a surface wind ⁵⁰⁴ stress in the presence of a horizontal buoyancy gradient is given by (Thomas 2005; Thomas and ⁵⁰⁵ Ferrari 2008),

$$J_z^F \sim \frac{\tau_w}{\rho h_{Ek}} \nabla b. \tag{29}$$

However, if A_{ν} is allowed to vary vertically, with depth scale $h_{A\nu}$, this scaling is modified to become,

$$J_z^F \sim \frac{\tau_w}{\rho h_{Ek}} \nabla b \left(1 + \frac{h_{Ek}}{4h_{Av}} \right). \tag{30}$$

This relationship is shown in figure 14, for a wind stress aligned with the front, and an exponential A_{ν} profile that decreases (increases), $h_{A\nu} > 0$ ($h_{A\nu} < 0$), with depth. As $|h_{A\nu}|$ approaches h_{Ek} the surface cross-front current is enhanced (reduced) for $h_{A\nu} > 0$ ($h_{A\nu} < 0$), modifying the frictional PV flux. A similar result can be easily derived for the influence of vertical structure of A_{ν} on the frictional PV flux associated with the frontal spin down by the geostrophic stress (Thomas and Ferrari 2008). Thus, the vertical structure of mixing is linked to the flux of vertical potential vorticity through its effect on the differential horizontal advection of buoyancy.

A conceptual example of how this might affect the ocean boundary layer is is found by considering the near-surface response to up-front and down-front winds. Down-front winds advect

dense water over light water, leading to gravitational instability, whereas up-front winds advect 517 light water over dense, enhancing stratification (Thomas and Lee 2005). Making the idealization 518 that down-front winds lead to a well-mixed layer with $A_{\nu} \sim \text{constant}$, whereas up-front winds lead 519 to a stratified near-surface layer with A_{ν} decreasing with depth, would imply the existence of an 520 asymmetry in PV fluxes between the two cases. Consequently, for the same wind stress and buoy-521 ancy gradient magnitudes, wind-driven frictional PV fluxes may be enhanced in up-front wind 522 conditions relative to down-front winds, providing a possible alternative route to the creation of 523 positively skewed PV distributions (Thomas 2007). 524

525 6. Summary

In this manuscript we present an approximate solution to the generalized Ekman (Cronin and 526 Kessler 2009), or Turbulent Thermal Wind (Gula et al. 2014; McWilliams et al. 2015), balance. 527 While this theory omits many aspects of boundary layer physics that are likely to be active in the 528 real ocean, the simplicity and generality of the solution provides a useful tool for gaining insight 529 into the underlying dynamics beyond that available from numerical methods. The full solution, 530 (7), given in terms of an integral over a Green's function, can be applied quite generally to a variety 531 of sources of near-surface shear, and further allows for arbitrary vertical structure in A_{ν} , subject to 532 the constraints imposed by the WKB method. Many existing modified Ekman theories can thus be 533 considered as particular cases of this solution, providing a framework for comparing their effects 534 on ageostrophic ocean currents. 535

Two important aspects of surface layer dynamics which are not as readily accommodated in this framework are time-dependence, and non-linearity. Observations suggest significant diurnal variatiability of near-surface shear (Price et al. 1986; Schudlich and Price 1998; Cronin and Kessler 2009; Smyth et al. 2013; Wenegrat and McPhaden 2015), which has been suggested as an expla-

nation of observed discrepancies with classic Ekman theory (Price and Sundermeyer 1999). How-540 ever, analysis of near-surface velocity observations appears to suggest that some of the observed 541 features which have been used to argue for the role of time-variability, such as a flattened spiral, can 542 also be very well explained by alternate mechanisms that do not invoke time-dependence (Lewis 543 and Belcher 2004; Polton et al. 2005; Cronin and Kessler 2009). Disentangling these effects using 544 observations is further complicated by measurement challenges, particularly for moored observa-545 tions which can be biased by surface waves (Rascle and Ardhuin 2009). A focus of future work 546 should be clarifying the relative contributions of, and interactions between, the diverse sources of 547 near-surface ageostrophic flow. 548

Non-linear effects may be of particular importance in examining sharp horizontal buoyancy 549 gradients (Stern 1965; Niiler 1969; Thomas and Rhines 2002; Mahadevan and Tandon 2006). 550 However, a range of modeling efforts which include more complete physics indicate that the basic 551 dynamical mechanisms discussed here continue to be of first order importance in the boundary 552 layer, even at high ε (Thompson 2000; Nagai et al. 2006; Ponte et al. 2013; Gula et al. 2014; 553 McWilliams et al. 2015). We also note that the work of Wu and Blumen (1982), and (Tan 2001), 554 can be considered as a blueprint for how the semi-geostrophic momentum approximation could be 555 incorporated into the solution given here. 556

Examining two limiting cases, the first for Stokes shear of shallow depth relative to h_{Ek} , and the second for a front much deeper than h_{Ek} , reveals the key underlying dynamics. Ocean sources of shear, in the presence of viscosity, act as sources of stress. These ocean sources of stress are, as a first approximation, independent of the surface wind stress, and are capable of driving their own ageostrophic flow, including creating a surface Ekman layer. The equivalency of the closed form solutions for the two limiting cases emphasizes how robust this interpretation of the underlying dynamics is, suggesting the same interpretation holds for the more general Green's ⁵⁶⁴ function solution (7), and highlighting a previously unnoted connection between the frictional ⁵⁶⁵ effects of surface waves and fronts.

The solutions presented here are unique in their ability to incorporate arbitrary vertical structure 566 in A_{ν} , which is motivated physically by modeling and direct turbulence measurements (Zikanov 567 et al. 2003; Kirincich 2013; Soloviev and Lukas 2014), and is shown here to lead to modifications 568 of both horizontal and vertical flows. Improved understanding of the spatial variability of mixing 569 is key to understanding and parameterizing these effects on boundary layer flow. Finally, it should 570 be emphasized that the various dynamical processes discussed here should not be considered as 571 the addition of new parameters to the Ekman problem, but rather as fundamental components of 572 the frictional response of the ocean boundary layer whose influence may be of the same order 573 of magnitude as the surface wind stress throughout large portions of the global oceans. The total 574 frictional ageostrophic response is a combination of a response to the surface wind, as in the classic 575 Ekman theory, and a response to ocean sources of shear. 576

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NCEP CFSR data is available at: http://rda.ucar.edu/pub/cfsr.html.

APPENDIX A

585

584

Derivation

Let $A_{\nu}(z) = A_{\nu 0} \psi(z)$, then, after non-dimensionalizing (3) as discussed in section (2), we have,

$$Ek\tau'' - \frac{i}{\psi(z)}\tau = \xi(z), \tag{A1}$$

587

$$\tau(h) = \tau_w, \tag{A2}$$

$$\tau(0) = 0, \tag{A3}$$

where primes indicate vertical derivatives and all variables are non-dimensional unless otherwise noted. The RHS of (A1) is given in terms of a generic inhomogeneous forcing function, $\xi(z)$, which could arise from geostrophic shear or Coriolis-Stokes shear as discussed in section 3. Solving first for the homogenous solution, and making the WKB assumption,

$$\tau \propto e^{(S_0 + S_1 \delta + S_2 \delta^2 + \dots)/\delta},\tag{A4}$$

593 gives,

$$Ek\left[\frac{S_0''}{\delta} + \frac{S_0'^2}{\delta^2} + \frac{2S_0'S_1'}{\delta} + S_1'' + S_1'^2\right]\tau - \frac{i}{\psi(z)}\tau = 0.$$
 (A5)

The distinguished limit for the parameter δ is therefore, $\delta \sim Ek^{\frac{1}{2}}$, and the balance conditions are given by,

$$O(\delta^{-2}): \quad S'_0 = \pm \sqrt{\frac{i}{\psi(z)}},\tag{A6}$$

596

$$O(\delta^{-1}): \quad S_1' = -\frac{S_0''}{2S_0'}.$$
 (A7)

Taking the positive root of
$$S'_0$$
 gives,

$$S_0 = \sqrt{i} \int_{-h}^{z} \psi(Z)^{-\frac{1}{2}} dZ,$$
 (A8)

598 and,

$$S_1 = \frac{1}{4} \log \psi(z). \tag{A9}$$

A similar argument is followed for the negative root, giving the two solutions to the ODE, which dimensionally are given by,

$$\tau(z) = C_1 A_{\nu}(z)^{\frac{1}{4}} e^{\theta(z)} + C_2 A_{\nu}(z)^{\frac{1}{4}} e^{-\theta(z)}, \tag{A10}$$

601 where,

$$\theta(z) = \sqrt{if} \int_{-h}^{z} A_{\nu}(Z)^{-\frac{1}{2}} dZ.$$
 (A11)

For the WKB approximation to hold, two conditions must be satisified (Bender and Orszag 1978),

$$\frac{Ek^{\frac{1}{2}}S_1}{S_0} \ll 1, \quad Ek^{\frac{1}{2}} \to 0,$$
 (A12)

604

$$Ek^{\frac{1}{2}}S_2 \ll 1, \quad Ek^{\frac{1}{2}} \to 0,$$
 (A13)

discussed further in Appendix B.

Variation of parameters gives the inhomogeneous portion of the solution,

$$\tau_p = -y_1 \int \frac{y_2 \xi(z)}{W(y_1, y_2)} dz + y_2 \int \frac{y_1 \xi(z)}{W(y_1, y_2)} dz.$$
(A14)

 $_{607}$ Where *W* is the Wronskian,

$$W(y_1, y_2) = y_1 y_2' - y_2 y_1' = -2\sqrt{if}.$$
(A15)

608 Thus,

$$\tau_p = A_v(z)^{\frac{1}{4}} e^{\theta(z)} \int \frac{e^{-\theta(z)} A_v(z)^{\frac{1}{4}} \xi(z)}{2\sqrt{if}} \, \mathrm{d}z - A_v(z)^{\frac{1}{4}} e^{-\theta(z)} \int \frac{e^{\theta(z)} A_v(z)^{\frac{1}{4}} \xi(z)}{2\sqrt{if}} \, \mathrm{d}z. \tag{A16}$$

⁶⁰⁹ Changing the limits of integration gives,

$$\tau_p = \int_0^z \frac{\sinh[\theta(z) - \theta(s)] A_\nu(z)^{\frac{1}{4}} A_\nu(s)^{\frac{1}{4}} \xi(s)}{\sqrt{if}} \, \mathrm{d}s. \tag{A17}$$

⁶¹⁰ So, the total solution, before application of the boundary conditions, is given by,

$$\tau_t(z) = C_1 A_\nu(z)^{\frac{1}{4}} e^{\theta(z)} + C_2 A_\nu(z)^{\frac{1}{4}} e^{-\theta(z)} + \int_0^z \frac{\sinh[\theta(z) - \theta(s)] A_\nu(z)^{\frac{1}{4}} A_\nu(s)^{\frac{1}{4}} \xi(s)}{\sqrt{if}} \, \mathrm{d}s.$$
(A18)

⁶¹¹ Applying the surface BC gives,

$$\tau_{w} = C_{1}A_{v}(0)^{\frac{1}{4}}e^{\theta(0)} + C_{2}A_{v}(0)^{\frac{1}{4}}e^{-\theta(0)}, \qquad (A19)$$

612 therefore,

$$\tau_{t}(z) = 2C_{1}A_{\nu}(z)^{\frac{1}{4}}e^{\theta(0)}sinh\left[\theta(z) - \theta(0)\right] + \tau_{w}\left(\frac{A_{\nu}(z)}{A_{\nu}(0)}\right)^{\frac{1}{4}}e^{\theta(0) - \theta(z)} + \int_{0}^{z}\frac{sinh\left[\theta(z) - \theta(s)\right]A_{\nu}(z)^{\frac{1}{4}}A_{\nu}(s)^{\frac{1}{4}}\xi(s)}{\sqrt{if}}\,\mathrm{d}s.$$
(A20)

⁶¹³ The lower BC gives,

$$0 = 2C_1 A_{\nu}(-h)^{\frac{1}{4}} e^{\theta(0)} \sinh\left[\theta(-h) - \theta(0)\right] + \tau_w \left(\frac{A_{\nu}(-h)}{A_{\nu}(0)}\right)^{\frac{1}{4}} e^{\theta(0) - \theta(-h)} + \int_0^{-h} \frac{\sinh\left[\theta(-h) - \theta(s)\right] A_{\nu}(-h)^{\frac{1}{4}} A_{\nu}(s)^{\frac{1}{4}} \xi(s)}{\sqrt{if}} \, \mathrm{d}s. \quad (A21)$$

Following Hidaka (1955), we multiply equation (A20) by $sinh[-\theta(0)]$, equation (A21) by $-\left(\frac{A_{\nu}(z)}{A_{\nu}(-h)}\right)^{\frac{1}{4}}sinh[\theta(z)-\theta(0)]$, and add them, giving,

$$\tau(z) = \tau_w \left(\frac{A_v(z)}{A_v(0)}\right)^{\frac{1}{4}} \frac{\sinh[\theta(z)]}{\sinh[\theta(0)]} + \int_0^z \frac{\sinh[\theta(z) - \theta(s)]\sinh[\theta(0)]\xi(s)A_v(s)^{\frac{1}{4}}A_v(z)^{\frac{1}{4}}}{(if)^{\frac{1}{2}}\sinh[\theta(0)]} ds - \int_0^{-h} \frac{\sinh[\theta(z) - \theta(0)]\sinh[\theta(s)]\xi(s)A_v(s)^{\frac{1}{4}}A_v(z)^{\frac{1}{4}}}{(if)^{\frac{1}{2}}\sinh[\theta(0)]} ds.$$
(A22)

⁶¹⁶ This can be re-written as,

$$\tau(z) = \tau_w \left(\frac{A_v(z)}{A_v(0)}\right)^{\frac{1}{4}} \frac{\sinh\left[\theta(z)\right]}{\sinh\left[\theta(0)\right]} + \int_{-h}^0 G(z,s)\xi(s)\,\mathrm{d}s,\tag{A23}$$

617

$$\boldsymbol{\theta}(z) = \sqrt{if} \int_{-h}^{z} A_{\nu}(Z)^{-\frac{1}{2}} \,\mathrm{d}Z,\tag{A24}$$

$$G(z,s) = \begin{cases} \frac{\sinh[\theta(z)] \sinh[\theta(s) - \theta(0)] A_{\nu}(s)^{\frac{1}{4}} A_{\nu}(z)^{\frac{1}{4}}}{\sinh[\theta(0)] \sqrt{if}} & \text{if } s > z \\ \frac{\sinh[\theta(s)] \sinh[\theta(z) - \theta(0)] A_{\nu}(s)^{\frac{1}{4}} A_{\nu}(z)^{\frac{1}{4}}}{\sinh[\theta(0)] \sqrt{if}} & \text{if } s < z. \end{cases}$$
(A25)

APPENDIX B

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Accuracy of approximate solution

The validity of the physical-optics WKB approximation requires the criteria (A12) and (A13) 621 be satisified (Bender and Orszag 1978). The relative error in the approximation will then be a 622 function of the small parameter $\delta \sim Ek^{1/2}$ and the first ignored term, S2, in (A4), which involves 623 both first and second derivatives of A_{ν} . Thus, errors will be a function of the Ekman number, Ek, 624 as well as the particular vertical structure of A_{ν} . Anecdotally, the WKB solution (7) has proven 625 extremely accurate across a wide-range of vertical structures of A_{ν} , and values of Ek, considered 626 in developing the model (see also Grisogono 1995; Berger and Grisogono 1998). However, to 627 better illustrate the accuracy of the approximate solution we consider the relative error associated 628 with 3 idealized forms of A_{ν} (figure B1). 629

⁶³⁰ Case I is a simple exponentially decaying profile,

$$A_{\nu}(z) = A_{\nu 0} e^{\frac{z}{0.125}},\tag{B1}$$

⁶³¹ chosen for its analytic simplicity, and consistency with observations (Peters et al. 1988; Dillon
⁶³² et al. 1989). Case II is a linearly decaying profile,

$$A_{\nu}(z) = A_{\nu 0} \left(1 + \frac{\hat{z}}{1+\mu} \right),$$
 (B2)

where μ is a small value added to avoid a singularity at z = -h. Case III is a modified Gaussian profile (Parmhed et al. 2005),

$$A_{\nu}(z) = A_{\nu 0} \phi \hat{z} e^{-\frac{1}{2} (\frac{\hat{z}}{0.25})^2}$$
(B3)

where, $\phi = e^{1/2}/0.25$, which approximates the polynomial profile of O'Brien (1970). This profile violates (A12) in a thin layer near z = 0, where $A_v \to 0$, and hence we evaluate the error associated with this profile only over the depth range where the WKB solution can be expected to be valid. Namely $\hat{z} \le z_p$, where z_p is given by Parmhed et al. (2005) as,

$$z_p = -\frac{1}{4} \left[W\left(\frac{2}{\sqrt{\phi}}\right) \right]^2,\tag{B4}$$

where *W* is the Lambert W function (Corless et al. 1996). For the values used here, $z_p = -0.06$, thus, rather than patching an additional inner solution for $\hat{z} > z_p$, we make the simplifying approximation of applying the surface boundary condition directly to the WKB solution at z_p , similar to the introduction of a roughness length scale (Madsen 1977), and equivalent in the error analysis to the requirement that any inner solution be exact.

To form an estimate of the relative error as a function of Ek, WKB solutions (7) are compared to numeric solutions, found using a shooting method, and the normalized maximum error identified in each vertical profile,

$$\widehat{\tau_{err}}(Ek) = \frac{max\{|\tau_{WKB}(z, Ek) - \tau_{num}(z, Ek)|\}}{max\{|\tau_{num}(z, Ek)|\}}.$$
(B5)

Results are plotted in figure B2, for wind stress only forcing (top), and solutions forced by a 647 vertically uniform buoyancy gradient (bottom). Also plotted for reference is the value of the small 648 parameter, $\delta \sim Ek^{1/2}$. Errors are generally small, and, for $Ek \leq 10^{-1}$, the only case with relative 649 errors exceeding 10% is the modified exponential profile, (B3). The errors associated with this 650 profile are strongly dependent on the choice of patching depth, z_p , rather than the overall vertical 651 structure, as can be anticipated through the logarithmic singularity evident in (A12). Hence caution 652 is required in applying (7) in cases where $A_v \rightarrow 0$ near a boundary. Despite the limitations of the 653 WKB method, its simplicity, and generality, argue its utility, particularly in comparison to the 654 often strict parameter requirements associated with other analytic solutions techniques. 655

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Parameter	Value	Physical Interpretation			
b_f	$-0.6g/ ho_0$	Buoyancy change across front			
b_b	bf/2	Buoyancy change across thermocline			
α	-1.185	Horizontal scaling factor			
ZO	0.75	Vertical position			
<i>x</i> ₀	1	Horizontal position			
d_0	0.125	Horizontal scale			
d_1	0.125	Vertical scale			
f	$6.88 \times 10^{-5} s^{-1}$	Coriolis frequency			
$A_{\nu 0}$	$3\times 10^{-2}m^2s^{-1}$	Eddy viscosity magnitude			
A_{vb}	$1\times 10^{-4}m^2s^{-1}$	Background viscosity			
h_0	0.84	Turbulent boundary layer depth			
δ_h	0.05	Across front change in boundary layer depth			
ζ_0	$5 imes 10^{-3}$	Regularization constant			

TABLE 1. Parameters for equations (19-23).

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FIG. 7. Across-front sections of the idealized two-dimensional front discussed in section 4, with contours of constant ρ indicated in solid black. Upper left: along-front geostrophic velocity, upper right: A_{ν} , lower panel: frictional ageostrophic overturning streamfunction (m^2s^{-1} , eq. 24), with contours of constant A_{ν} in gray.



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FIG. 11. Zonal, and temporal, averages of γ_{GEO} and γ_{CS} .



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